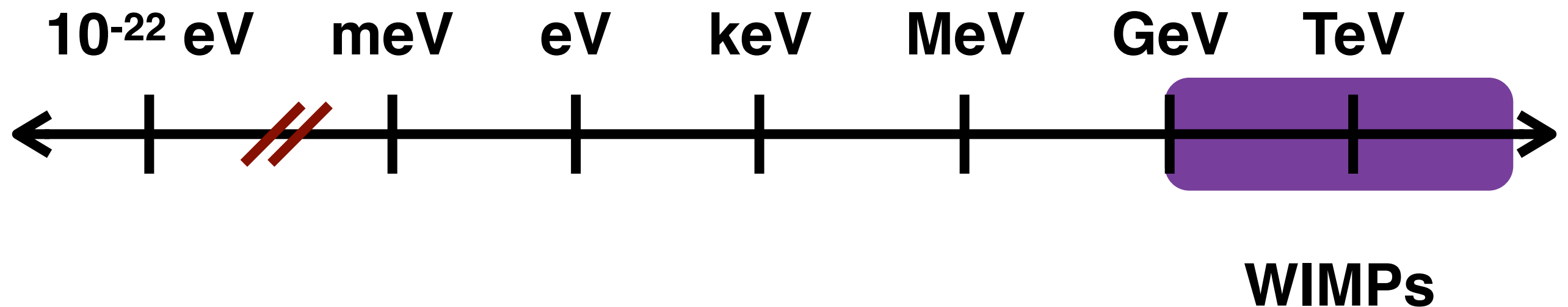


Dispatches from the sub-GeV Dark Matter Frontier

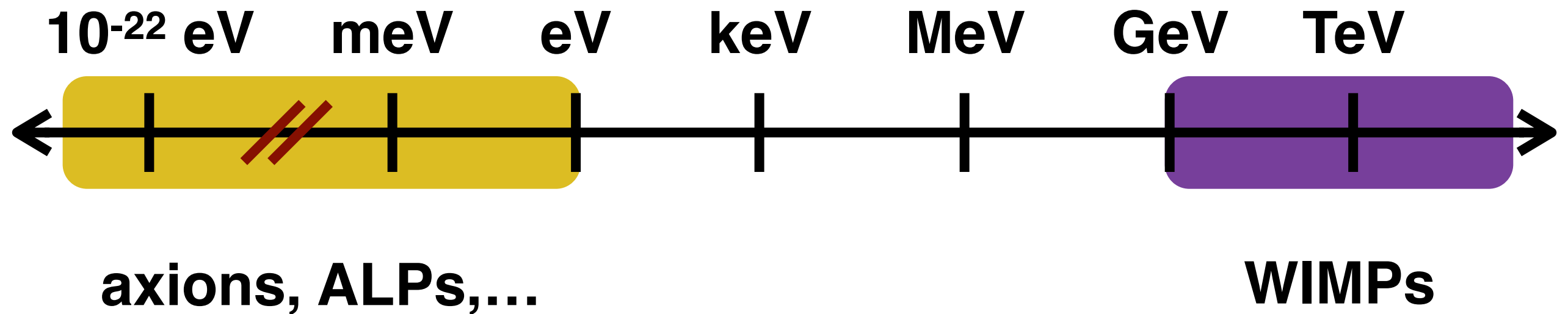
Tien-Tien Yu (University of Oregon)

Institute of Mathematics and Physics of the Universe — December 14, 2018

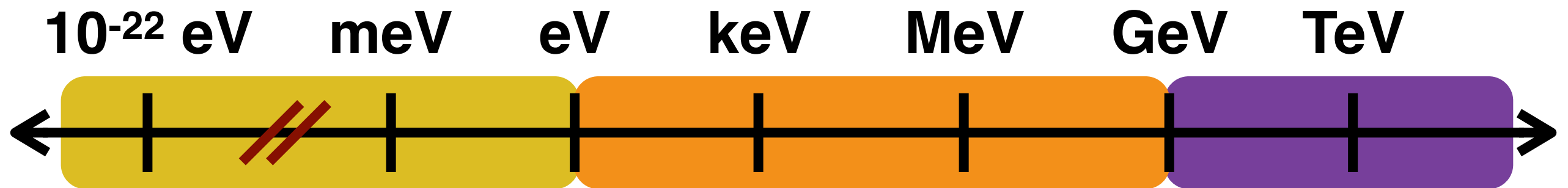
dark sector candidates



dark sector candidates

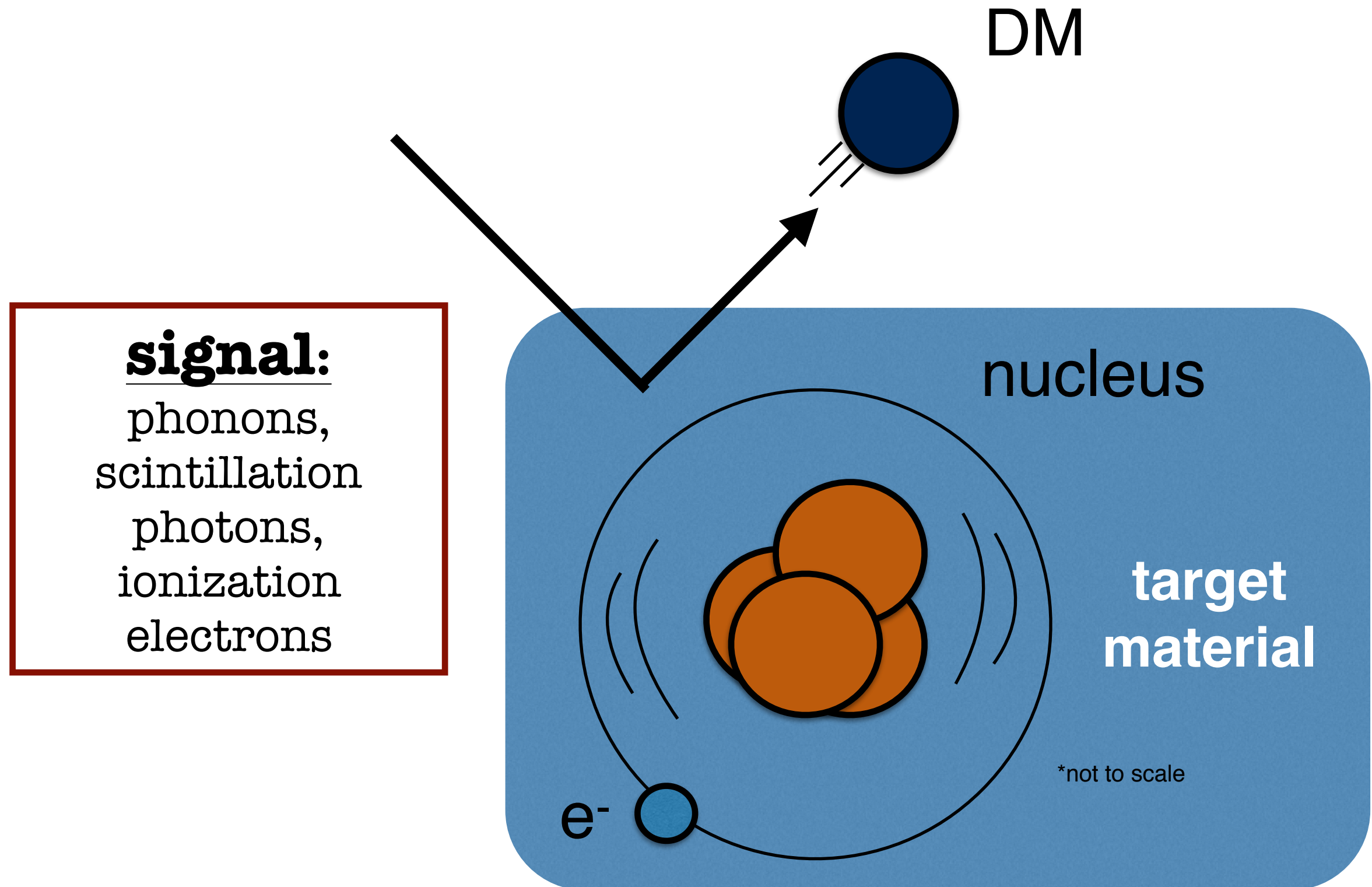


dark sector candidates

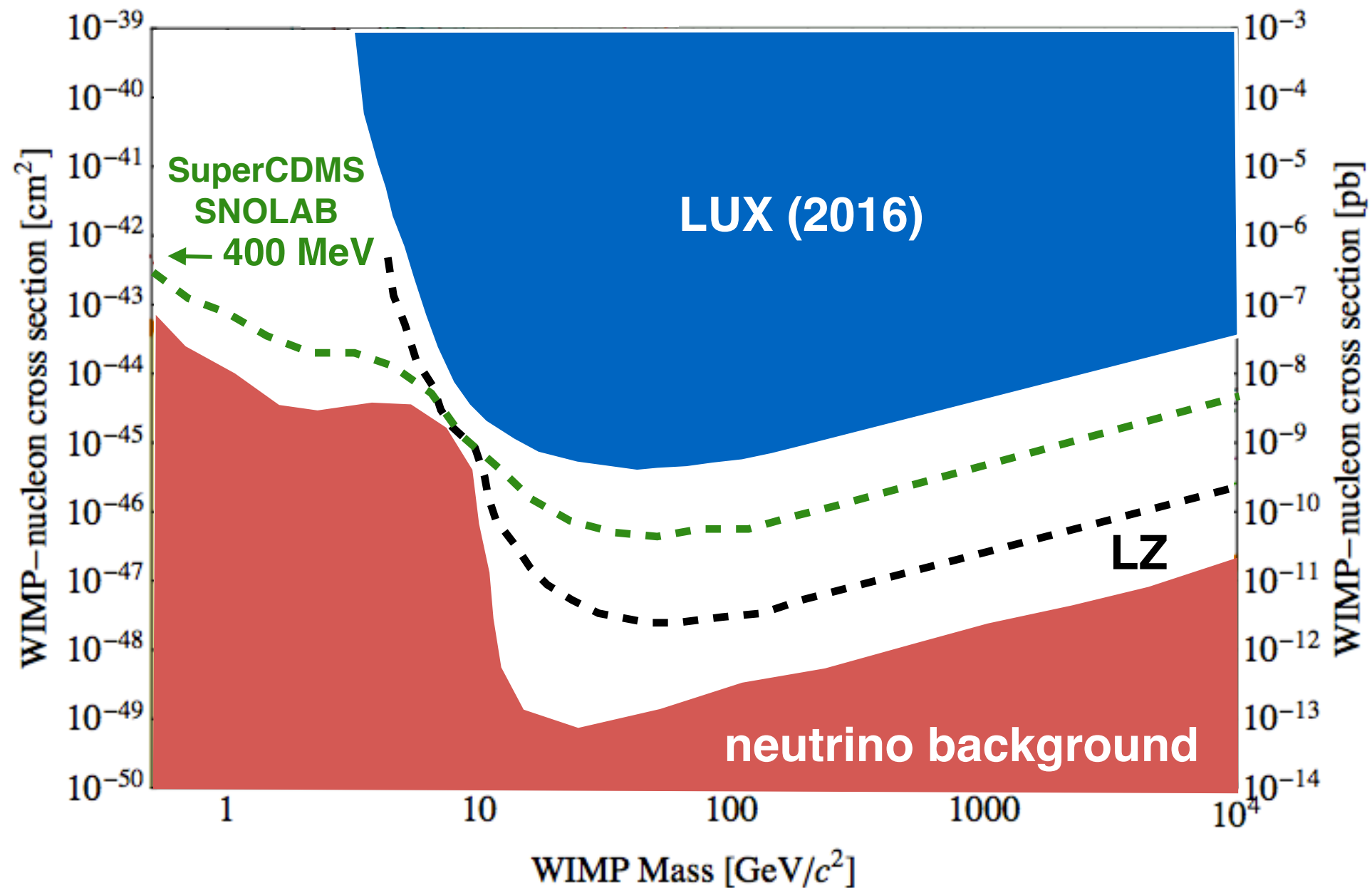


I will focus here

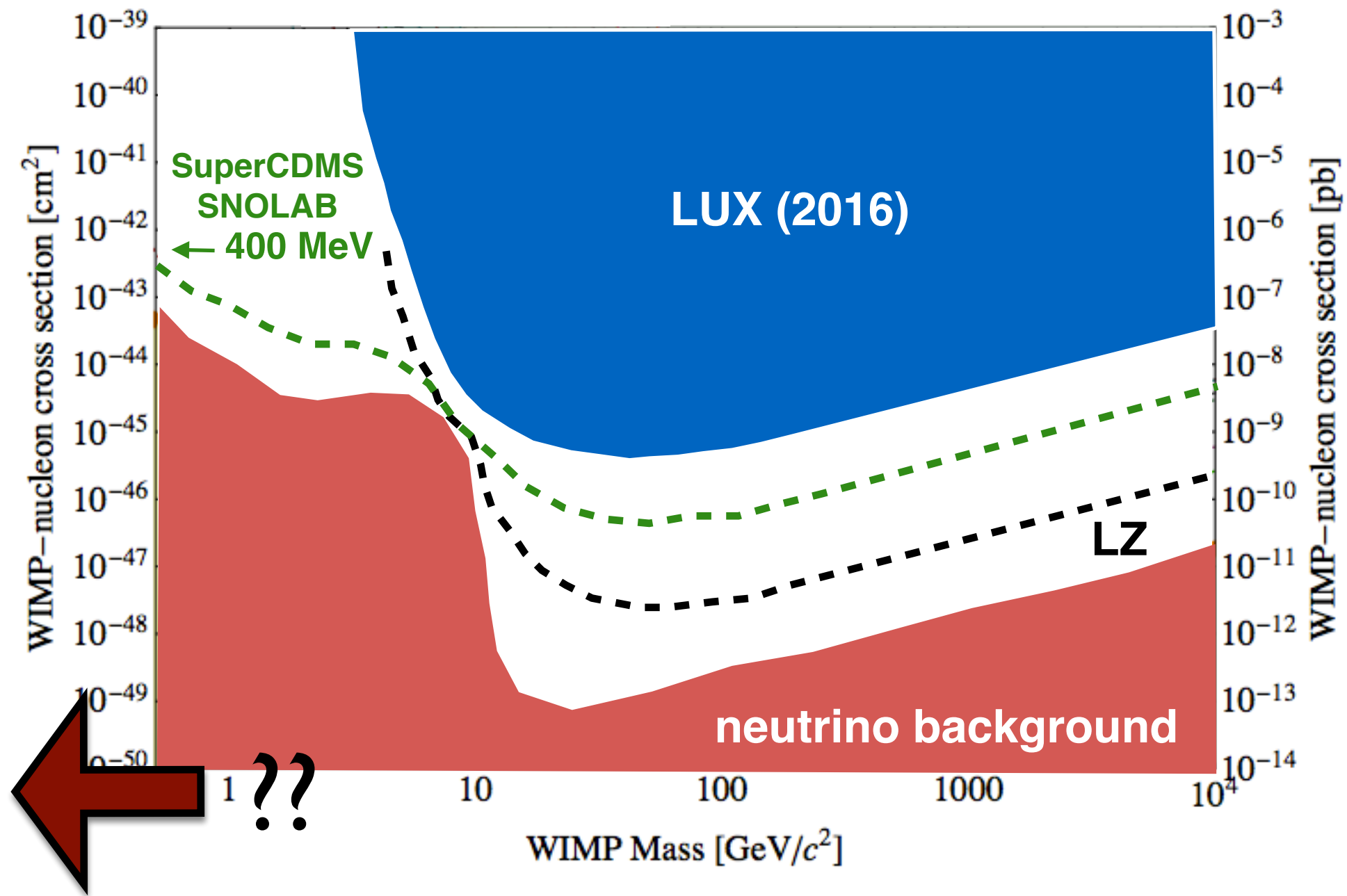
DM direct detection

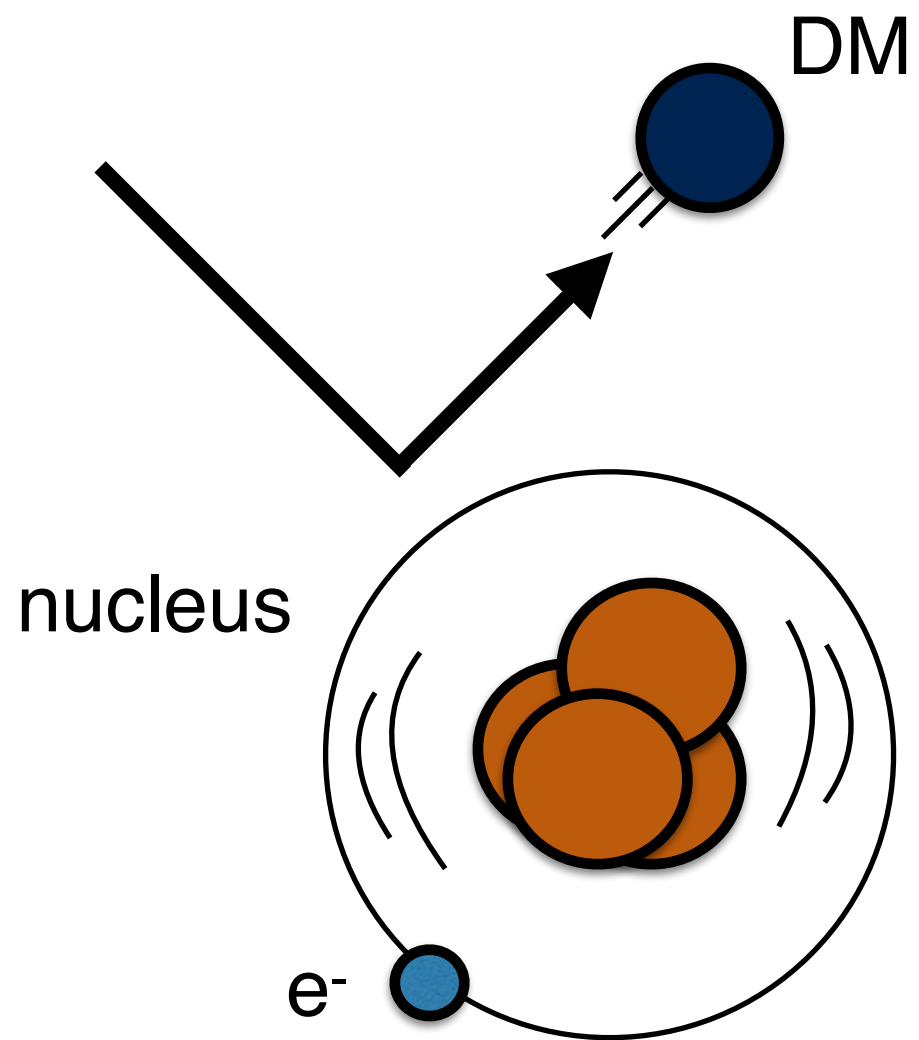


direct detection



direct detection

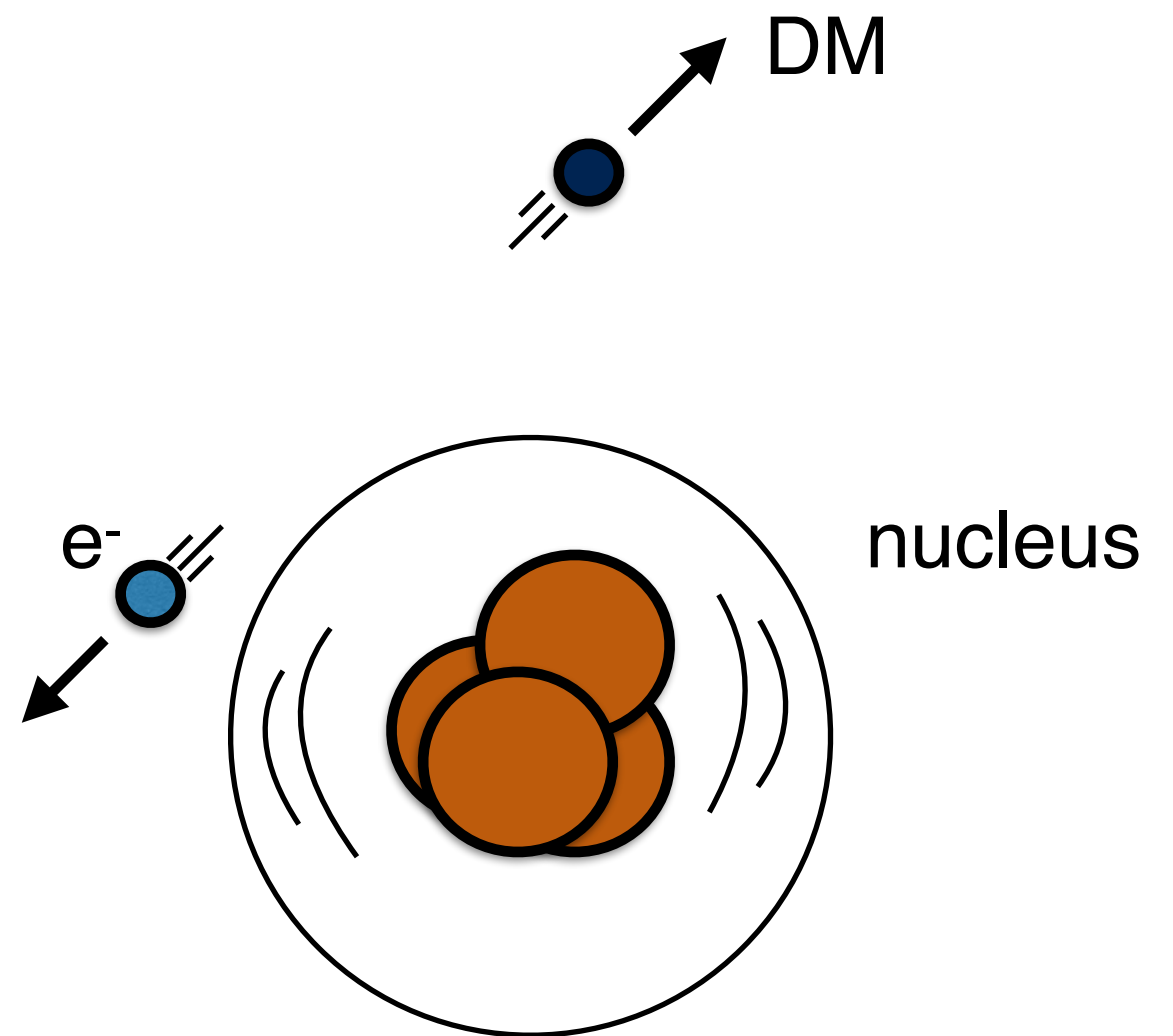




$$E_R = \frac{q^2}{2m_N}$$

$$\simeq 50 \text{ keV} \left(\frac{m_\chi}{100 \text{ GeV}} \right)^2 \left(\frac{100 \text{ GeV}}{m_N} \right)$$

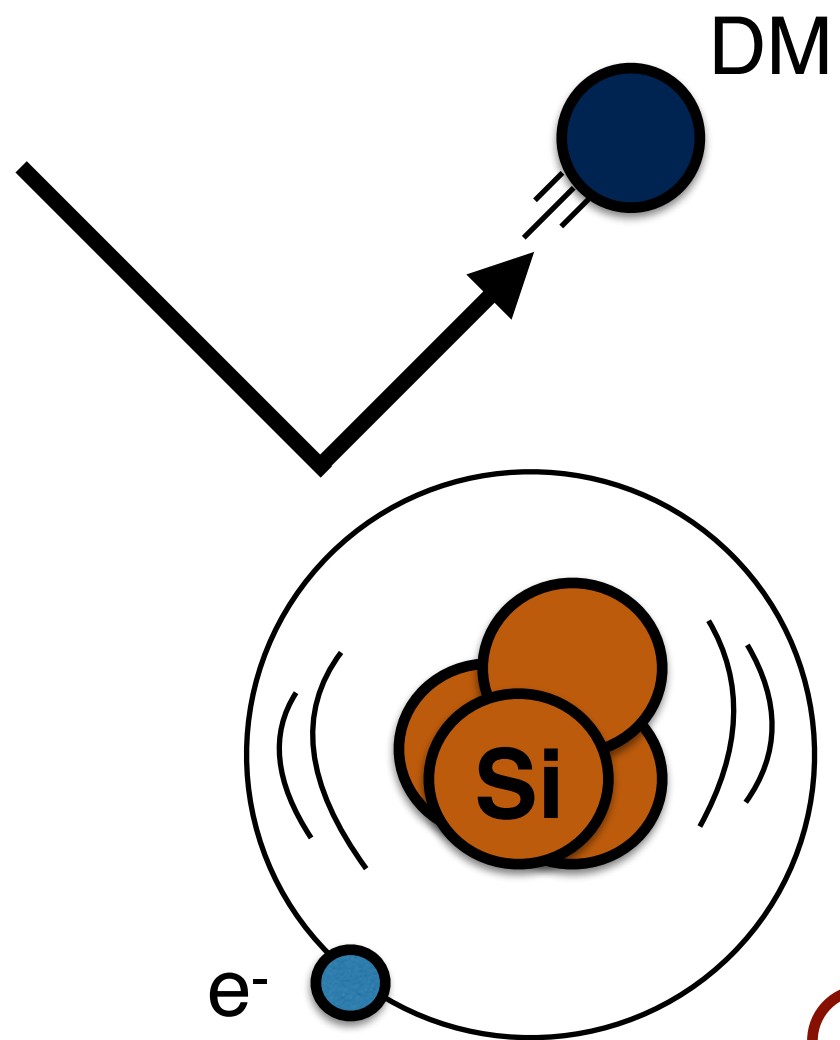
DM-nuclear scattering



$$E_R = \vec{q} \cdot \vec{v} - \frac{q^2}{2\mu_{\chi N}}$$

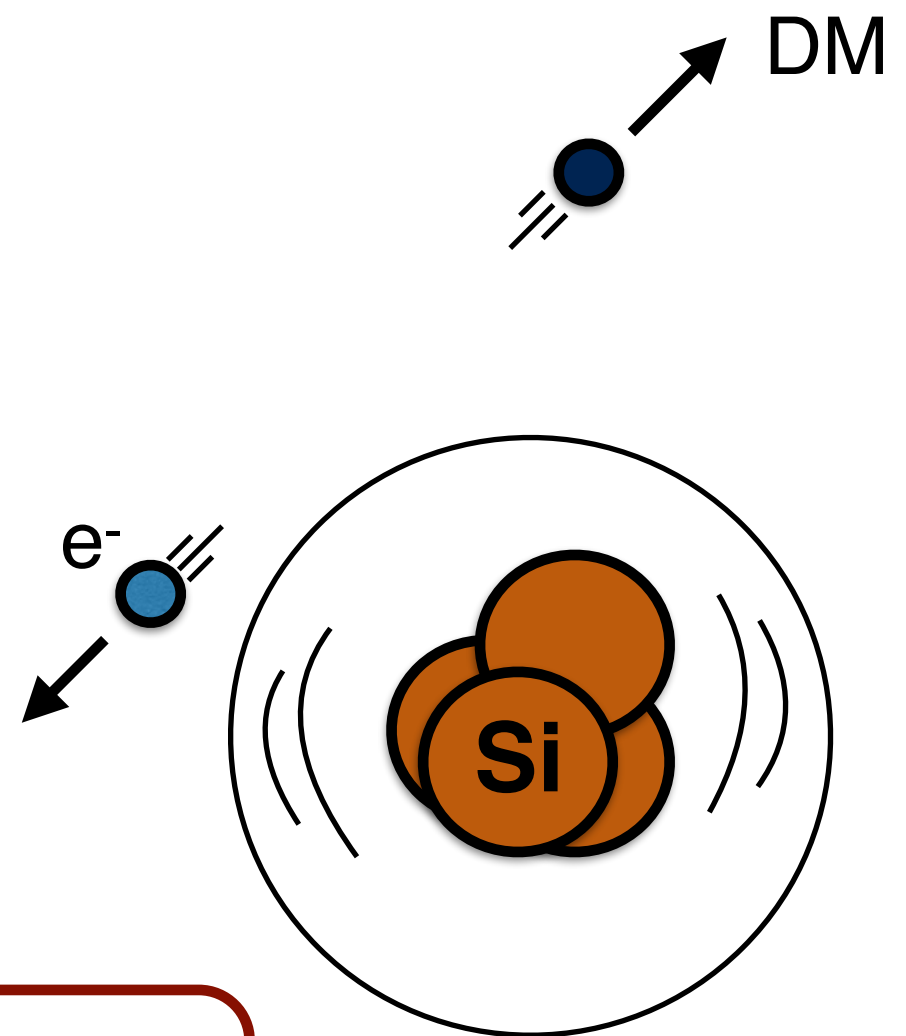
$$\sim \frac{1}{2} \text{ eV} \times \left(\frac{m_\chi}{\text{MeV}} \right)$$

DM-electron scattering



$$E_R \sim 0.1 \text{ eV}$$

DM-nuclear scattering



$$E_R \sim 50 \text{ eV}$$

DM-electron scattering

$$m_N = 28 \text{ GeV}$$

$$m_\chi = 100 \text{ MeV}$$

calculation

ingredients for rate

astrophysics

$$\frac{d\langle\sigma v\rangle}{d\ln E_R} = \frac{\bar{\sigma}_e}{8\mu_{\chi e}^2} \int q \, dq |f(k, q)|^2 |F_{DM}(q)|^2 \eta(v_{min})$$

$$\eta(v_{min}) = \int_{v_{min}} \frac{d^3v}{v} f_{MB}(\vec{v})$$

Maxwell-Boltzmann

$$v_{min} = \frac{E_R + E_B}{q} + \frac{q}{2m_\chi}$$

ingredients for rate

particle physics

$$\frac{d\langle\sigma v\rangle}{d\ln E_R} = \frac{\bar{\sigma}_e}{8\mu_{\chi e}^2} \int q \, dq |f(k, q)|^2 |F_{DM}(q)|^2 \eta(v_{min})$$

$$\bar{\sigma}_e = \frac{\mu_{\chi e}^2}{16\pi m_\chi^2 m_e^2} \overline{|\mathcal{M}_{\chi e}(q)|^2}_{q^2=\alpha^2 m_e^2}$$

$$F_{DM}(q) \simeq \begin{cases} 1 & \text{heavy mediator} \\ \frac{\alpha m_e}{q} & \text{electric dipole moment} \\ \frac{\alpha^2 m_e^2}{q^2} & \text{light mediator} \end{cases}$$

ingredients for rate

material dependent

$$\frac{d\langle\sigma v\rangle}{d\ln E_R} = \frac{\bar{\sigma}_e}{8\mu_{\chi e}^2} \int q \, dq |f(k, q)|^2 |F_{DM}(q)|^2 \eta(v_{min})$$

$$|f(k, q)|^2 = \left| \int d^3x \psi_f^*(\vec{x}) \psi_i(\vec{x}) e^{i\vec{q}\cdot\vec{x}} \right|^2$$

wave-function overlap between initial
and final electron states

probability of going from state i to i'

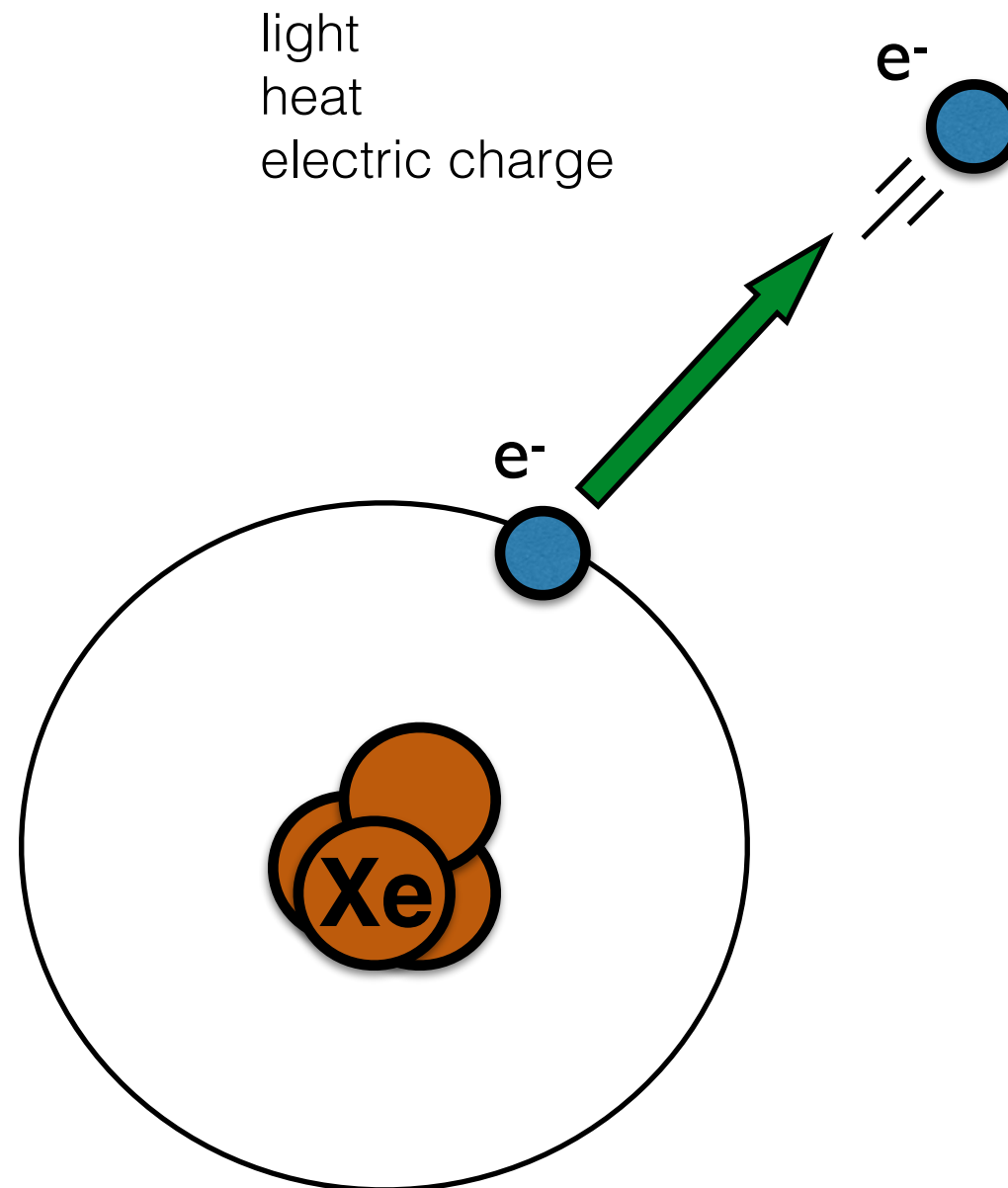
scattering rate

$$\frac{d\langle\sigma v\rangle}{d\ln E_R} = \frac{\overline{\sigma}_e}{8\mu_{\chi e}^2} \int q \, dq \underbrace{|f(k, q)|^2}_{\text{particle physics}} \underbrace{|F_{DM}(q)|^2}_{\text{astrophysics}} \underbrace{\eta(v_{min})}_{\text{detector}}$$

$$R = N_T \frac{\rho_\chi}{m_\chi} \int_{E_{R,cut}} d\ln E_R \frac{d\langle\sigma v\rangle}{d\ln E_R}$$

R = number of events/time/volume

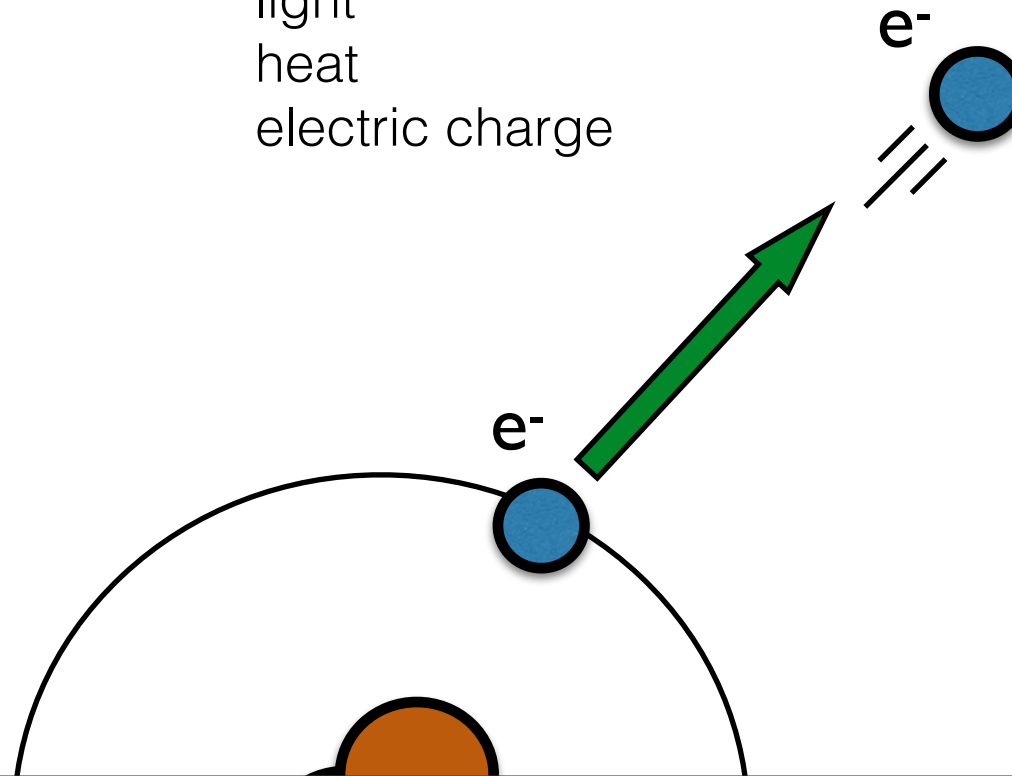
**liquid
xenon**



$$|f(k, q)|^2 = \left| \int d^3x \psi_f^*(\vec{x}) \psi_i(\vec{x}) e^{i\vec{q} \cdot \vec{x}} \right|^2$$

**liquid
xenon**

light
heat
electric charge

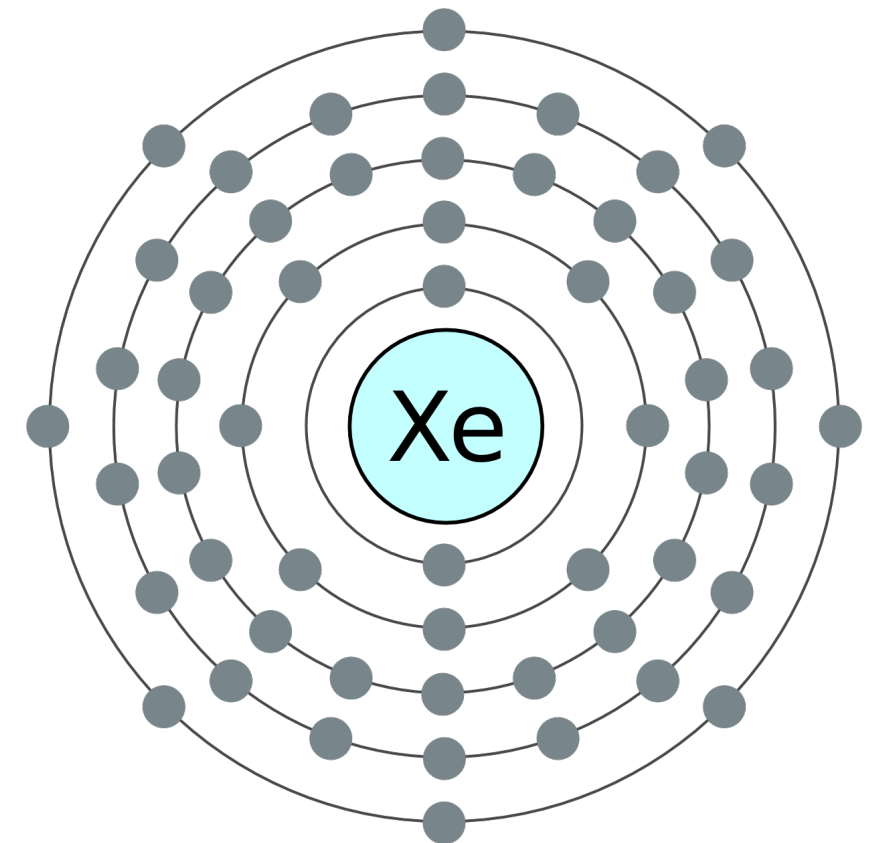
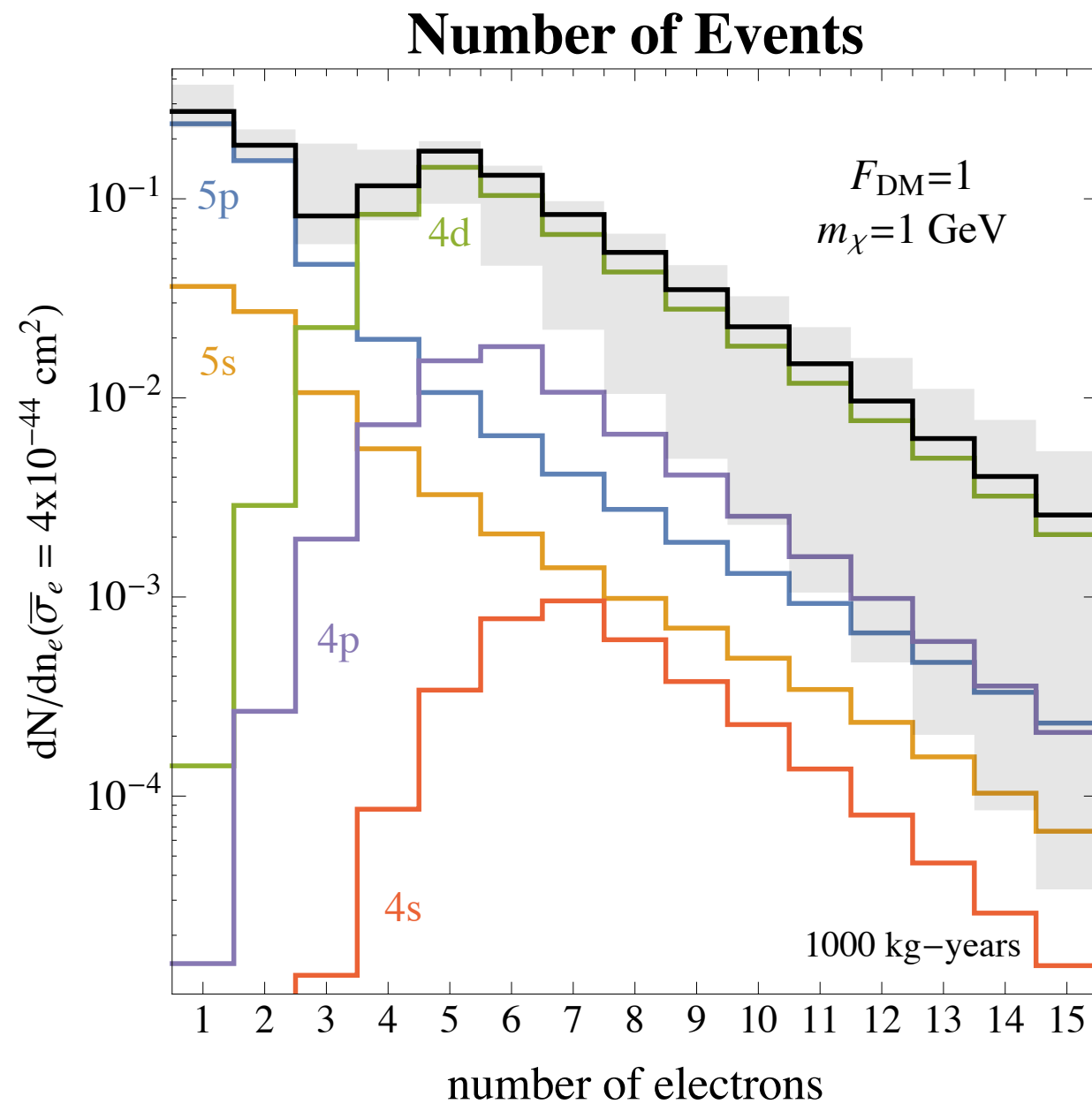


**treat the target electrons as single-particle states
bound in isolated xenon atoms**

$$|f(k, q)|^2 = \left| \int d^3x \psi_f^*(\vec{x}) \psi_i(\vec{x}) e^{i\vec{q} \cdot \vec{x}} \right|^2$$

XENON

54: Xenon



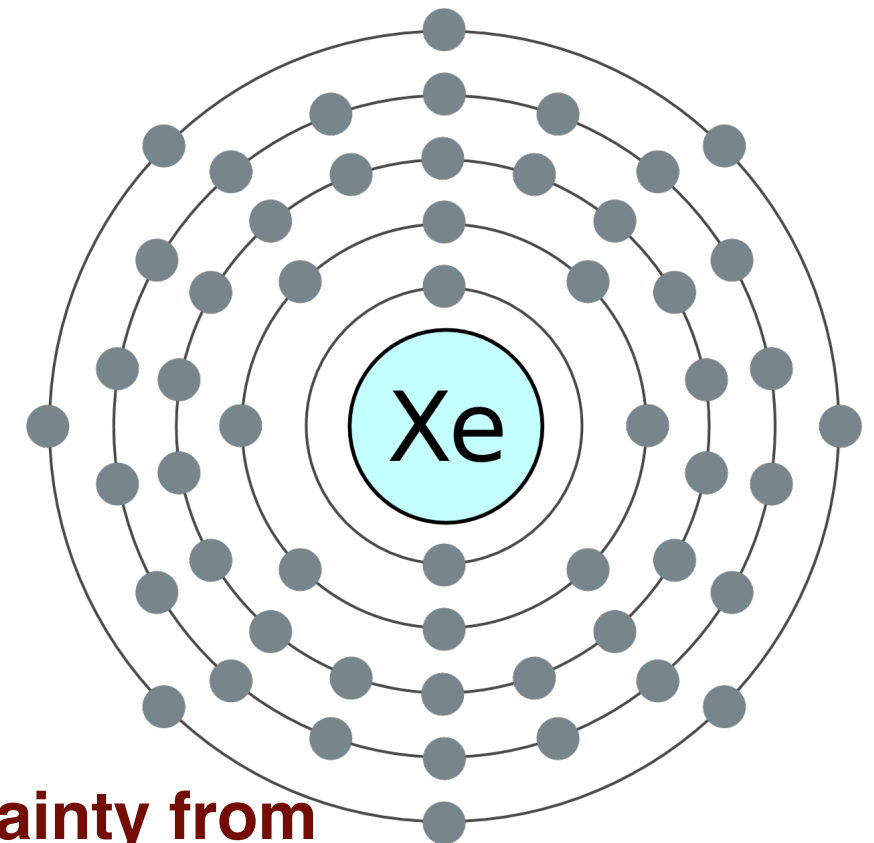
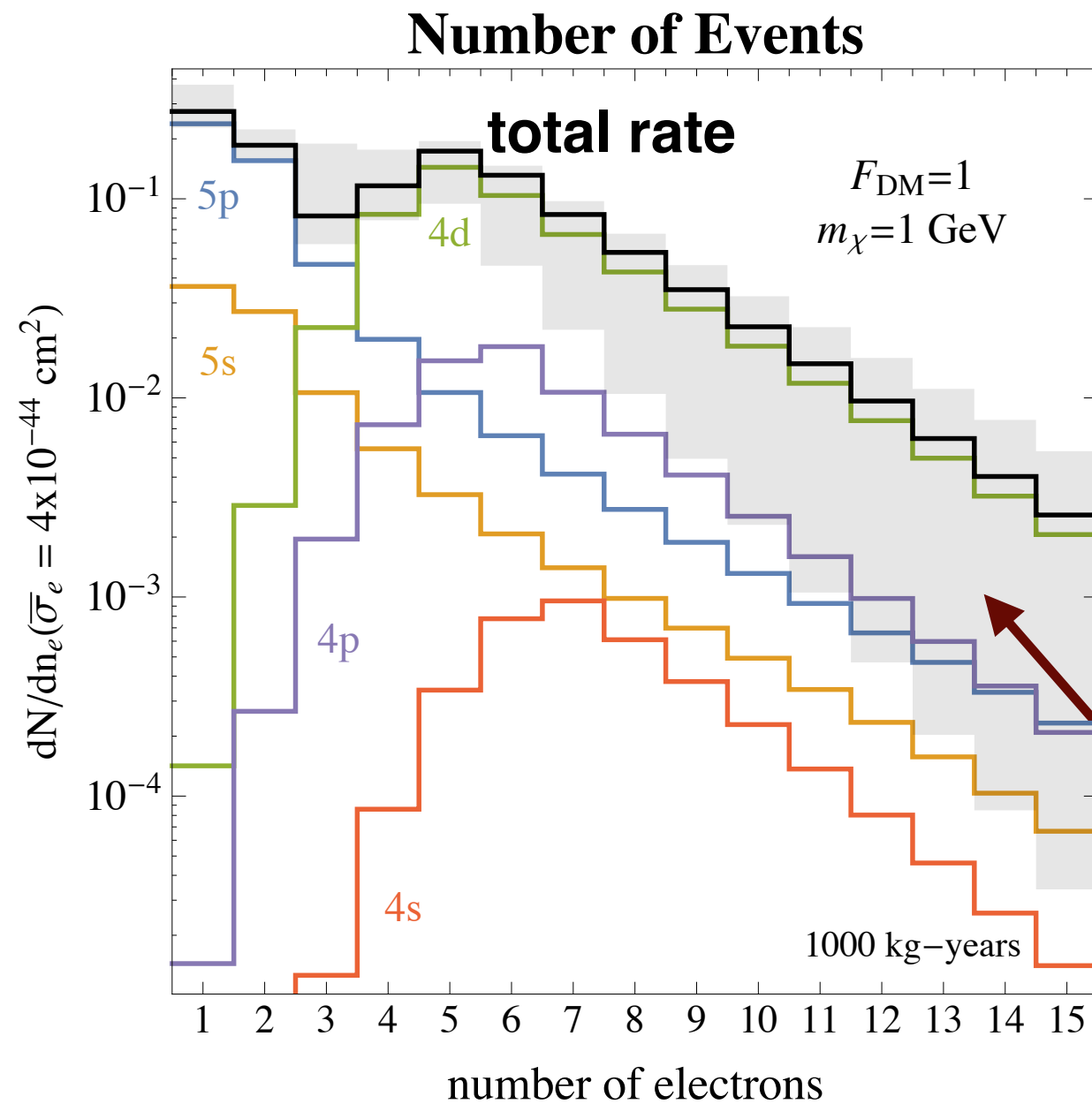
http://commons.wikimedia.org/wiki/Category:Electron_shell_diagrams

Essig, Volansky, TTY [1703.00910]

Electron configuration:
[Ar] 3d¹⁰ 4s² 4p⁶ 4d¹⁰ 5s² 5p⁶

XENON

54: Xenon

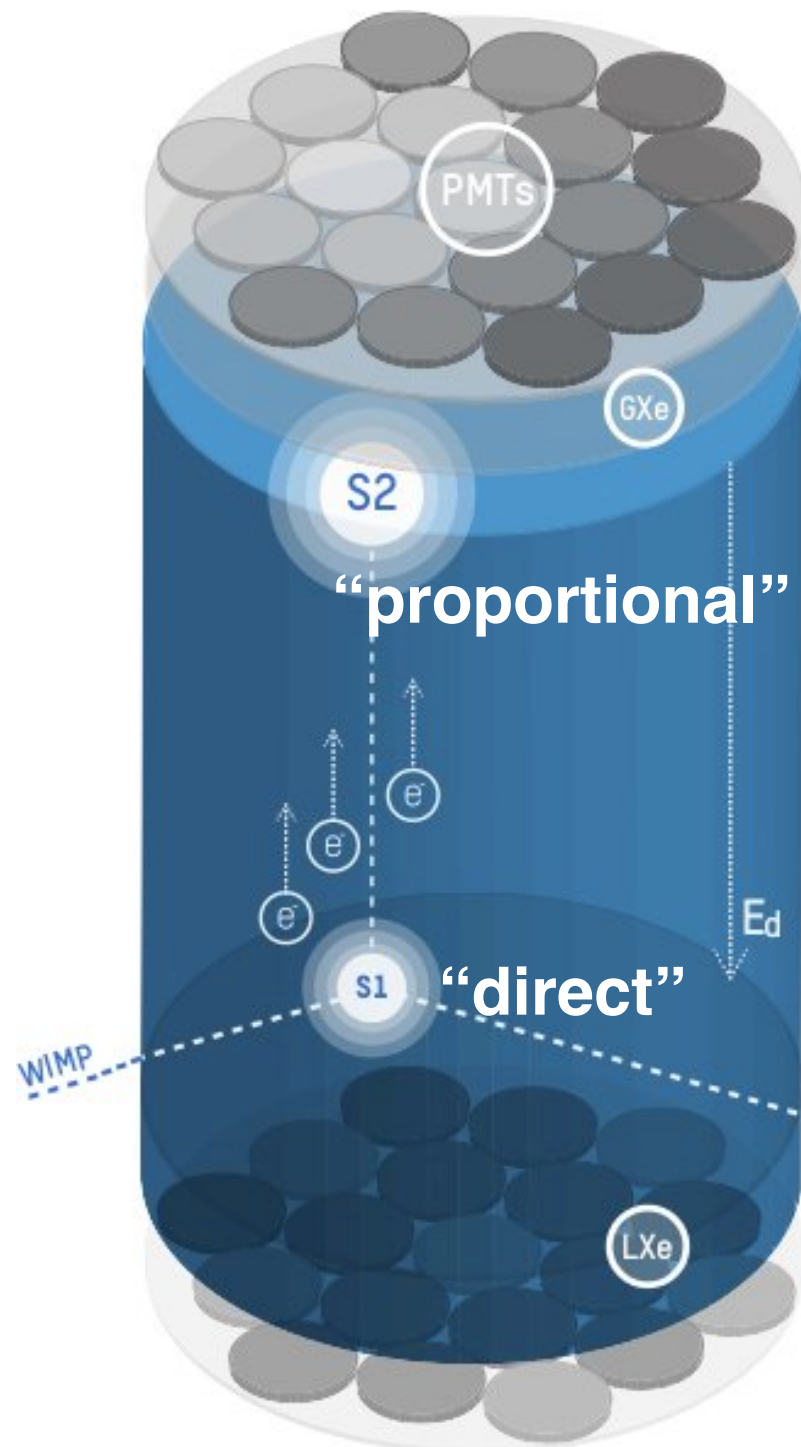


**uncertainty from
secondary ionization
modeling**

Essig, Volansky, TTY [1703.00910]

Electron configuration:
[Ar] 3d10 4s2 4p6 4d10 5s2 5p6

a XENON detector



i.e. XENON10, XENON100, XENON1T, LUX

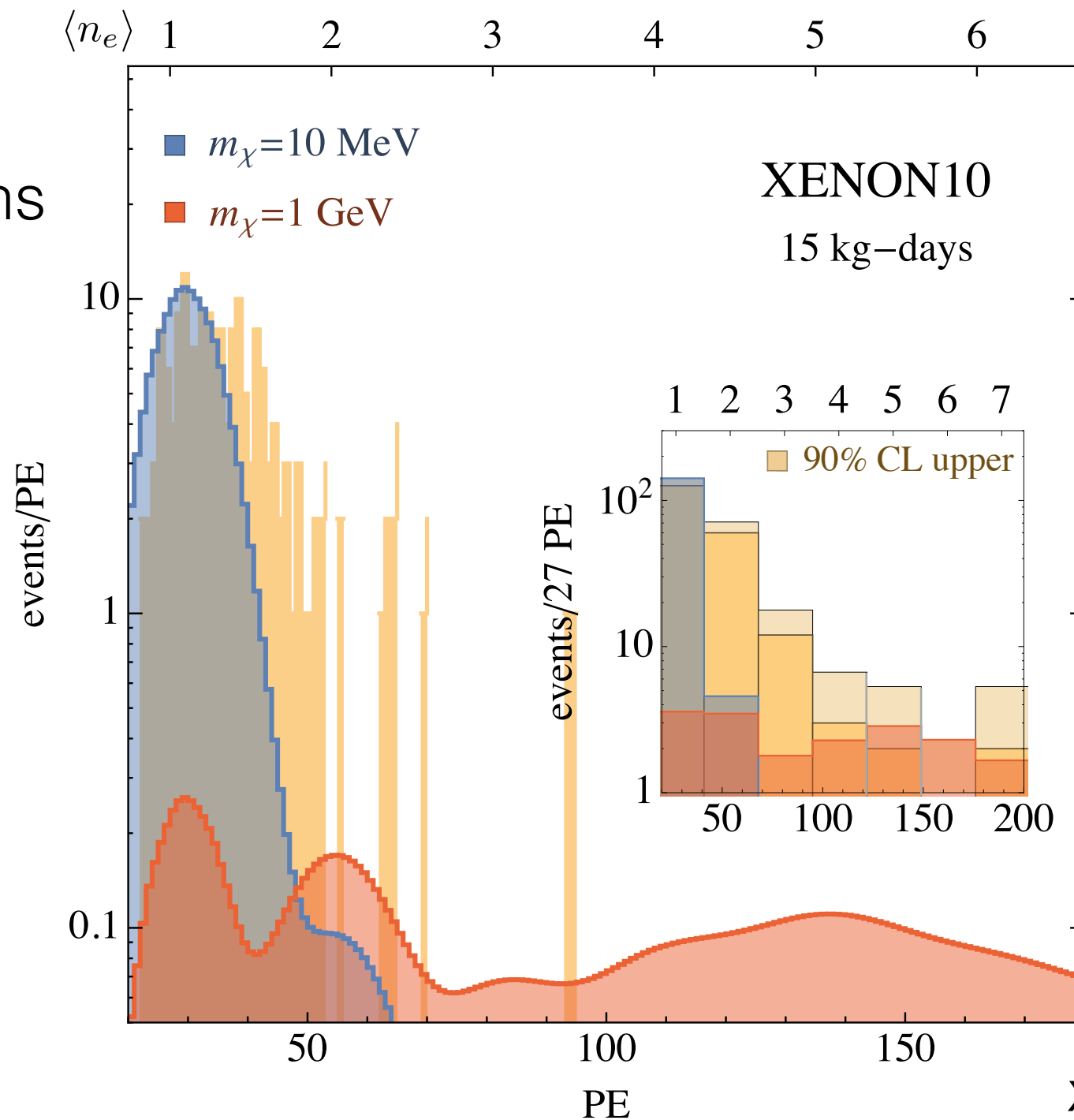
DM-electron scattering
=
S2 only signal

measures **Photo**Electrons

*can also do this with LAr detectors like DarkSide

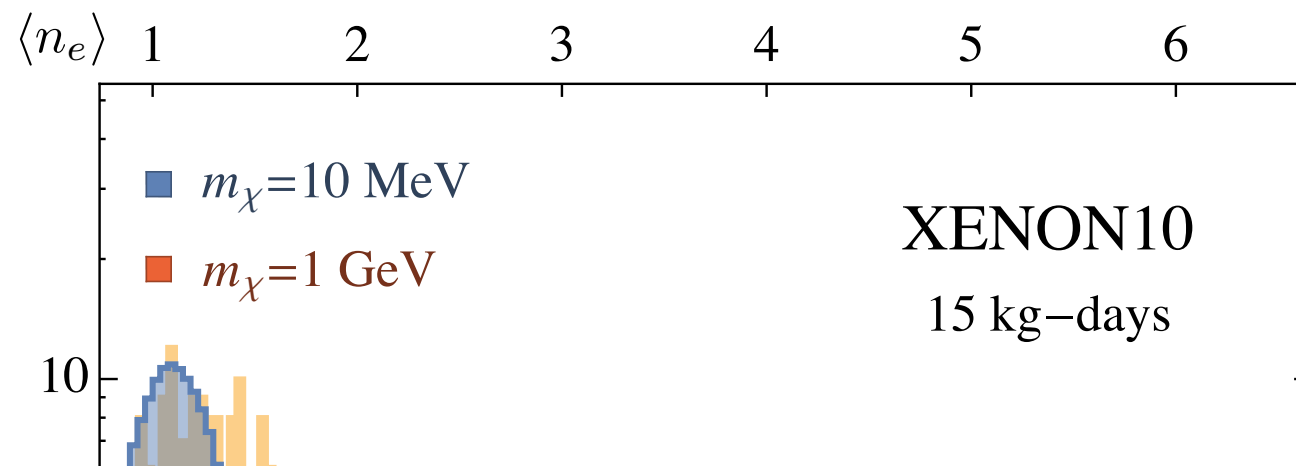
XENON10

1 electron =
27 **P**hoto**E**lectrons

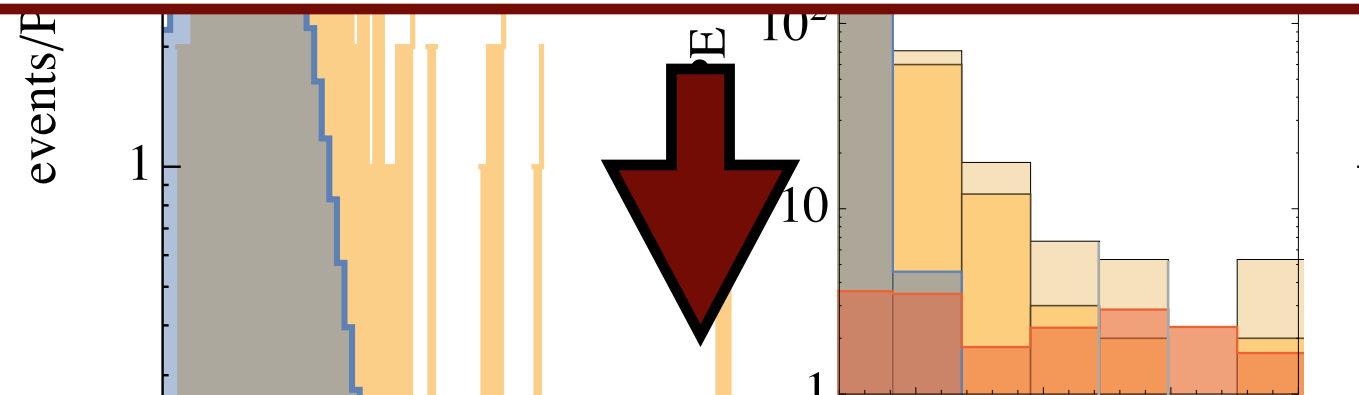


XENON10

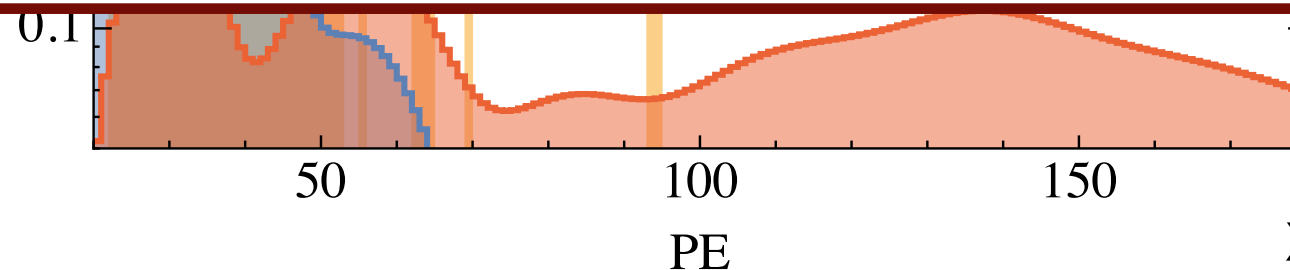
1 electron =
27 **P**hoto**E**lectrons



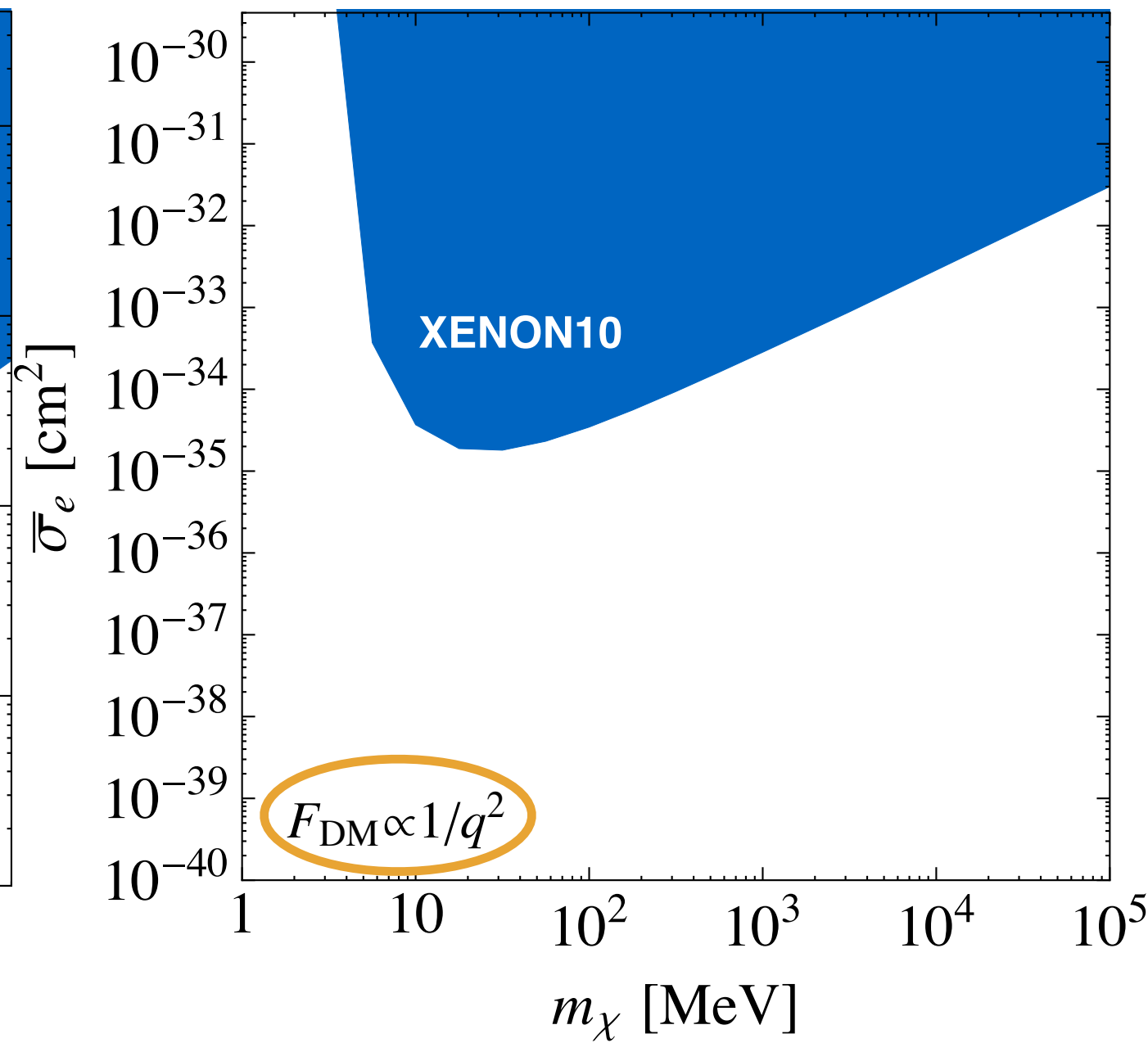
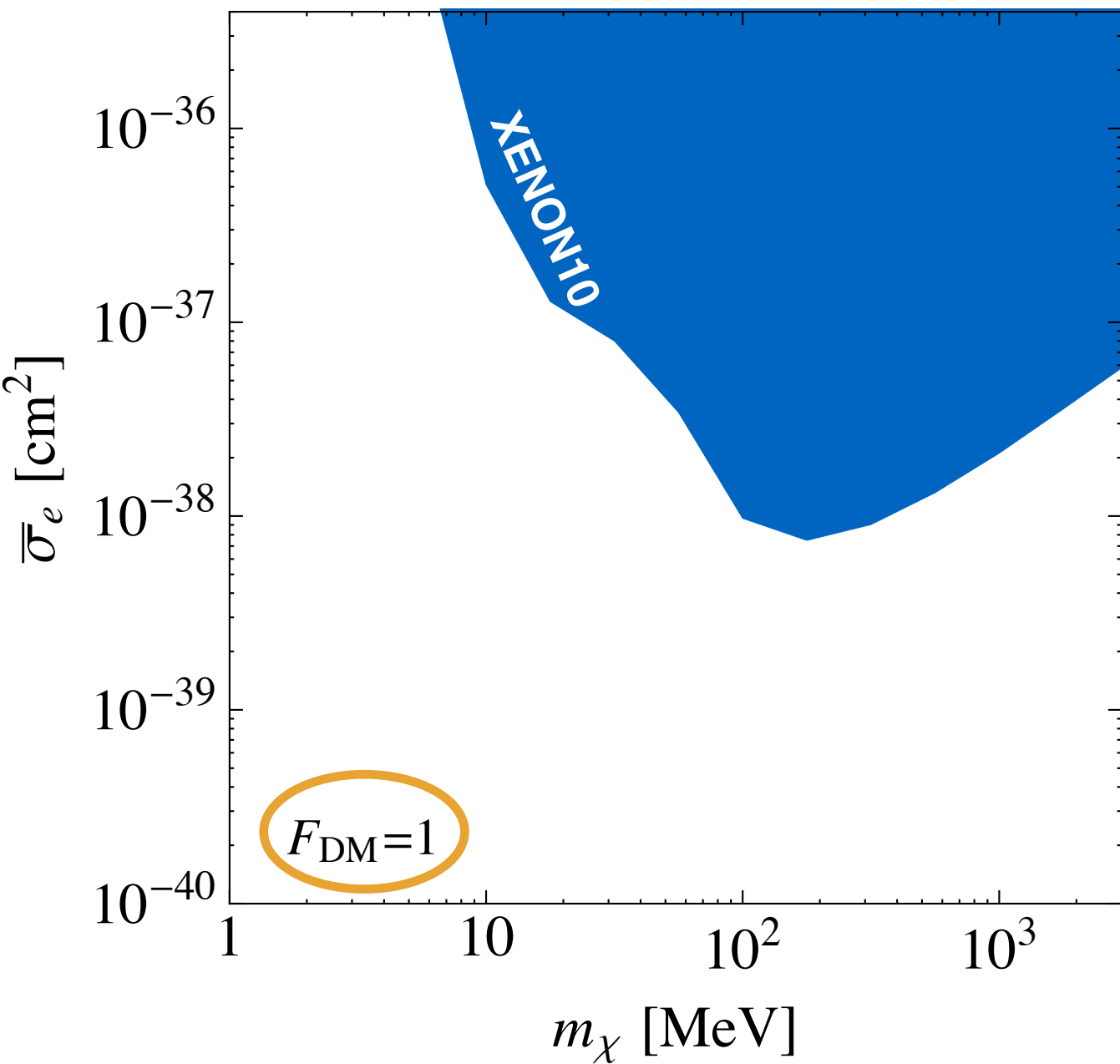
signal < Number of observed events



constrain size of DM-electron scattering cross-section

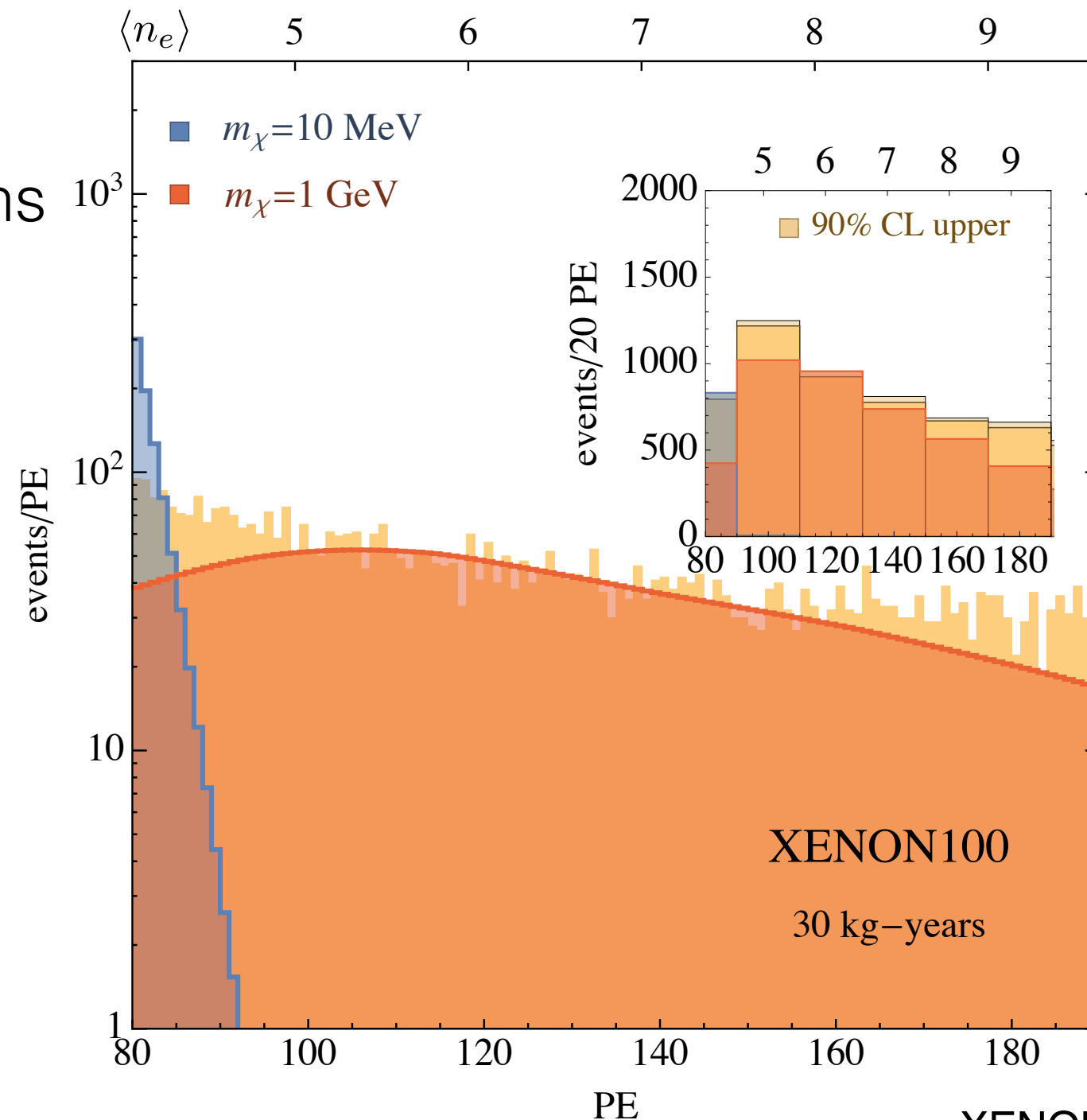


XENON

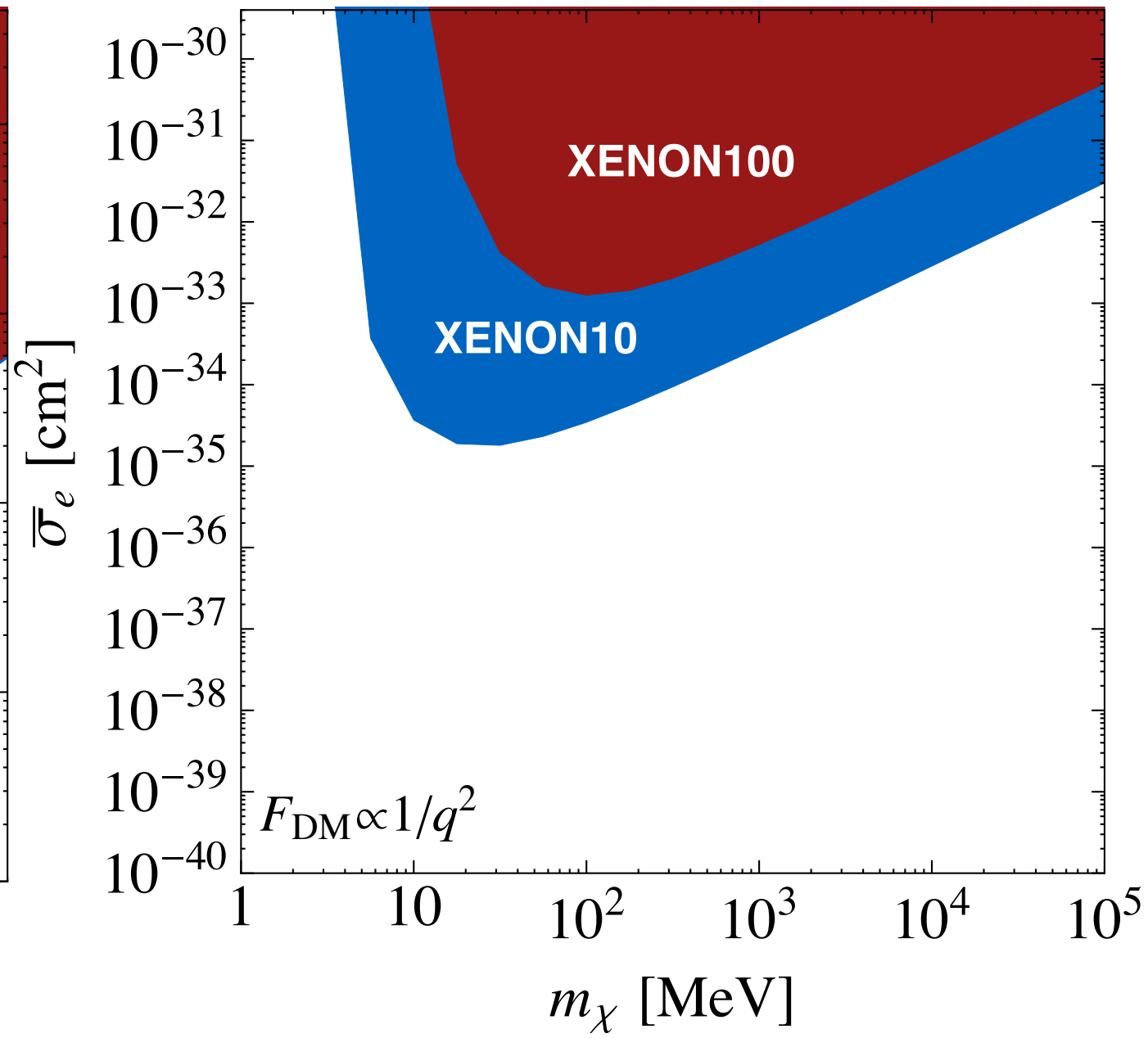
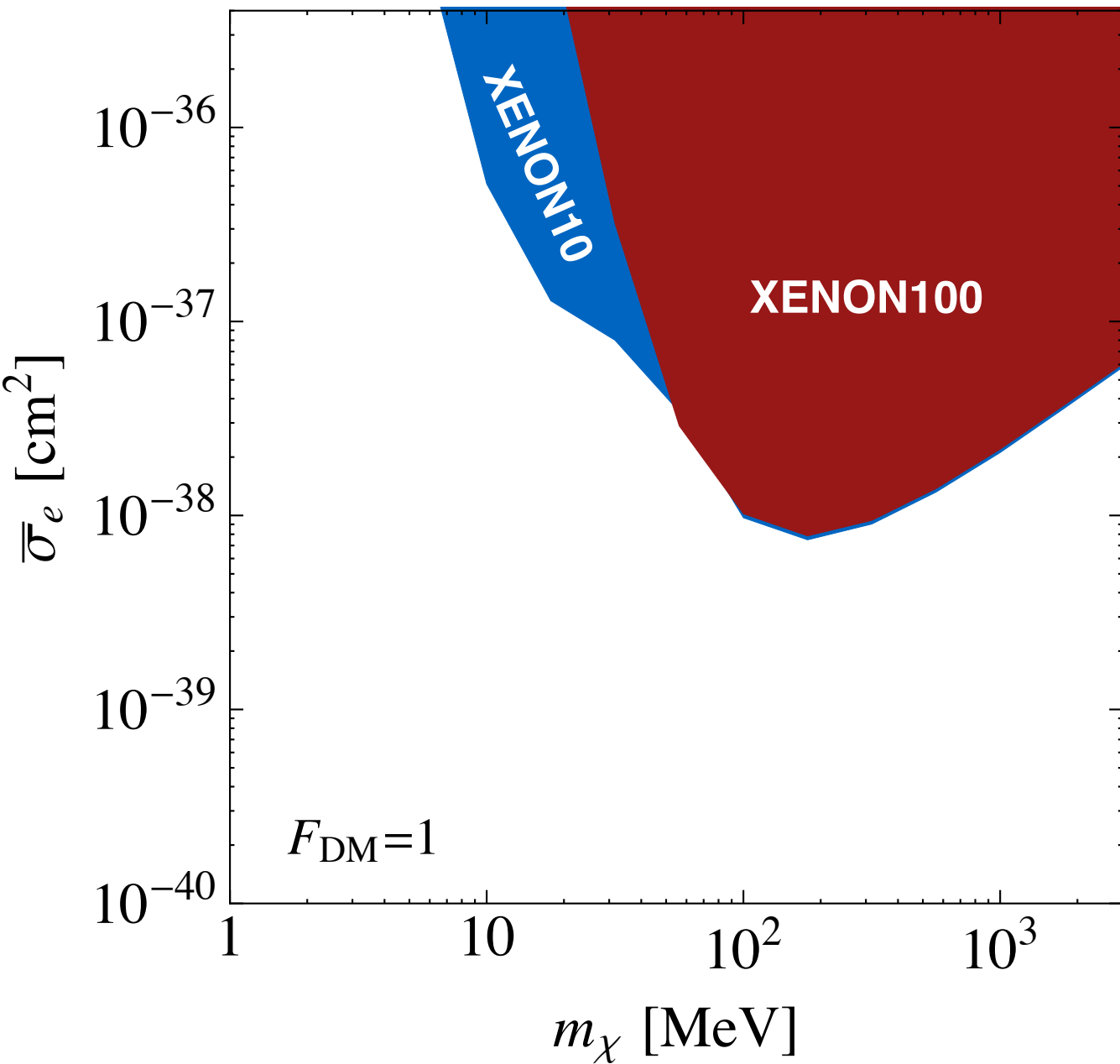


XENON100

1 electron =
20 **P**hoto**E**lectrons

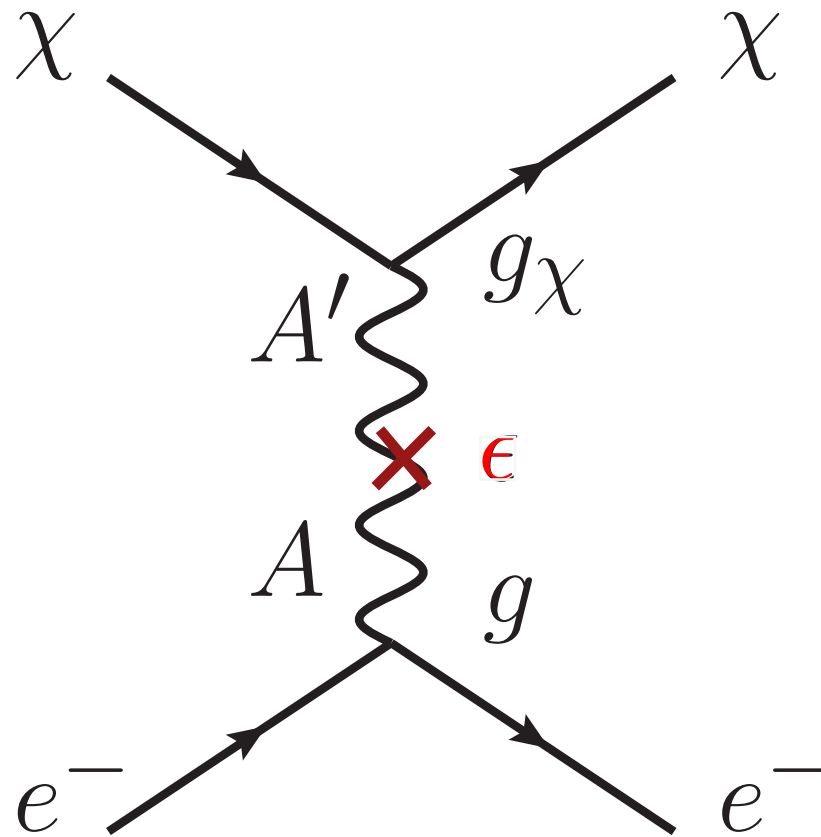


XENON



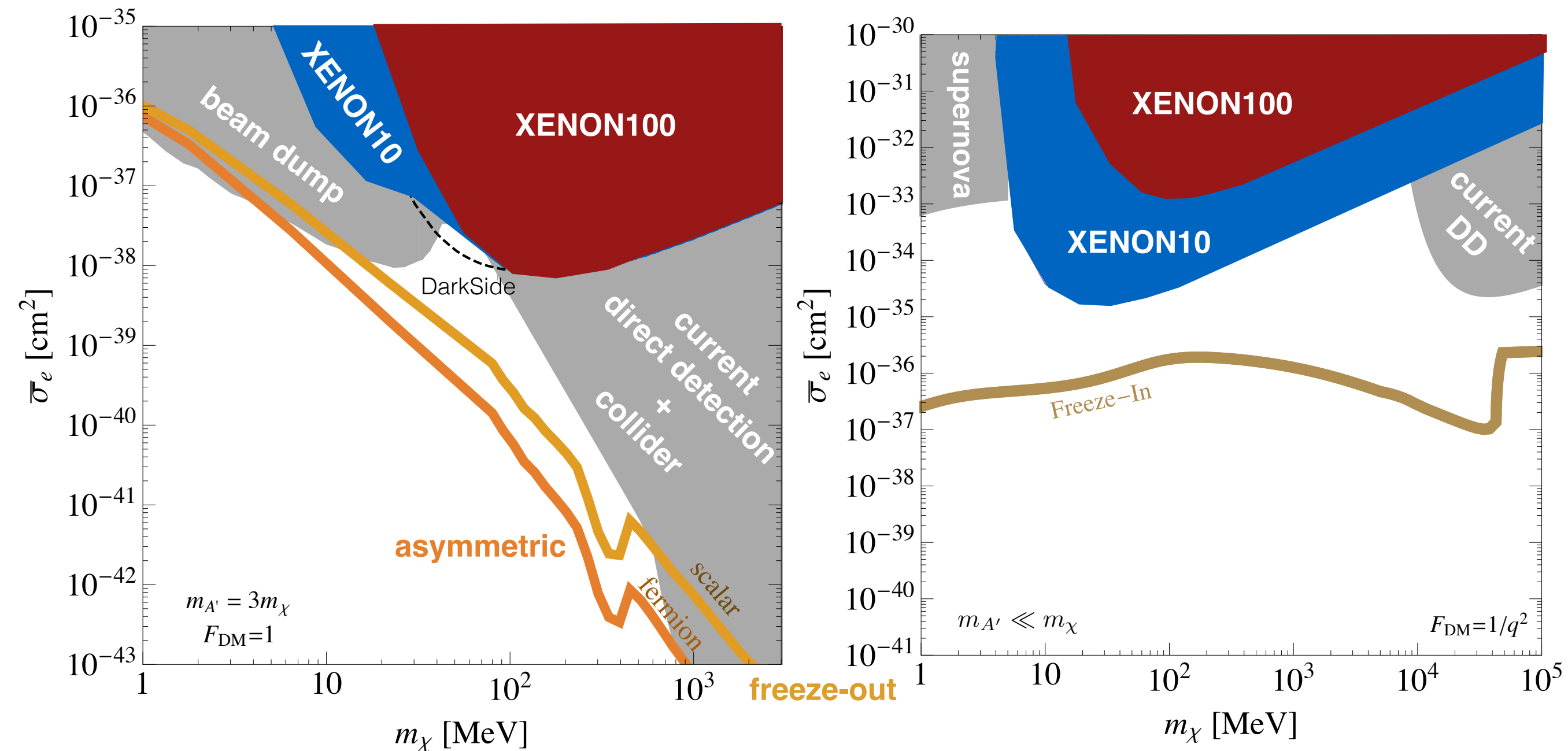
a model: dark photon

$$\mathcal{L} \supset -\frac{1}{4}F'^{\mu\nu}F'_{\mu\nu} - \frac{\epsilon}{2}F^{\mu\nu}F'_{\mu\nu} + \frac{1}{2}m_{A'}^2 A'^{\mu}A'_{\mu}$$



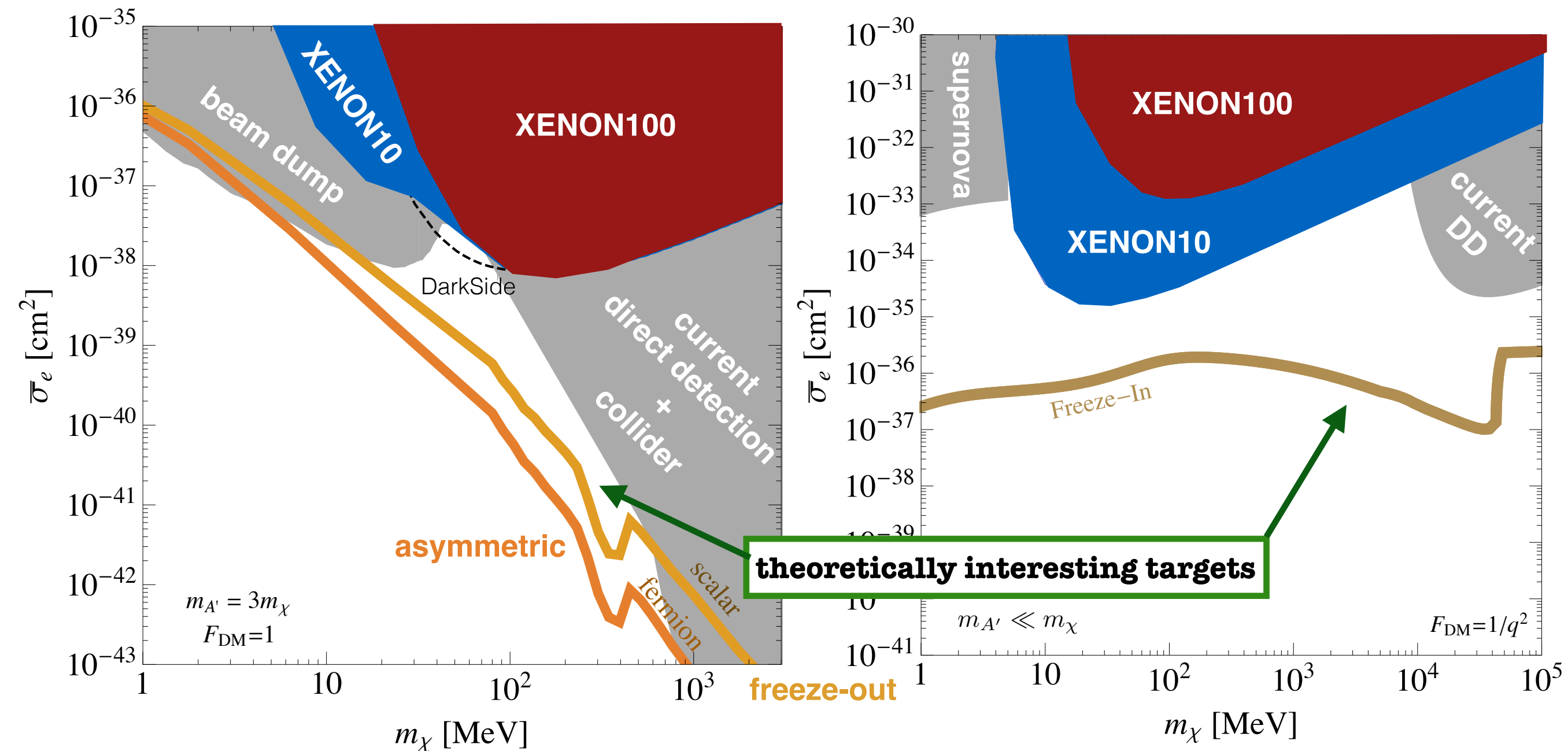
$$F_{DM}(q) = \frac{m_{A'}^2 + \alpha^2 m_e^2}{m_{A'}^2 + q^2} \simeq \begin{cases} 1, & m_{A'} \gg \alpha m_e \\ \frac{\alpha^2 m_e^2}{q^2}, & m_{A'} \ll \alpha m_e \end{cases}$$

dark photon



Essig, Volansky, TTY [1703.00910], DarkSide Collaboration [1802.06998]

dark photon



challenges

- detector specific backgrounds
 - e^- gets trapped in liquid-gas interface and is later released
- ionization energy (12.1 eV) limits DM mass reach to few MeV

challenges

- detector specific background

- e- gets ionized at liquid-gas interface and is later re-

have different detector setup

- ionization energy (12.1 eV) limits DM mass reach to few MeV

challenges

- detector specific background
- e- gets stuck in liquid-gas interface is later released
- ionization energy is a few MeV

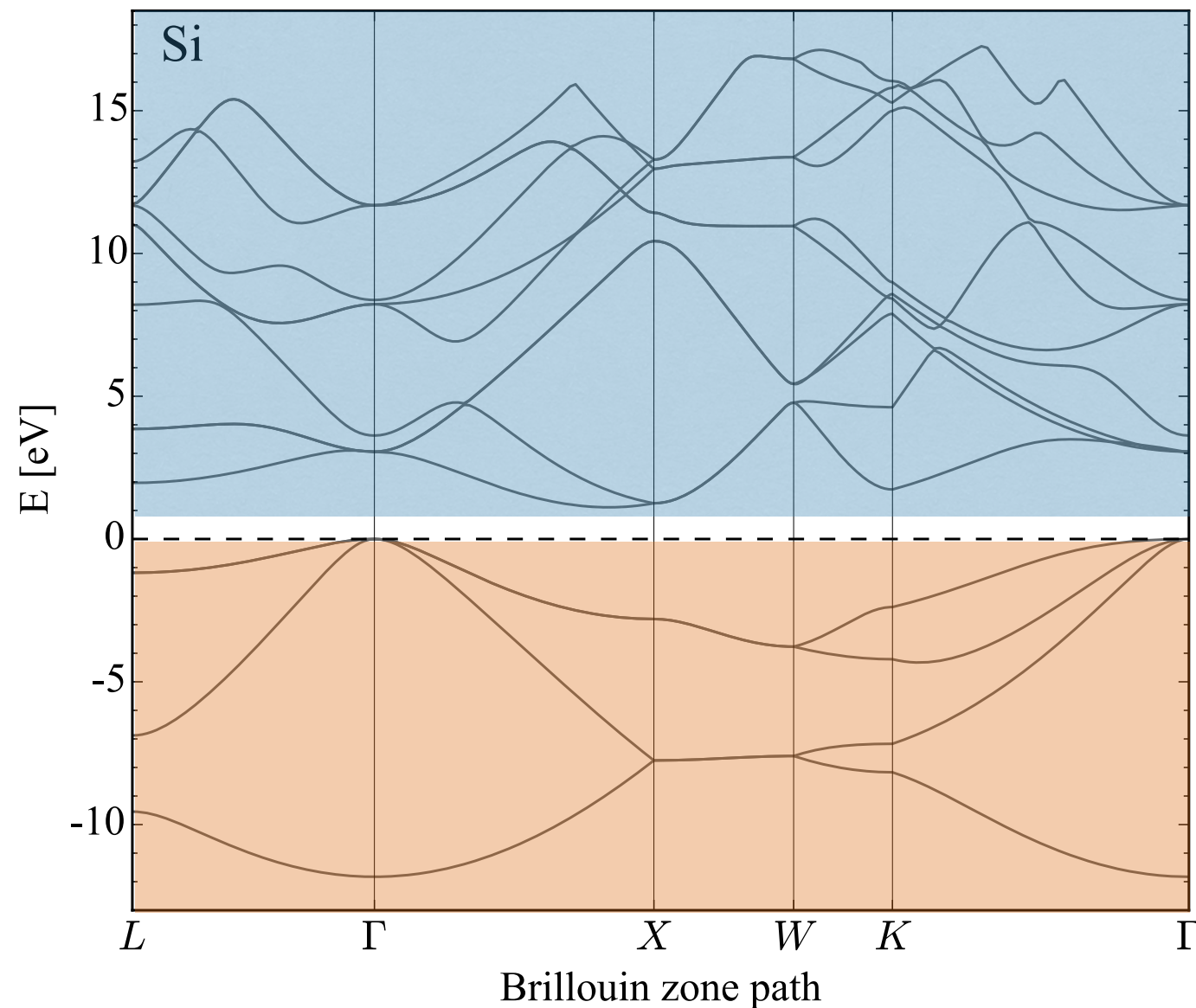
have different detector setup

find a material with smaller ionization energy

electron energy

- noble gases: $\Delta E_e \sim 10 \text{ eV}$
- semiconductors: $\Delta E_e \sim 1 \text{ eV}$

semiconductor targets



**empty conduction
band**

band gap

**filled valence
band**

Essig, Fernandez-Serra, Mardon, Soto, TTY [1509.01598] JHEP 1605 (2016) 046

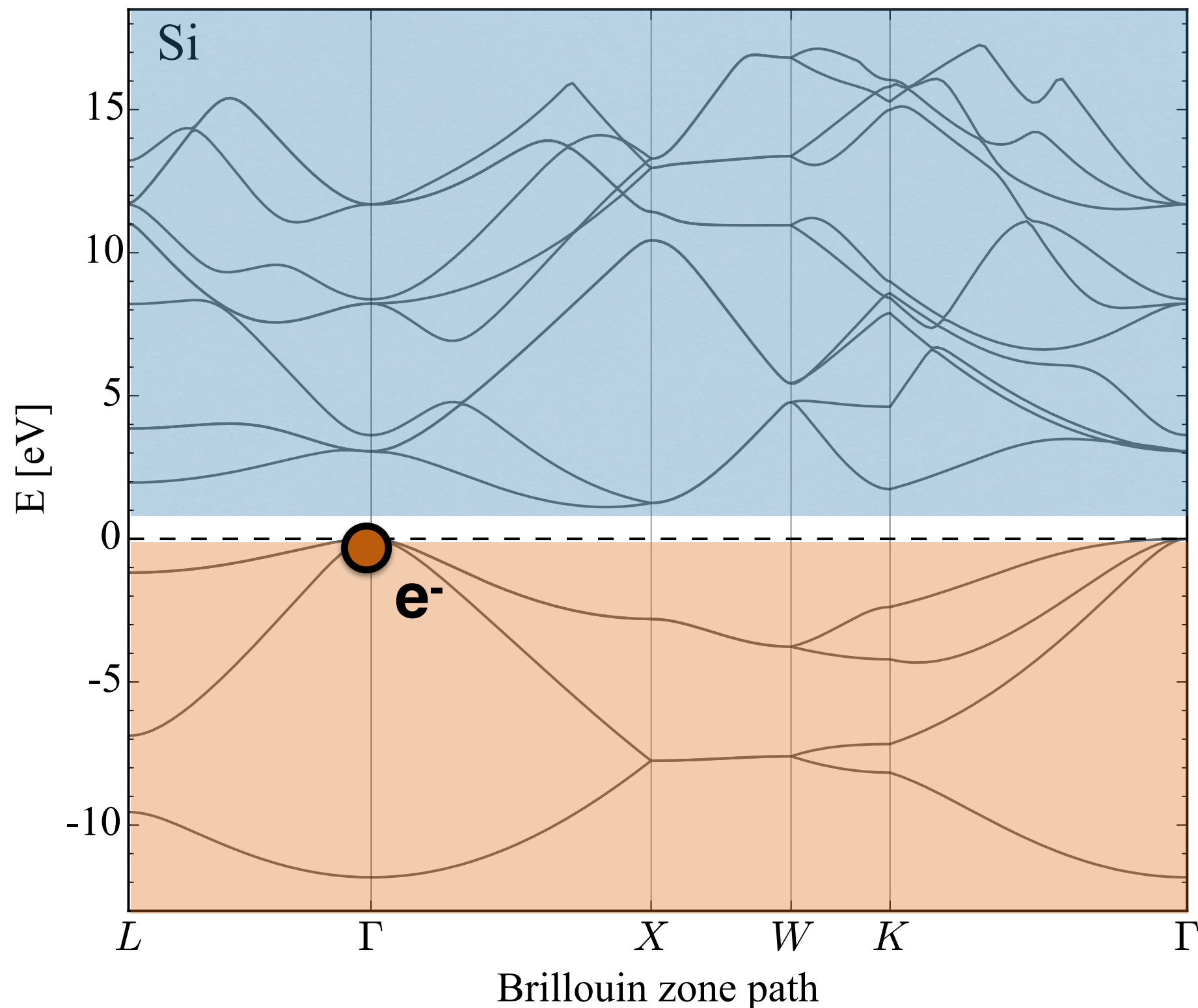
see also:

Essig, Mardon, Volansky [1108.5383] Phys.Rev. D85 (2012) 076007

Graham, Kaplan, Rajendran, Walters [1203.2531] Phys.Dark Univ. 1 (2012) 32-49

Lee, Lisanti, Mishra-Sharma, Safdi [1508.07361] Phys.Rev. D92 (2015) 083517

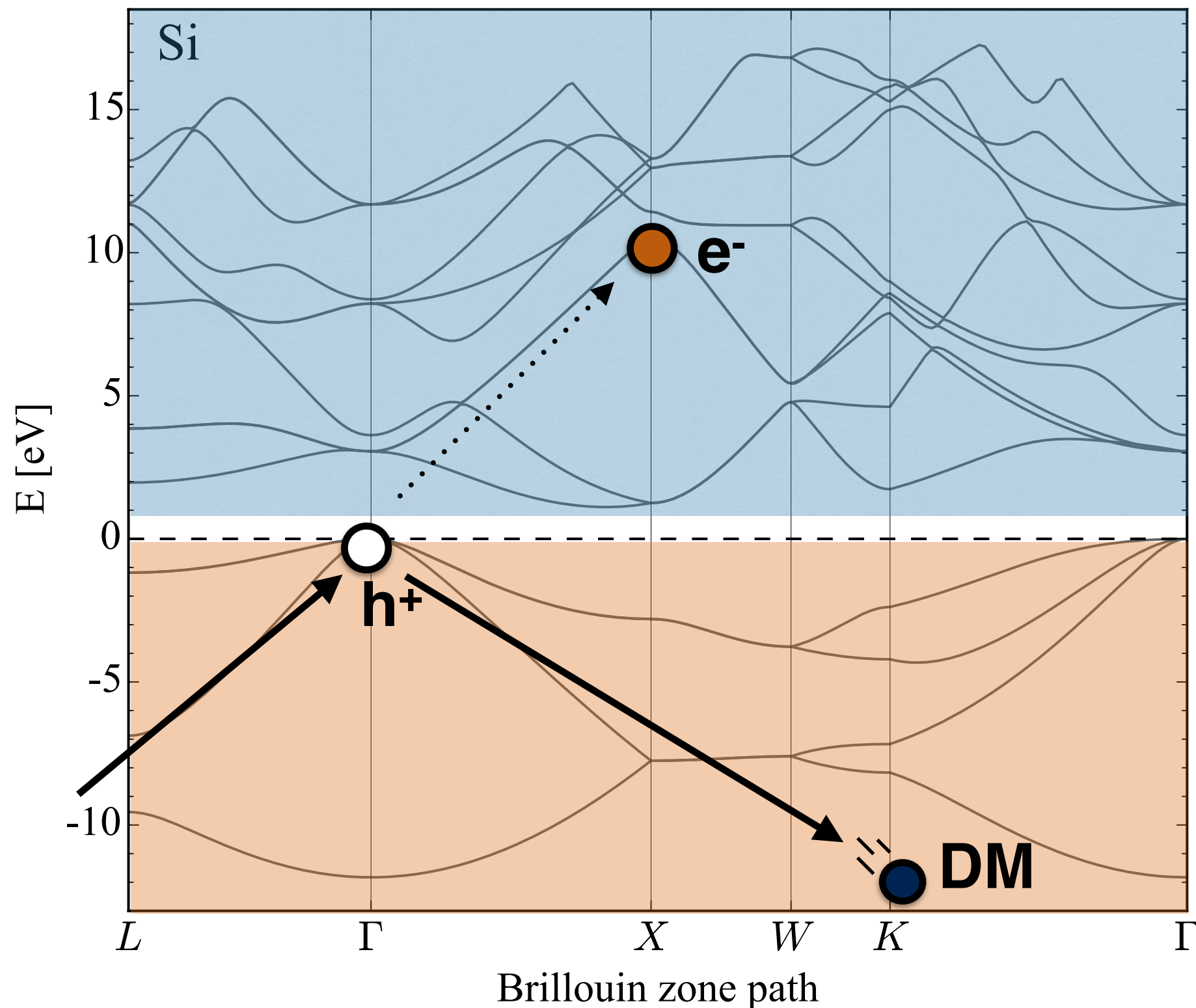
semiconductor targets



band gap [eV]

Ge	0.67
Si	1.1
GaAs	1.5
NaI	5.9
CsI	6.4

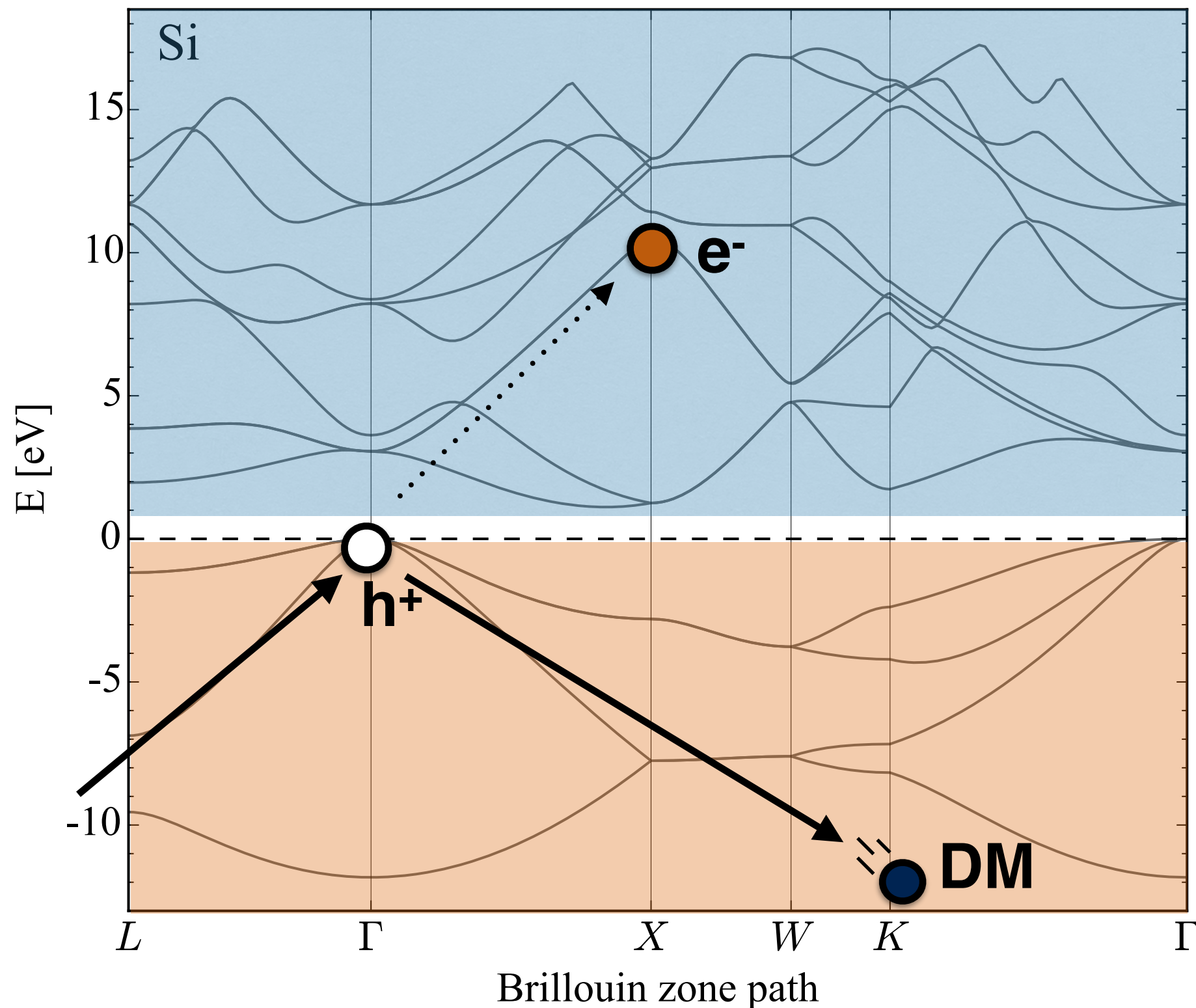
semiconductor targets



band gap [eV]

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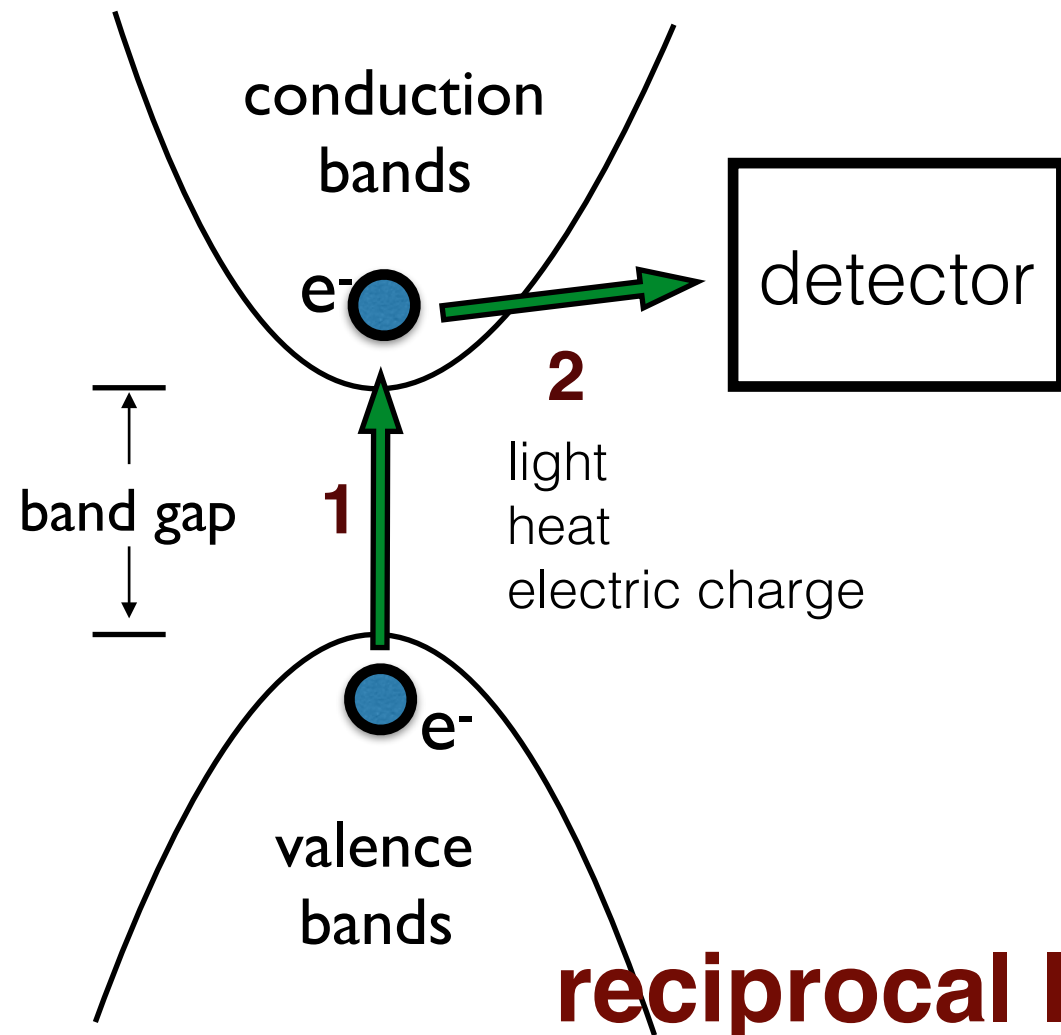
semiconductor targets



**apply an
electric field
and extract
the electron(s)
“ionization”**

**e^-/h^+ recombine
to produce
photon(s)
“scintillation”**

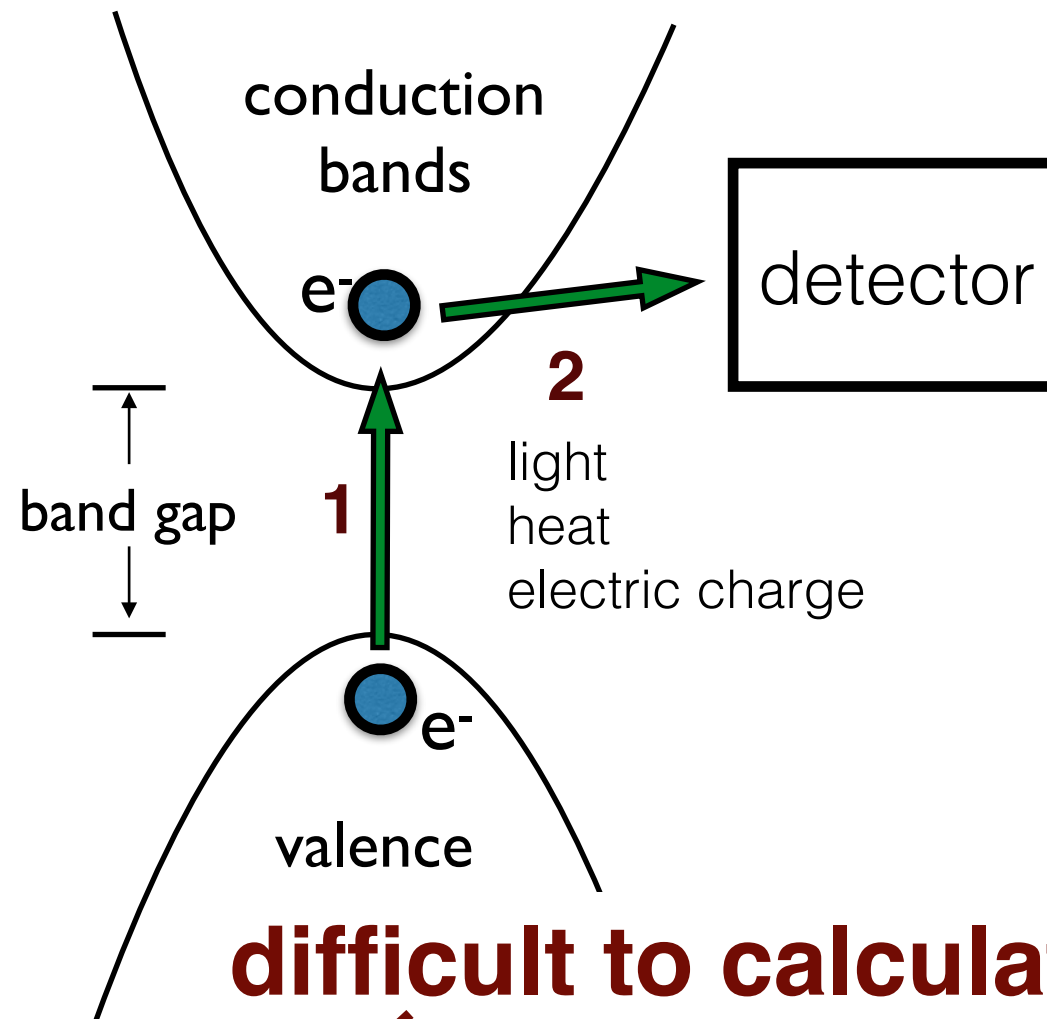
solid



$$\left| f_{i \rightarrow i'}(\vec{q}, \vec{k}) \right|^2 = \left| \sum_G \psi_{i'}^*(\vec{k} + \vec{G} + \vec{q}) \psi_i(\vec{k} + \vec{G}) \right|^2$$

**electrons in a solid are part
of a complicated, many-body system**

solid



difficult to calculate these

$$\left| f_{i \rightarrow i'}(\vec{q}, \vec{k}) \right|^2 = \left| \sum_G \psi_{i'}^*(\vec{k} + \vec{G} + \vec{q}) \psi_i(\vec{k} + \vec{G}) \right|^2$$

**electrons in a solid are part
of a complicated, many-body system**

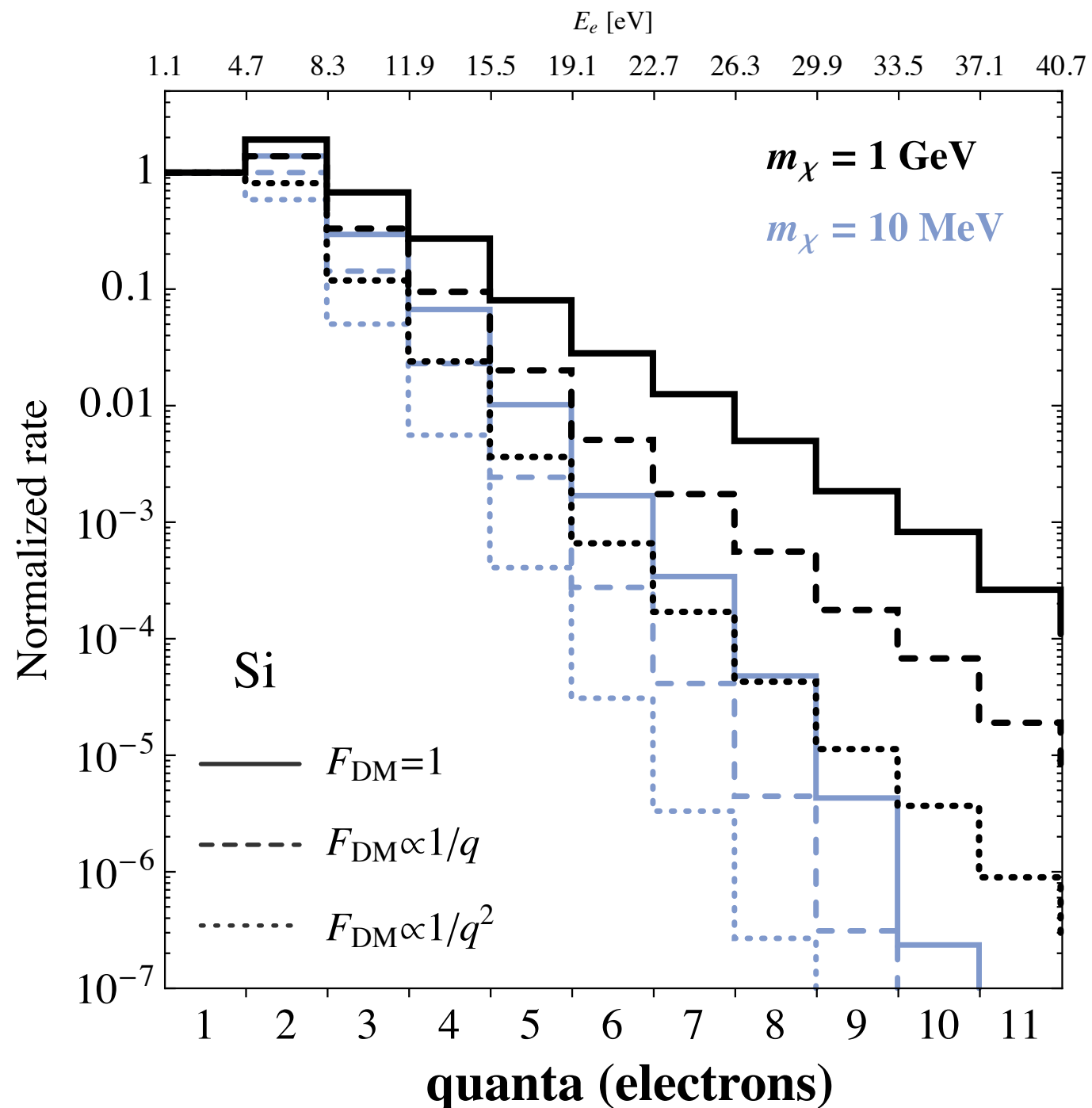


<http://www.quantum-espresso.org/>

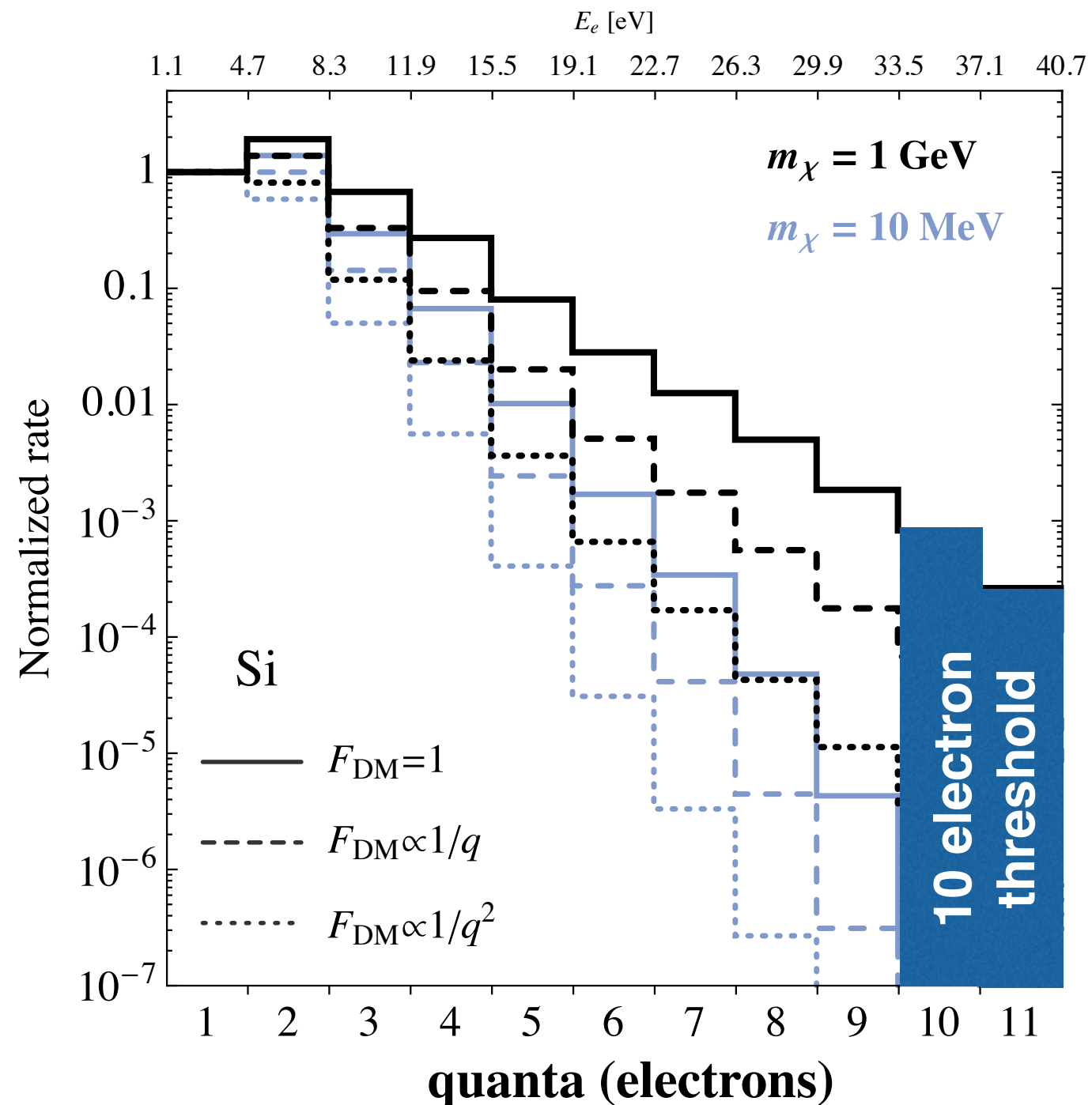
- **open source code for calculating electronic structure**
- **module **QEdark** calculates form factor:**

<http://ddldm.physics.sunysb.edu/>

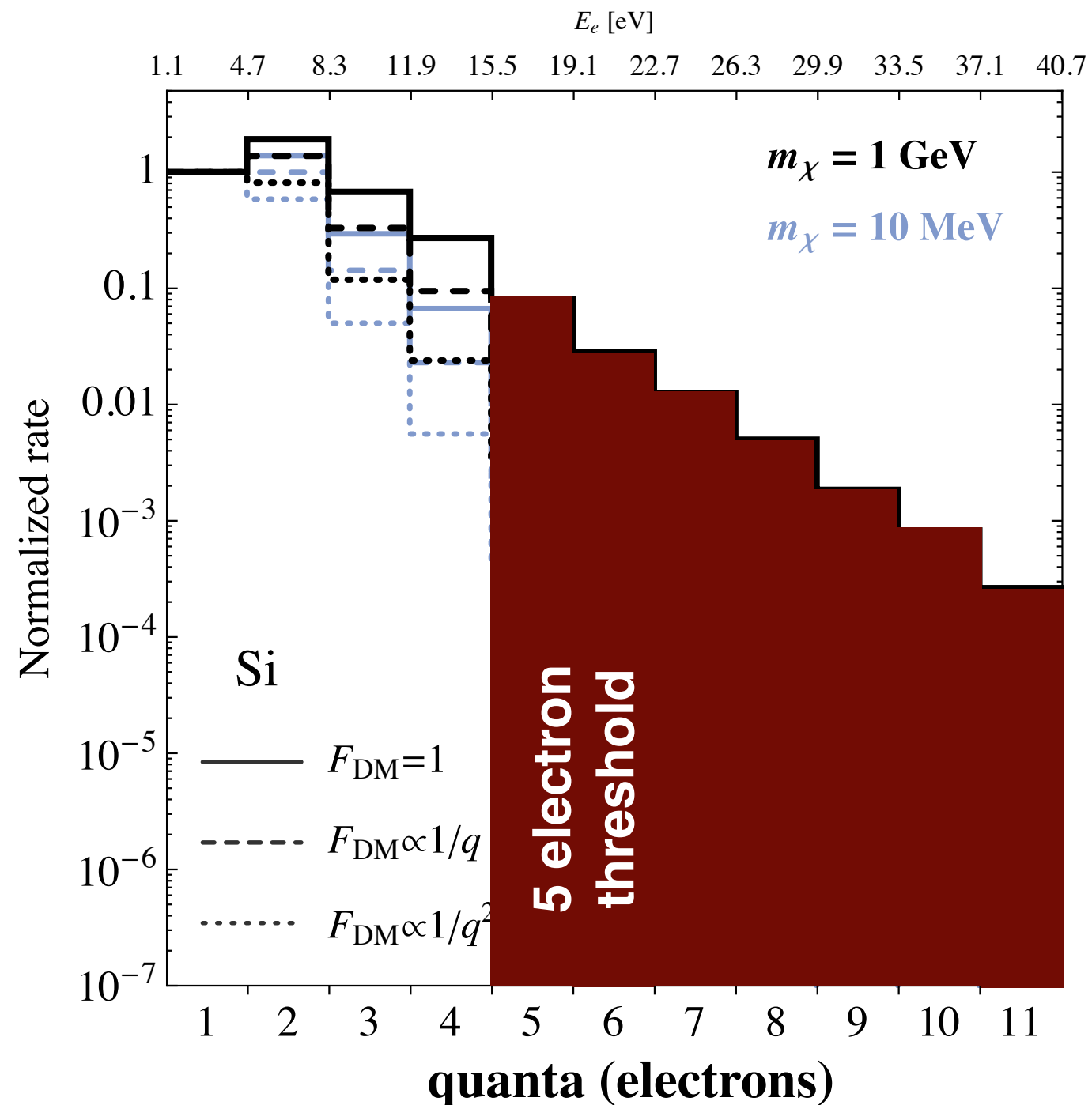
threshold dependence



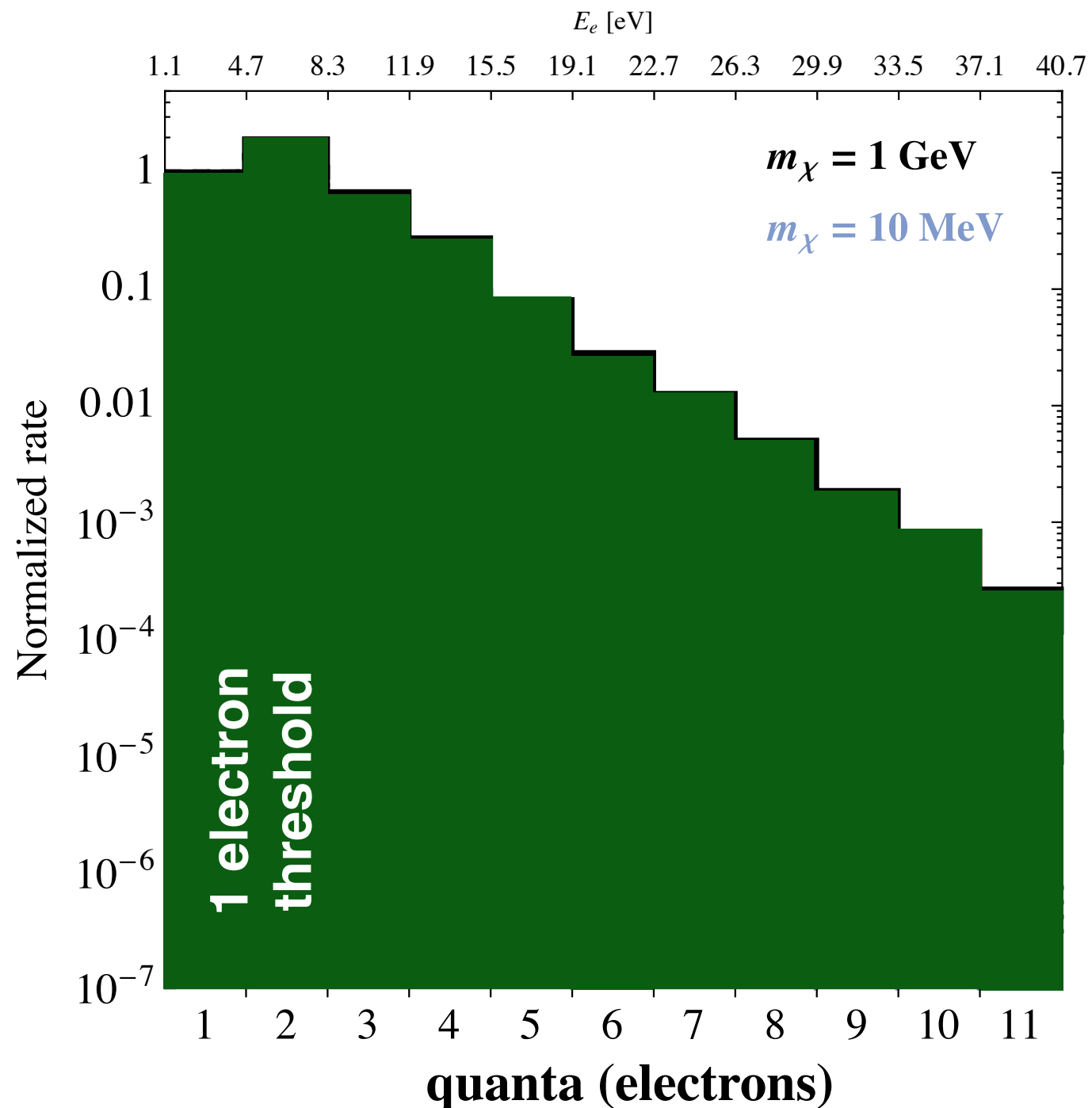
threshold dependence



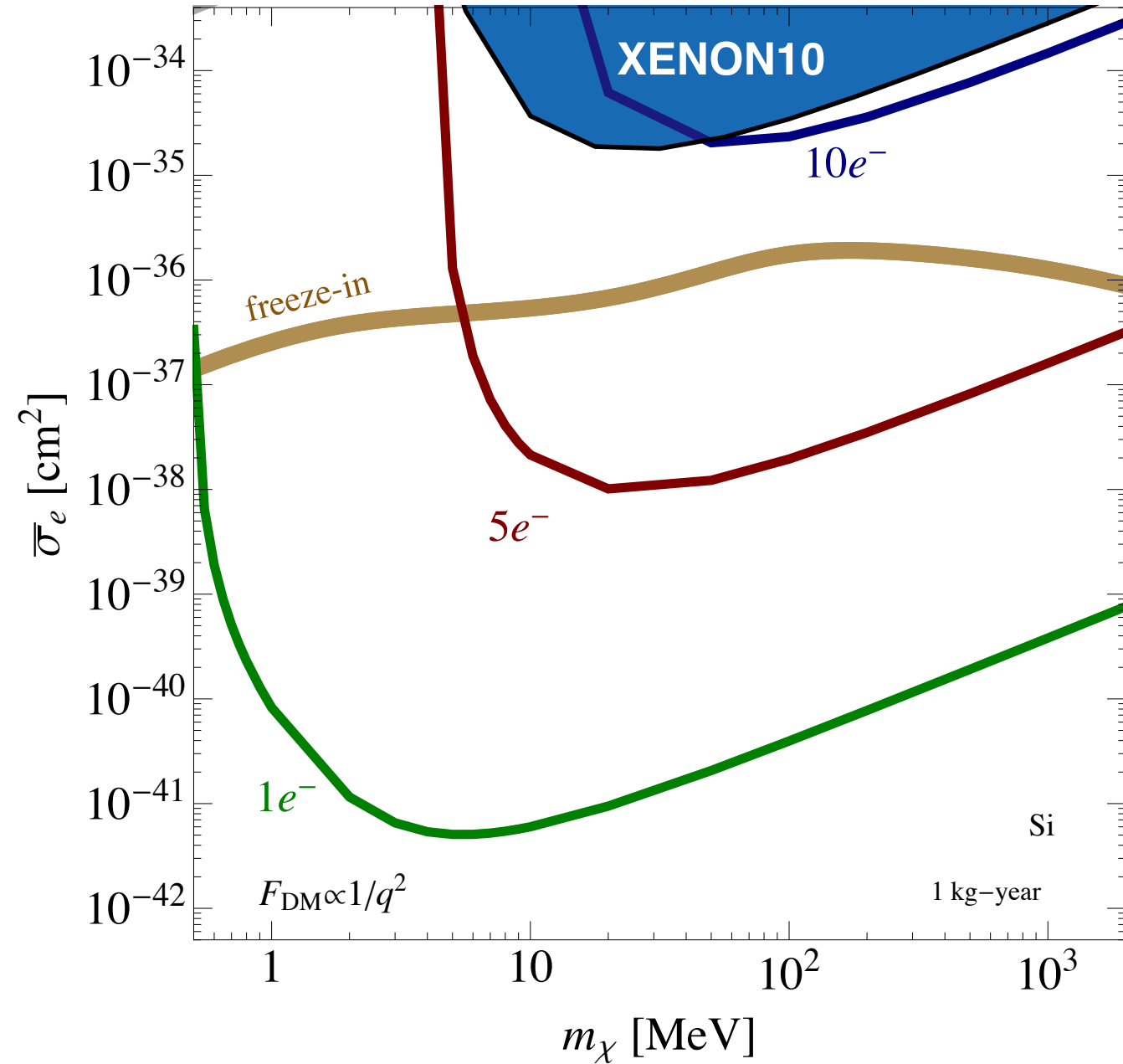
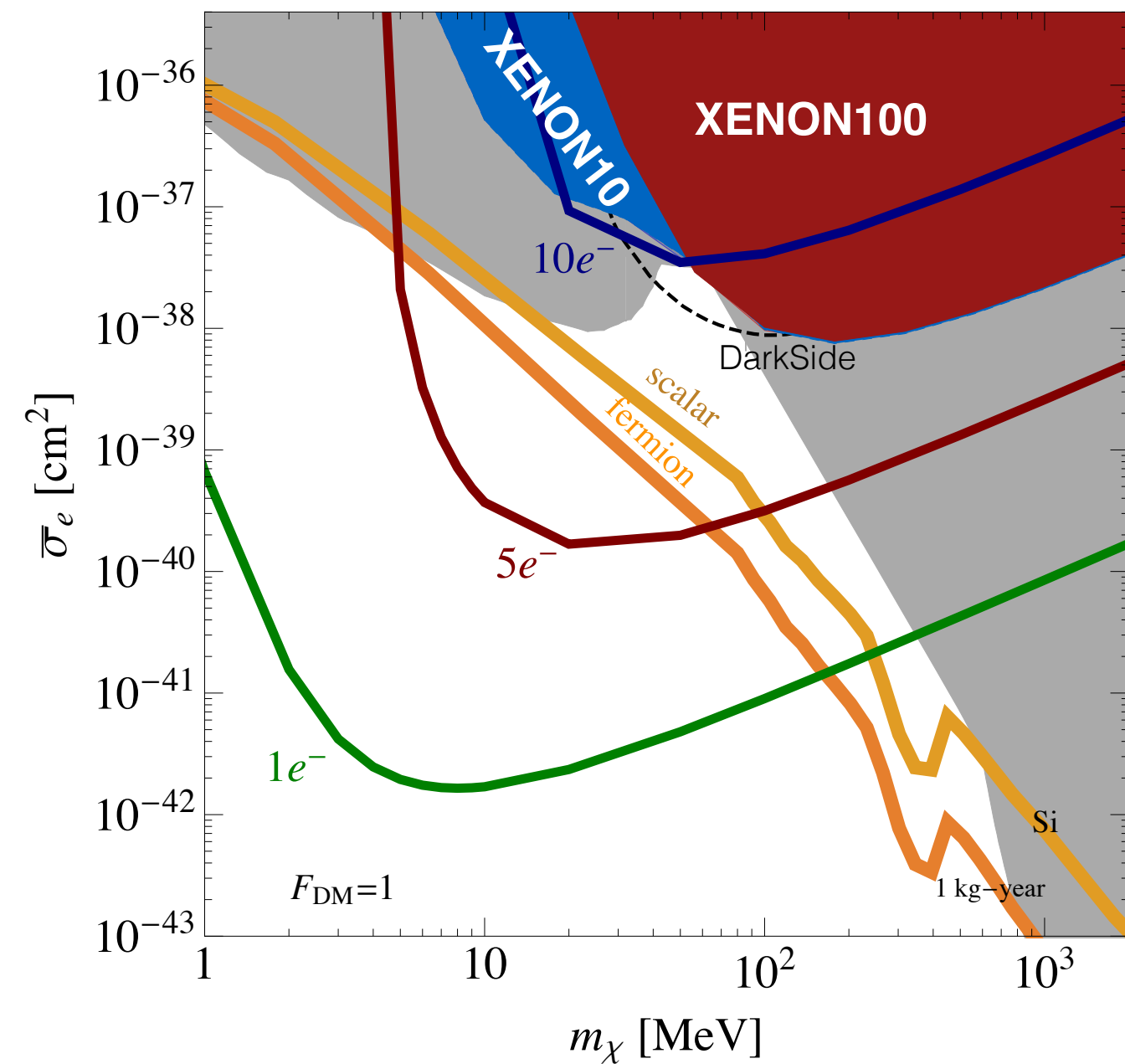
threshold dependence



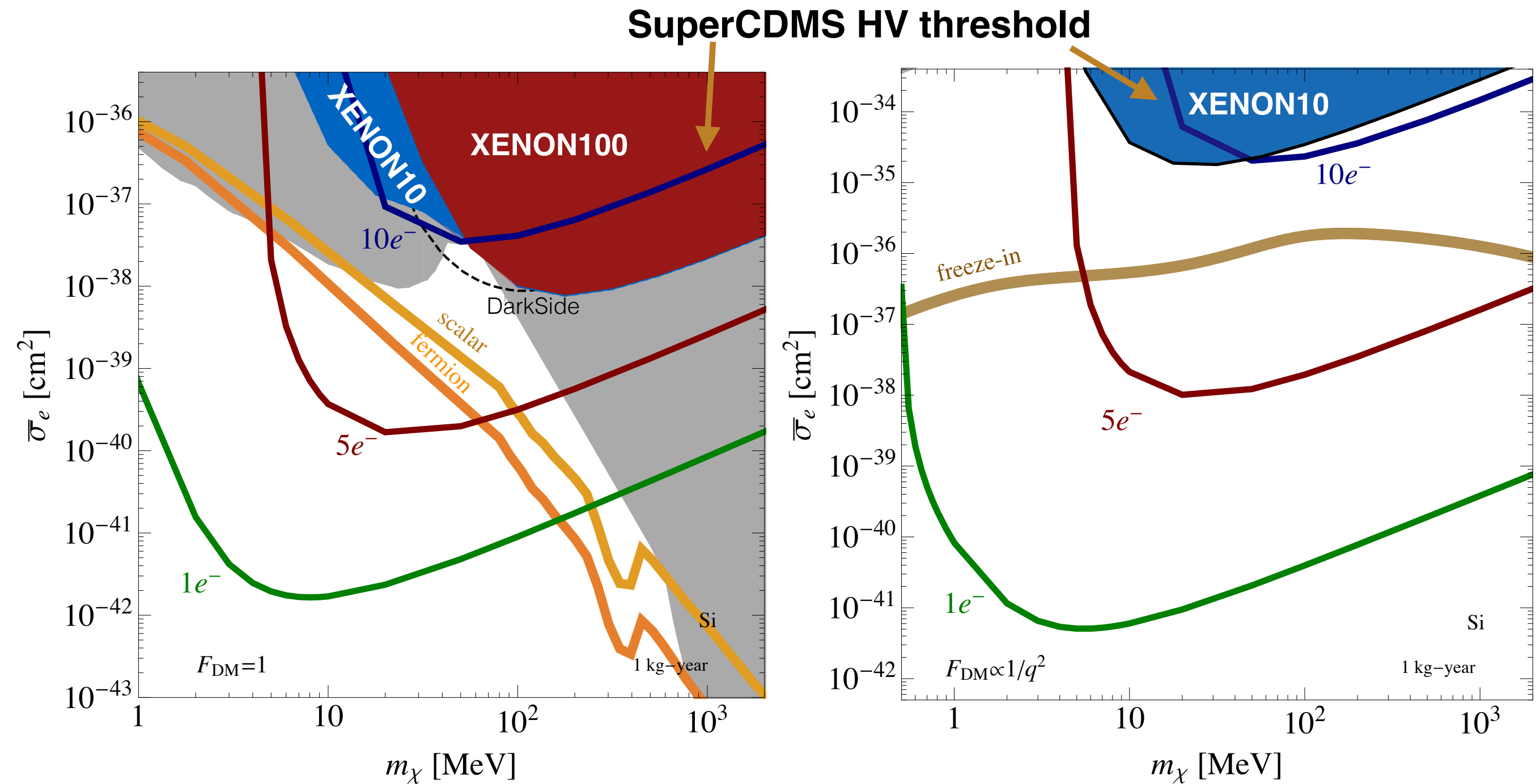
threshold dependence



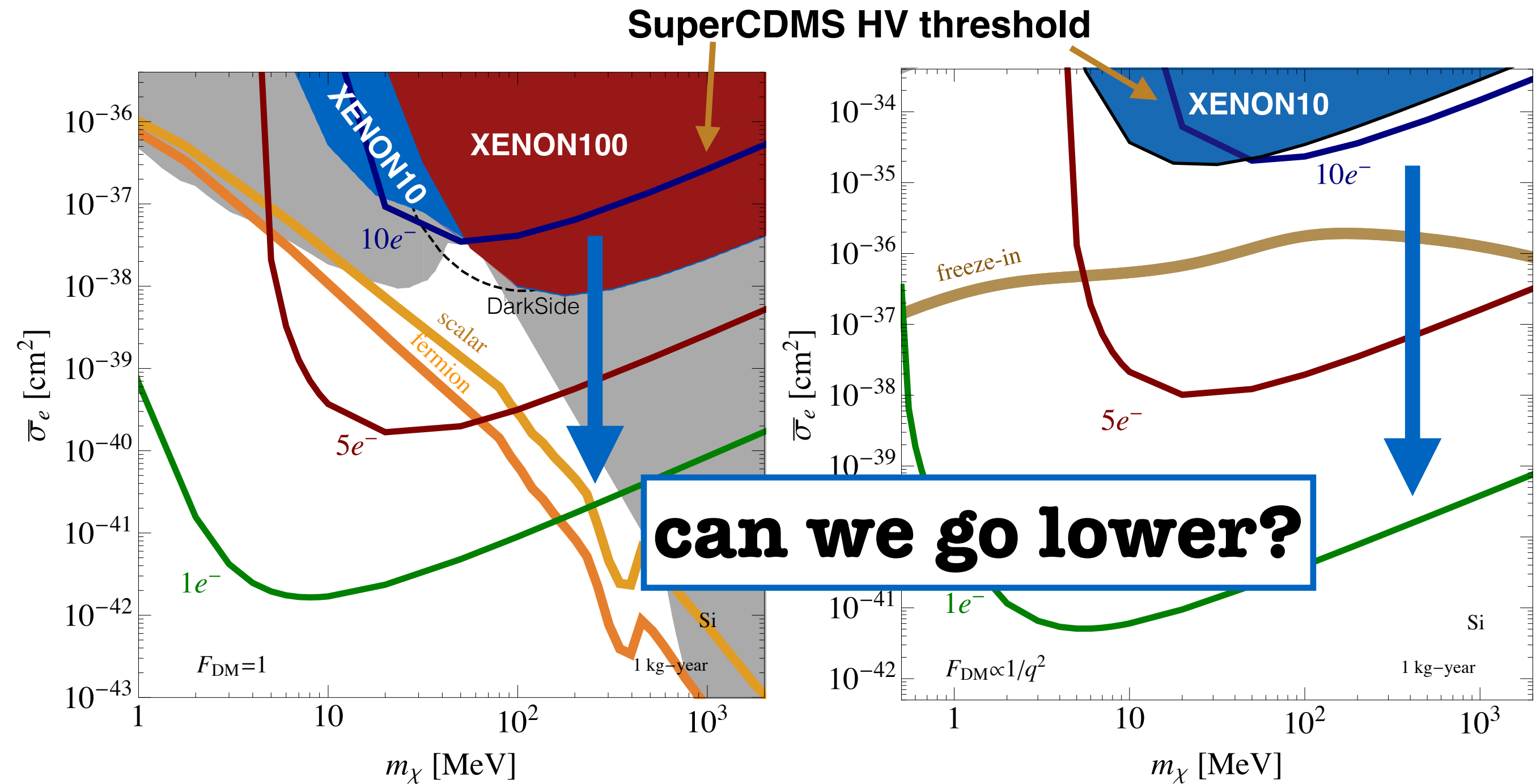
threshold dependence



threshold dependence



threshold dependence



SENSEI

Sub-Electron-Noise Skipper CCD Experimental Instrument

- fully-depleted 200 micrometer silicon CCD detector
- 4126 x 866 pixels
- operated at 140K
- currently at 1 gram, proposed to 100 grams
- skipper technology: measure charge/pixel multiple times



Comisión Nacional
de Energía Atómica



Fermilab



Stony Brook
University

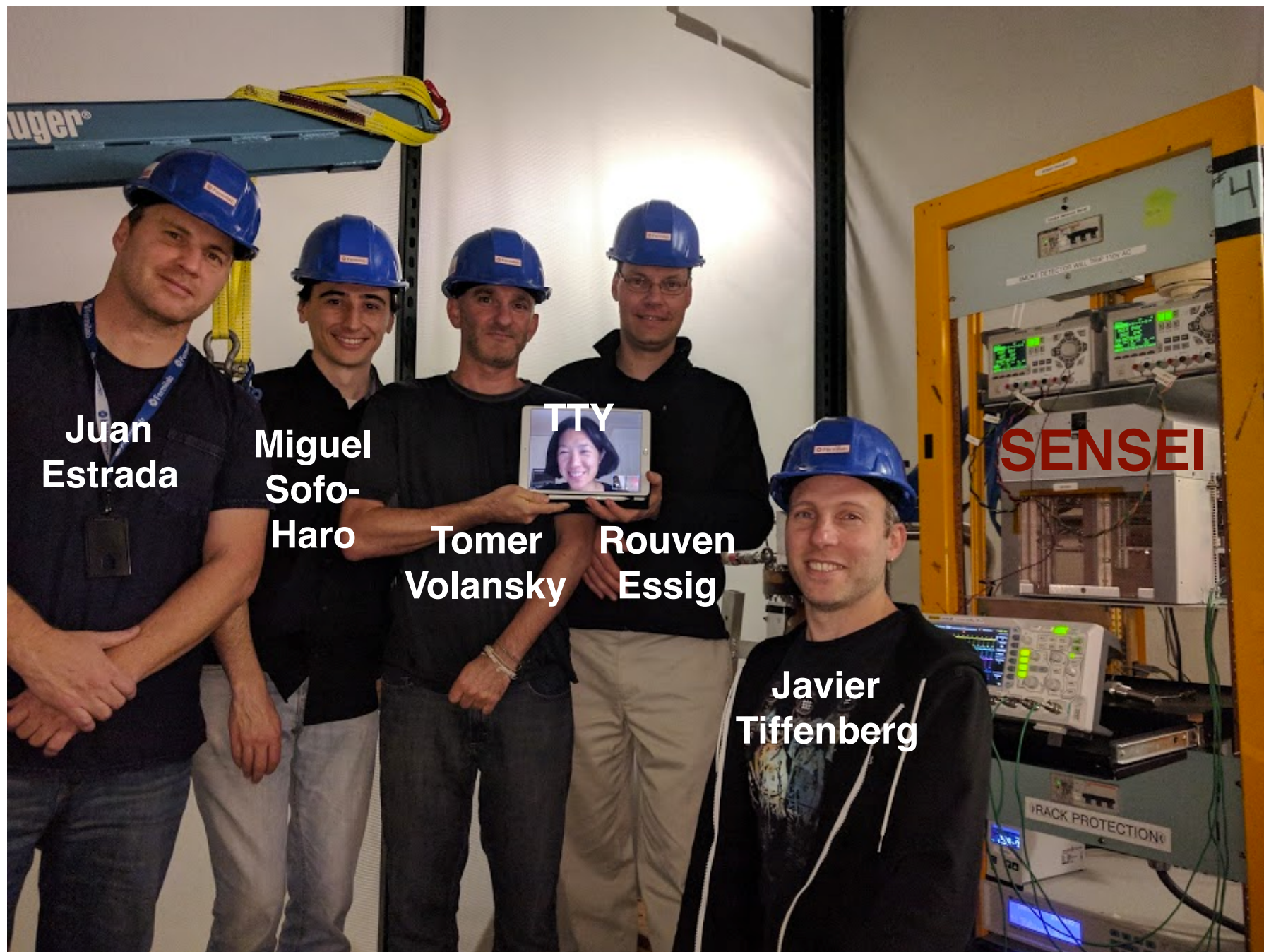




SENSEI



Sub-**E**lectron-**N**oise **S**kipper CCD **E**xperimental **I**nstrument



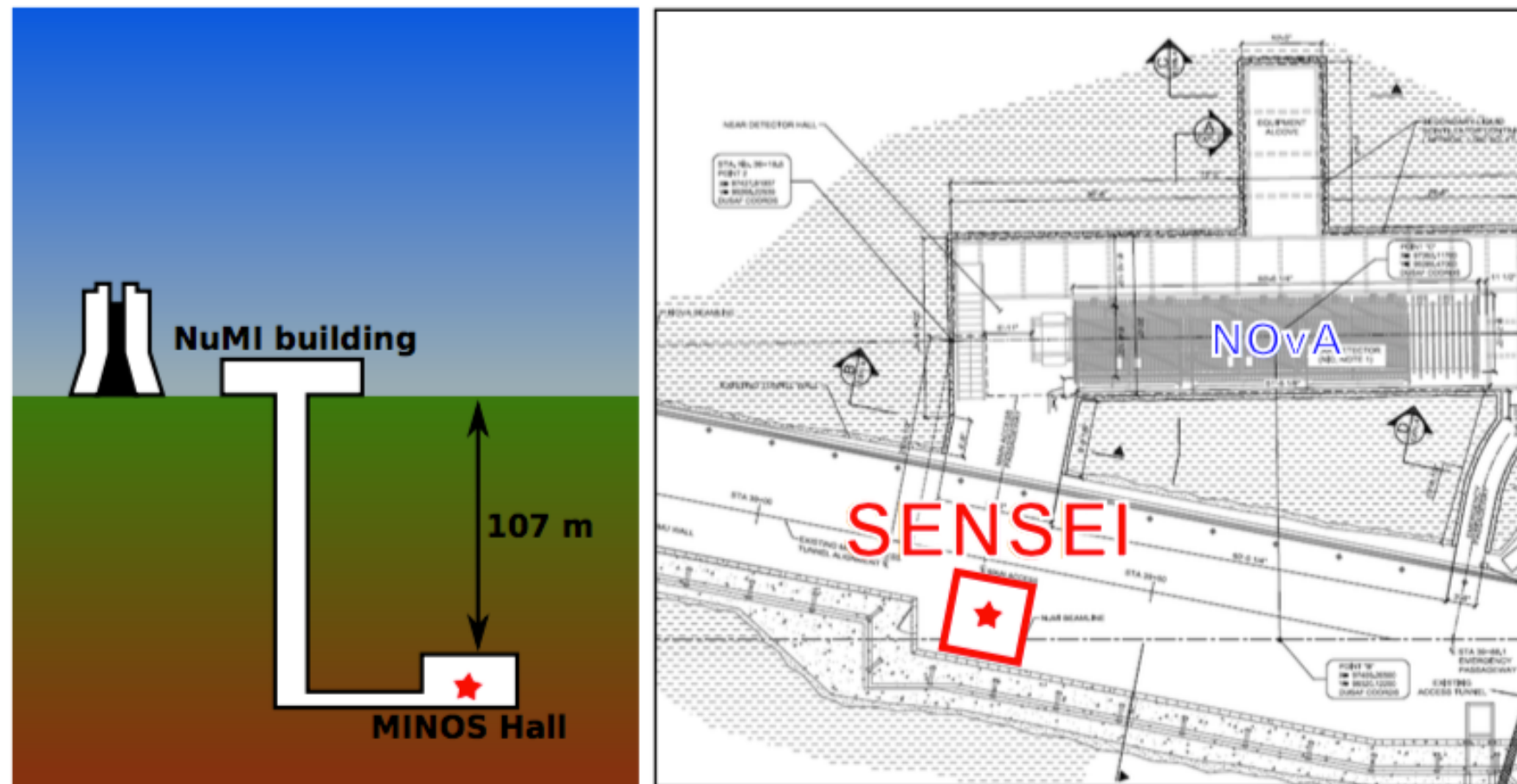
- + Guillermo Fernandez Moroni
- + Michael Crisler
- + Erez Etzion
- + Liron Barak
- + students + postdocs

SENSEI

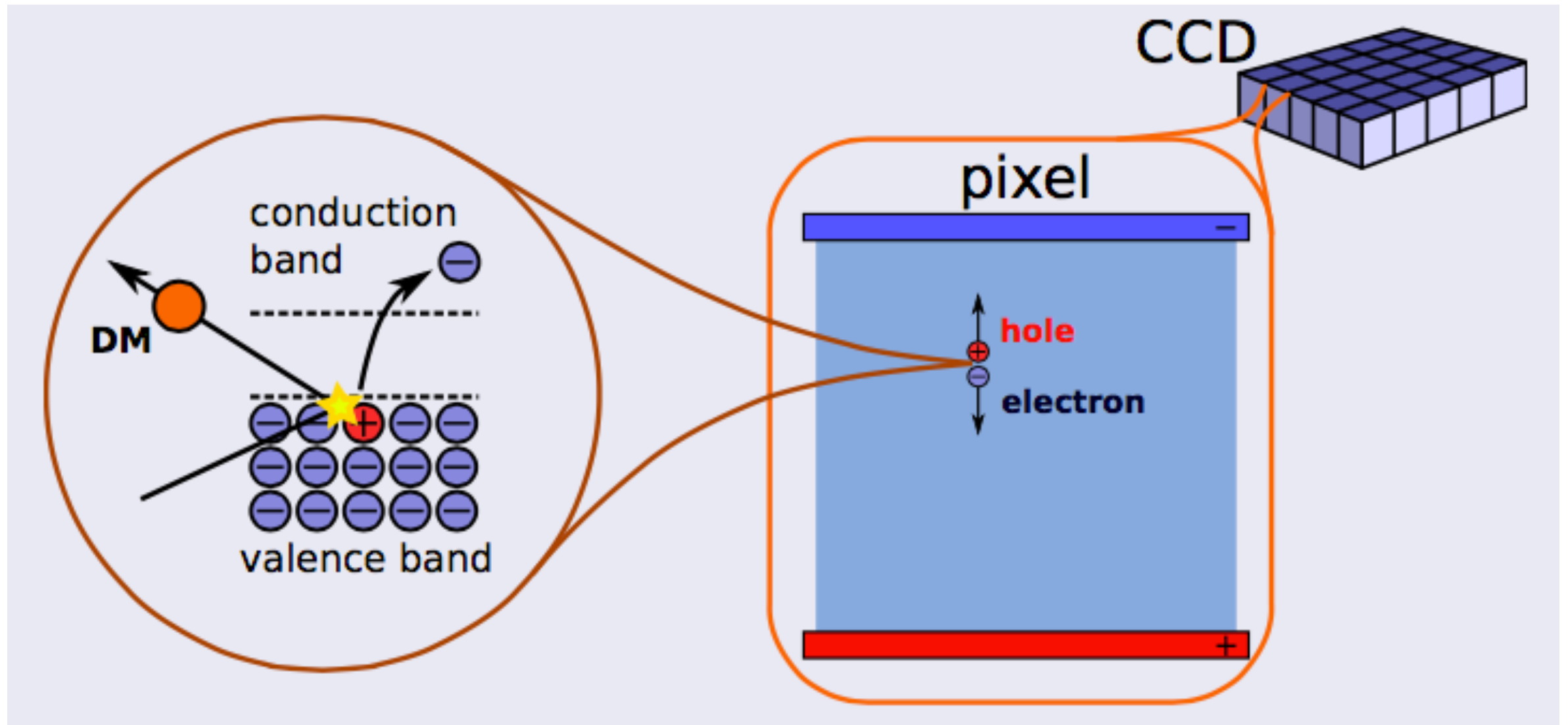
Sub-Electron-Noise Skipper CCD Experimental Instrument

Whats next: Installation @MINOS & low radiation package

Technology demonstration: installation at shallow underground site



silicon CCD detector

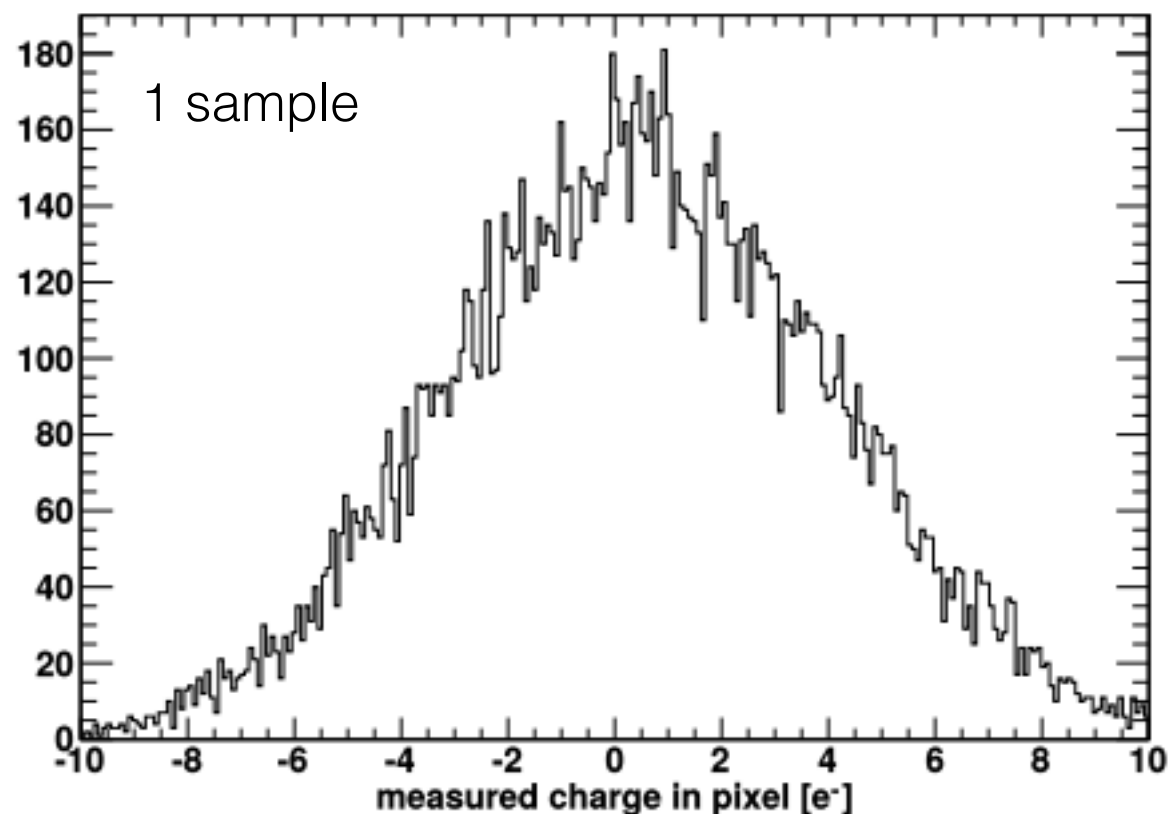


requires very low noise!
two sources: readout, dark current

skipper readout

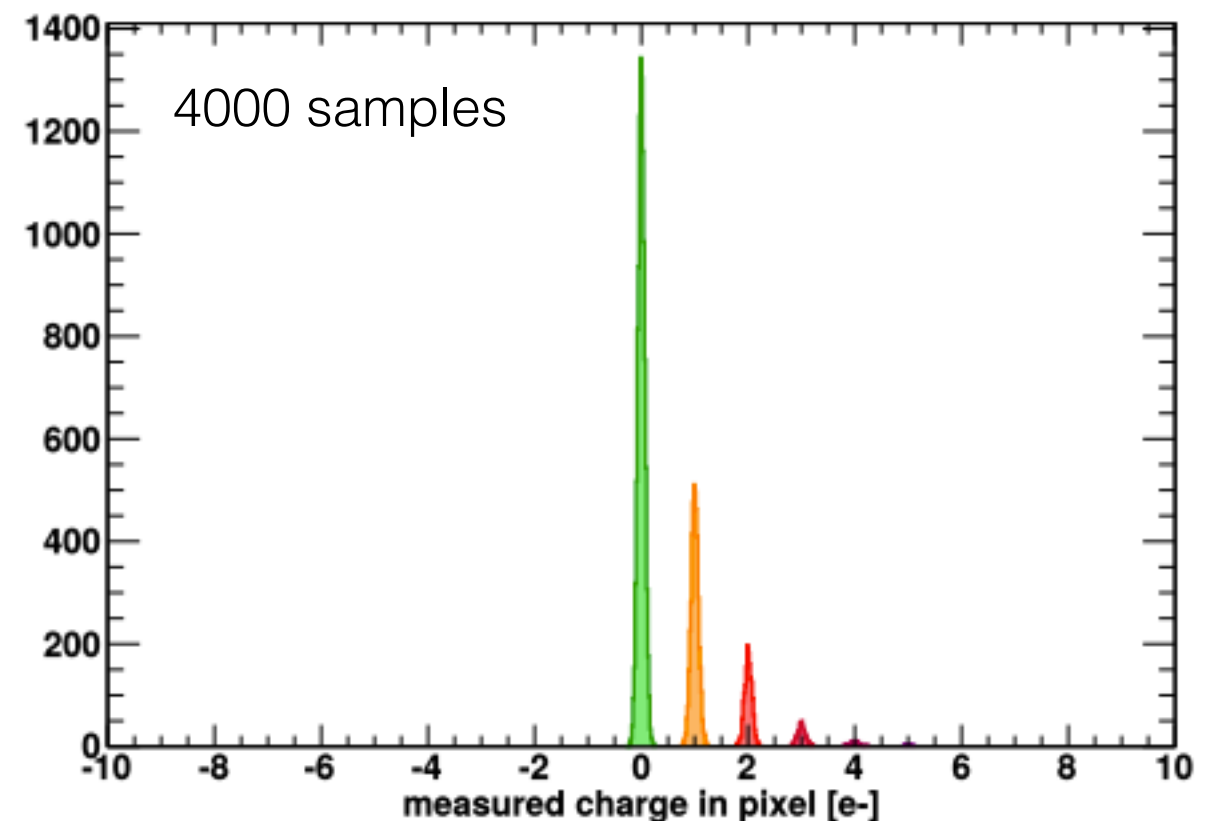
standard CCD

Readout-noise: 3.5 e RMS



skipper CCD

Readout-noise: 0.06 e RMS



Tiffenberg, Sofo-Haro, Drlica-Wagner, Essig,
Guardincerri, Holland, Volansky, TTY
[1706.00028]

skipper readout

standard CCD

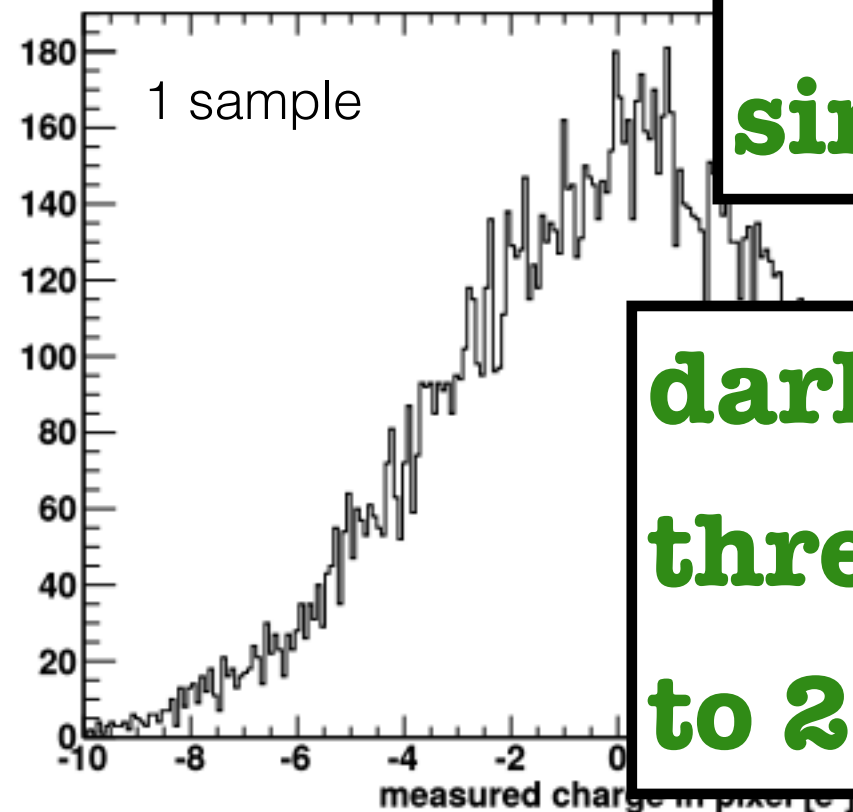
skipper CCD

Readout-noise:

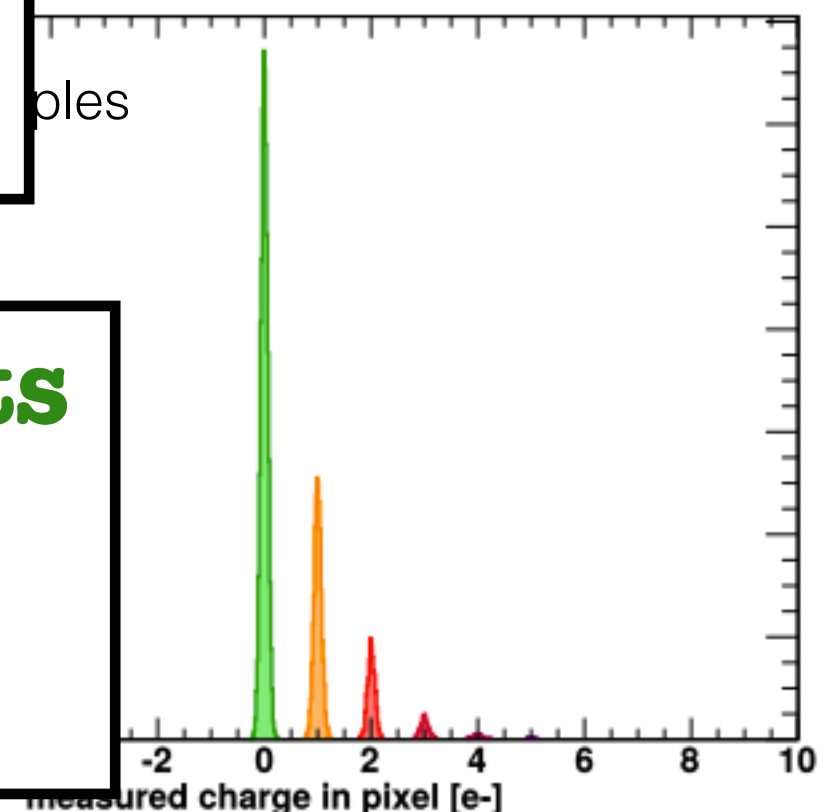
can count

single electrons!

ut-noise: 0.06 e RMS

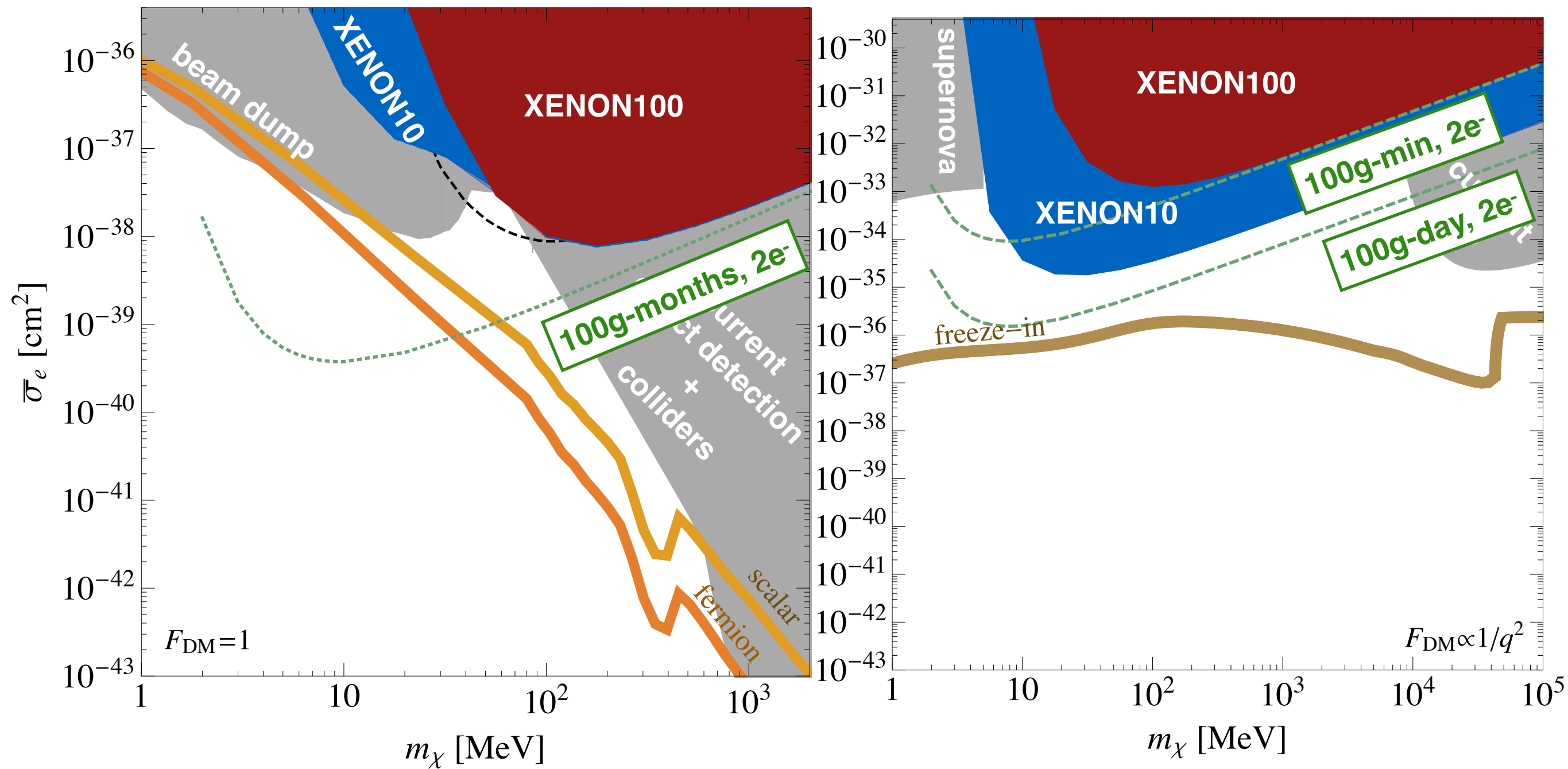


**dark current limits
threshold
to 2 electrons**

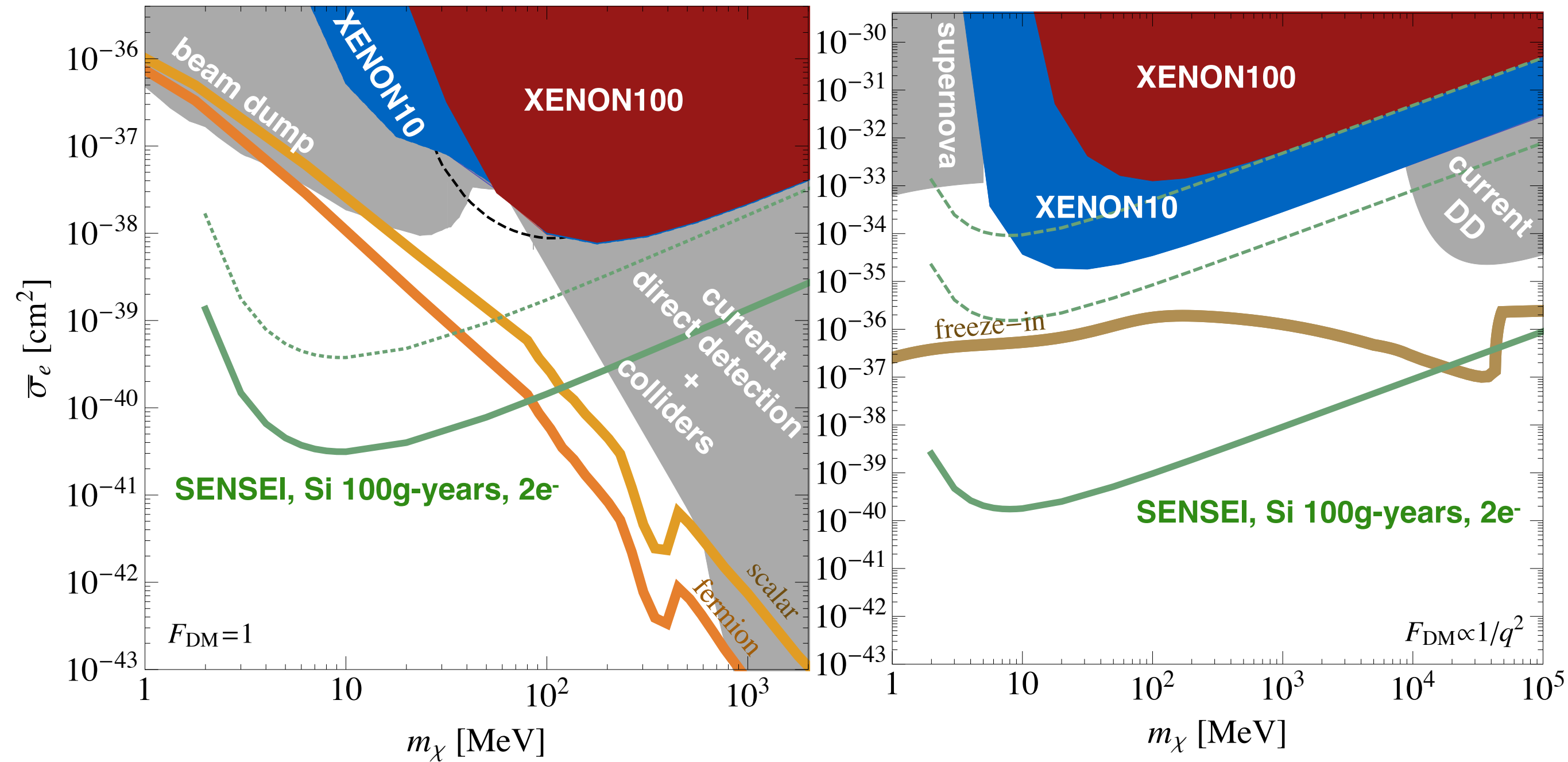


Tiffenberg, Sofo-Haro, Drlica-Wagner, Essig,
Guardincerri, Holland, Volansky, TTY
[1706.00028]

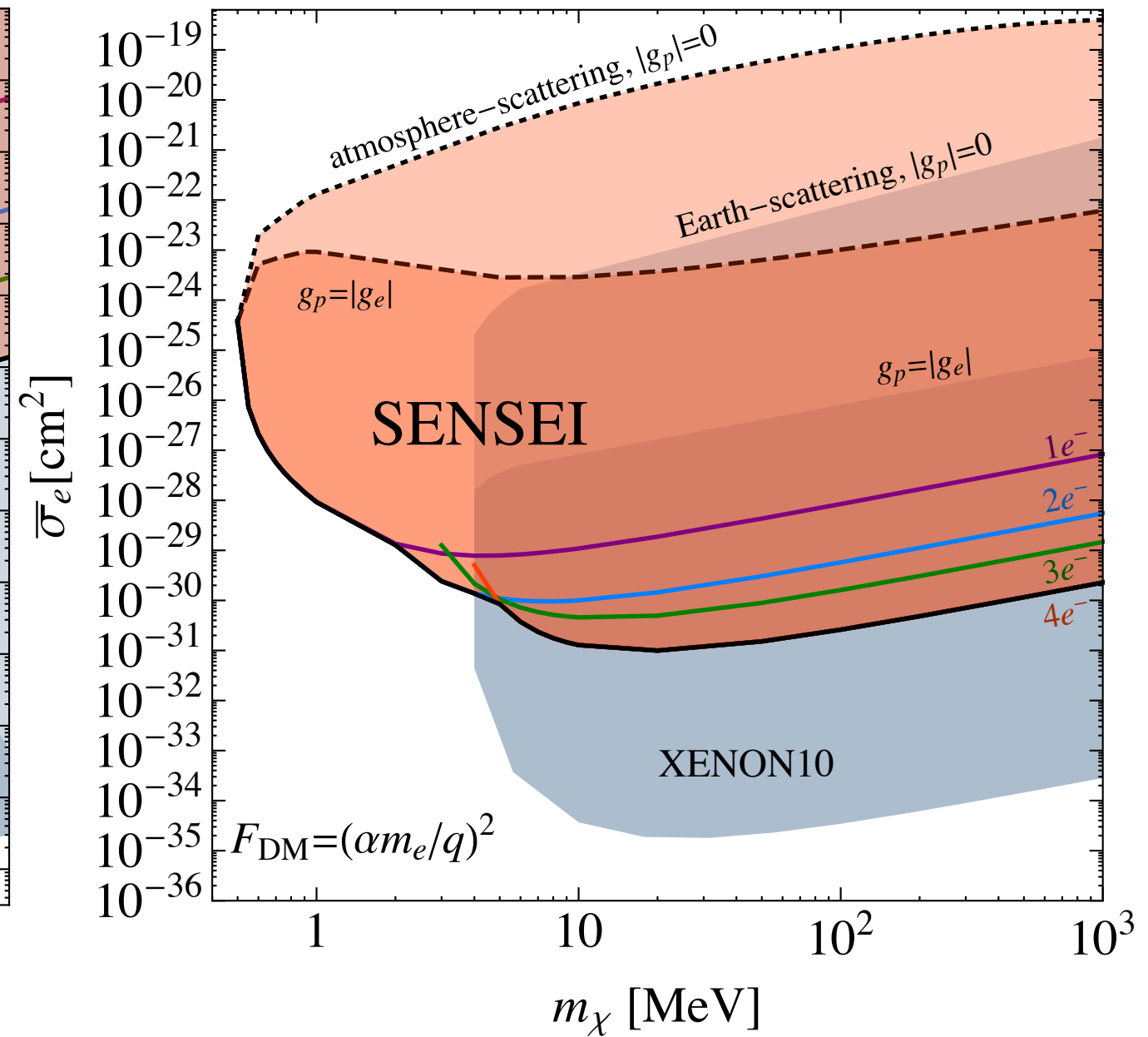
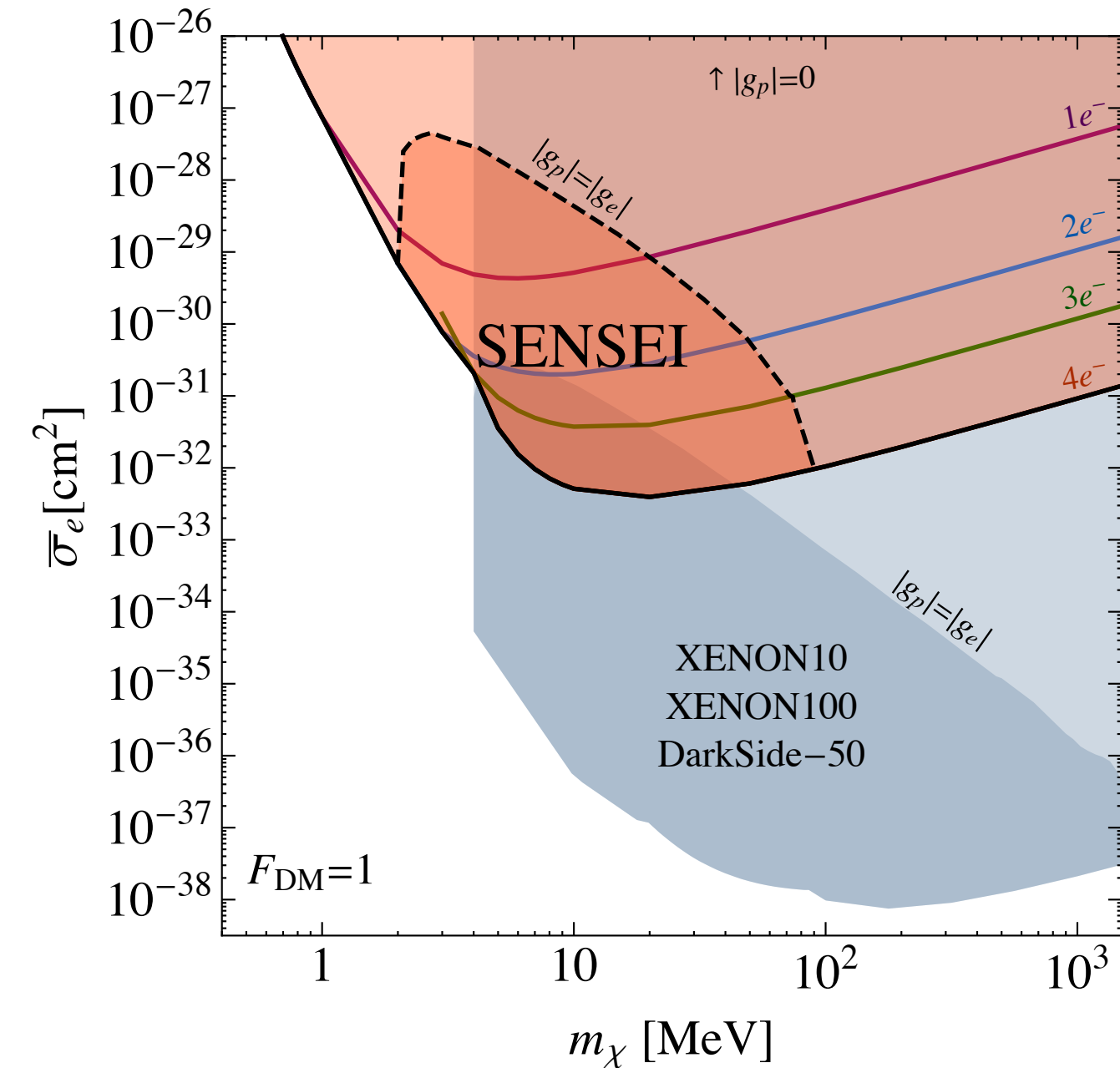
SENSEI projections



SENSEI projections

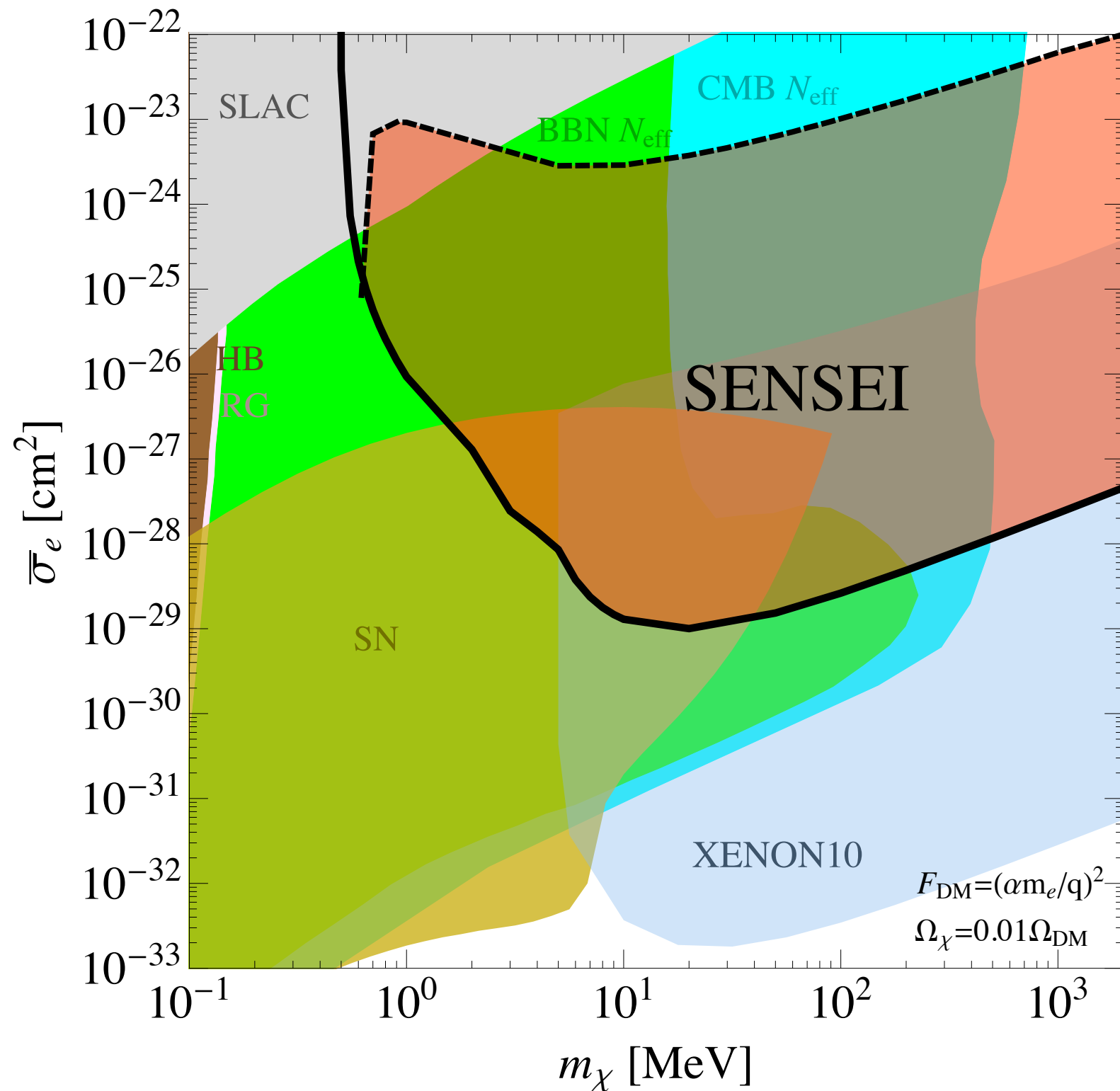


surface run at FNAL



0.019 gram-days

surface run at FNAL



future of SENSEI

- new results with 0.188 g-days of exposure coming very soon!
- begin deployment of 10 g detector at SNOlab early 2019
- ultimate science goal of 100g-years of exposure expected by end of 2021