

# Progress of The Daya Bay Reactor Neutrino Experiment

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On Behalf of the Daya Bay Collaboration

大亚湾反应堆中微子实验站

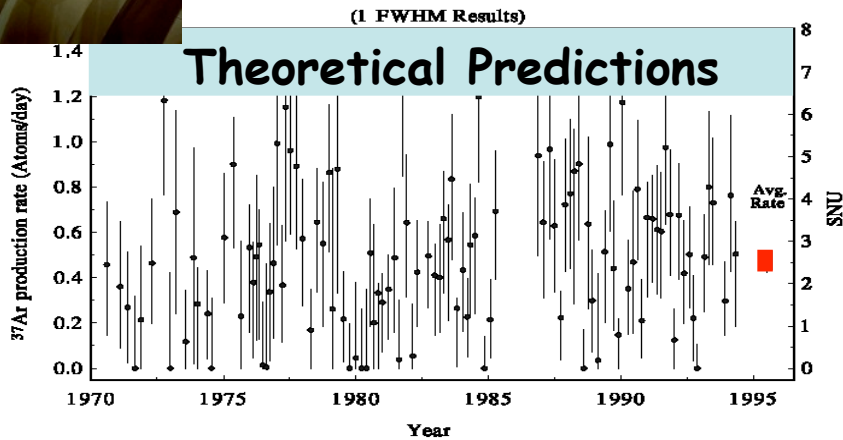
Daya Bay Reactor Neutrino Experiment Station

Seminar at IPMU, Tokyo, Japan  
27 June 2019

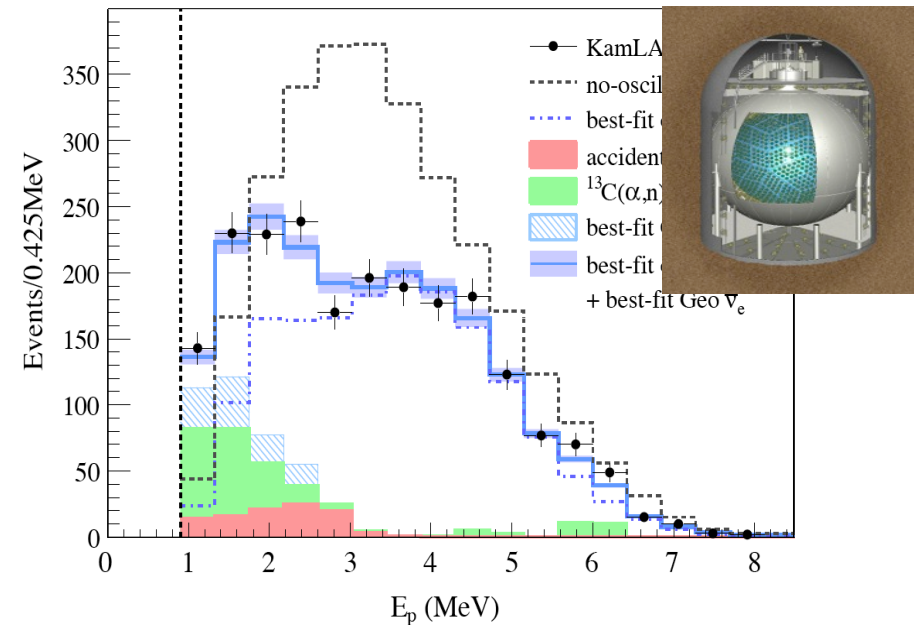
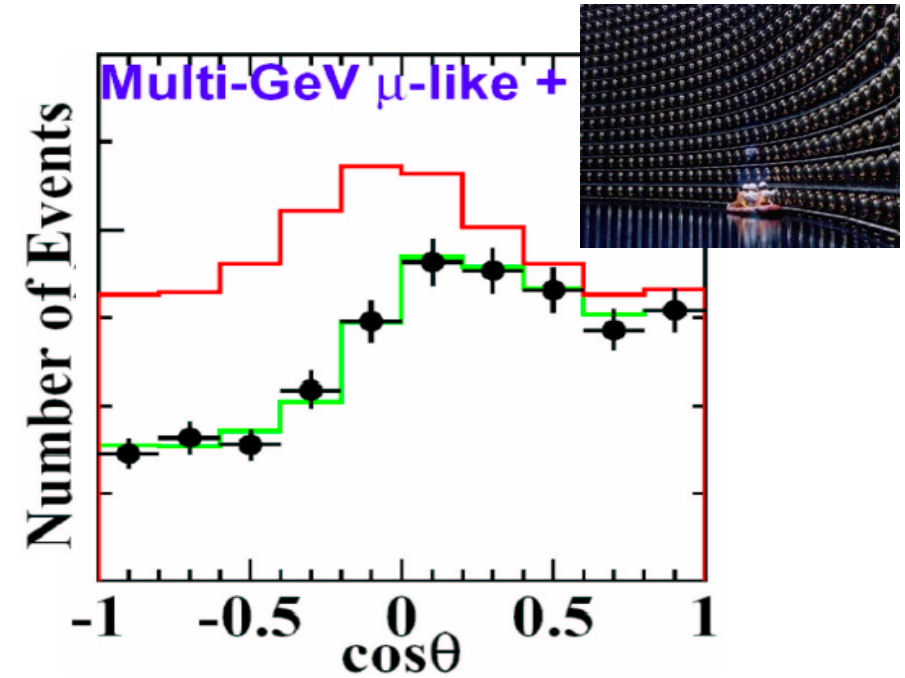
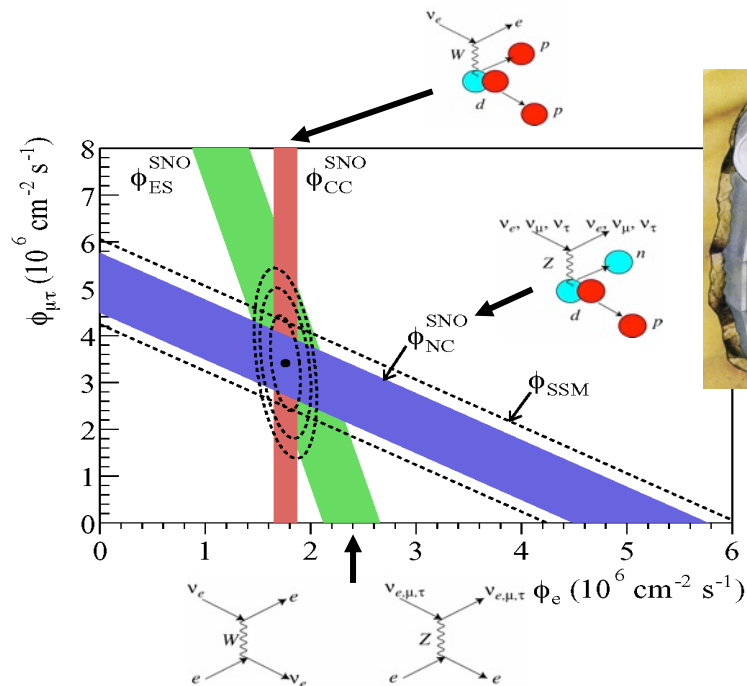
# Outline

- Introduction
  - Motivation
  - The Daya Bay experiment
- Neutrino oscillation
  - Latest measurements of  $\theta_{13}$  and  $\Delta m^2_{ee}$  using nGd samples
  - Search for a light sterile neutrino
- Absolute measurement of reactor antineutrinos
  - Flux
  - Energy spectrum

# Discoveries of Neutrino Oscillation



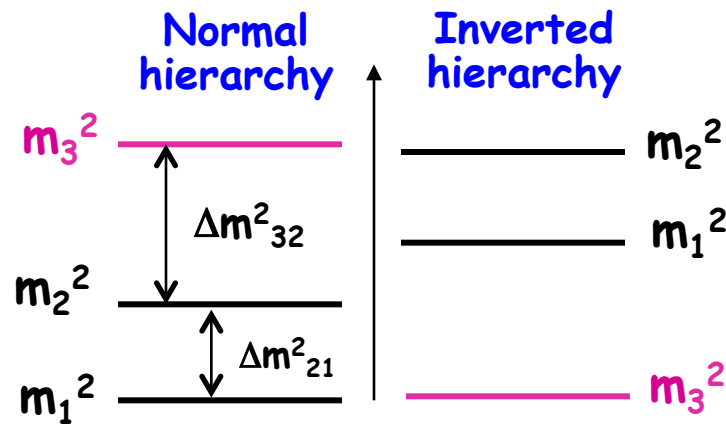
$$1 \text{ SNU} = 10^{-36} \text{ interaction/atom/s}$$



# Neutrino Mixing Circa 2011

$$U_{PMNS} = \begin{pmatrix} \cos\theta_{12} & \sin\theta_{12} & 0 \\ -\sin\theta_{12} & \cos\theta_{12} & 0 \\ 0 & 0 & 1 \end{pmatrix} \begin{pmatrix} \cos\theta_{13} & 0 & e^{-i\delta}\sin\theta_{13} \\ 0 & 1 & 0 \\ -e^{i\delta}\sin\theta_{13} & 0 & \cos\theta_{13} \end{pmatrix} \begin{pmatrix} 1 & 0 & 0 \\ 0 & \cos\theta_{23} & \sin\theta_{23} \\ 0 & -\sin\theta_{23} & \cos\theta_{23} \end{pmatrix}$$

$\theta_{12} = 33^\circ \pm 1^\circ$ 
 $\theta_{13} \text{ and } \delta ?$ 
 $\theta_{23} \approx 42^\circ \pm 3^\circ$



Which one ?

$$|\Delta m_{31}^2| = |\Delta m_{32}^2| \pm |\Delta m_{21}^2| \approx |\Delta m_{32}^2|$$

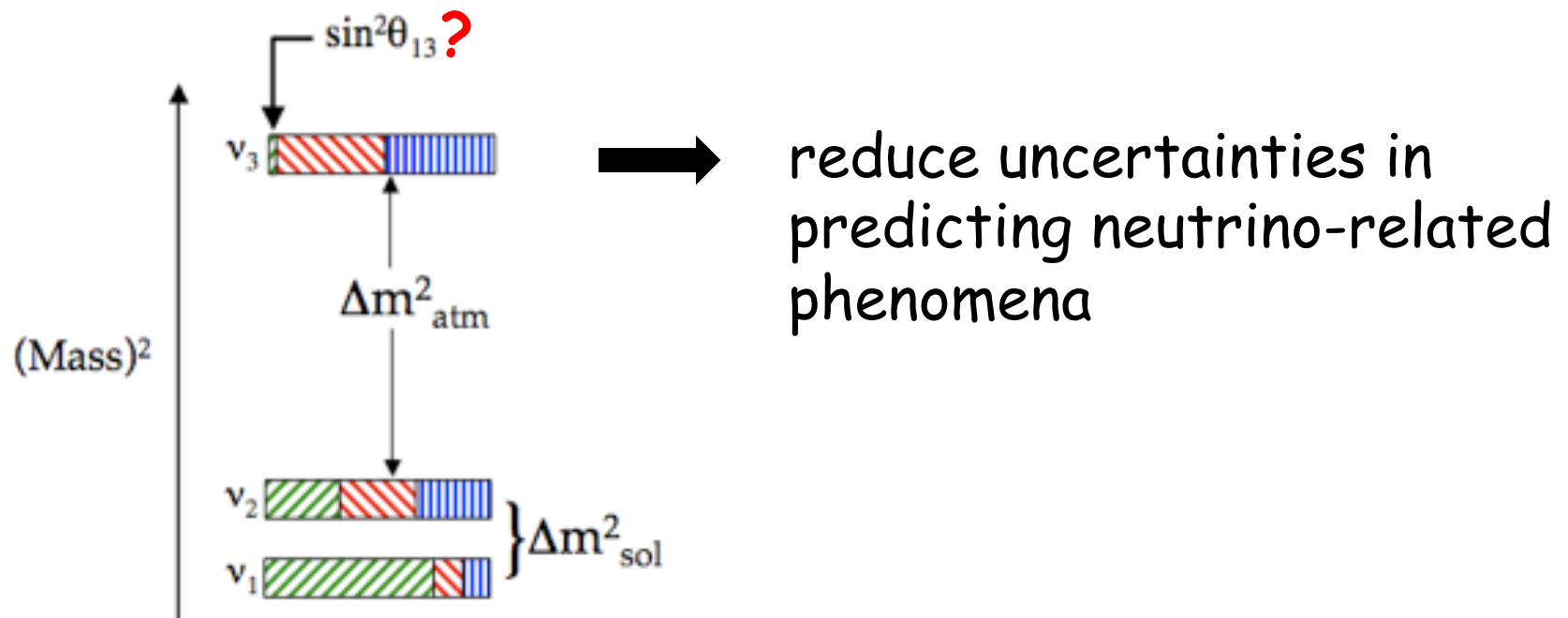
$\uparrow$ 
 $\uparrow$

$(2.45 \pm 0.09) \times 10^{-3} \text{ eV}^2$ 
 $(7.6 \pm 0.2) \times 10^{-5} \text{ eV}^2$



## Significance of Knowing $\theta_{13}$

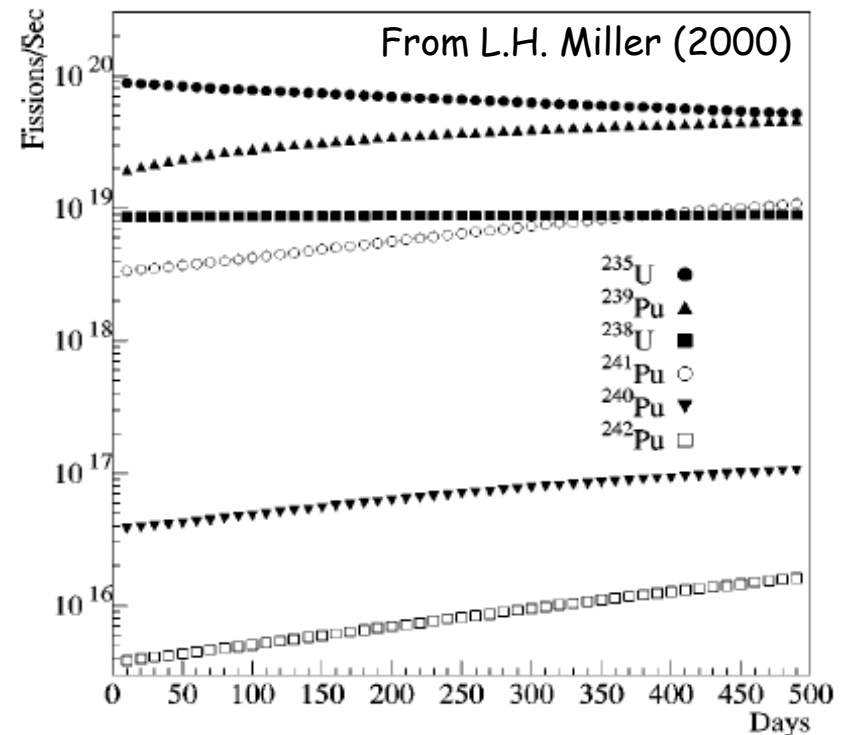
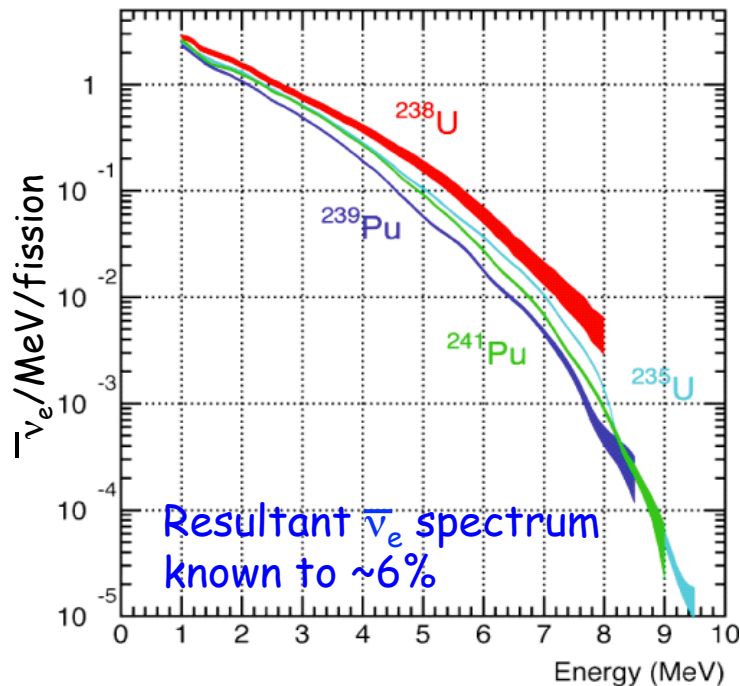
- Complete the determination of the mixing matrix
  - guide model-building
- Determine  $\nu_e$  fraction of  $\nu_3$



- $\theta_{13}$  is the gateway to CP violation in the neutrino sector:  
 $P(\nu_\mu \rightarrow \nu_e) - P(\bar{\nu}_\mu \rightarrow \bar{\nu}_e) \propto \sin 2\theta_{13} \cos \theta_{13} \sin \delta$

# Production of Reactor $\bar{\nu}_e$

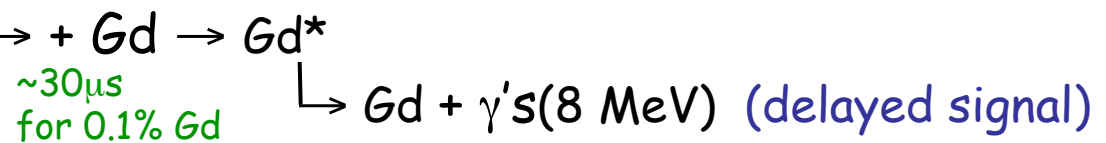
- A pure, intense source of low-energy  $\bar{\nu}_e$ :  
 $3 \text{ GW}_{\text{th}}$  generates  $6 \times 10^{20} \bar{\nu}_e$  per sec



- $\bar{\nu}_e$  related to  $^{235}\text{U}$ ,  $^{239}\text{U}$ , and  $^{241}\text{Pu}$  :
  - measure  $\beta$  spectrum using thermal neutron induced fission on the isotope
  - convert  $\beta$  spectrum to  $\bar{\nu}_e$  spectrum
- $\bar{\nu}_e$  related to  $^{238}\text{U}$  :
  - $\bar{\nu}_e$  spectrum is based on calculation, now measurement as well.
- Uncertainty in  $\bar{\nu}_e$  yield, ~2%, due to
  - Thermal power (0.5%)
  - Sampling of fuel
  - Analysis of fractions of isotopes in samples

# Detecting Reactor $\bar{\nu}_e$

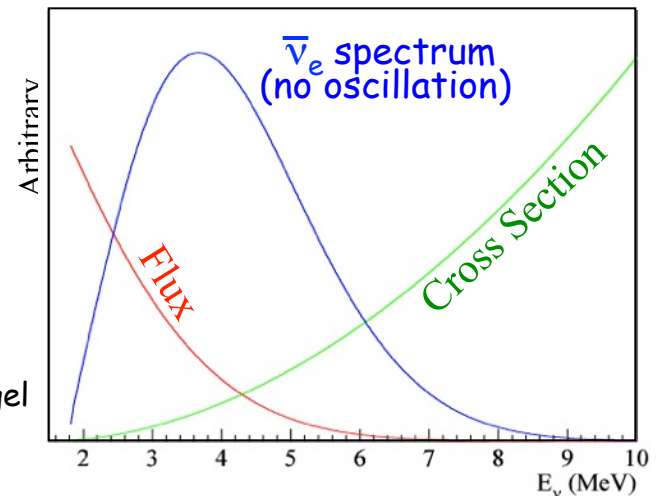
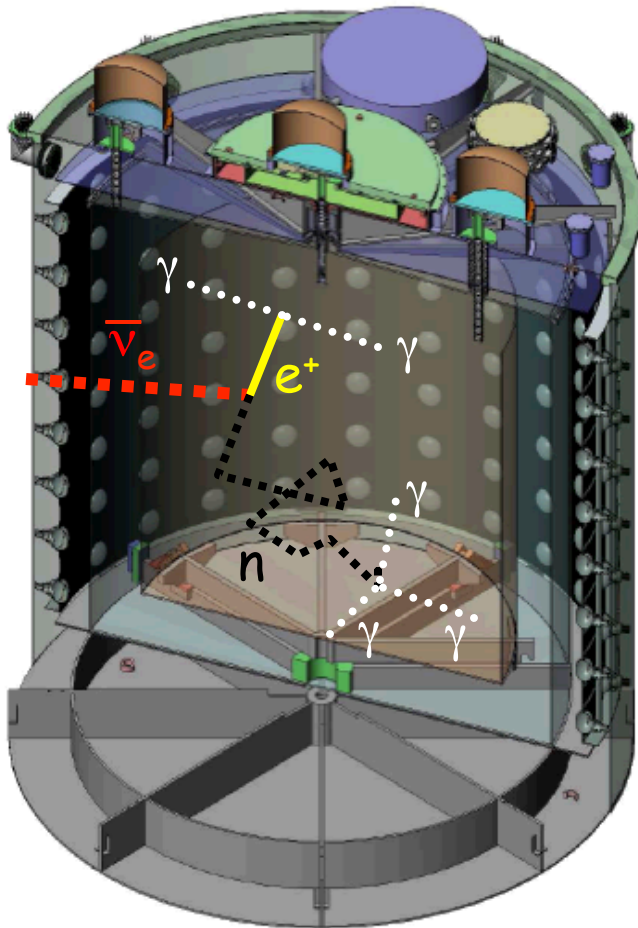
- Detect inverse  $\beta$ -decay reaction (IBD) in liquid scintillator:



- Time- and energy-tagged signal is a good tool to suppress background events.
- Energy of  $\bar{\nu}_e$  is given by:

$$E_{\nu} \approx T_{e^+} + T_n + (m_n - m_p) + m_{e^+} \approx T_{e^+} + 1.8 \text{ MeV}$$

tens of keV

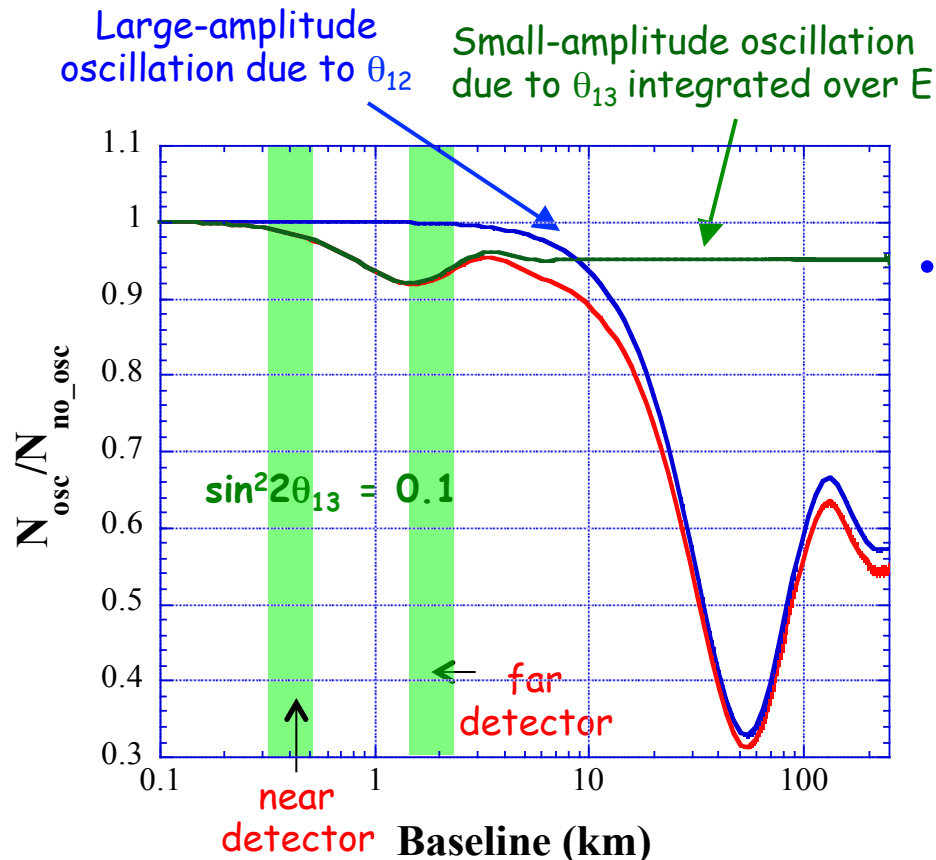


From Bemporad, Gratta and Vogel

# Determining $\theta_{13}$ With Reactor $\bar{\nu}_e$

- Look for disappearance of electron antineutrinos from reactors:

$$P(\bar{\nu}_e \rightarrow \bar{\nu}_e) = 1 - \sin^2 2\theta_{13} \sin^2 \left( \frac{\Delta m_{ee}^2 L}{4E} \right) + \cos^4 \theta_{13} \sin^2 2\theta_{12} \sin^2 \left( \frac{\Delta m_{21}^2 L}{4E} \right)$$



$$\sin^2 \left( \frac{\Delta m_{ee}^2 L}{4E} \right) \leftrightarrow \cos^2 \theta_{12} \sin^2 \left( \frac{\Delta m_{31}^2 L}{4E} \right) + \sin^2 \theta_{12} \sin^2 \left( \frac{\Delta m_{32}^2 L}{4E} \right)$$

- Perform relative measurement, for a given  $E$  :

$$\frac{R_{Far}}{R_{Near}} = \left( \frac{N_{Far}}{N_{Near}} \right) \left( \frac{\epsilon_{Far}}{\epsilon_{Near}} \right) \left( \frac{L_{Near}}{L_{Far}} \right)^2 \left( \frac{P_{Far}}{P_{Near}} \right)$$

$\bar{\nu}_e$  rate      number of protons      detection efficiency       $1/r^2$       yield  $\sin^2 2\theta_{13}$

Correlated errors are cancelled & enable precise measurement.



# The Daya Bay Collaboration

~200 members from 42 institutions



## Asia (23)

Beijing Normal Univ., CGNPG, CIAE, CQU,  
Dongguan Univ. Tech., ECUST, IHEP,  
Nankai Univ., NCEPU, NCTU, Nanjing Univ.,  
NUDT, Shandong Univ., Shanghai Jiaotong Univ.,  
Shenzhen Univ., Tsinghua Univ., USTC,  
Xi'an Jiaotong Univ., Zhongshan Univ.,  
Chin. Univ. of Hong Kong, Univ. of Hong Kong,  
Nat. Taiwan Univ., Nat. Chiao Tung Univ.,  
National United Univ.

## Europe (2)

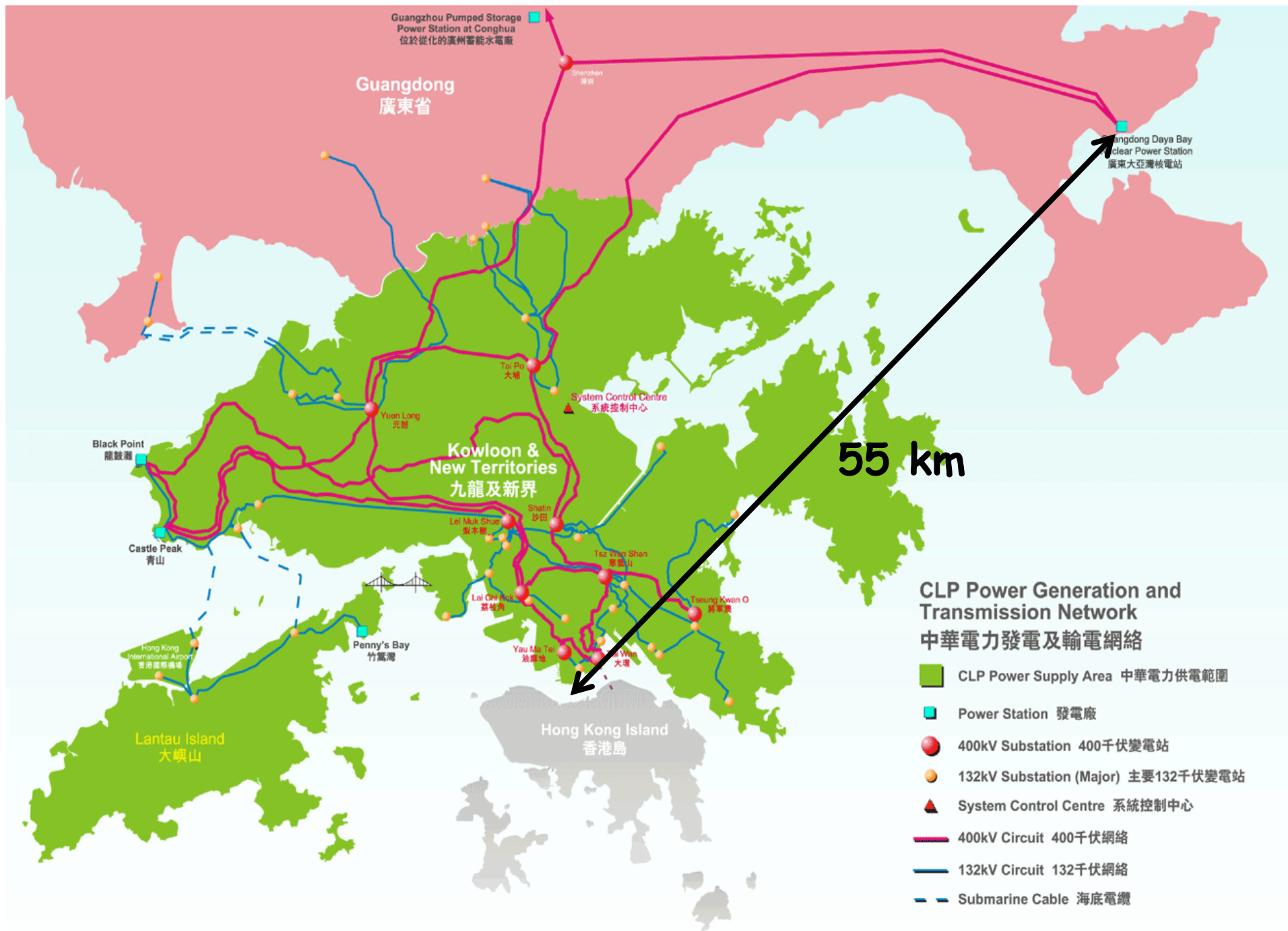
JINR, Dubna, Russia  
Charles University, Czech Republic

## North America (16)

BNL, Illinois Inst. Tech.,  
Iowa State Univ., LBNL, Princeton,  
RPI, Siena, Temple, UC-Berkeley,  
UC-Irvine, Univ. of Cincinnati,  
Univ. of Wisconsin-Madison,  
Univ. of Illinois-Urbana-Champaign,  
Virginia Tech., William & Mary, Yale

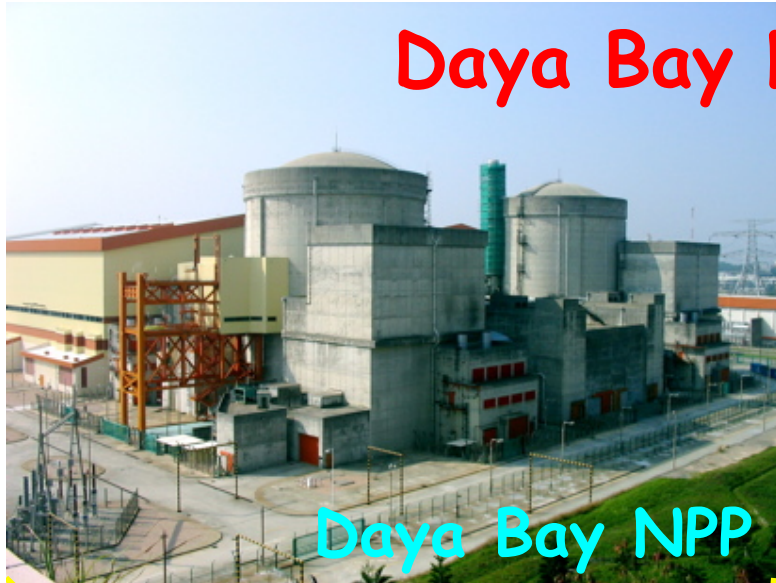
## South America (1)

Cath. Univ. of Chile

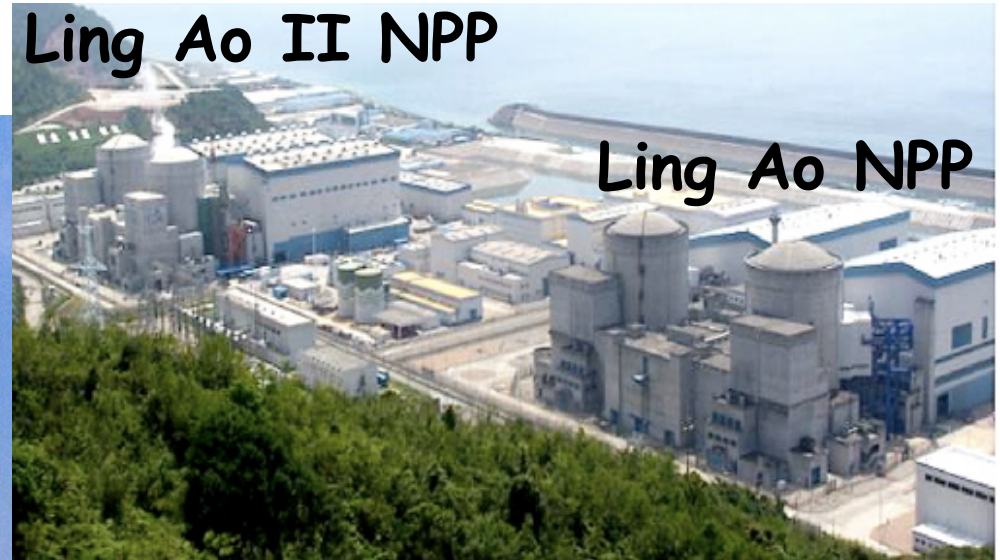




# Daya Bay Nuclear Power Complex



Ling Ao II NPP



Maximum thermal power:  $6 \times 2.9 \text{ GW}_{\text{th}}$







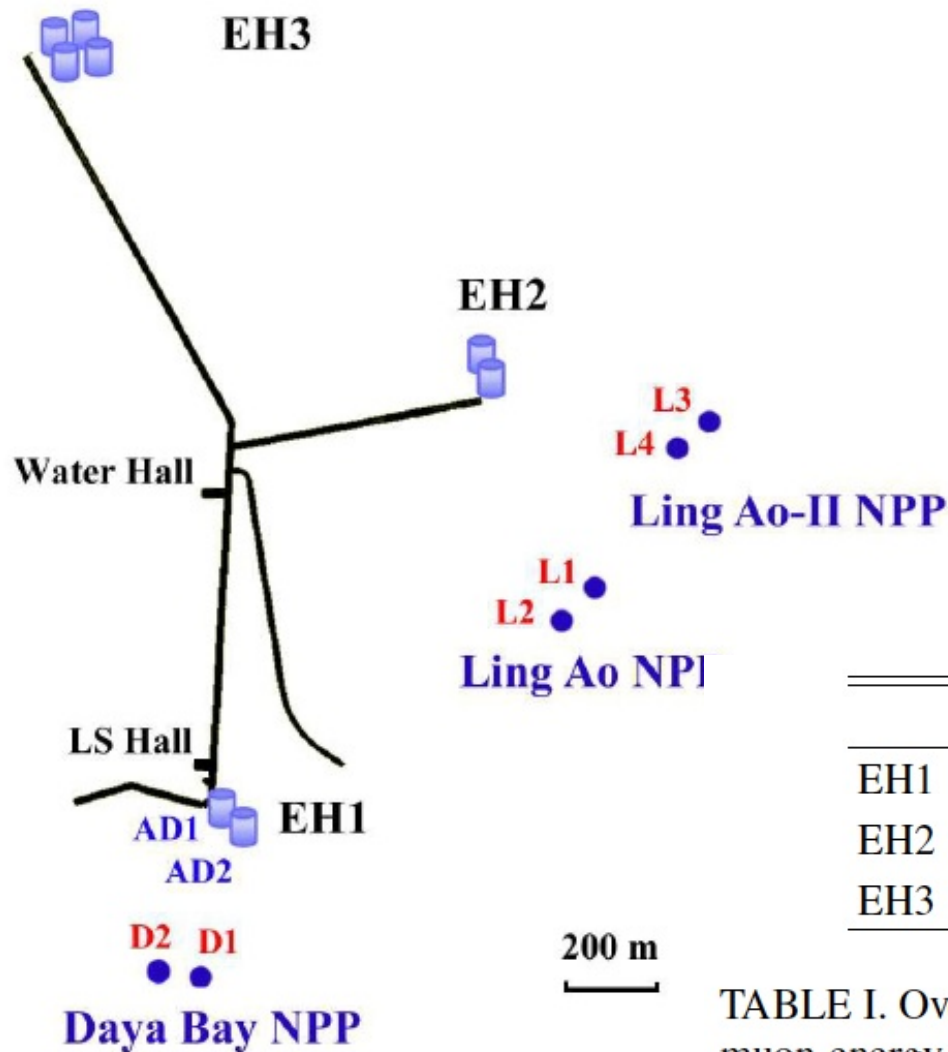
## Baselines

Detailed Survey:

- GPS above ground
- Total Station underground
- Final precision: 18mm

Validation:

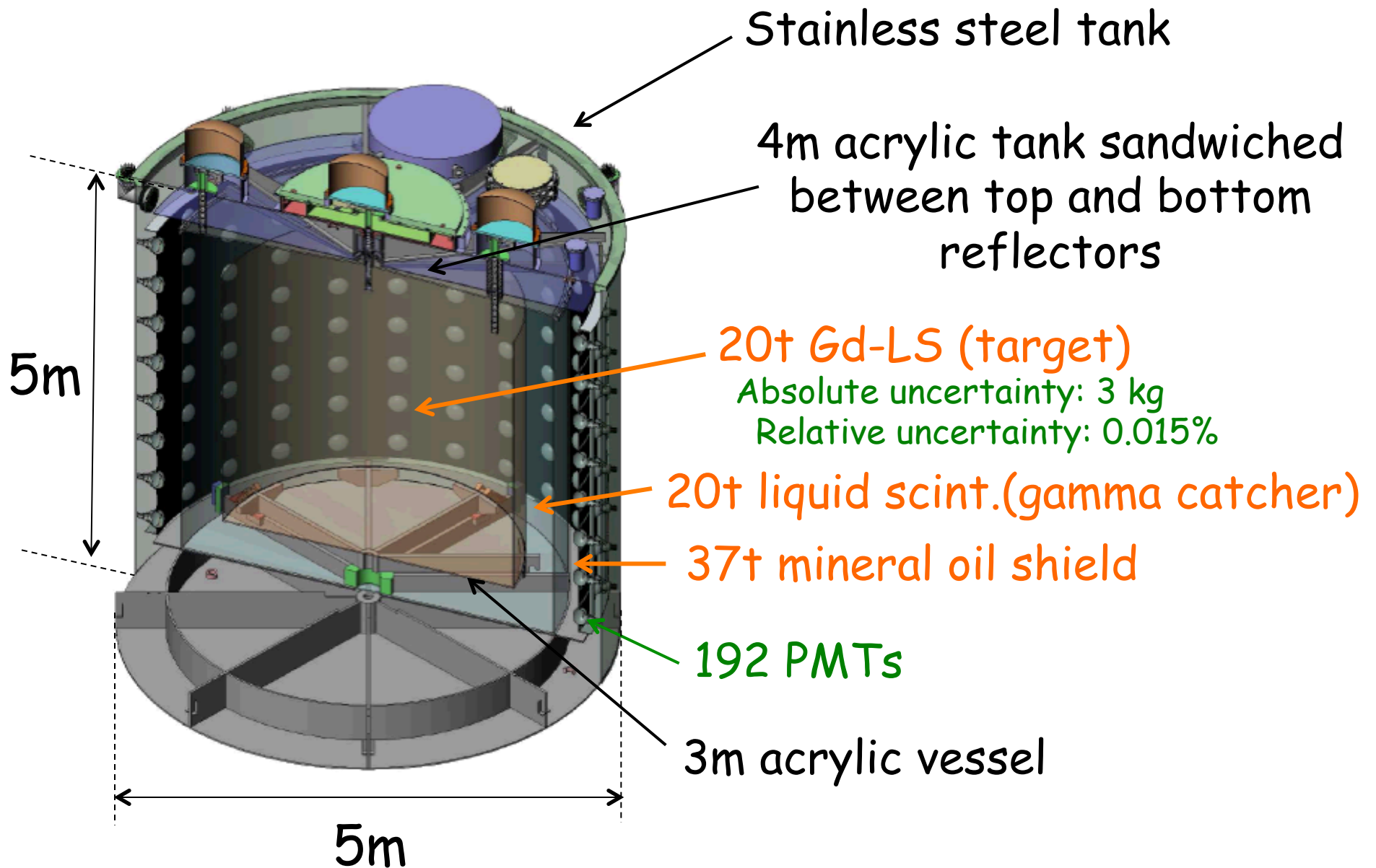
- 3 independent calculations
- Cross-check survey
- Consistent with power plant and design plans



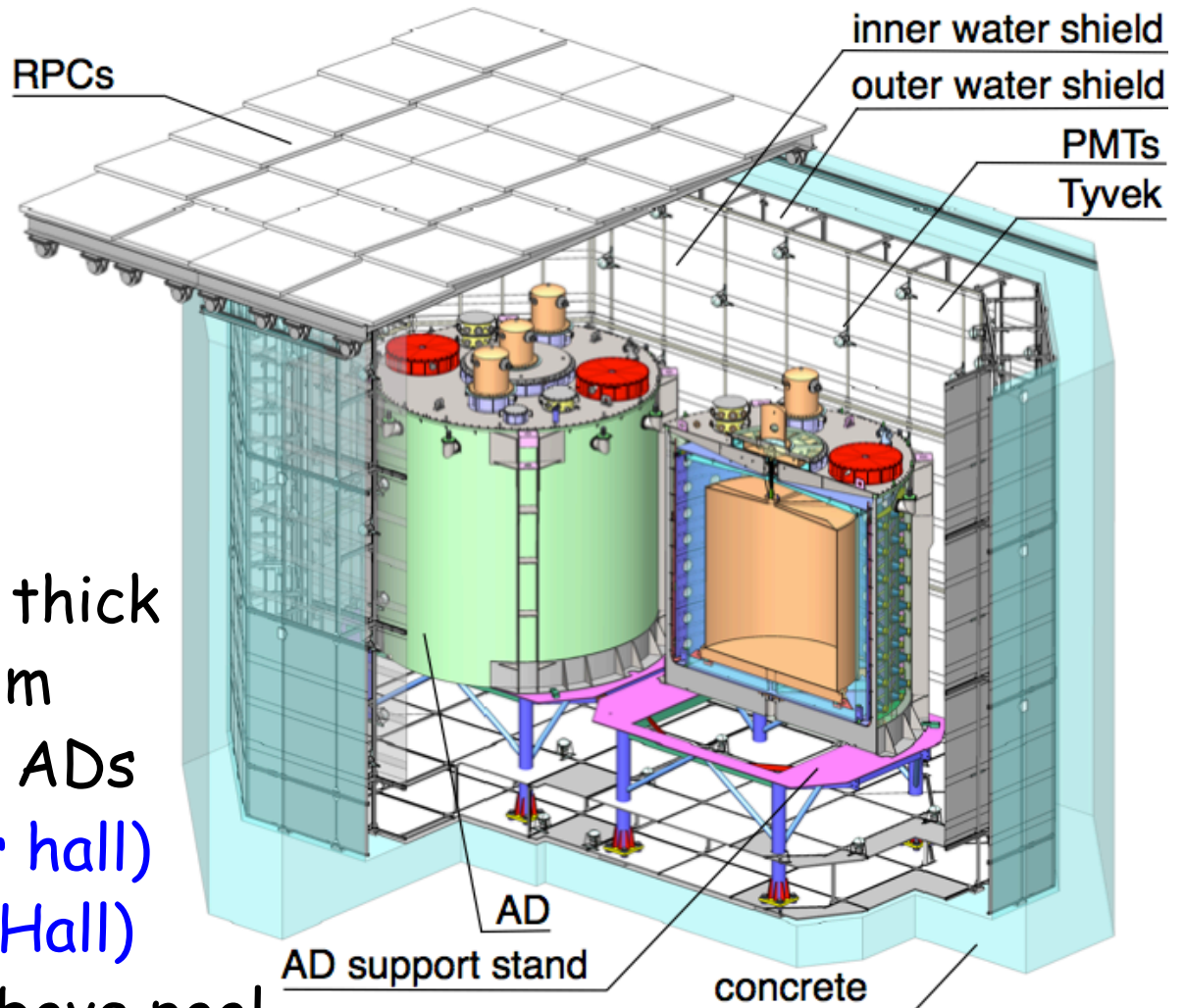
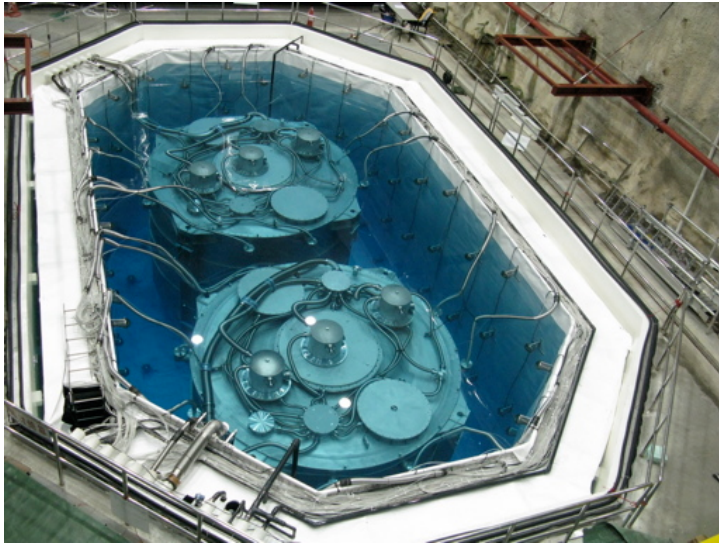
	Overburden	$R_\mu$	$E_\mu$	D1,2	L1,2	L3,4
EH1	250	1.27	57	364	857	1307
EH2	265	0.95	58	1348	480	528
EH3	860	0.056	137	1912	1540	1548

TABLE I. Overburden (m.w.e), muon rate  $R_\mu$  (Hz/m<sup>2</sup>), and average muon energy  $E_\mu$  (GeV) of the three EHs, and the distances (m) to the reactor pairs.

## Antineutrino Detector (AD)



# Muon Tagging System



- Outer Cherenkov: 1-m thick
- Inner Cherenkov: 1.5 m
- 2.5 m of water above ADs
  - 288 8" PMTs (near hall)
  - 384 8" PMTs (Far Hall)
- 4-layer RPC modules above pool
  - 54 modules (near hall)
  - 81 modules (Far Hall)

## Operation of Daya Bay

Date	Operation
24 December 2011	Data taking with 6 ADs EH1: 2 ADs EH2: 1 AD EH3: 3 ADs
28 July – 19 October 2012	Special calibration runs Installation of last 2 ADs
19 October 2012	Data taking with 8 ADs
20 Dec 2016 – 26 Jan 2017	Special calibration runs EH1 AD1 used for LS studies
26 January 2017	Data taking with 7 ADs EH1: 1 ADs EH2: 2 AD EH3: 4 ADs



# Triggers & Their Performance

Discriminator threshold:

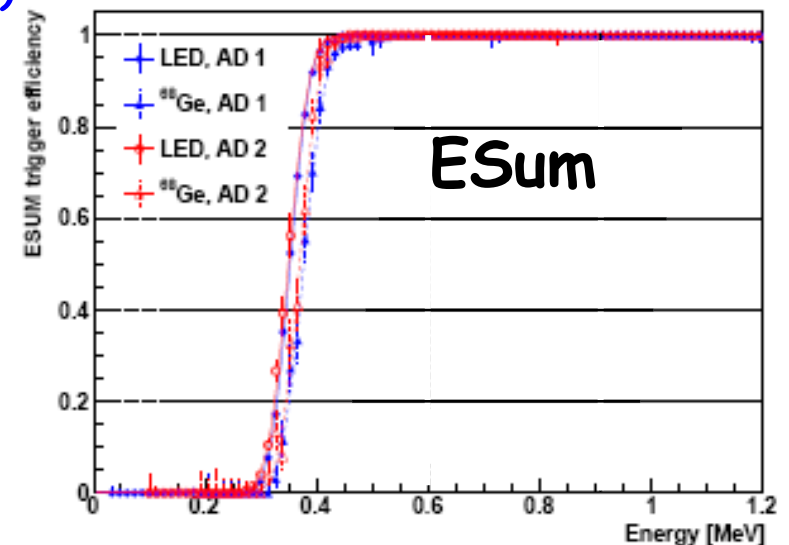
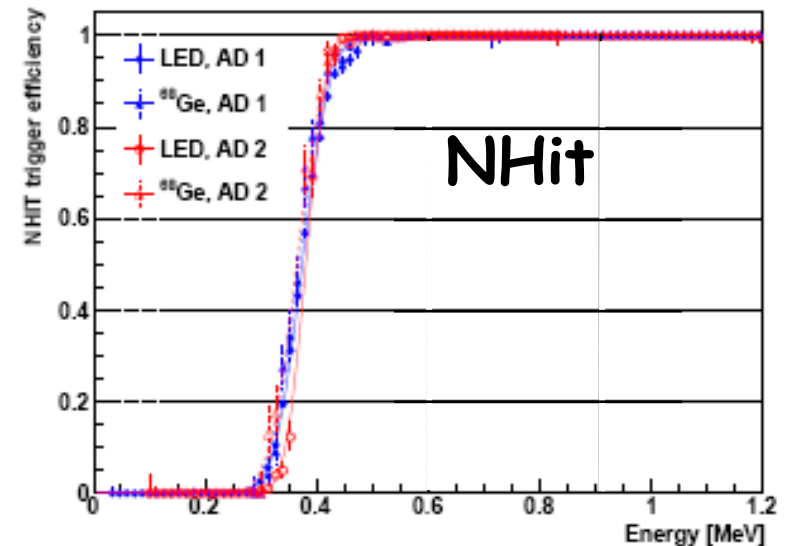
- $\sim 0.25$  p.e. for PMT signal

Triggers:

- AD:  $\geq 45$  PMTs (digital trigger)  
 $\geq 0.4$  MeV (analog trigger)
- Inner Water Cherenkov:  $\geq 6$  PMTs
- Outer Water Cherenkov:  $\geq 7$  PMTs (near)  
 $\geq 8$  PMTs (far)

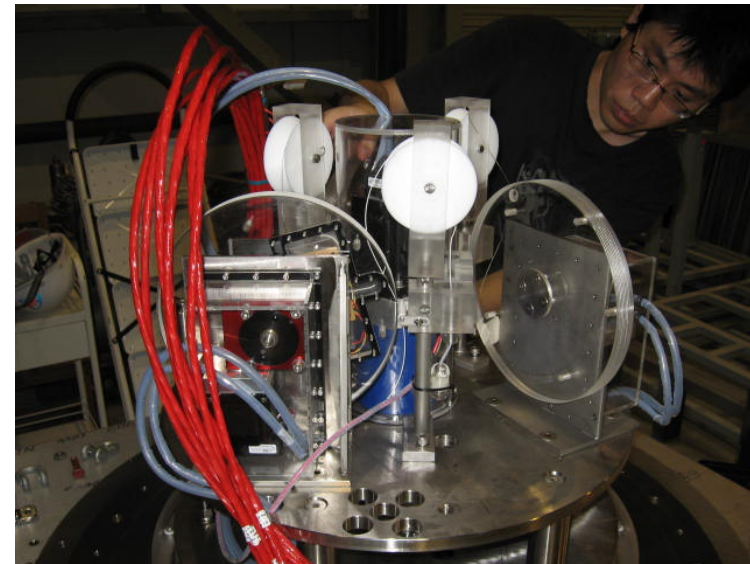
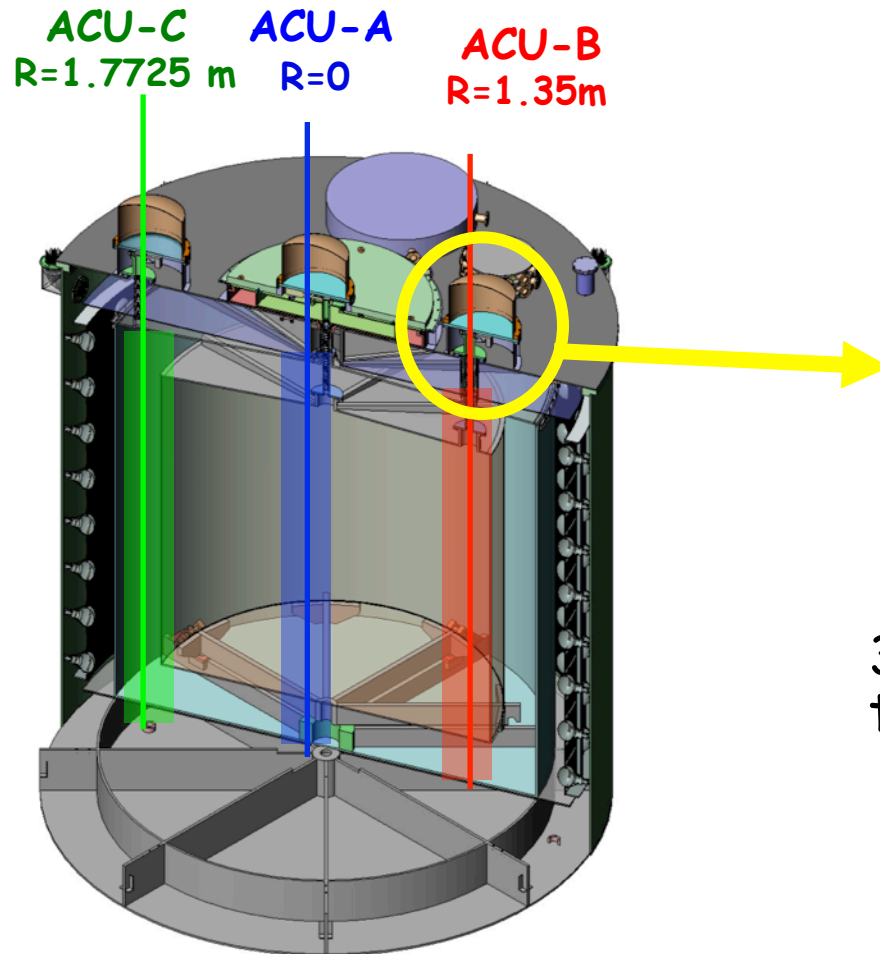
Trigger rate:

- AD:  $< 280$  Hz
- Inner Water Cherenkov:  $< 250$  Hz
- Outer Water Cherenkov:  $< 330$  Hz
- Periodic: 10 Hz



# Calibration System of ADs

3 Automatic Calibration Units (ACUs) on each detector



3 sources for each z-axis on a turntable (position accuracy < 5 mm):

- $^{68}\text{Ge}$  ( $2 \times 0.511$  MeV  $\gamma$ 's; 10 Hz)
- $^{241}\text{Am}$ - $^{13}\text{C}$  neutron source (3.5 MeV n without  $\gamma$ ; 0.7 Hz)
- $^{60}\text{Co}$  (1.173+1.332 MeV  $\gamma$ 's; 100 Hz)
- LED diffuser ball (500 Hz)

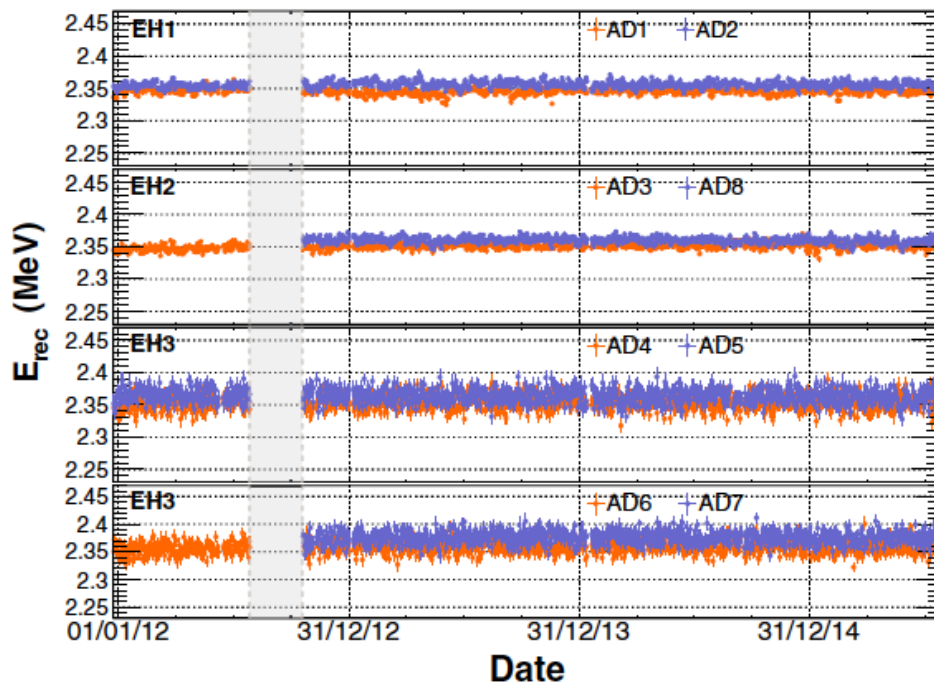
Three axes: center, edge of target, middle of gamma catcher

# Stability of Calibration

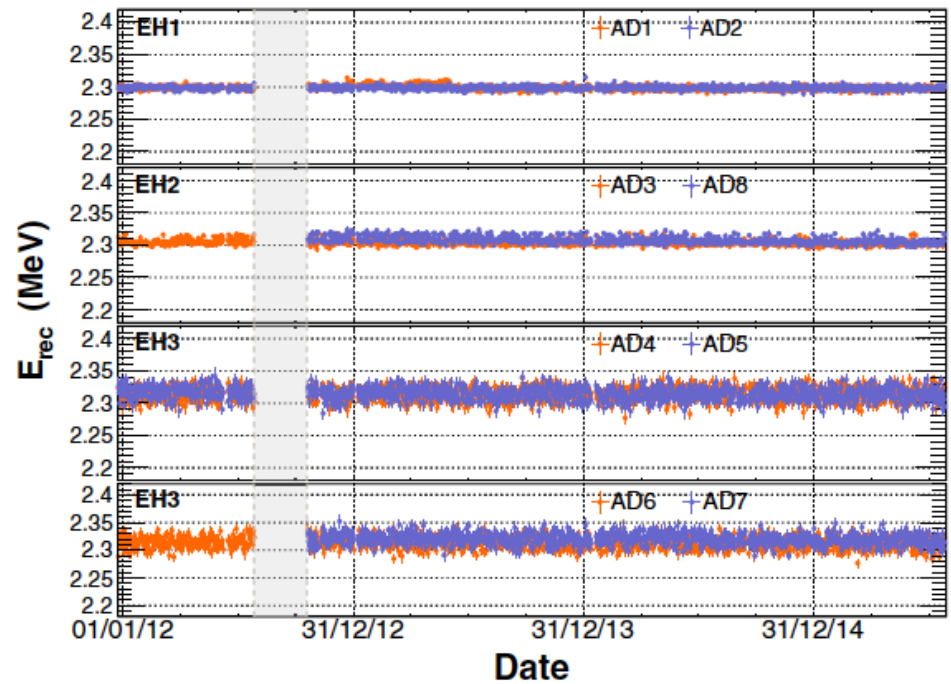
Method A: base on weekly deployed  $^{60}\text{Co}$  and Am-C sources in ACUs

Method B: use spallation neutrons daily

Method A



Method B

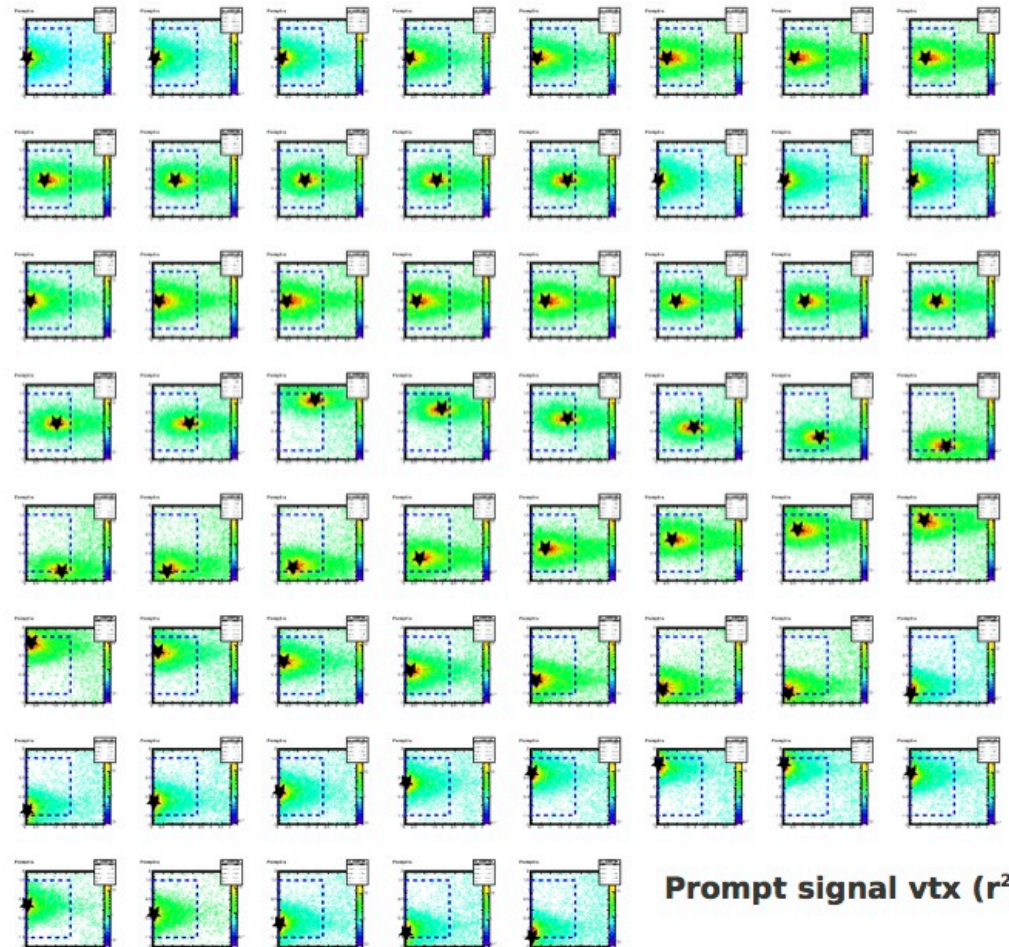


Reconstructed energy of IBD n-Gd capture is stable to better than 0.2 %



# Improving Non-uniformity Correction

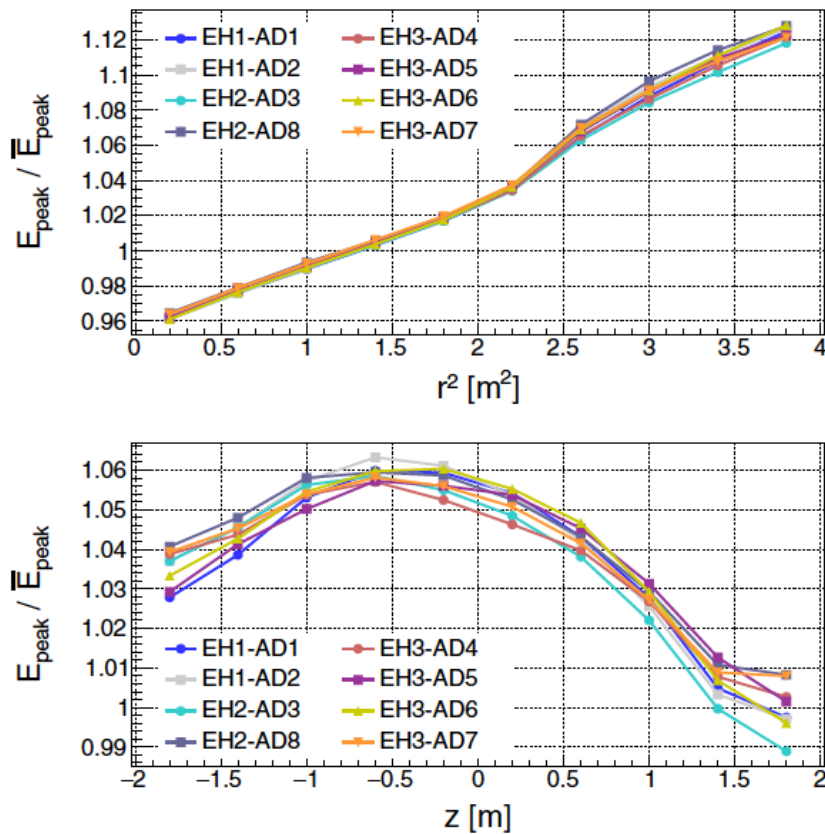
- Use the Manual Calibration Unit with a Pu-C/<sup>60</sup>Co source deployed in AD1 to map out the response throughout the fiducial volume in  $(r, \varphi, z)$ .



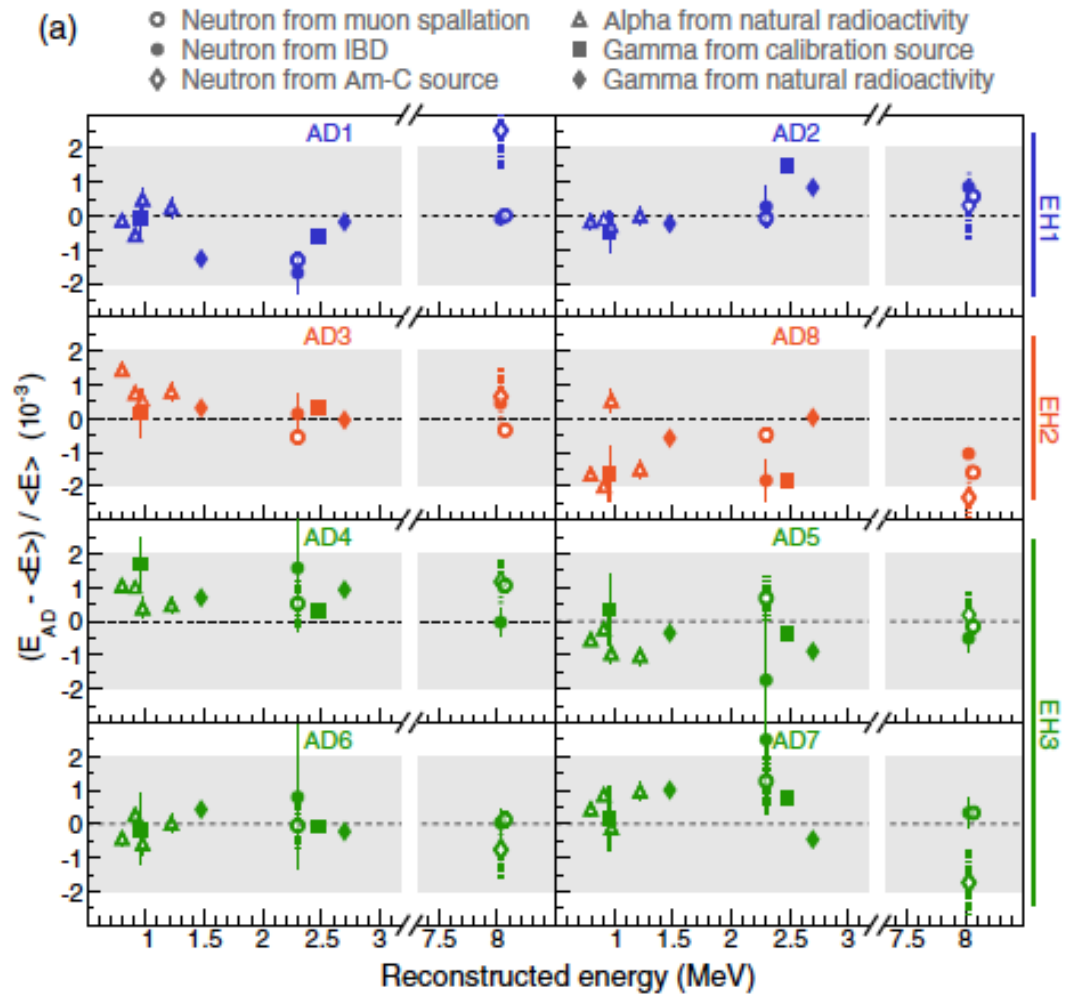
Prompt signal vtx ( $r^2, z$ )



# Energy Calibration

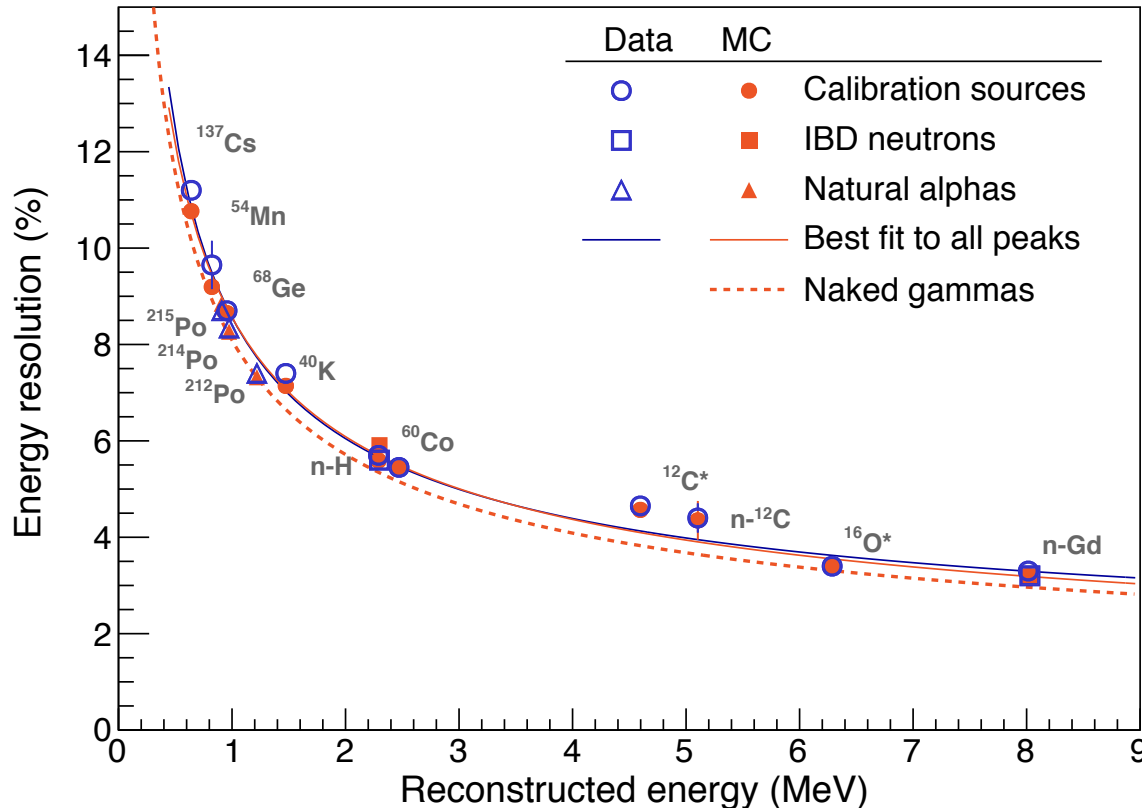


Position non-uniformity  
has been corrected for.



Relative uncertainty of  $<0.2\%$   
between detectors.

# Energy Resolution



- Calibrated primarily using  $\gamma$  sources

Parametrization:

$$\frac{\sigma_E}{E} = \sqrt{a^2 + \frac{b^2}{E} + \frac{c^2}{E^2}}$$

where

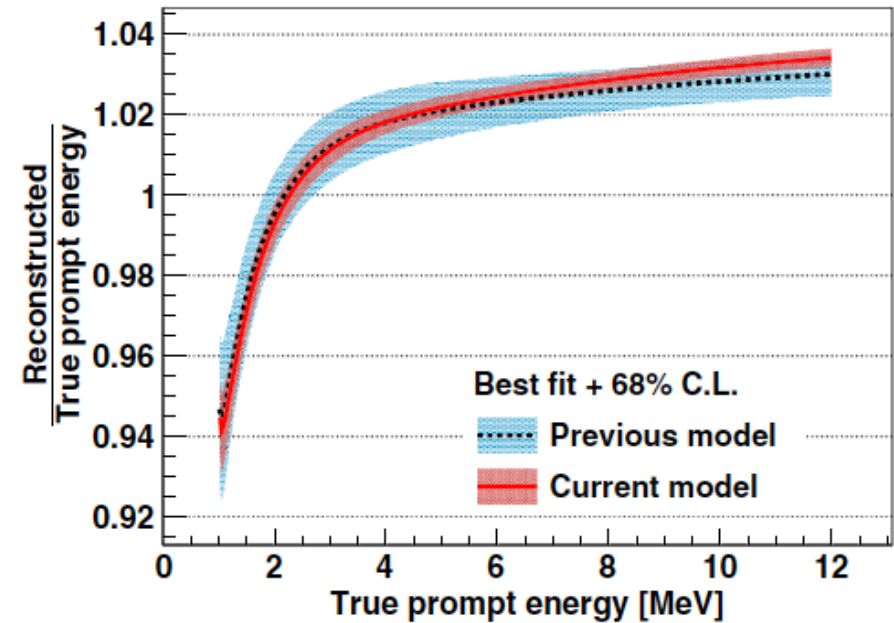
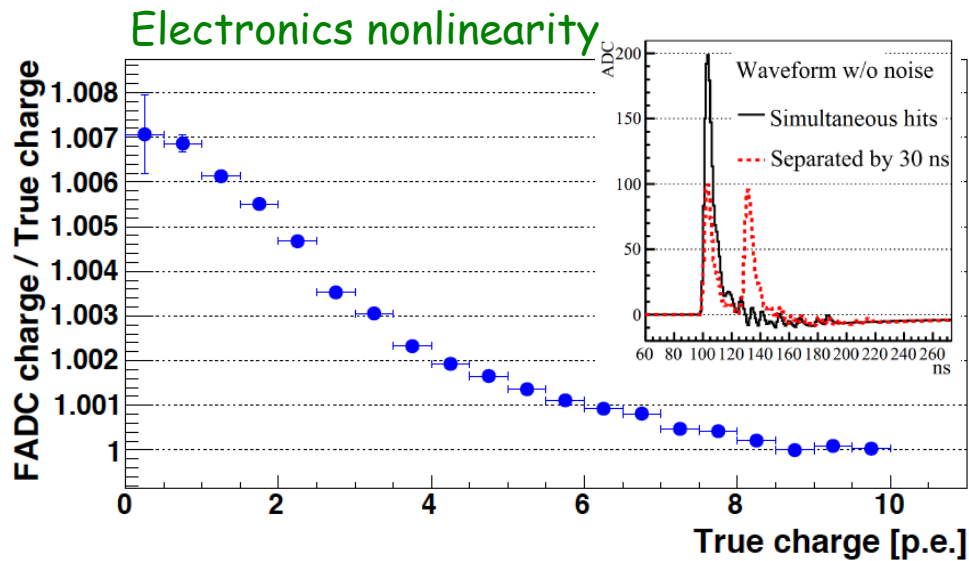
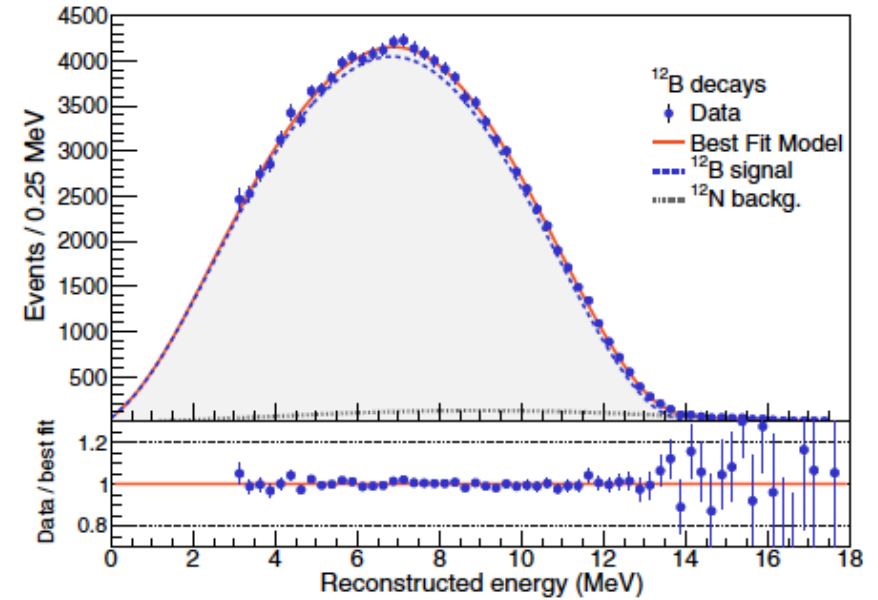
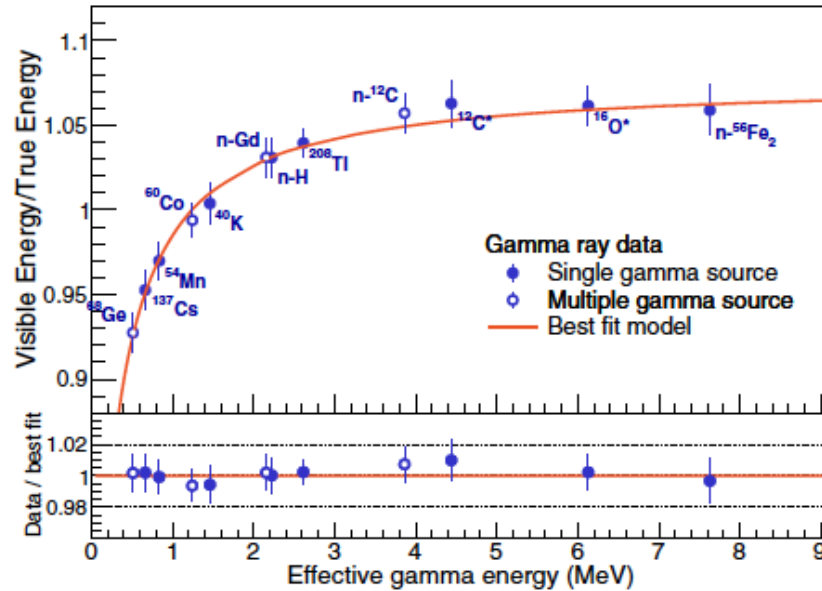
a : Spatial/temp. resolution

b : Photon statistics

c : Dark noise (constant)

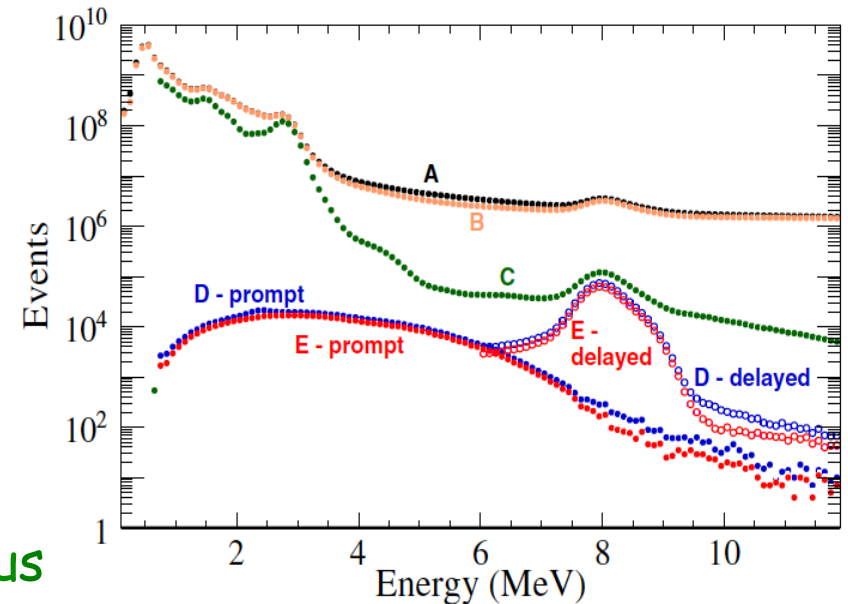
- Radioactive sources placed at the detector center
- Additional data from IBD and spallation neutrons, uniformly distributed in LS
- Alpha particles used as cross-check
  - Larger uncertainties due to different response from electronics

# Energy Nonlinearity

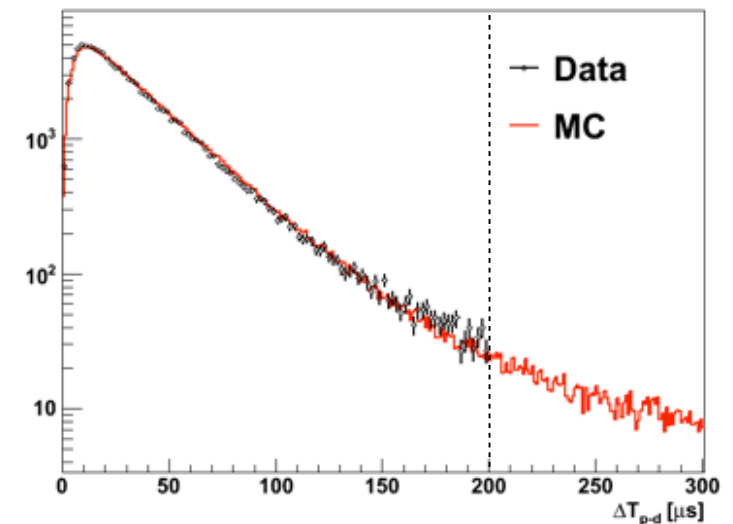


# IBD Selection with n-Capture on Gd

- Reject flashing PMTs
- Veto muons:
  - Pool muon: reject 0.6 ms
  - AD muon ( $>20$  MeV): reject 1 ms
  - AD muon ( $>2.5$  GeV): reject 1 s
- Prompt energy:  $0.7 \text{ MeV} < E_p < 12 \text{ MeV}$
- Delayed energy:  $6 \text{ MeV} < E_d < 12 \text{ MeV}$
- Neutron capture time :  $1 \mu\text{s} < \Delta t < 200 \mu\text{s}$
- Multiplicity cut: select isolated energy-pairs

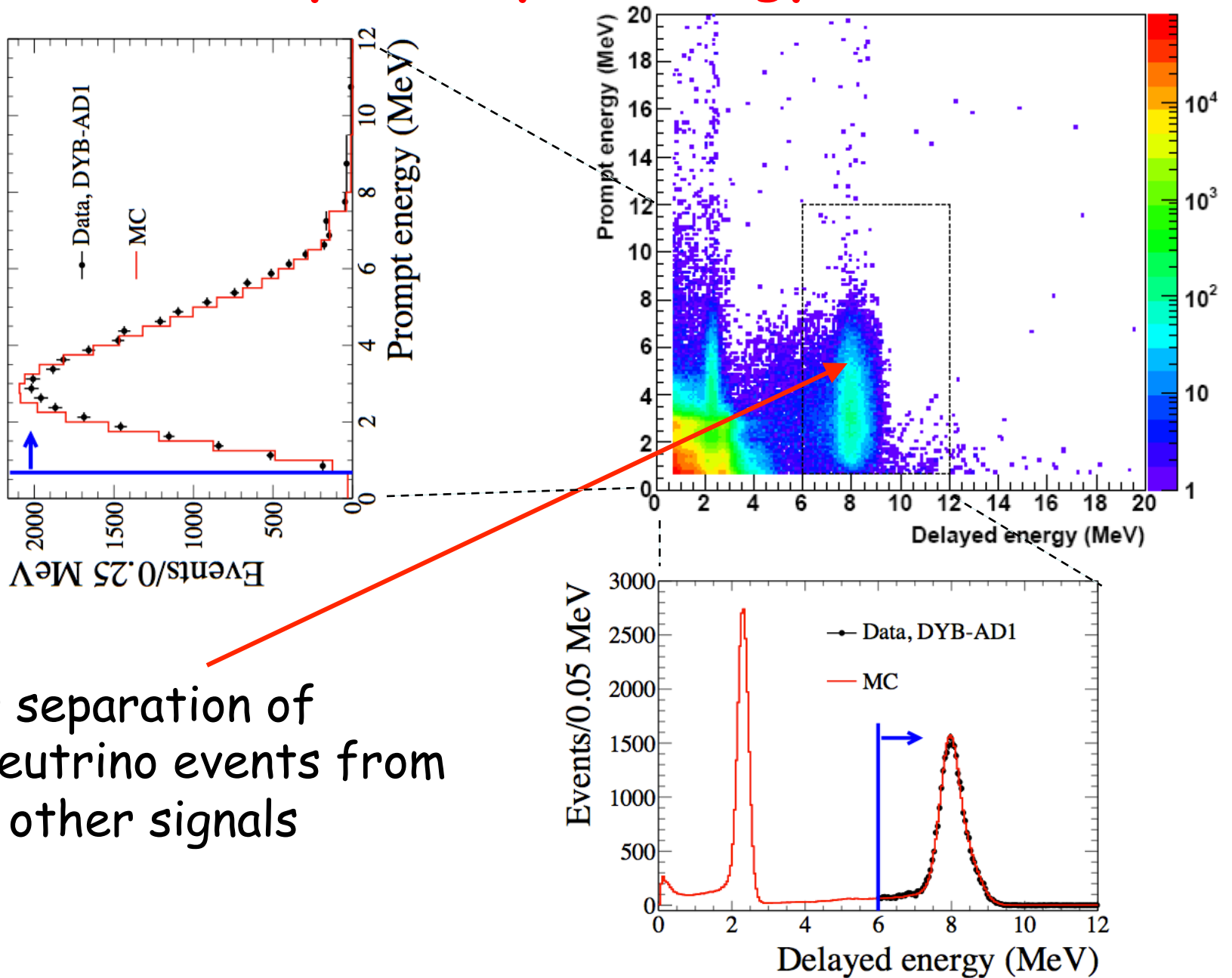


	Efficiency	Correlated	Uncorrelated
Target protons	-	0.92%	0.03%
Flasher cut	99.98%	0.01%	0.01%
Delayed energy cut	92.7%	0.97%	0.08%
Prompt energy cut	99.8%	0.10%	0.01%
Multiplicity cut		0.02%	0.01%
Capture time cut	98.7%	0.12%	0.01%
Gd capture fraction	84.2%	0.95%	0.10%
Spill-in	104.9%	1.00%	0.02%
Livetime	-	0.002%	0.01%
Combined	80.6%	1.93%	0.13%



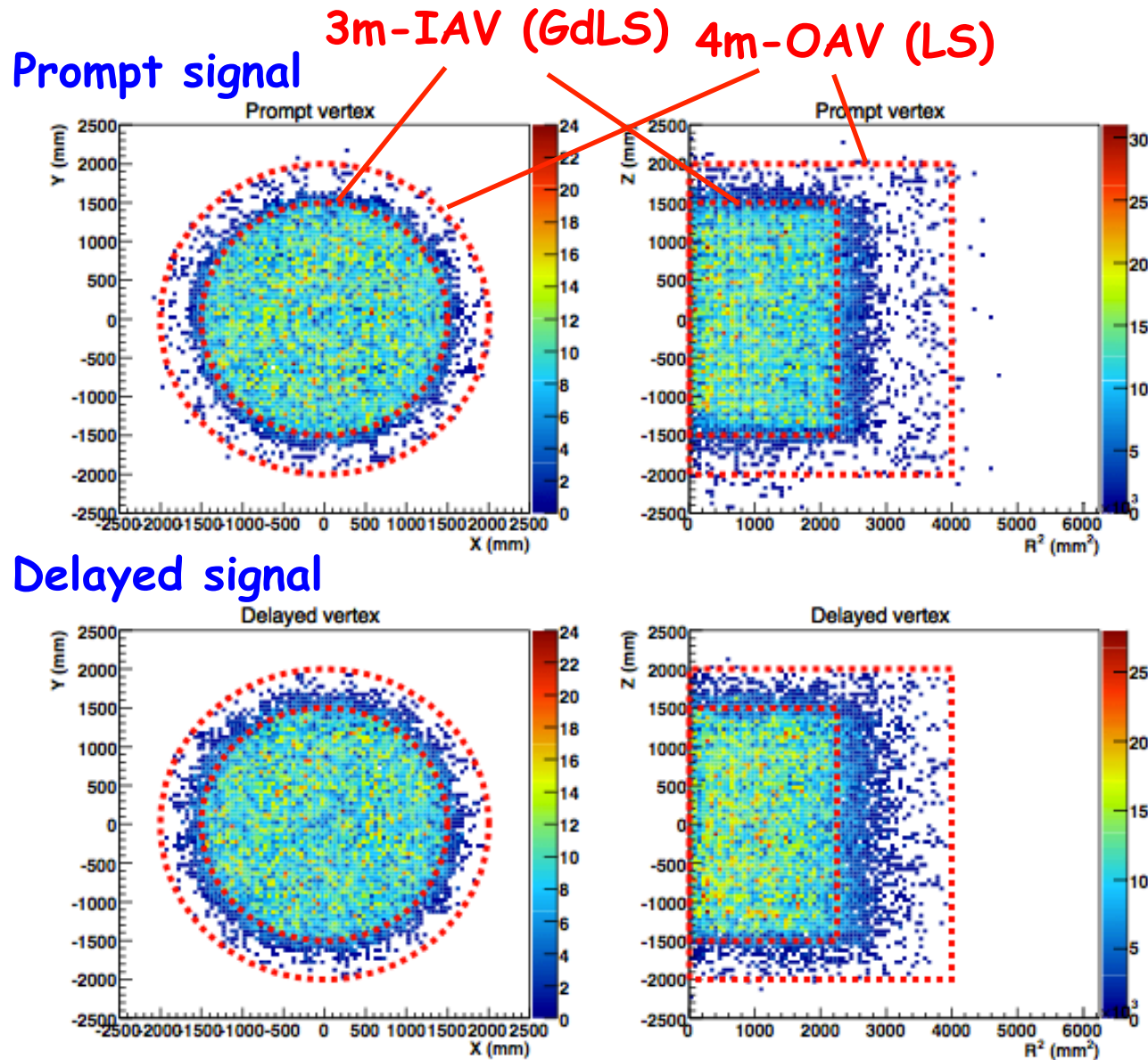


# Prompt/Delayed Energy



Clear separation of antineutrino events from most other signals

# Spatial Distributions of IBD candidates

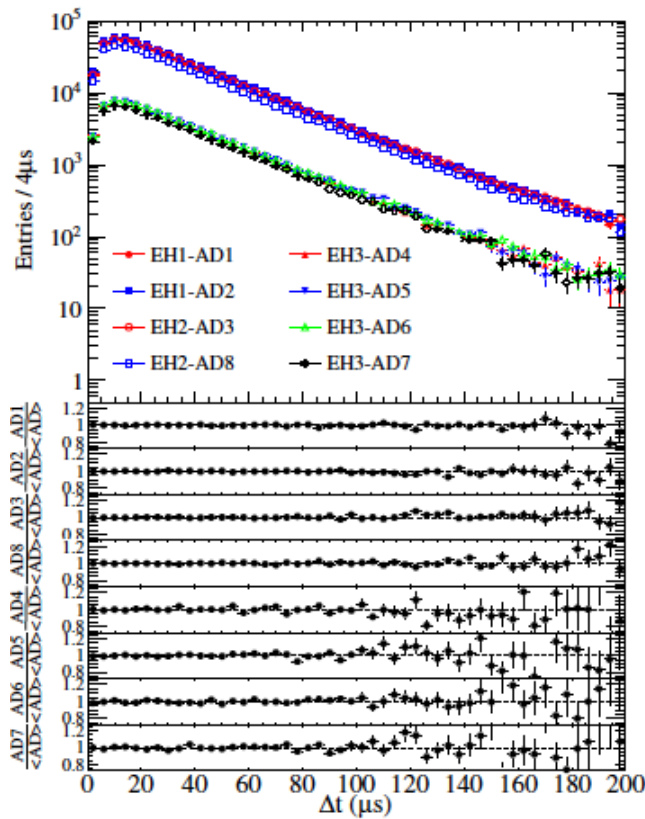


Real data  
EH1-AD1

- After applying all IBD selection cuts.
- Vertices from IBD candidates are uniformly distributed within 3m-IAV.

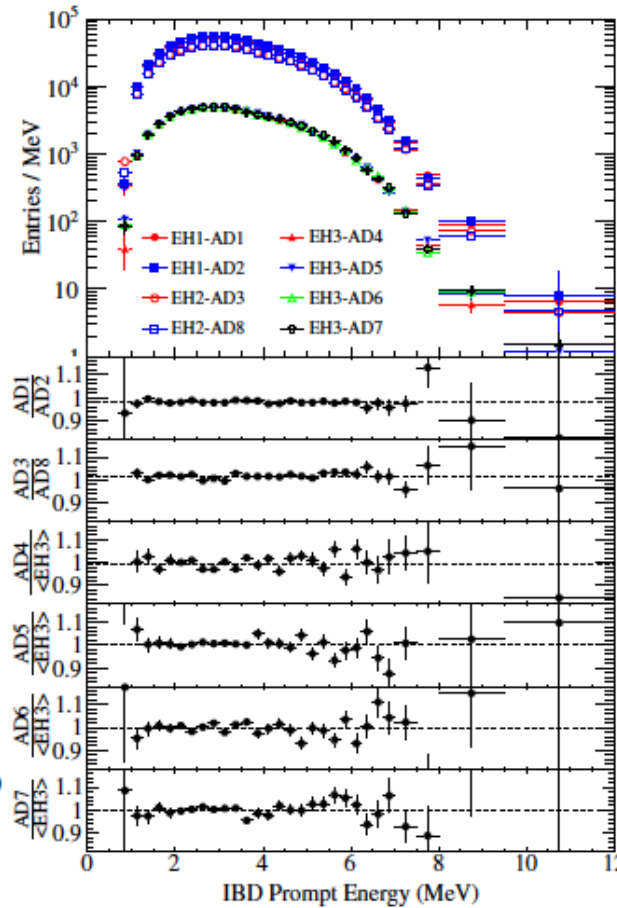
# Relative Performance of ADs

## IBD capture time

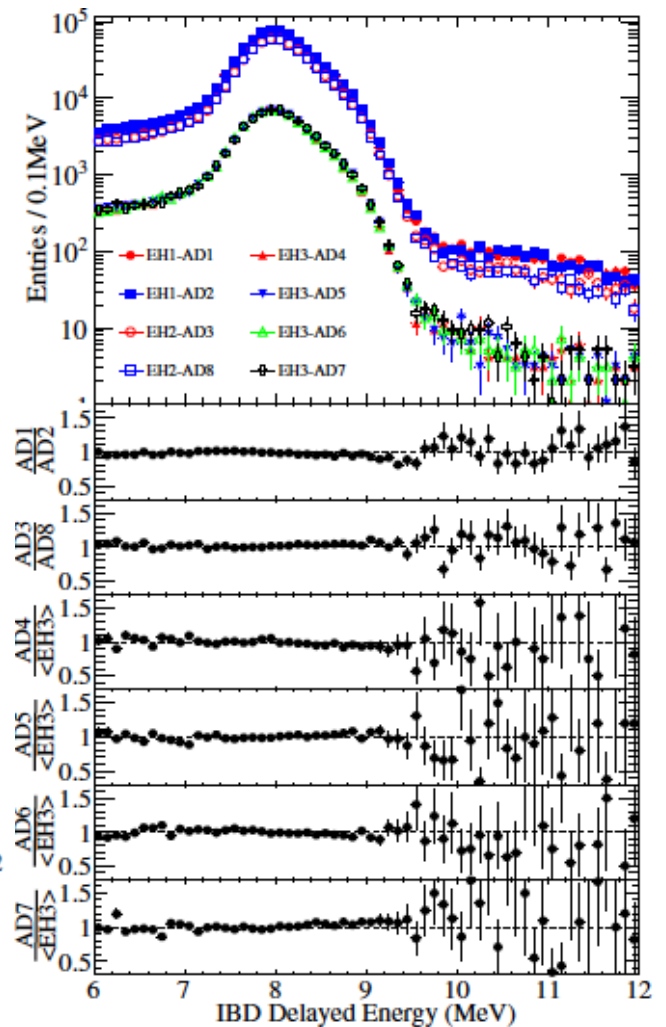


$\tau_{cap} \sim 29 \mu s$

## Prompt-energy spectra



## Delayed-energy spectra



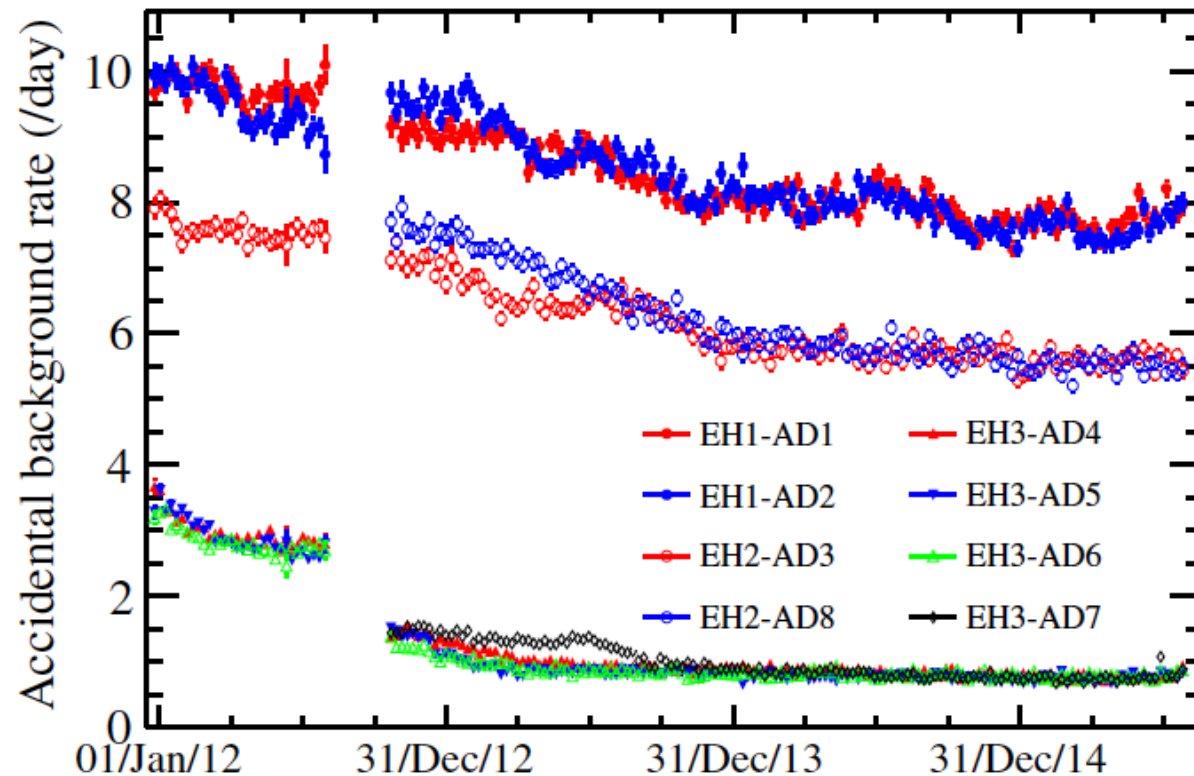
## Remaining Background

- Uncorrelated background
  - **Accidentals**: two uncorrelated events 'accidentally' pass the cuts and mimic IBD event.
- Correlated background
  - **Muon spallation products**
    - Fast neutron
    - ${}^9\text{Li}/{}^8\text{He}$
  - Correlated signals from  ${}^{241}\text{Am}-{}^{13}\text{C}$  source
  - ${}^{13}\text{C}(\alpha, n){}^{16}\text{O}$



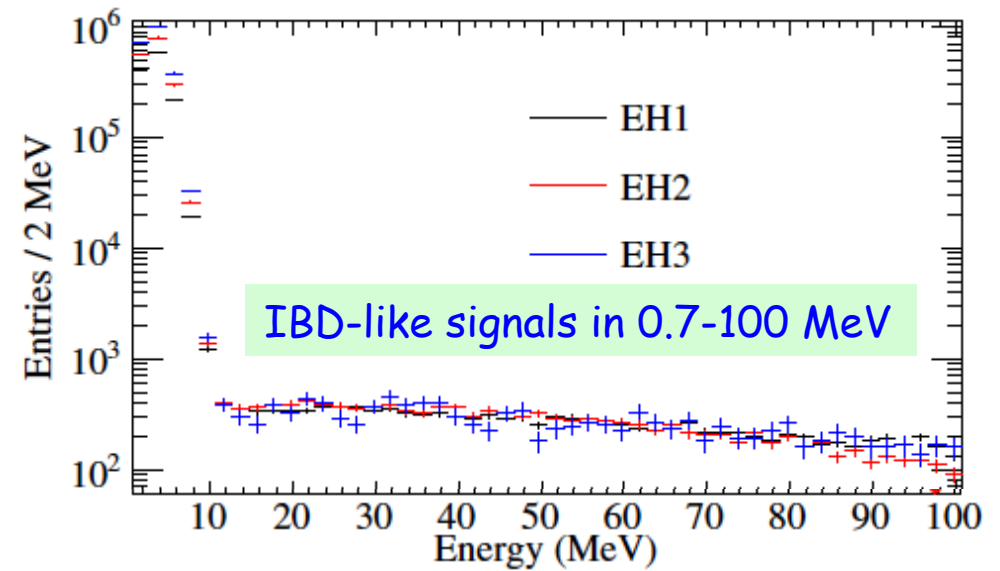
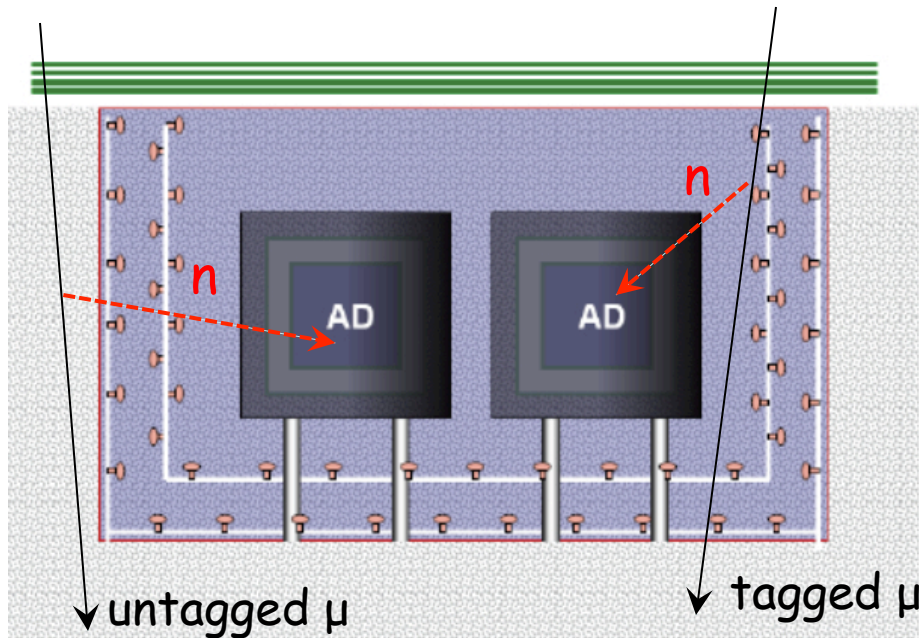
## Background: Accidentals

Two uncorrelated single signals mimic an antineutrino signal  
Rate and spectrum can be accurately predicted from singles data.



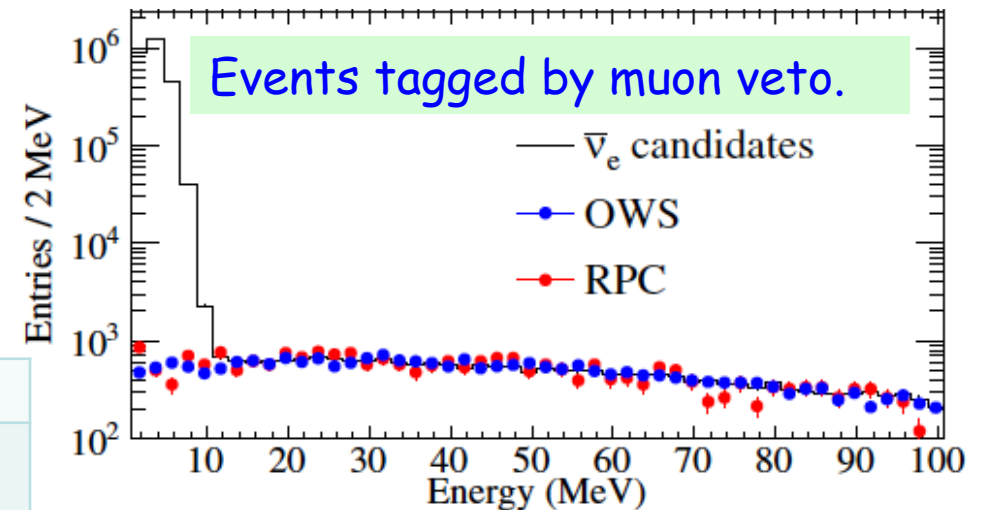
	EH1-AD1	EH1-AD2	EH2-AD3	EH2-AD8	EH3-AD4	EH3-AD5	EH3-AD6	EH3-AD7
Accidental rate(/day)	$8.46 \pm 0.09$	$8.46 \pm 0.09$	$6.29 \pm 0.06$	$6.18 \pm 0.06$	$1.27 \pm 0.01$	$1.19 \pm 0.01$	$1.20 \pm 0.01$	$0.98 \pm 0.01$

# Background: Fast Neutrons



Can mimic the IBD signal:

- Prompt: Neutron collides/stops in target
- Delayed: Neutron captures on Gd



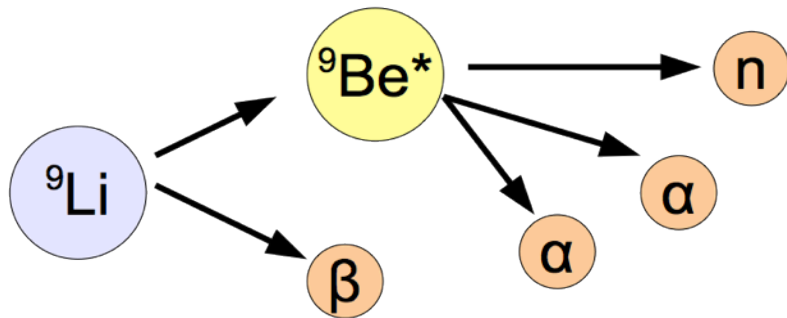
	EH1	EH2	EH3
Fast-n (/AD/day)	0.79±0.10	0.57±0.07	0.05±0.01

# Background: $^9\text{Li}/^8\text{He}$ $\beta$ -n Decays

- Generated by cosmic rays
- Long-lived
- Mimic antineutrino signal

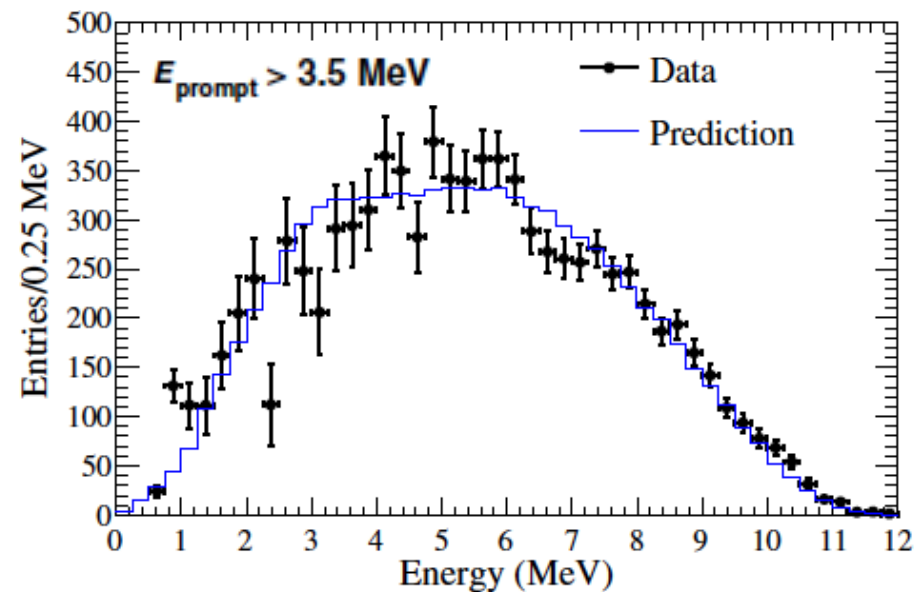
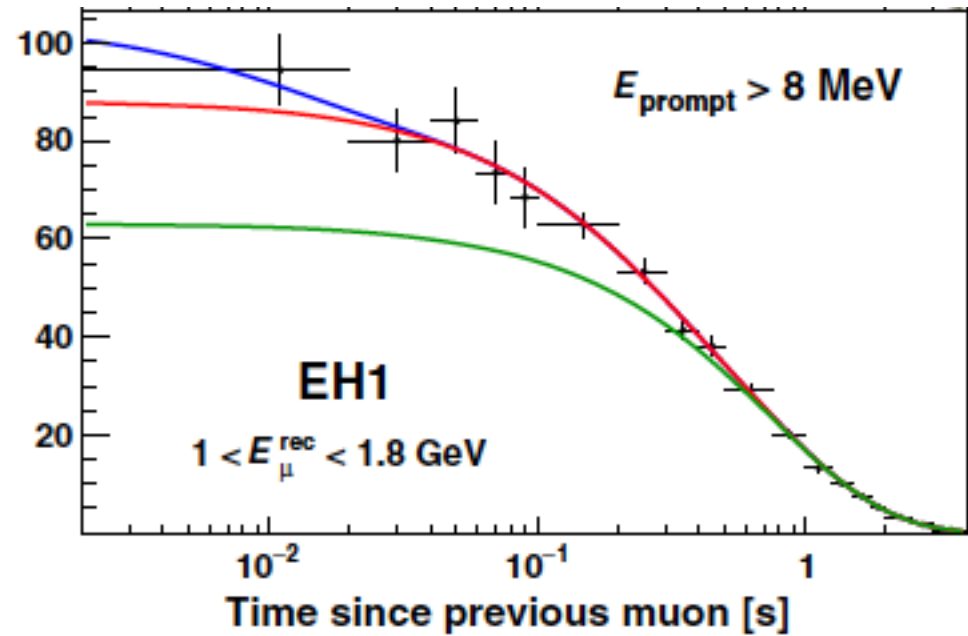
## $\beta$ -n decay:

- Prompt:  $\beta$ -decay
- Delayed: neutron capture



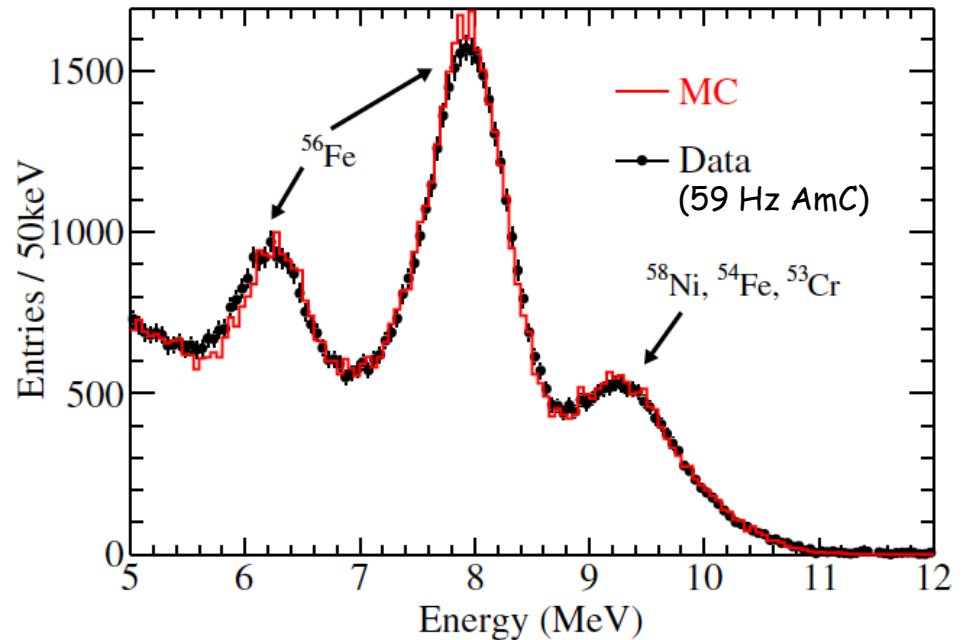
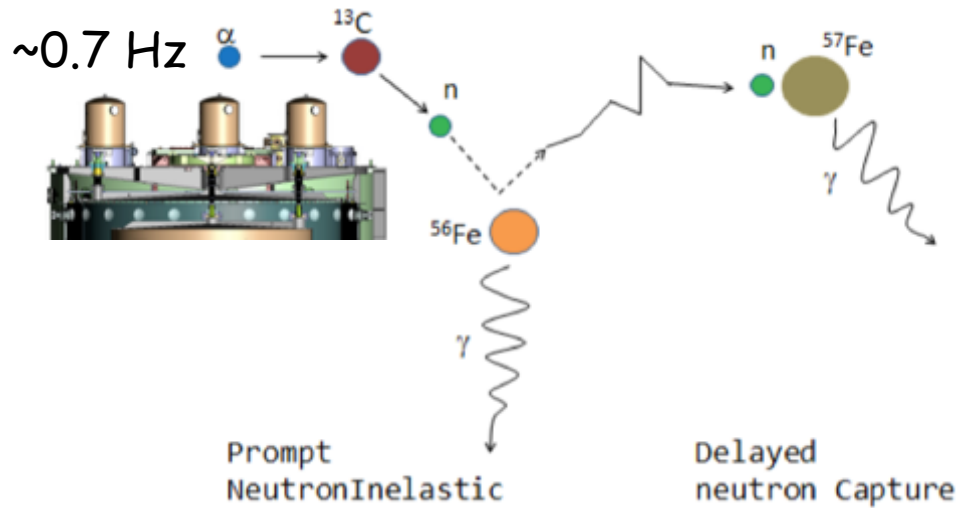
$^9\text{Li}$ :  $\tau_{1/2} = 178 \text{ ms}$ ,  $Q = 13.6 \text{ MeV}$

$^8\text{He}$ :  $\tau_{1/2} = 119 \text{ ms}$ ,  $Q = 10.6 \text{ MeV}$

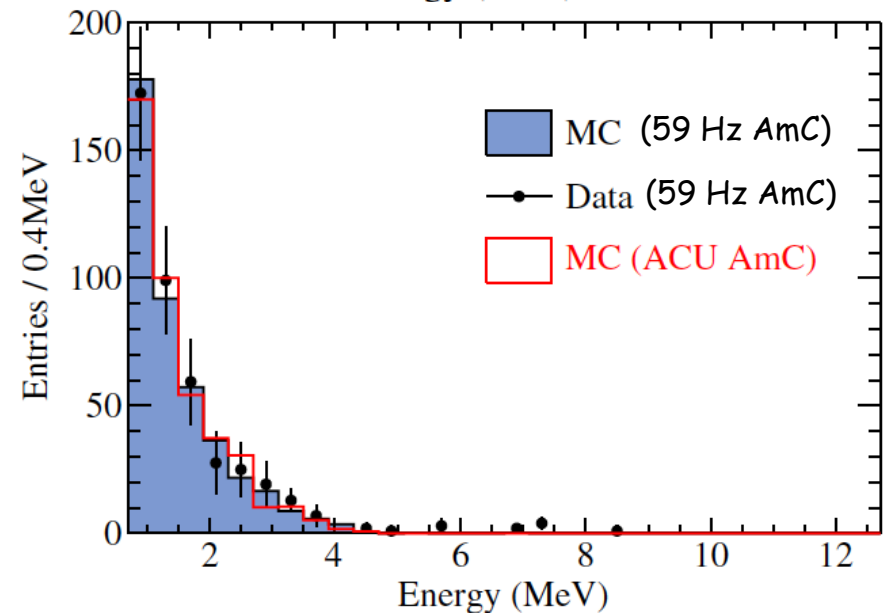


	EH1	EH2	EH3
$^9\text{Li}/^8\text{He}$ (/AD/day)	$2.38 \pm 0.66$	$1.59 \pm 0.49$	$0.19 \pm 0.08$

# $^{241}\text{Am}-^{13}\text{C}$ Source Background



- Placed a ~59 Hz Am-C source on top of EH3 AD5 for 10 days in summer of 2012 to understand this background.
- Removed this source from ACU-B and ACU-C from all ADs in EH3 in summer of 2012.





	EH1-AD1	EH1-AD2	EH2-AD3	EH2-AD8	EH3-AD4	EH3-AD5	EH3-AD6	EH3-AD7
$^{13}\text{C}(\alpha, n)^{16}\text{O}$ rate(/day)	0.08±0.04	0.07±0.04	0.05±0.03	0.07±0.04	0.05±0.03	0.05±0.03	0.05±0.03	0.05±0.03

## Summary of Signal and Background

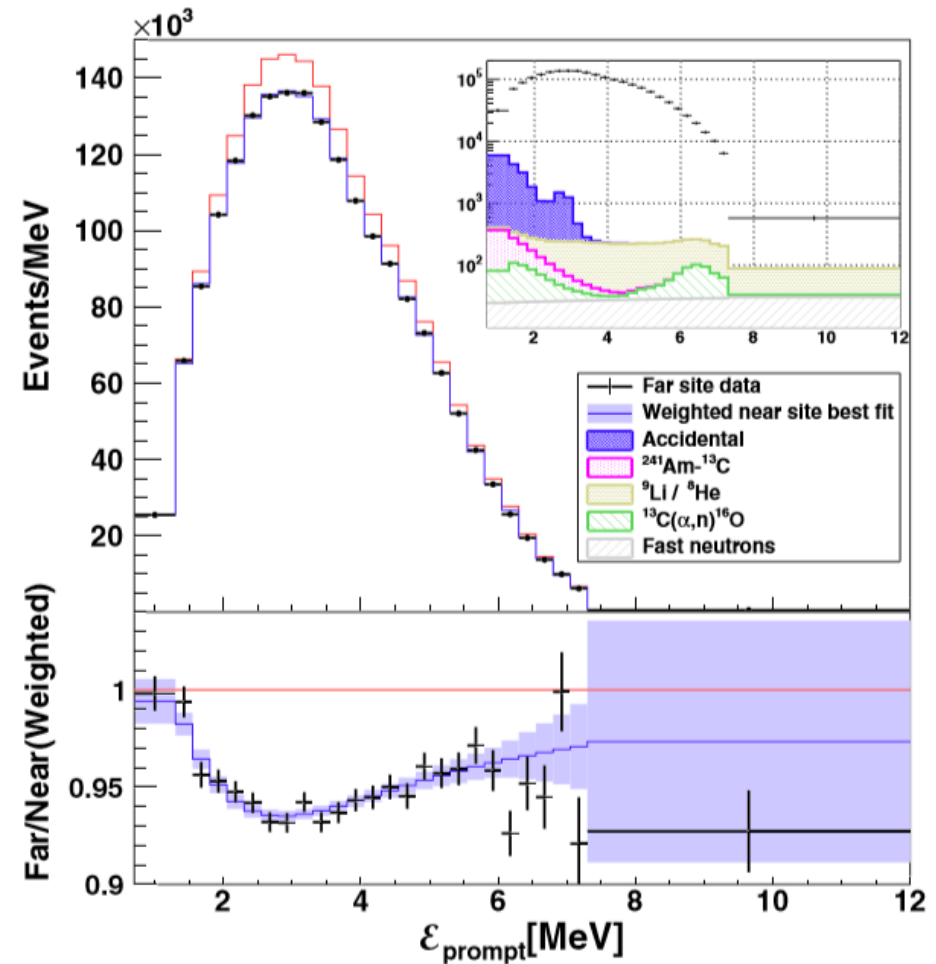
- Obtained  $\sim 4 \times 10^6$  IBD events in 1958 days with
  - 6 ADs (217 days)
  - 8 ADs (1524 days)
  - 7 ADs (217 days)

	EH1		EH2		EH3			
	AD1	AD2	AD3	AD8	AD4	AD5	AD6	AD7
$\bar{\nu}_e$ candidates	830 036	964 381	889 171	784 736	127 107	127 726	126 666	113 922
DAQ live time (days)	1536.621	1737.616	1741.235	1554.044	1739.611	1739.611	1739.611	1551.945
$\varepsilon_\mu \times \varepsilon_m$	0.8050	0.8013	0.8369	0.8360	0.9596	0.9595	0.9592	0.9595
Accidentals ( $\text{day}^{-1}$ )	$8.27 \pm 0.08$	$8.12 \pm 0.08$	$6.00 \pm 0.06$	$5.86 \pm 0.06$	$1.06 \pm 0.01$	$1.00 \pm 0.01$	$1.03 \pm 0.01$	$0.86 \pm 0.01$
Fast neutron ( $\text{AD}^{-1} \text{ day}^{-1}$ )	$0.79 \pm 0.10$		$0.57 \pm 0.07$		$0.05 \pm 0.01$			
$^9\text{Li}/^8\text{He}$ ( $\text{AD}^{-1} \text{ day}^{-1}$ )	$2.38 \pm 0.66$		$1.59 \pm 0.49$		$0.19 \pm 0.08$			
Am-C correlated ( $\text{day}^{-1}$ )	$0.17 \pm 0.07$	$0.15 \pm 0.07$	$0.14 \pm 0.06$	$0.13 \pm 0.06$	$0.06 \pm 0.03$	$0.05 \pm 0.02$	$0.05 \pm 0.02$	$0.04 \pm 0.02$
$^{13}\text{C} (\alpha, n) ^{16}\text{O}$ ( $\text{day}^{-1}$ )	$0.08 \pm 0.04$	$0.06 \pm 0.03$	$0.04 \pm 0.02$	$0.06 \pm 0.03$	$0.04 \pm 0.02$	$0.04 \pm 0.02$	$0.04 \pm 0.02$	$0.04 \pm 0.02$
$\bar{\nu}_e$ rate ( $\text{day}^{-1}$ )	$659.36 \pm 1.00$	$681.09 \pm 0.98$	$601.83 \pm 0.82$	$595.82 \pm 0.85$	$74.75 \pm 0.23$	$75.19 \pm 0.23$	$74.56 \pm 0.23$	$75.33 \pm 0.24$

- Consistent rates for side-by-side detectors

# Prompt-energy Spectrum in 2018

- 3.5 millions inverse beta decay candidates (IBDs) were detected in the near halls.
- 0.5 million IBDs were observed in the far hall.
- Daily rate was  $\sim 2500$  IBDs in the near halls and  $\sim 300$  IBDs in the far hall.
- $\leq 2\%$  backgrounds.



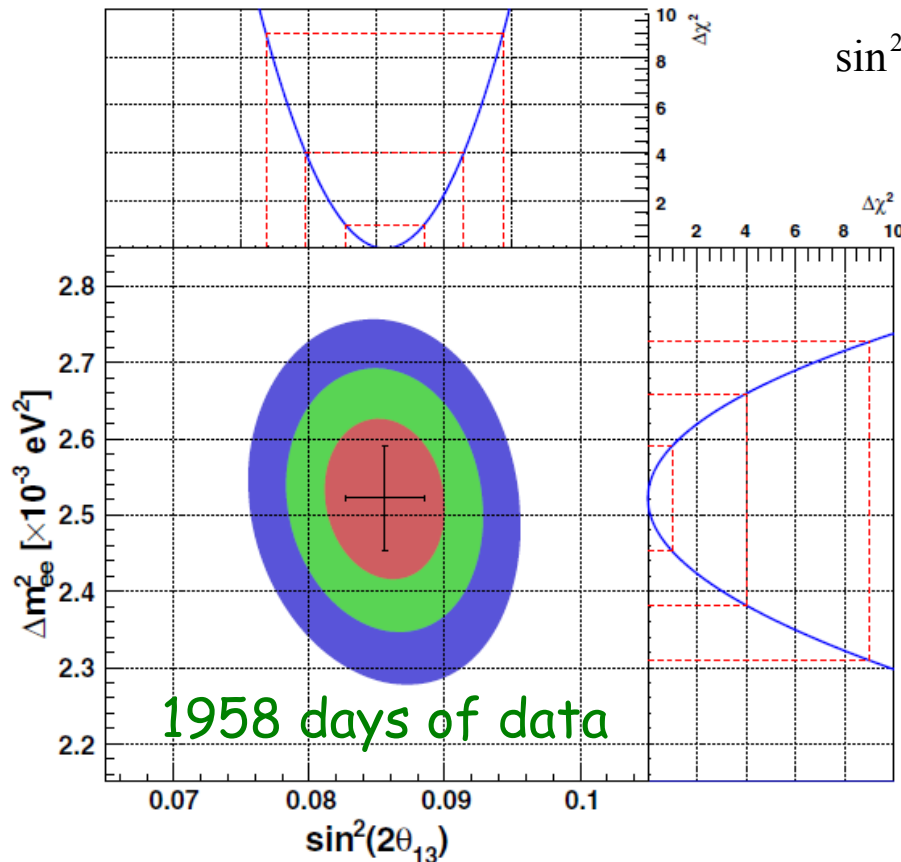
1958 days of data



# Most Precise $\sin^2 2\theta_{13}$ & $\Delta m_{ee}^2$

$$P(\bar{\nu}_e \rightarrow \bar{\nu}_e) = 1 - \sin^2 2\theta_{13} \sin^2 \left( \frac{\Delta m_{ee}^2 L}{4E} \right) + \cos^4 \theta_{13} \sin^2 2\theta_{12} \sin^2 \left( \frac{\Delta m_{21}^2 L}{4E} \right)$$

$$\sin^2 \left( \frac{\Delta m_{ee}^2 L}{4E} \right) \leftrightarrow \cos^2 \theta_{12} \sin^2 \left( \frac{\Delta m_{31}^2 L}{4E} \right) + \sin^2 \theta_{12} \sin^2 \left( \frac{\Delta m_{32}^2 L}{4E} \right)$$



$$\chi^2/\text{NDF} = 148/154$$

PRL 121 (2018) 241805

$$\sin^2 2\theta_{13} = 0.0856 \pm 0.0029$$

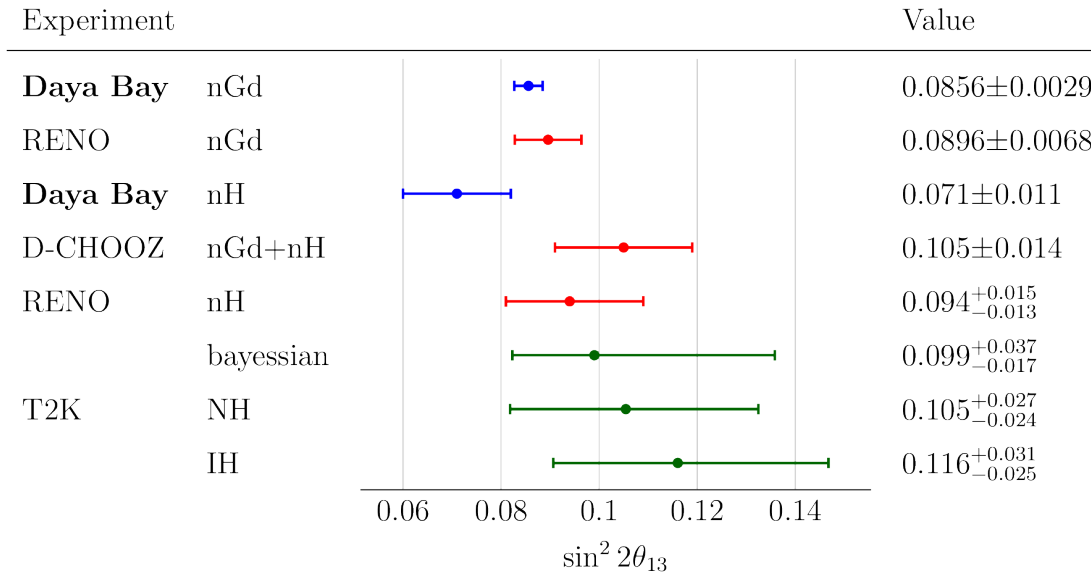
$$|\Delta m_{ee}^2| = (2.522^{+0.068}_{-0.070}) \times 10^{-3} \text{ eV}^2$$

$$\Delta m_{32}^2 = +(2.471^{+0.068}_{-0.070}) \times 10^{-3} \text{ eV}^2 \text{ (NH)}$$

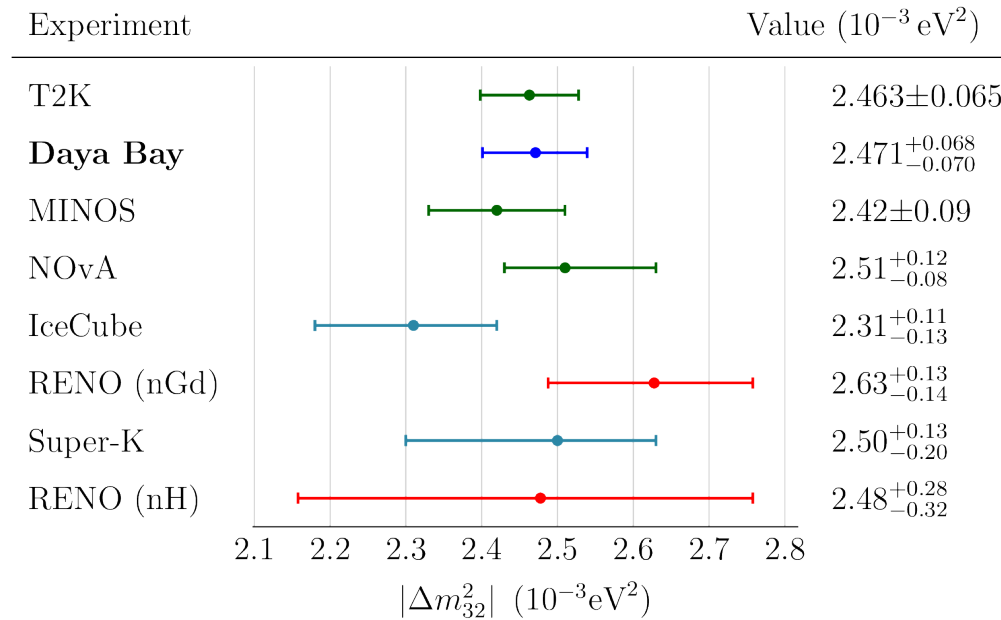
$$\Delta m_{32}^2 = -(2.575^{+0.068}_{-0.070}) \times 10^{-3} \text{ eV}^2 \text{ (IH)}$$

# $\sin^2 2\theta_{13}$ & $\Delta m^2_{32}$ : Global Landscape

$\sin^2 2\theta_{13}$

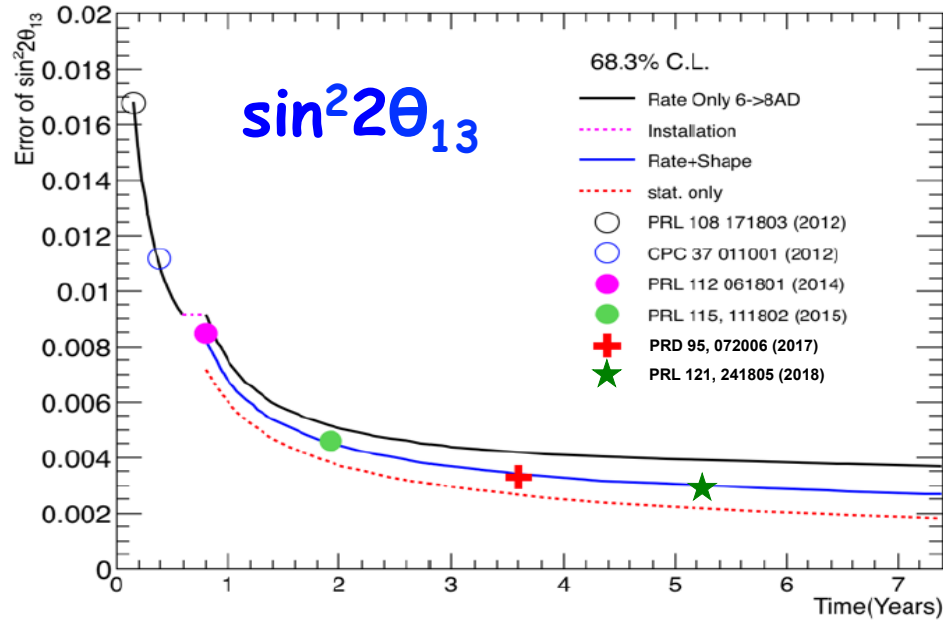


$\Delta m^2_{32}$   
Normal mass  
hierarchy

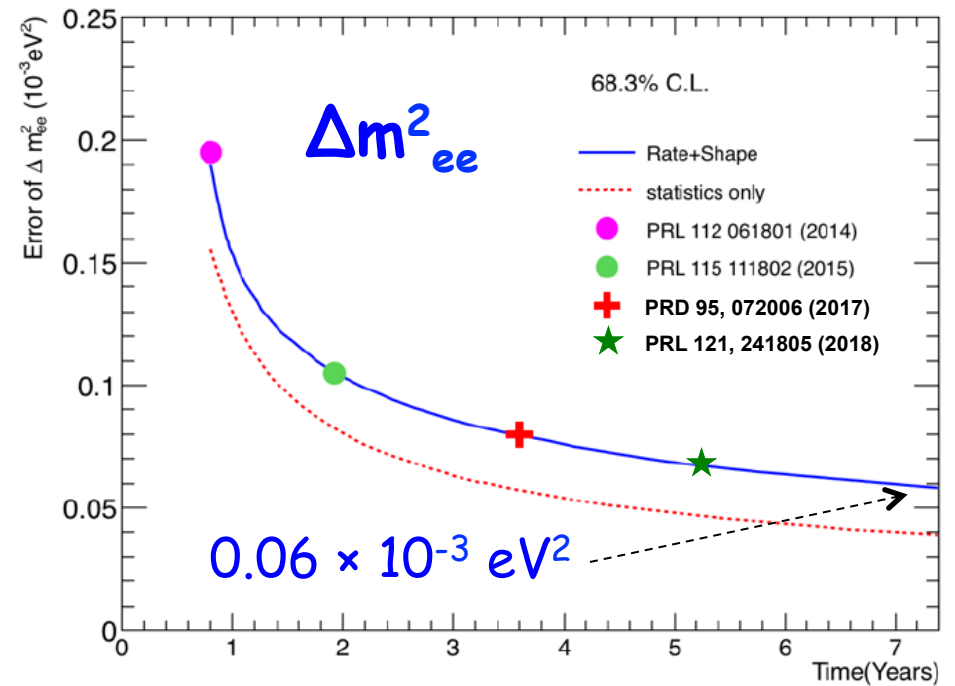


$|\Delta m^2_{32}|$  obtained with  $\nu_e$  &  $\nu_\mu$  agree, supporting 3-flavor paradigm

# Outlook of $\sin^2 2\theta_{13}$ & $\Delta m^2_{ee}$



0.0025

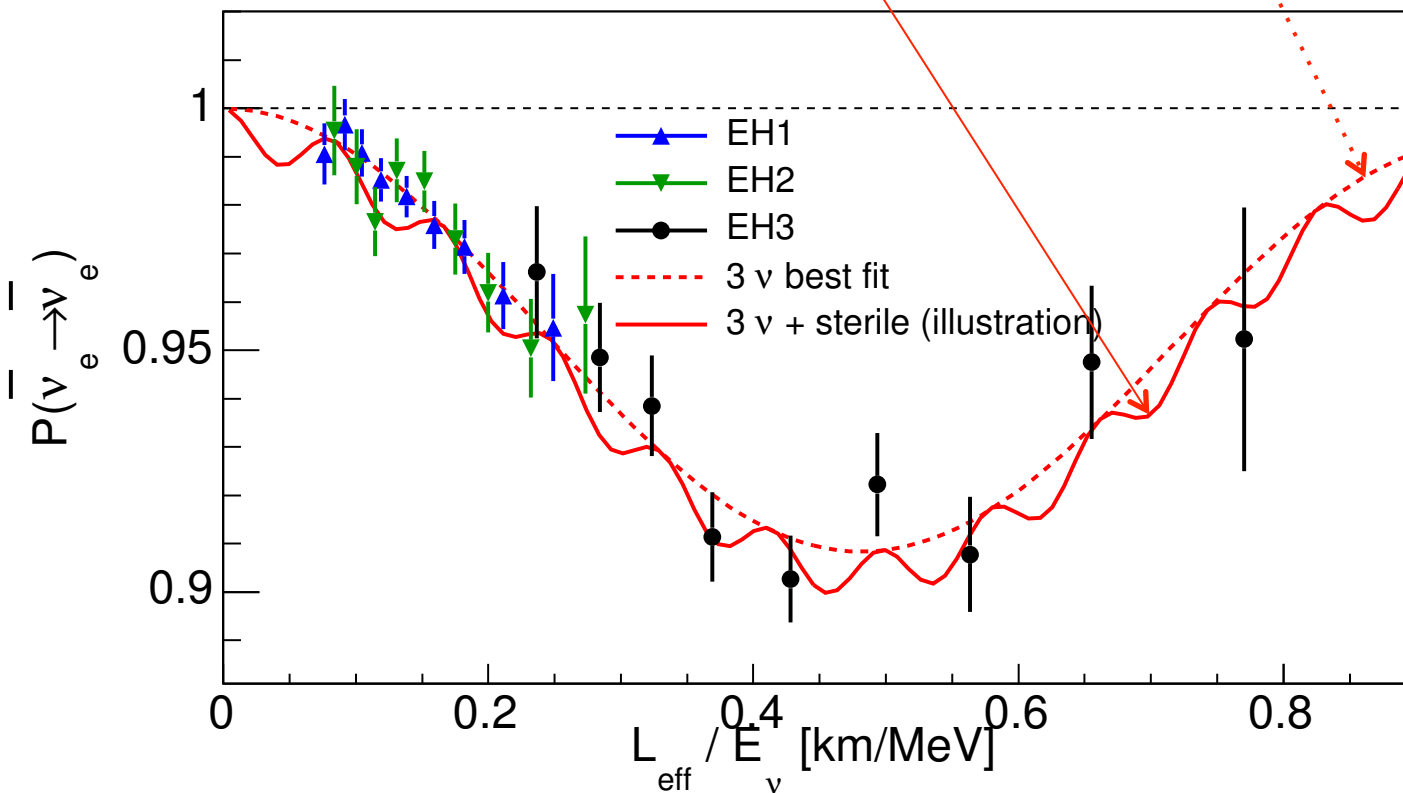




# Sterile-active Neutrino Oscillation

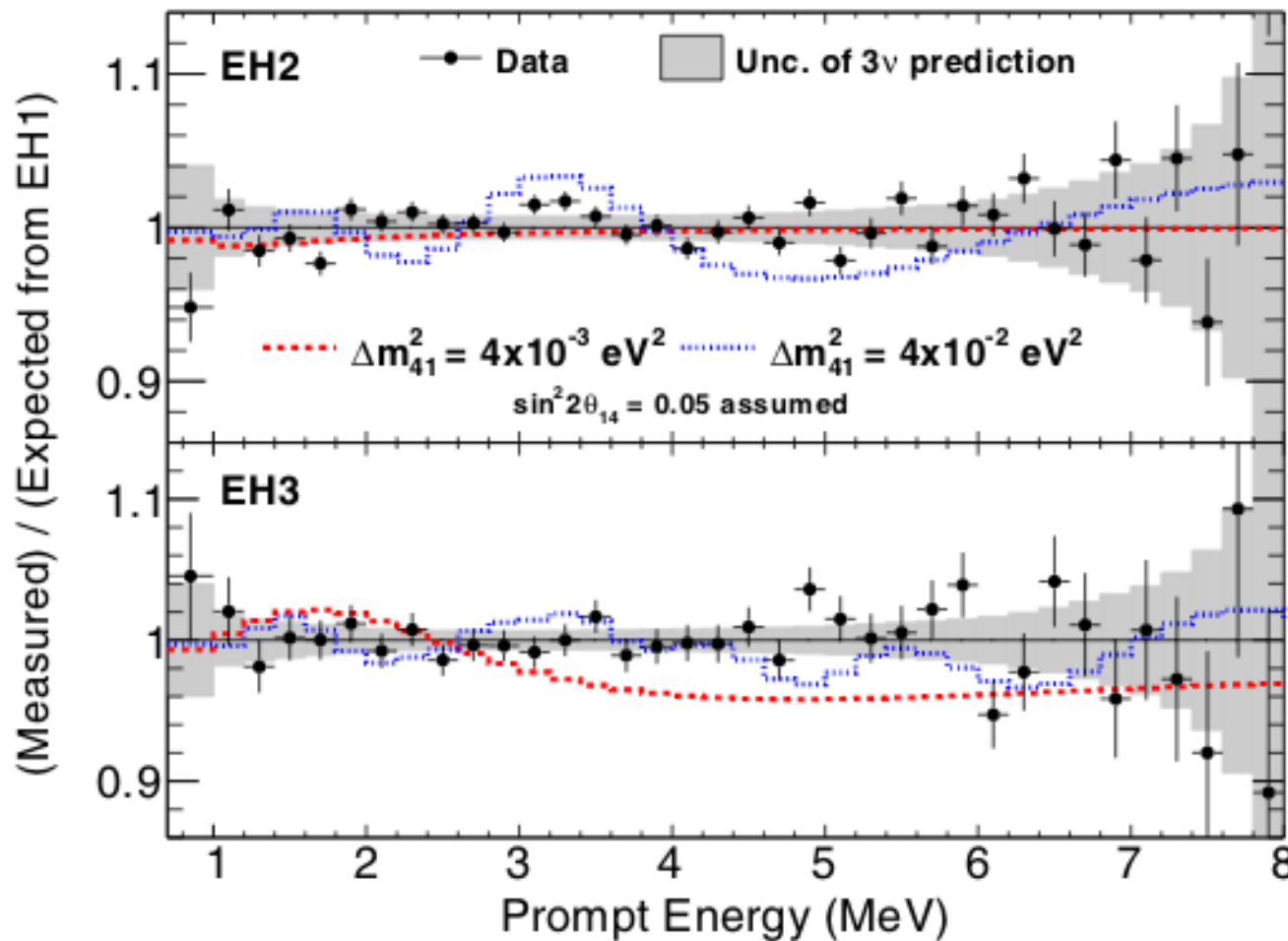
- Assume mixing of 3 flavours of active neutrinos with 1 sterile neutrino.
- Survival probability of reactor antineutrino is modified:

$$P(\bar{\nu}_e \rightarrow \bar{\nu}_e) \cong 1 - \sin^2 2\theta_{14} \sin^2 \left( \frac{\Delta m_{41}^2 L}{4E_\nu} \right) - \cos^4 \theta_{14} \sin^2 2\theta_{13} \sin^2 \left( \frac{\Delta m_{ee}^2 L}{4E_\nu} \right)$$

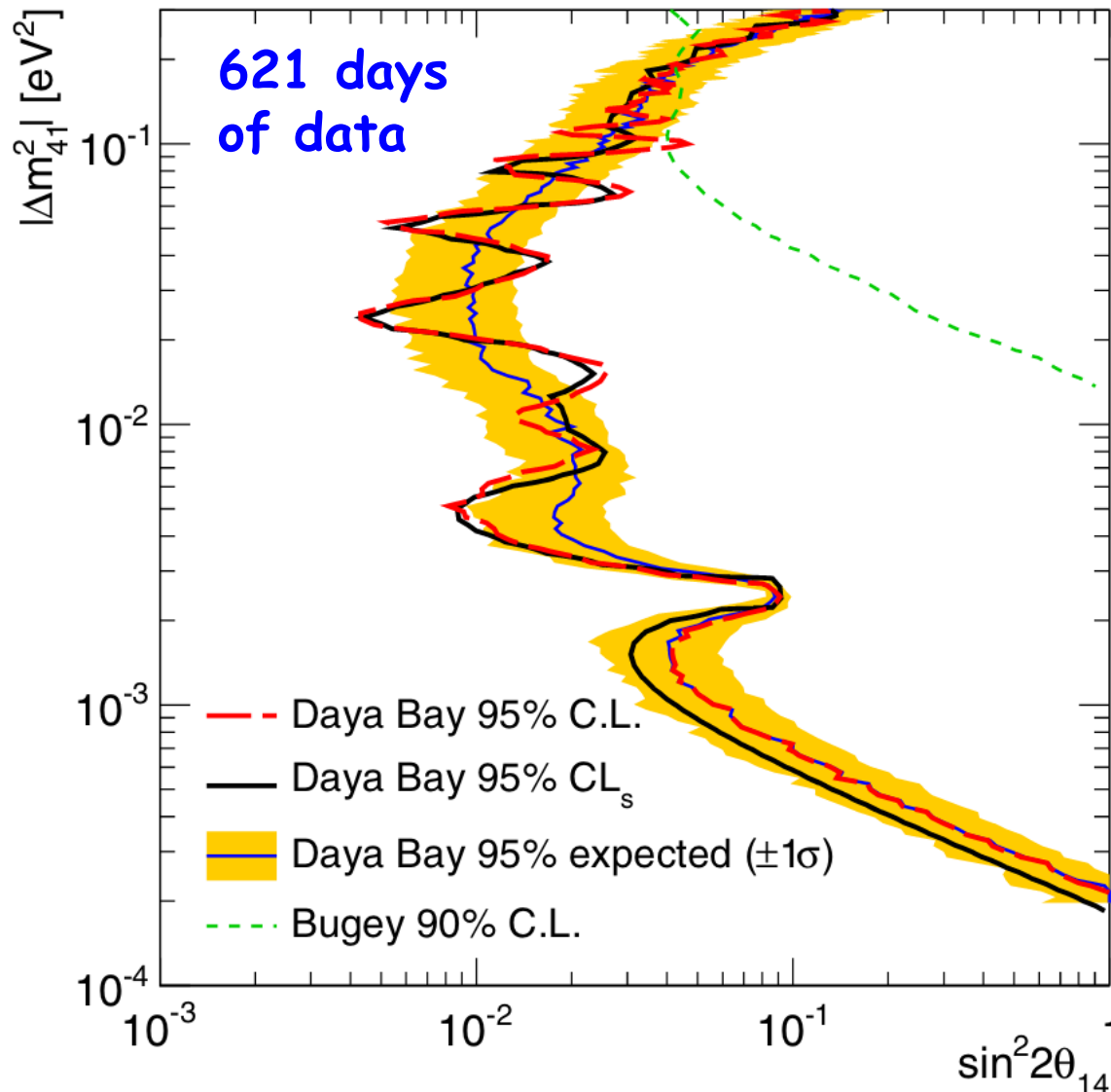


# Look For Additional Spectral Distortion

- Multiple baselines and detectors
  - cover a broad mass range to search for sterile neutrino
  - relative measurement of energy spectra reduces systematic errors



# Limit on A Light Sterile Neutrino

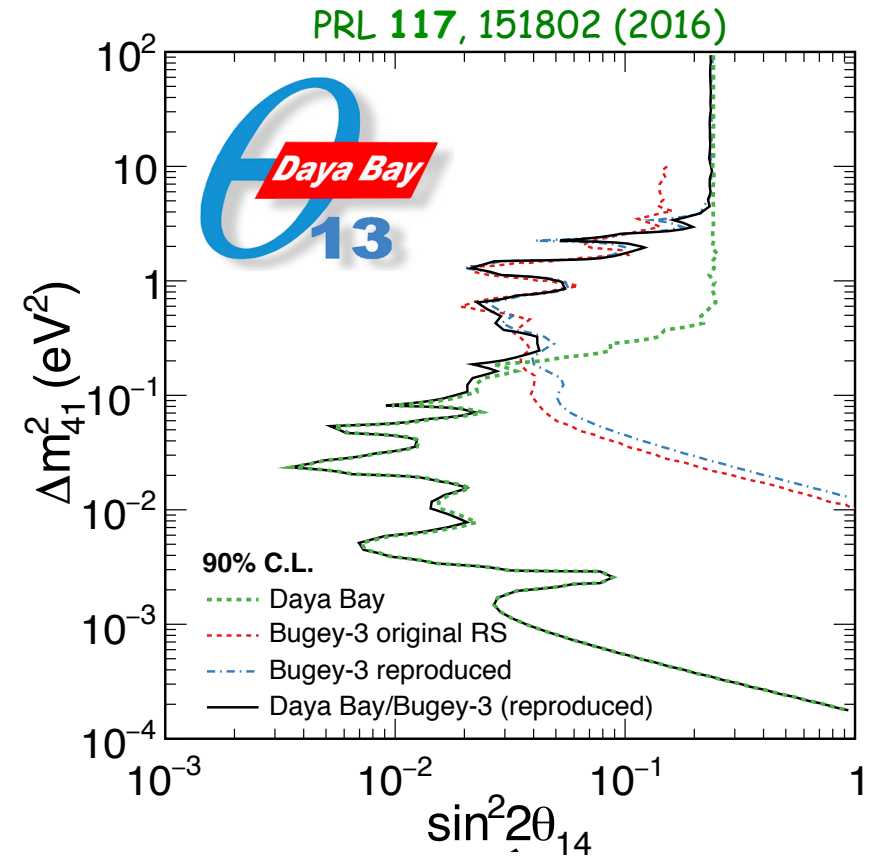
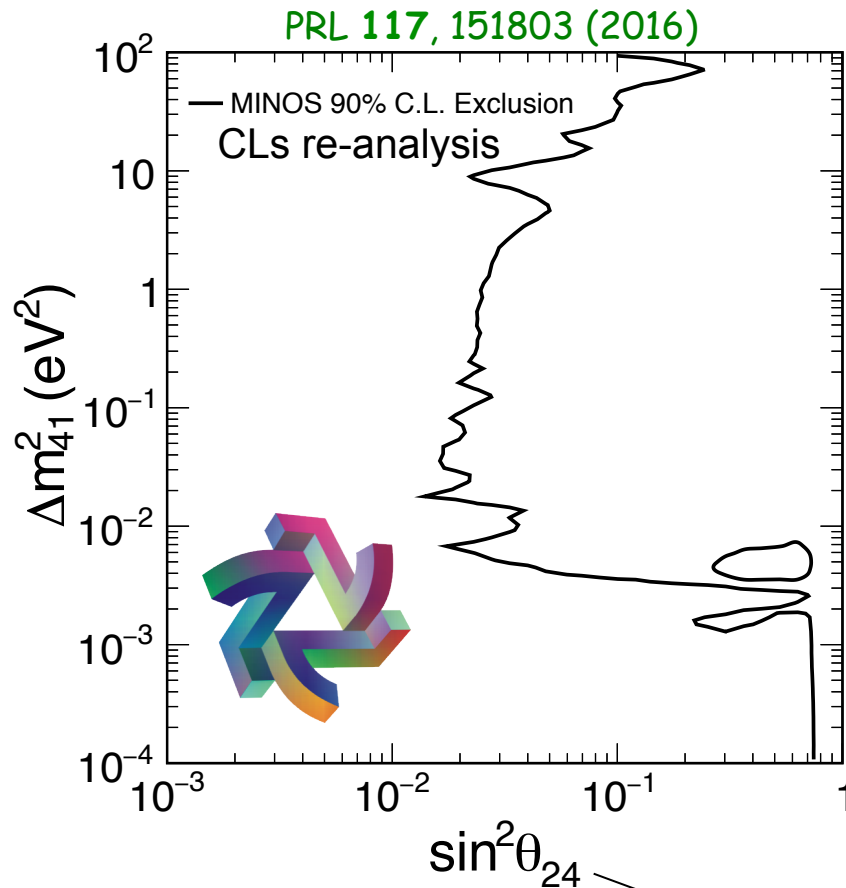


- Two analyses yield consistent results
- Results are insensitive to combinations of mass hierarchies for active and sterile neutrinos
- Provide the best limit in the  $|\Delta m^2_{41}| < 0.1 \text{ eV}^2$  region
- Region of  $|\Delta m^2_{41}| < 0.3 \text{ eV}^2$  is insensitive to the reactor antineutrino flux model



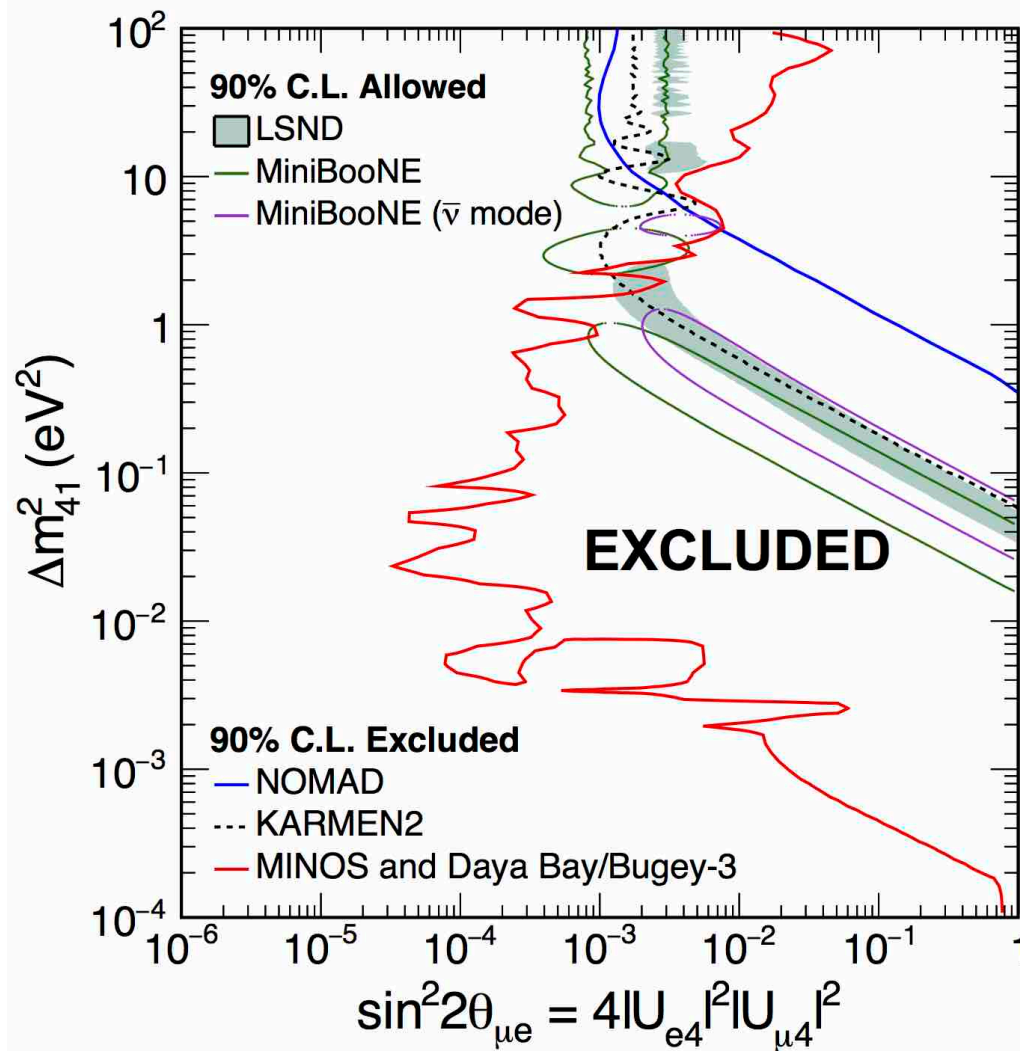
# Sterile Neutrinos: MINOS+Daya Bay+Bugey-3

- $\nu_\mu$  to  $\nu_e$  appearance results from LSND and MiniBooNE hinted  $\sin^2 2\theta_{\mu e} > 0$ .



$$4|U_{\mu 4}|^2 |U_{e 4}|^2 = \sin^2 \theta_{24} \sin^2 2\theta_{14} \equiv \sin^2 2\theta_{\mu e}$$

# Sterile Neutrinos: MINOS+Daya Bay+Bugey-3

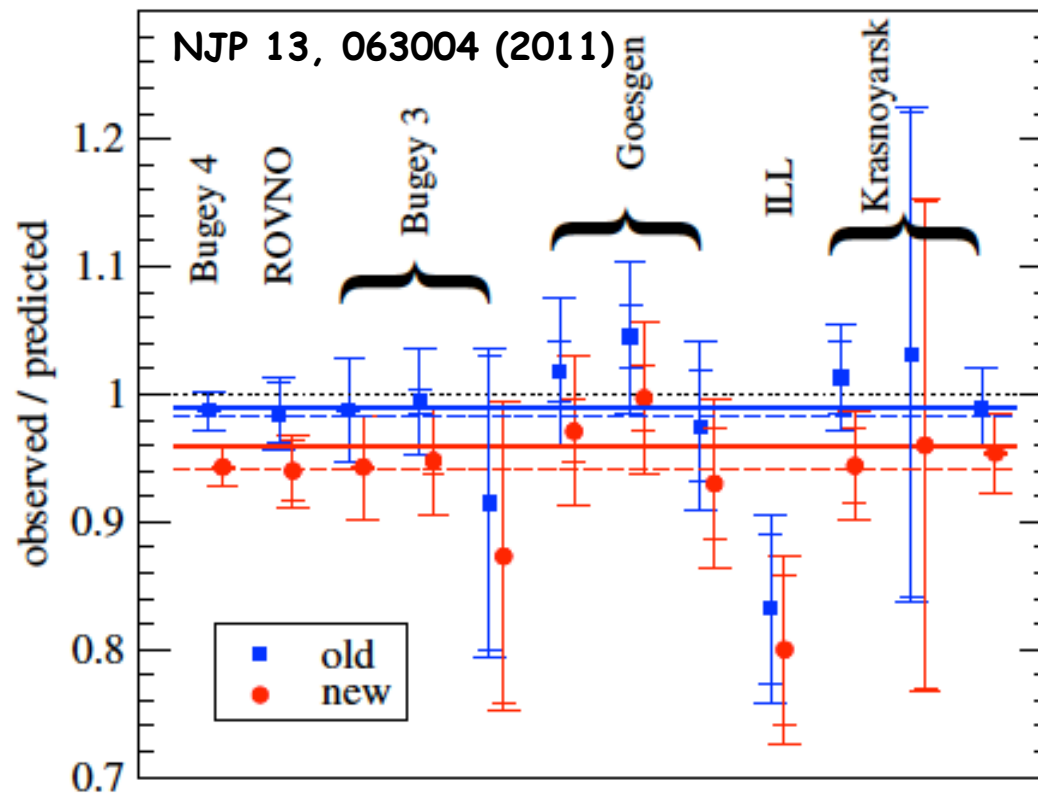


PRL 117, 151801 (2016)

- Allowed region of LSND and MiniBooNE is excluded for  $\Delta m^2_{41} < 0.8 \text{ eV}^2$  (95% c.l.)

# Reactor Antineutrino Anomaly

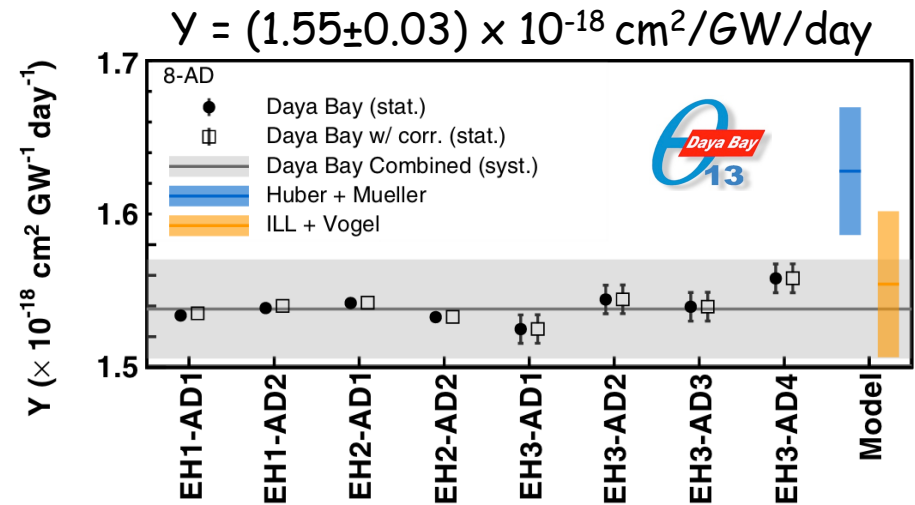
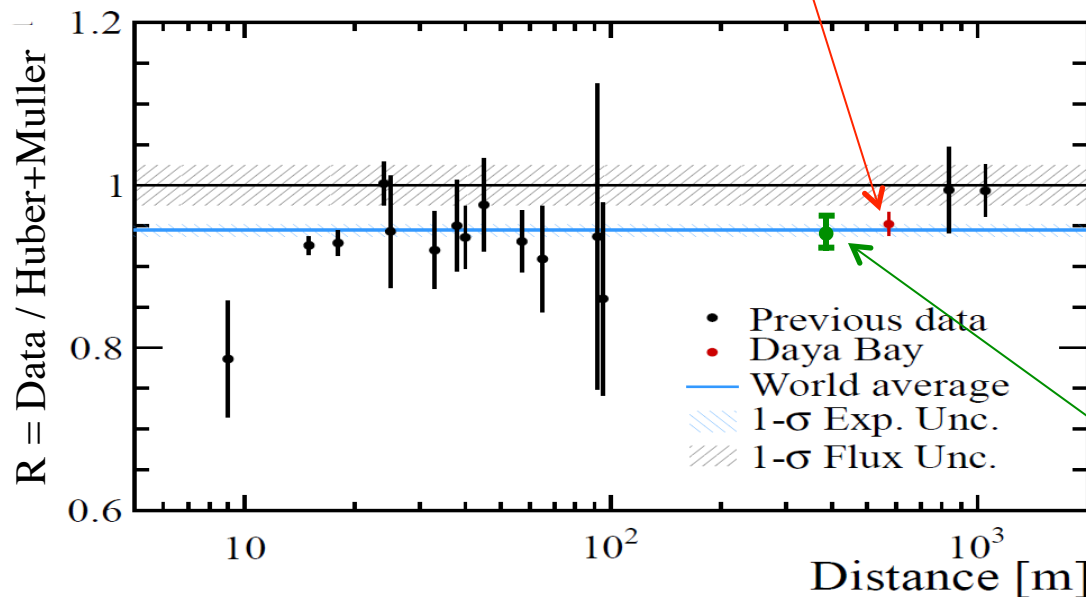
- Reactor antineutrino flux at short distance is  $\sim 6\%$  smaller
  - New calculations yielded 3% more flux
    - Mention et al., PRD **83**, 054615 (2011) and update (2012)
    - Huber PRC **84**, 024617 (2011)
  - Included contributions from long-lived isotopes
  - Measured neutron lifetime has decreased, leading to larger  $\sigma(\text{IBD})$ .



# Reactor $\bar{\nu}_e$ : Absolute Flux

Daya Bay (1230 days) :  
 $R = 0.952 \pm 0.014 \pm 0.023$

arXiv: 1808.10836



Previous average:

←  $R = 0.942 \pm 0.009 \text{ (exp)}$   
 $\pm 0.025 \text{ (model)}$

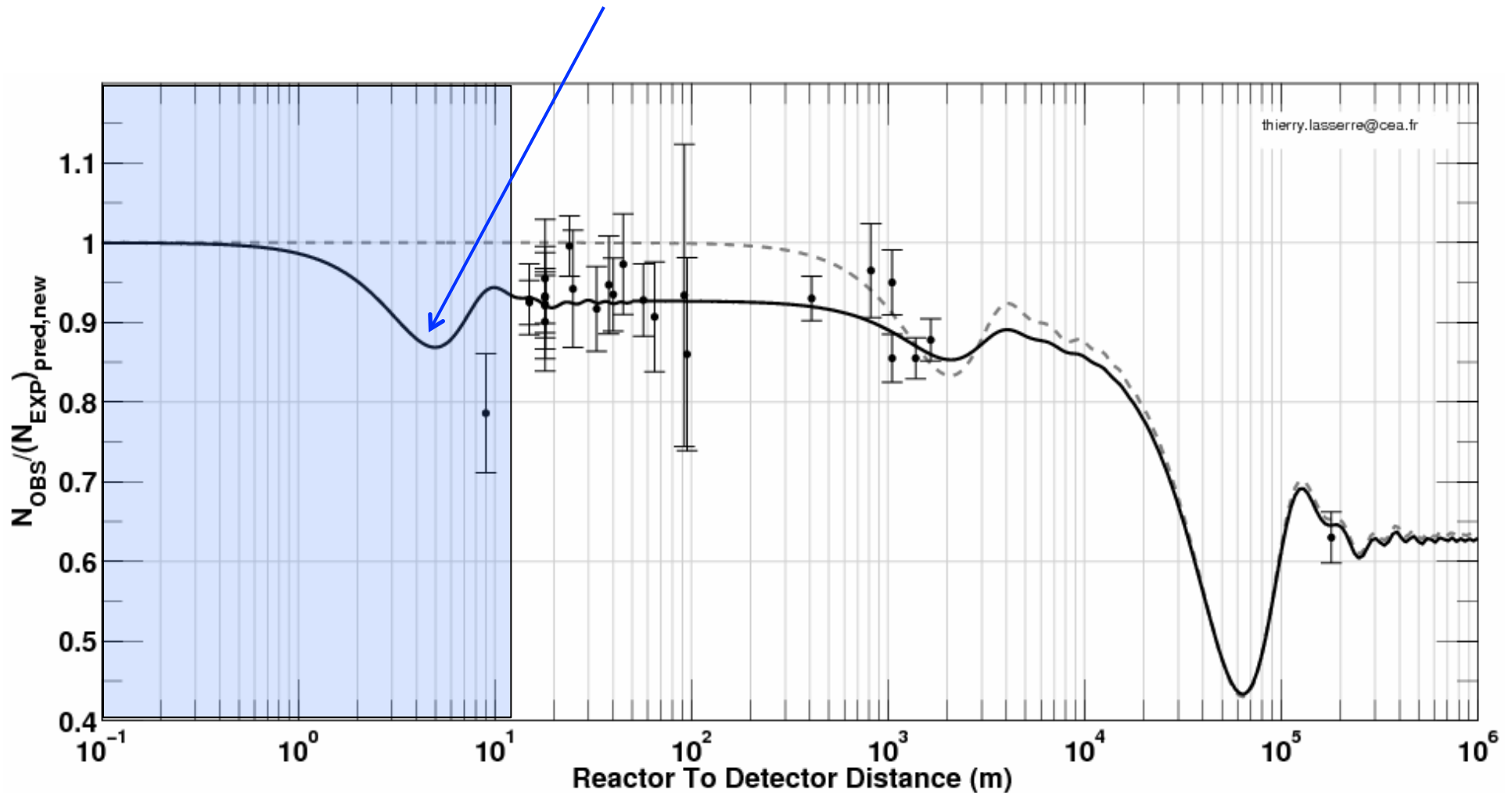
RENO (500 days) :  
 $R = 0.946 \pm 0.021 \text{ (prelim.)}$

- Reactor  $\bar{\nu}_e$  anomaly due to
  - new physics ?
  - incomplete theoretical understanding ? [see Huber's Neutrino 2016 talk]

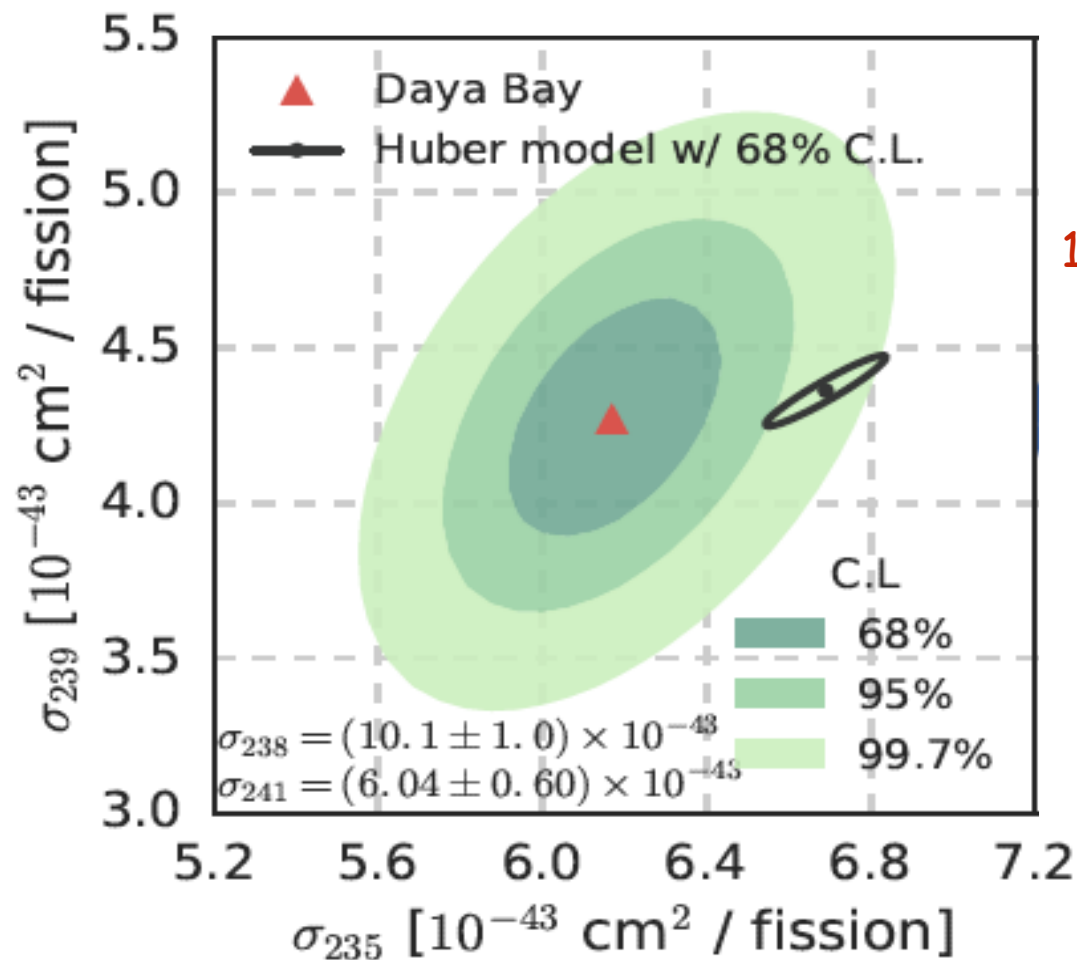


# Sterile Neutrino As A Solution

Reactor anti-neutrino anomaly may be due to sterile-active neutrino oscillation with  $\Delta m^2 \sim 1 \text{ eV}^2$ :



# IBD Yields Per Fission



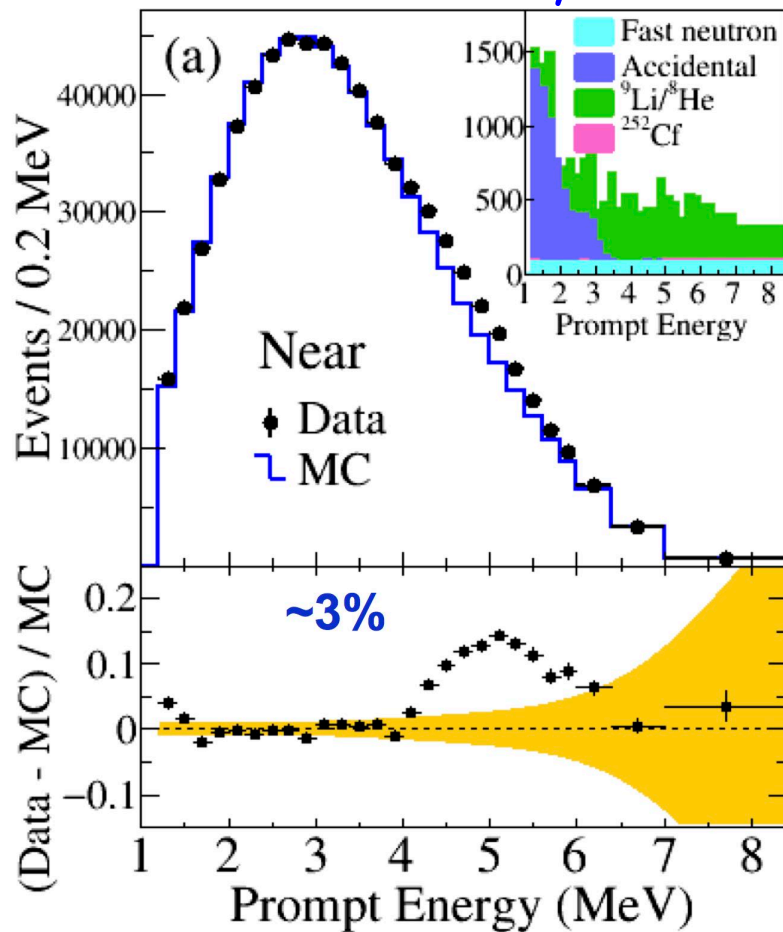
1230 days, near detectors  
PRL 118 (2017) 251801

- 7.8% overestimation of predicted anti-neutrino flux from  $^{235}\text{U}$
- $^{235}\text{U}$  could be the key source of the reactor anti-neutrino anomaly.

# Excess Near 5 MeV in Spectrum

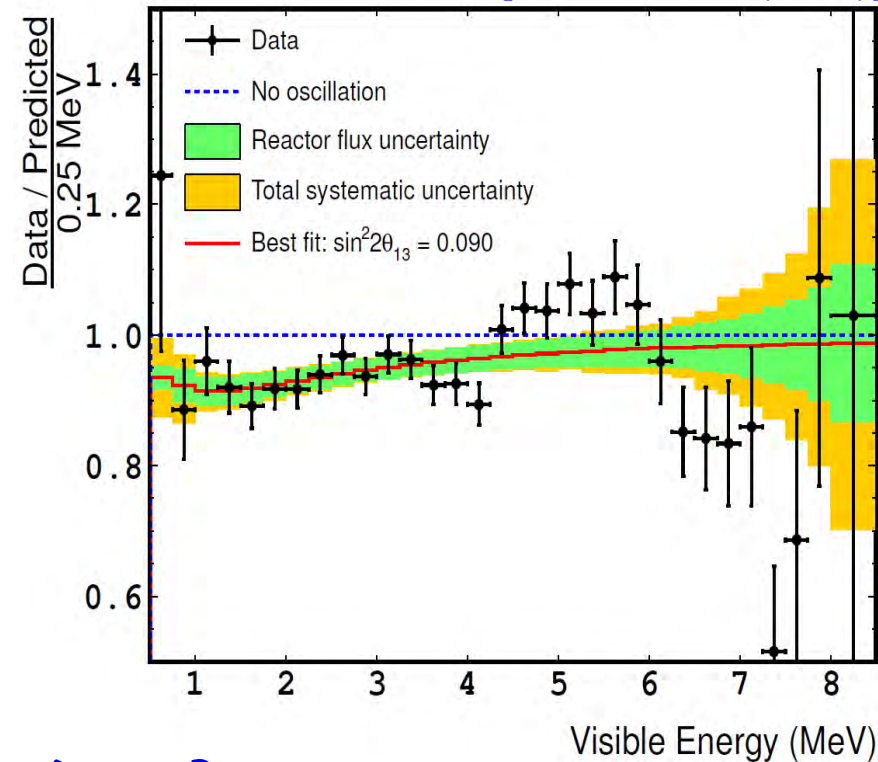
Prediction: Huber+Mueller

RENO Preliminary

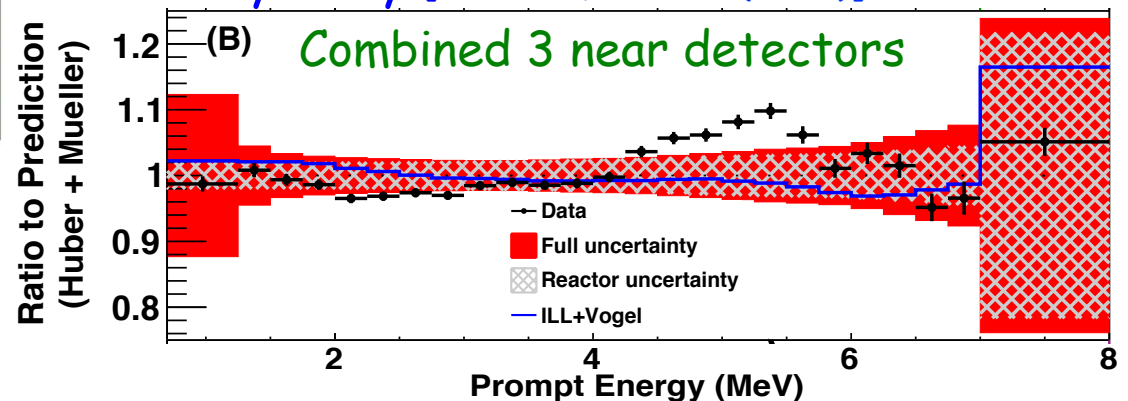


Excess tracks reactor power.

Double Chooz [JHEP 10, 086(2014)]



Daya Bay [PRL 116, 061801 (2016)]



## Summary

- Daya Bay continues to be the leading reactor-based neutrino experiment that
  - has acquired the largest sample of reactor antineutrinos
  - provides the most precise measurement of
    - $\sin^2 2\theta_{13}$  and  $\Delta m^2_{ee}$
    - absolute reactor antineutrino spectrum
  - has the best sensitivity of searching for a light sterile neutrino with  $\sim 10^{-3} \text{ eV}^2 < \text{mass} < \sim 10^{-1} \text{ eV}^2$
  - yields interesting results on other topics.
- Daya Bay will stop data taking by the end of 2020
  - Precision in  $\sin^2 2\theta_{13} \approx 0.0025$ 
    - Remains the most precise in the foreseeable future
  - Precision in  $\Delta m^2_{ee} \approx 0.06 \times 10^{-3} \text{ eV}^2$



Thank You