# HobEx The Habitable Exoplanet Observatory

Exploring New Worlds – Understanding Our Universe

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### Hober of The US Astrophysics Decadal Survey

#### New Worlds, New Horizons

#### For large space-based telescopes –

- Astro2000: James Webb Space Telescope (JWST) ~2021.
- Astro2010: Wide Field Infrared Survey Telescope (WFIRST), ~2025 (Jason Rhodes talk 4pm Sep 25<sup>th</sup>).
- Astro 2020: Habitable Exoplanet Observatory (HabEx; UV–near-IR) If prioritized, mid-to-late 2030's.
  - The other studies are:
    - Origins Space Telescope (OST; Far-IR)
    - Large UV, Optical, IR Telescope (LUVOIR; UV-near-IR)
    - Lynx (X-ray)

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### The HabEx Final Report

- 3.5 years
- 178 authors
- 9 architectures
- 552 pages





Exploring New Worlds, Understanding Our Universe

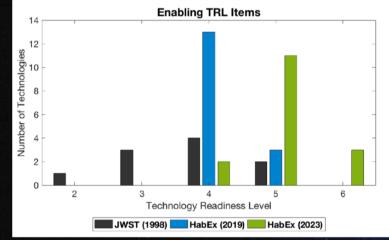
www.jpl.nasa.gov/habex/pdf/HabEx-Final-Report-Public-Release.pdf

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# Study Philosophy

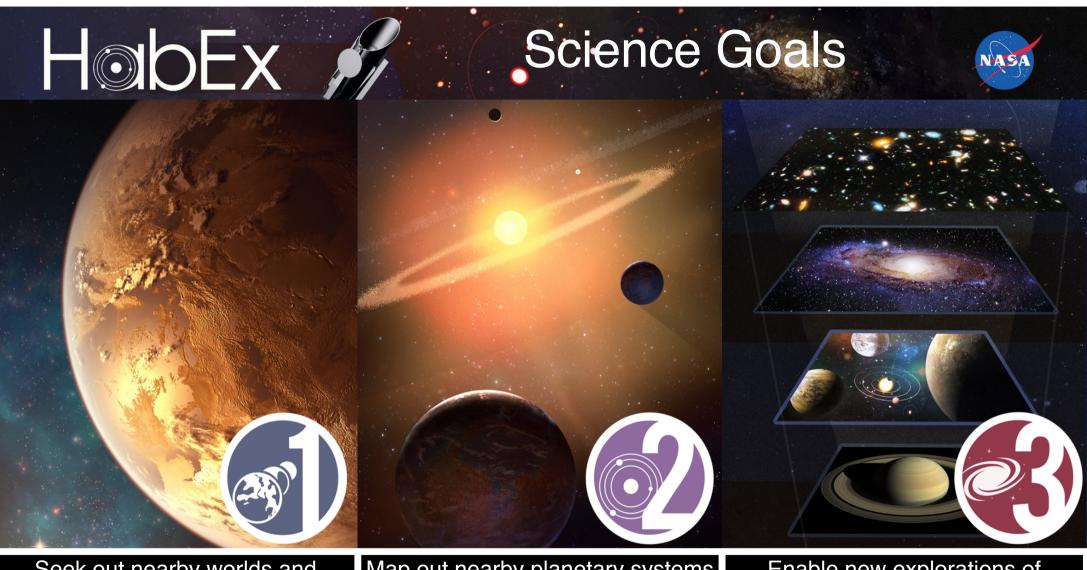
NASA: "Develop an exoplanet direct imaging mission"

- HabEx Team:
- Maximize
  - Science return for both
    - Exoplanet direct imaging
    - Astrophysics & cosmology



#### HabEx Team: *Minimize*

- Cost
- Risk
- Development schedule



Seek out nearby worlds and explore their habitability

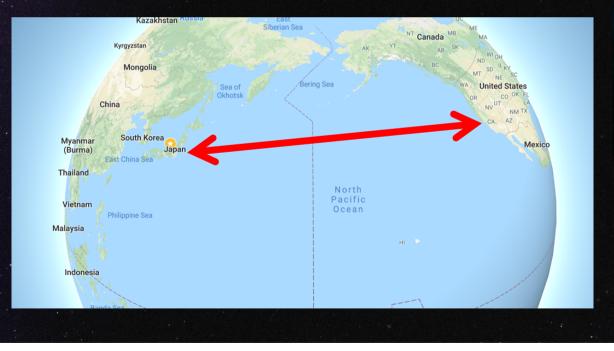
Map out nearby planetary systems and understand their diversity.

Enable new explorations of systems in the UV to near-IR

### **Exoplanet Direct Imaging**

Seeing an exoplanet around a star is like trying to see a firefly near a spotlight in Los Angeles... when you are standing in Tokyo!

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Seeing an Earth-like exoplanet in the habitable zone around a sunlike star is like trying to see a firefly near ONE THOUSAND spotlights in Los Angeles... when you are in standing in Tokyo!

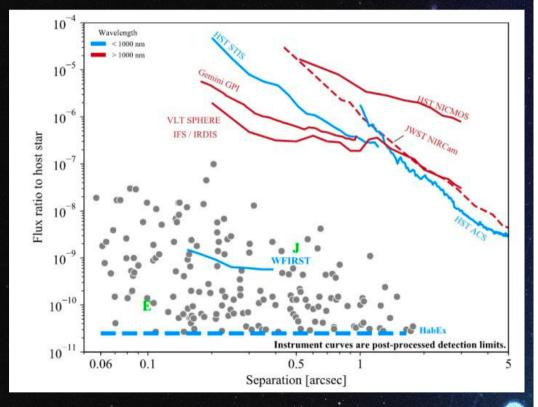
#### High Contrast Imaging and Spectroscopy

Contrast Ratio (exoplanet light to star light)

#### 10<sup>-5</sup> 1 part in 10,000

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- What we can get from ground-based coronagraphs now
- 1 part in 1,000,000,000
- What WFIRST's coronagraph is being designed to achieve
- **x 10<sup>-10</sup>** 4 parts in 10,000,000,000
- What we have demonstrated in a lab for WFIRST
- 10<sup>-10</sup> 1 part in 10,000,000,000
  - What HabEx's coronagraph is being designed to achieve (and what we need to see another Earth)



#### We have already found a lot of Exoplanets



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#### Exoplanets:

Cumulative Detections by Discovery Year

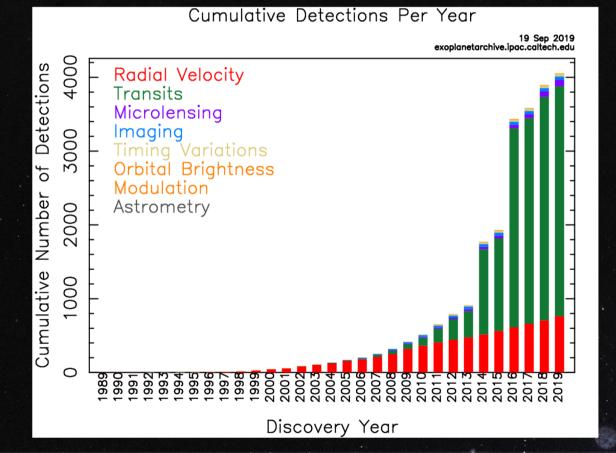
1989-2018

Plots generated Sept. 27, 2018

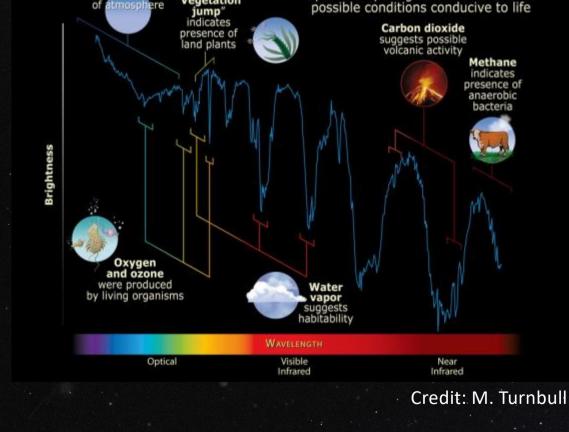
#### We have already found a lot of Exoplanets

#### 4055 as of 5 days ago!

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But we really want to find "Earth 2.0" "Blue of the sky" measures total amount of atmosphere" "Negetation indicates presence of



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### **Habitable Zone**

тоо нот

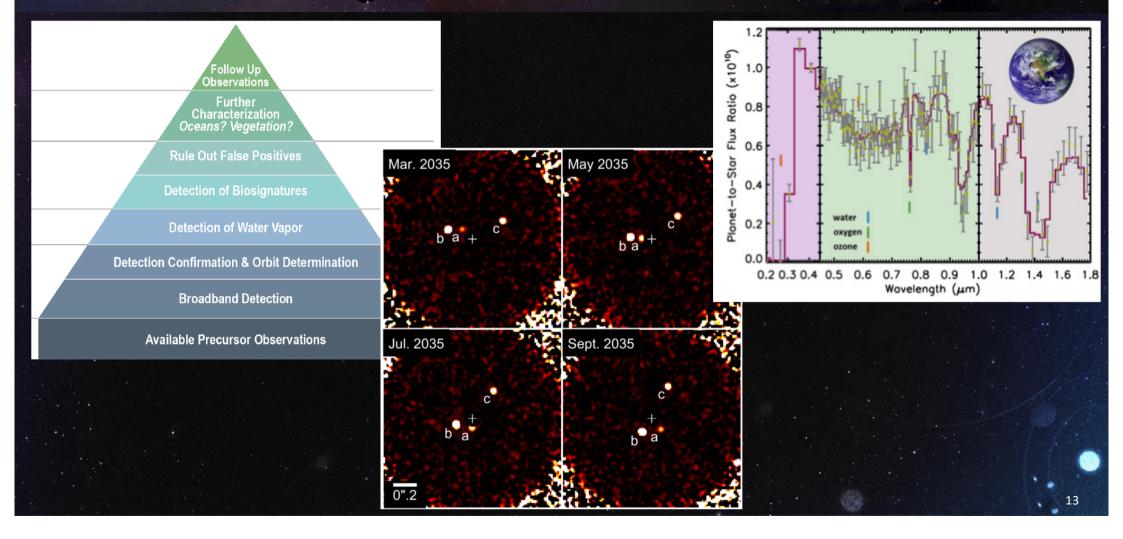
#### JUST RIGHT

TOO COLD

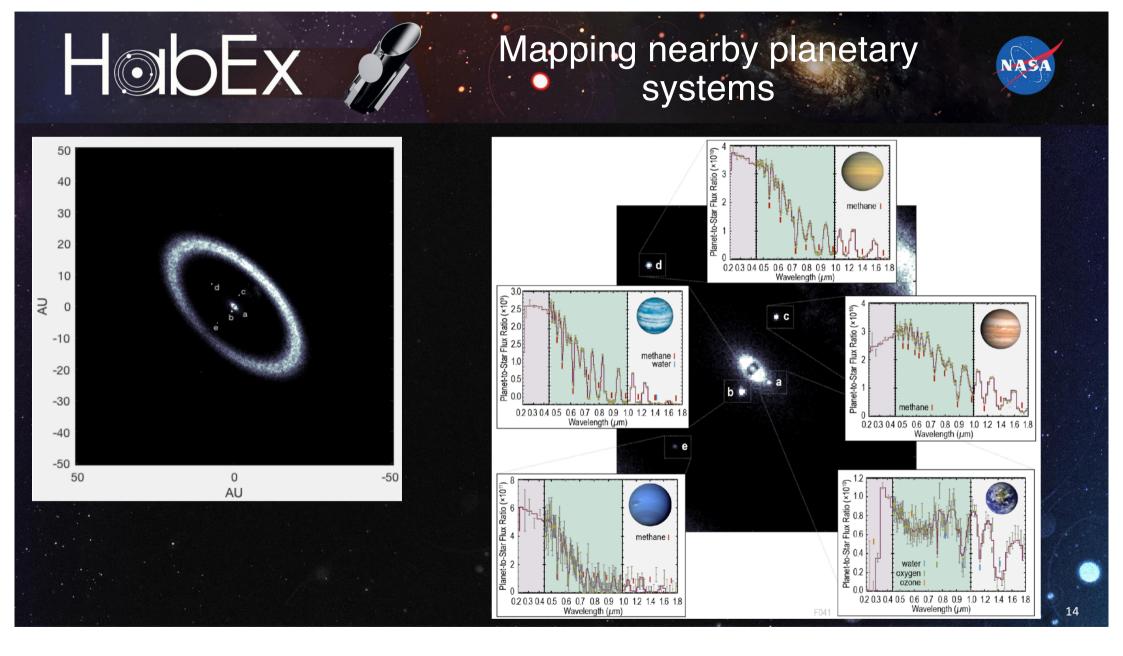
Planet size: 1-2x Earth



#### Seeking Potentially Habitable Worlds



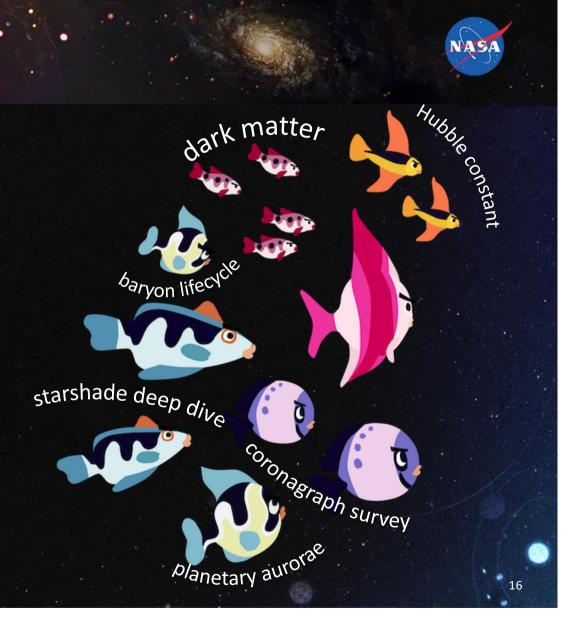
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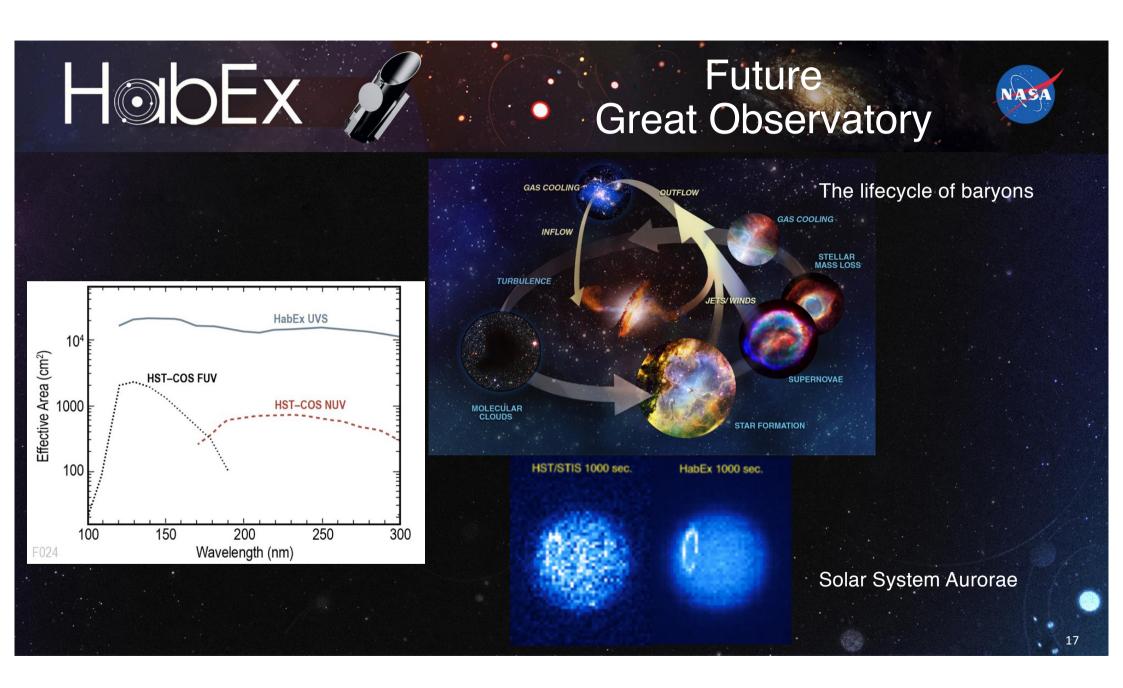


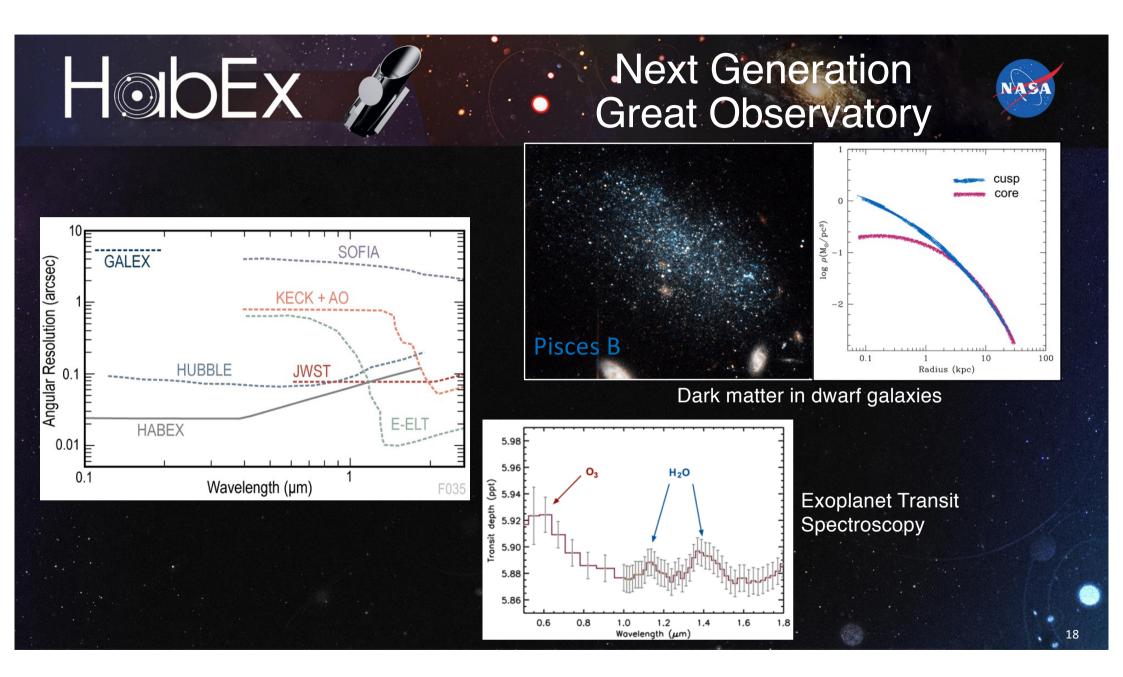
# Hodbex M What you're thinking... ctoplanet Direct Imagingo 60 Scienco

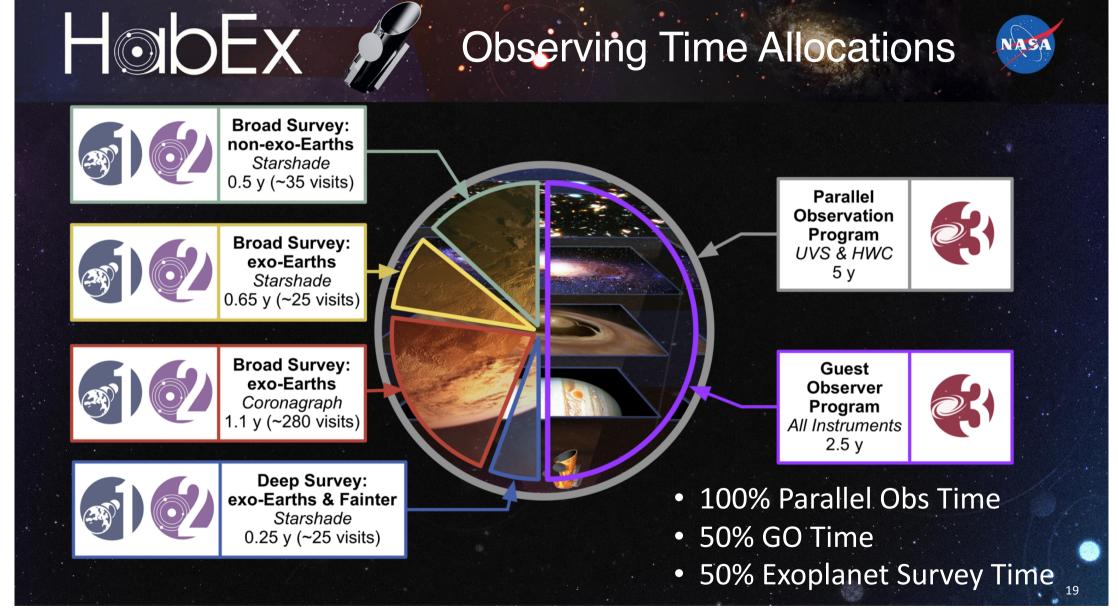
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The truth: 50% of HabEx's primary 5-year mission is dedicated to Guest Observer Science



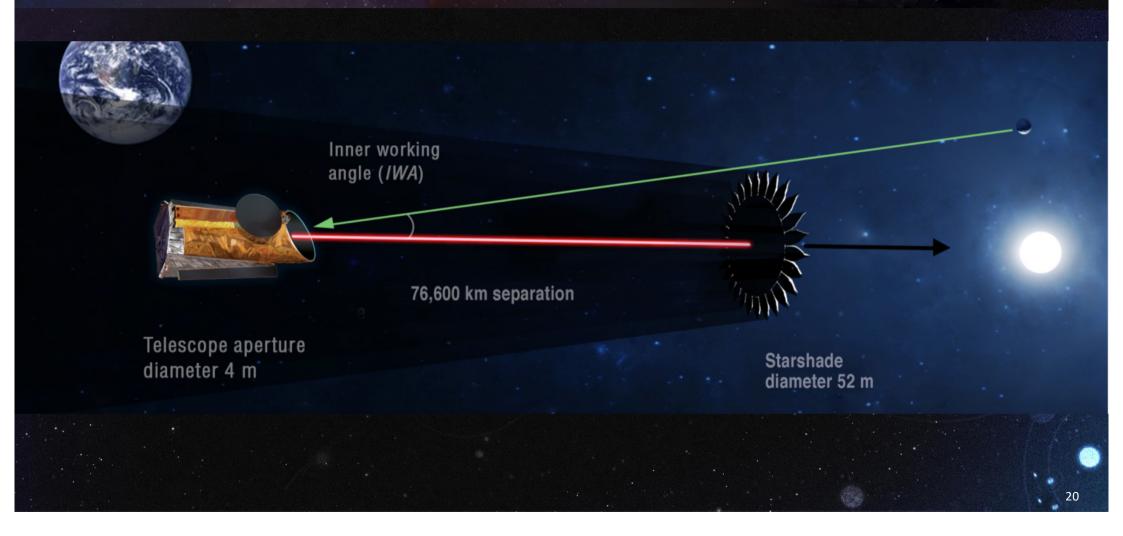






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### **Baseline Architecture**



### **Baseline Architecture**



Habitable Exoplanet Observatory (HabEx)

							<b>,</b> ,	
				Mission Duration	5 years (10	0 years consumables	)	
	Spacecraft	<b>Optical Telesc</b>	ope Assembly	Orbit	Earth-Sun L2 Halo			
	opuocolult	ephoni lolooo	Telescope Aperture Cover	Aperture	4 m unobs	cured		
				Telescope Type	Off-axis three-mirror anastigmat			
		Baffle Tube with		Primary Mirror	4 m monolith; glass-ceramic substrate; AI+MgF2 coating			
Solar Panels	nels	Forward Scarf		Instruments (4)	Exoplanet science: Coronagraph, Starshade Observatory science: UV Spectrograph, Workhorse Camera			
				Attitude Control	Slewing: hydrazine thrusters; Pointing: microthrusters			
						HST	HabEx	
			econdary Mirror ower Structure	Aperture		2.4 m obscured	4.0 m unobscured	
				Diffraction Limit		500 nm	400 nm	
	e Ring Primary Mirror						20 min (typical),	
	Assembly			Slew Rate (180 de	g)	~30 min (max)	5 min (max)	
Microthru	usters					5 mas (typical),		
in 8 locat	ions Integrated Science			<b>Pointing Accuracy</b>	,	2 mas (best)	0.7 mas	
No.	Instrument Module (ISIM)			<b>Spatial Resolution</b>	า	50 mas	25 mas	
	& Payload Radiators			Effective Area* (@	🦻 200nm)	700 cm <sup>2</sup>	10,000 cm <sup>2</sup>	
				Micro-shutters		No	Yes	
				Serviceable		Yes/Astronaut	Yes/Robotic	
						ture multiplied by	throughput and	
				quantum efficien	~ /			

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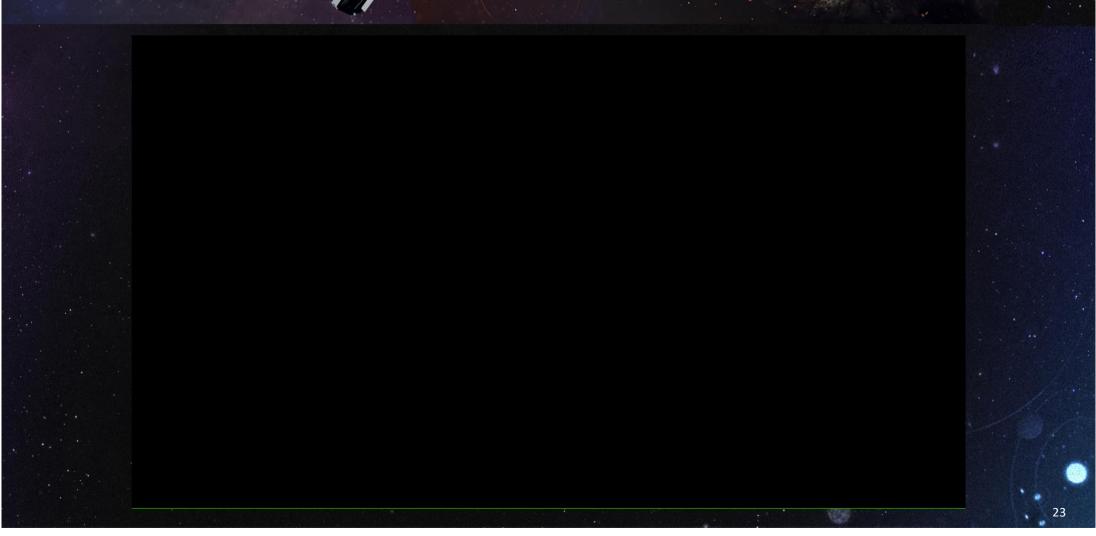
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### Starshade

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### Starshade



# HabEx Instruments

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	Coronagraph (HCG)	Starshade (SSI)	Workhorse Camera (HWC)	UV Spectograph (UVS)
Purpose	Exoplanet imaging and characterization	Exoplanet imaging and characterization	Multipurpose, wide-field imaging camera and spectograph for observatory science	High-resolution, UV imaging and spectroscopy for observatory science
Instrument Type	Vector Vortex charge 6 coronagraph with: - Raw contrast: $2.5 \times 10^{-10}$ at the IWA - $\Delta$ mag limit = $26.5$ - $20\%$ instantaneous bandwidth - Imager and spectograph	52 m diameter starshade occulter with: - 76,600 km separation (Visible) - Raw contrast: 1 x $10^{-10}$ at the IWA - $\Delta$ mag limit = 26.5 - 107% instantaneous bandwidth - Imager and spectograph	Imager and spectograph	High-resolution imager and spectrograph
Channels	Visible: 0.45–0.975 μm - Imager + IFS with <i>R</i> = 140 Near-IR: 0.975–1.8 μm - Imager + IFS with <i>R</i> = 40	UV: 0.2–0.45 μm - Imager + grism with <i>R</i> = 7 Visible: 0.45–0.975 μm - Imager + IFS with <i>R</i> = 140 Near-IR: 0.975–1.8 μm - Imager + IFS with <i>R</i> = 40	Visible: 0.37–0.975 μm - Imager + grism with <i>R</i> = 1,000 Near-IR: 0.95–1.8 μm - Imager + grism with <i>R</i> = 1,000	UV: 115–320 nm (with 115–370 nm available at <i>R</i> ≤ 1,000) <i>R</i> = 60,000; 25,000; 12,000; 6,000; 3,000; 1,000; 500; imaging
Field of View	IWA: 2.4 $\lambda/D$ = 62 mas at 0.5 $\mu$ m OWA: 32 $\lambda/D$ = 830 mas at 0.5 $\mu$ m	IWA: 58 mas at 0.3–1.0 μm OWA: 6 arcsec (Vis. broadband imaging) OWA: 1 arcsec (Visible IFS)	3 x 3 arcmin <sup>2</sup>	3 x 3 arcmin <sup>2</sup>
Features	64 x 64 deformable mirrors (2) Low-order wavefront sensing and control	Formation flying, sensing, and control	Microshutter array for multi-object spectroscopy - 2 x 2 array, 171 x 365 apertures	Microshutter array for multi-object spectroscopy - 2 x 2 array, 171 x 365 apertures

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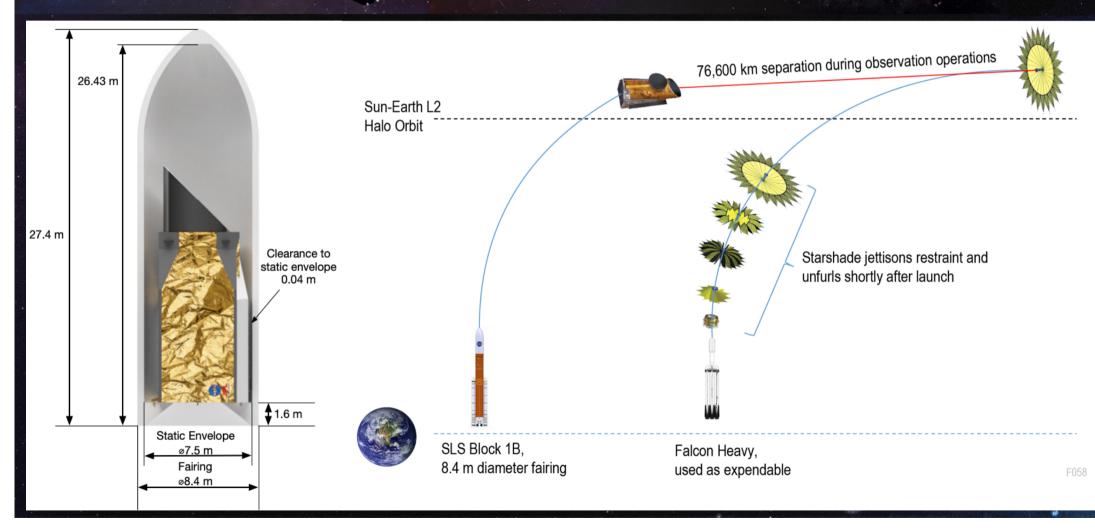
NASA

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## Launch

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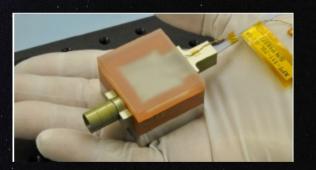
# HobEx R Enabling Technologies



WFIRST Coronagraph Instrument Testbed



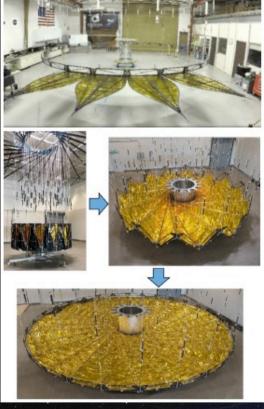
4.2m mirror for the Southern Astrophysical Research Telescope



Planar lightwave circuit beam launcher for laser metrology



LISA-Pathfinder colloidal microthrusters



Starshade 10m perimeter truss deployment tests

HabEx is the only study with all technologies at TRL 4 or higher.

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### Architecture Trade Space

		Starlight Suppression Method									
		H (Hybrid)		S (Starshade-on	lly)	C (Coronagraph-only)					
Diameter	4.0m	Planets: Effective collecting area: Spatial resolution: TRL-4 technologies: Cost (\$FY20):	186 10,000 cm <sup>2</sup> 25 mas 13 \$6.8B	Planets: Effective collecting area: Spatial resolution: TRL-4 technologies: Cost (\$FY20):	144 13,000 cm <sup>2</sup> 25 mas 9 \$5.7B	Planets: Effective collecting area: Spatial resolution: TRL-4 technologies: Cost (\$FY20):	119 10,000 cm <sup>2</sup> 25 mas 10 \$4.8B				
Aperture	3.2m	Planets: Effective collecting area: Spatial resolution: TRL-4 technologies: Cost (\$FY20):	110 6,400 cm <sup>2</sup> 31 mas 12 \$5.7B	Planets: Effective collecting area: Spatial resolution: TRL-4 technologies: Cost (\$FY20):	123 8,200 cm <sup>2</sup> 31 mas 9 \$5.0B	Planets: Effective collecting area: Spatial resolution: TRL-4 technologies: Cost (\$FY20):	86 6,400 cm <sup>2</sup> 31 mas 9 \$3.7B				
Telescope	2.4m	Planets: Effective collecting area: Spatial resolution: TRL-4 technologies: Cost (\$FY20):	79 2,300 cm <sup>2</sup> 42 mas 11 \$4.8B	Planets: Effective collecting area: Spatial resolution: TRL-4 technologies: Cost (\$FY20):	69 3,000 cm <sup>2</sup> 42 mas 8 \$4.0B	Planets: Effective collecting area: Spatial resolution: TRL-4 technologies: Cost (\$FY20):	28 2,300 cm <sup>2</sup> 42 mas 8 \$3.1B				

HabEx Science		HabEx Mission Architectures														
	Goals & Objectives		4H	4S	4C	3.2H	3.2S	3.2C	2.4H	2.4S	2.4C	HabEx Architectures				
<b>A</b>	01	Exo-Earth candidates around nearby sunlike stars?														
Habitable Exoplanets	<b>O</b> 2	Water vapor in rocky exoplanet atmospheres?														
	<b>O</b> 3	Biosignatures in rocky exoplanet atmosphere?														
Ha	04	Surface liquid water on rocky exoplanets?										All architectures include the HWC and UV				
6)	<b>O</b> 5	Architectures of nearby planetary systems?										H: Hybrid – Starshade and Coronagraph C: Coronagraph				
etary,	<b>O</b> 6	Exoplanet atmospheric variations in nearby systems?														
xoplanetary Systems	07	Water transport mechanisms in nearby planetary systems?										S: Starshade				
Exor	<b>O</b> 8	Debris disk architectures in nearby planetary systems?														
	<b>O</b> 9 L	Lifecycle of baryons?														
	010	Sources of reionization?														
	011	Origins of the elements?										Green: Meets baseline requirements				
Science	012	Discrepancies in measurements of the cosmic expansion rate?										Yellow: Meets threshold requirements				
	013	The nature of dark matter?										Orange: Below threshold requirements				
vator	014	Formation and evolution of globular clusters?														
Observatory	015	Habitable conditions on rocky planets around M-dwarfs?														
õ	016	Mechanisms responsible for transition disk architectures?														
	017	Physics driving star-planet interactions, <i>e.g.</i> auroral activity?										28				

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- HabEx briefing to the Decadal Survey Nov 19/20
- Decadal Survey recommendations released in 2021
- What would YOU do with HabEx?! Alina.A.Kiessling@jpl.nasa.gov

