

Holographic Entanglement Entropy (Introduction)

(with H. Casini, M. Huerta, J. Hung, A. Sinha, M. Smolkin & A. Yale)
(arXiv:1101.5813, arXiv:1102.0440, arXiv:1109.0???)

1. Entanglement Entropy

- condensed matter
- quantum information
- black hole microphysics

2. AdS/CFT correspondence

(gauge/gravity duality)

- string theory
- quantum gravity

3. Holographic Entanglement Entropy

- proposal by Ryu & Takayanagi (2006)

4. Two Recent Developments:

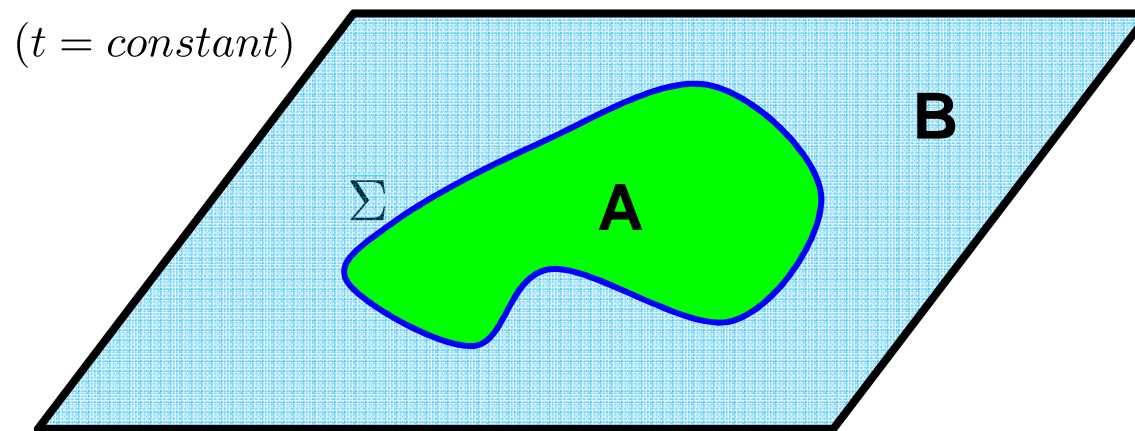
- precise connection between EE and central charges
- derivation of holographic EE for special geometries

5. Summary:

- holographic EE provides framework where we can learn about properties of both EE and quantum gravity

Entanglement Entropy

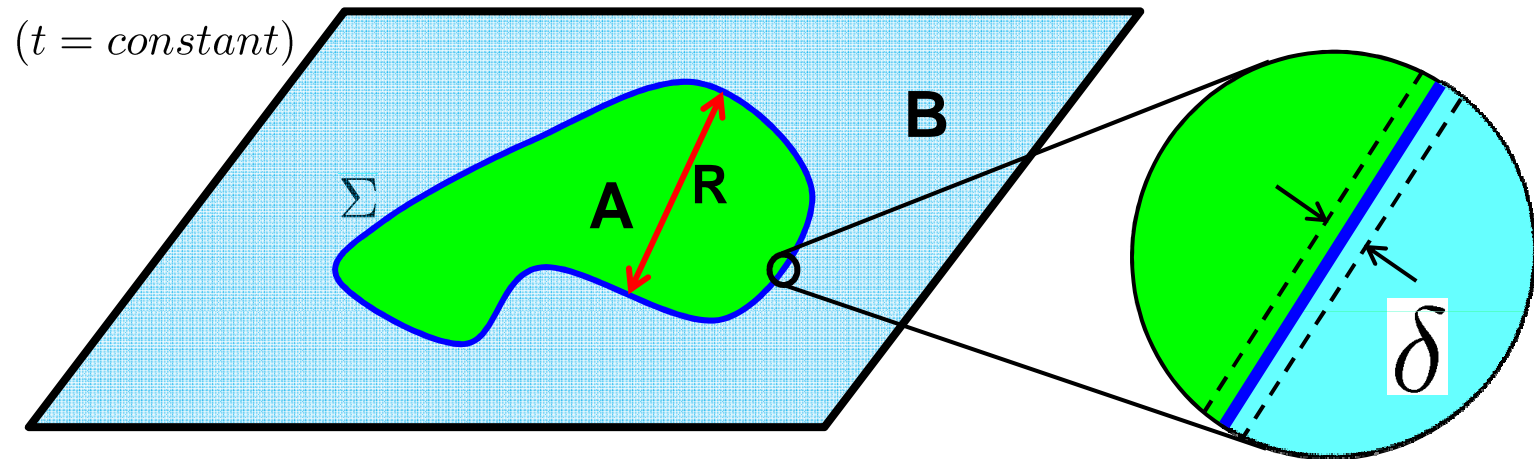
- what is entanglement entropy?
very general tool; divide quantum system into two parts and use entropy as measure of correlations between subsystems
 - in QFT, typically introduce a (smooth) boundary **or entangling surface** Σ which divides the space into two separate regions
 - integrate out degrees of freedom in “outside” region
 - remaining dof are described by a density matrix ρ_A
- calculate **von Neumann entropy**: $S_{EE} = -\text{Tr} [\rho_A \log \rho_A]$



Entanglement Entropy

- remaining dof are described by a density matrix ρ_A

→ calculate von Neumann entropy: $S_{EE} = -\text{Tr} [\rho_A \log \rho_A]$



- result is UV divergent!
- must regulate calculation: $\delta = \text{short-distance cut-off}$

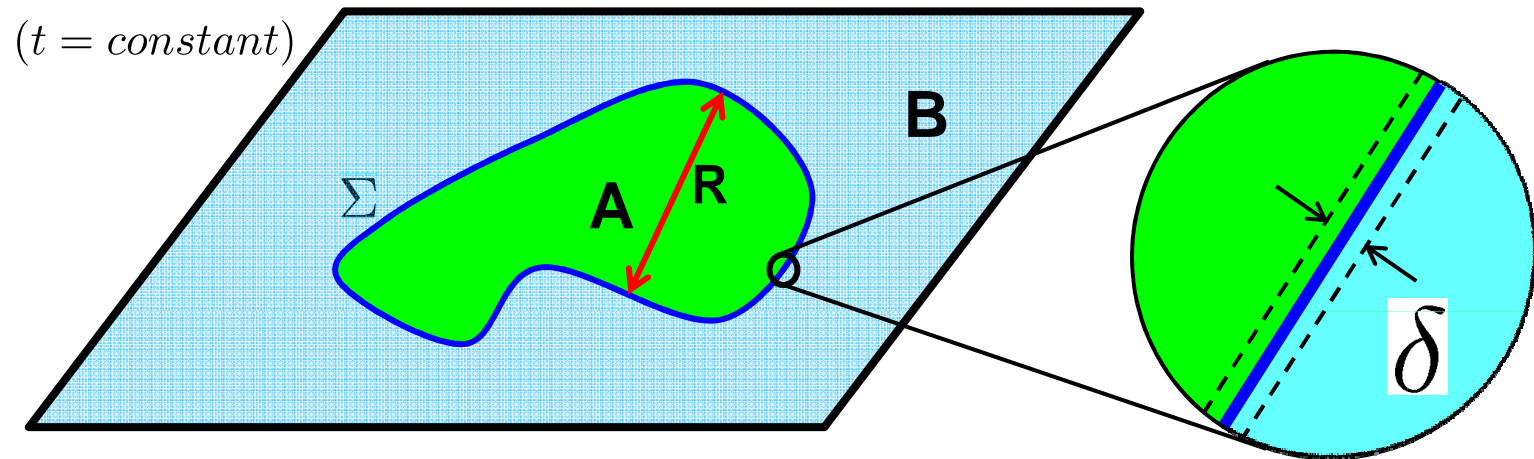
$$S = c_0 \frac{R^{d-2}}{\delta^{d-2}} + c_2 \frac{R^{d-4}}{\delta^{d-4}} + \dots \quad d = \text{spacetime dimension}$$

- careful analysis reveals geometric structure, eg, $S = \tilde{c}_0 \frac{\mathcal{A}_\Sigma}{\delta^{d-2}} + \dots$

Entanglement Entropy

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- leading coefficients sensitive to details of regulator, eg, $\delta \rightarrow 2\delta$
- find universal information characterizing underlying QFT in subleading terms, eg, $S = \dots + c_d \log(R/\delta) + \dots$

More general comments on **Entanglement Entropy**:

- nonlocal quantity which is (at best) very difficult to measure
→ no accepted experimental procedure
- in condensed matter theory: diagnostic to characterize quantum critical points or topological phases (eg, quantum hall fluids)
- in quantum information theory: useful measure of quantum entanglement (a computational resource)
- **black hole microphysics**: leading term obeys “area law” $S \simeq c_0 \frac{A_\Sigma}{\delta^{d-2}}$
→ suggested as origin of black hole entropy (eg, $\delta \simeq \ell_P$)
(Bombelli, Koul, Lee & Sorkin `86; Srednicki; Frolov & Novikov; Callan & Wilczek; Susskind;)
- recently considered in **AdS/CFT correspondence**

(Ryu & Takayanagi `06)

AdS/CFT correspondence:

anti-de Sitter space

quantum gravity

- negative cosmological constant
- $d+1$ spacetime dimensions



conformal field theory

quantum field theory

- no scale (at quantum level)
- d spacetime dimensions
- **no gravity!**



holography

Favorite example:

Type IIB superstrings
on $AdS_5 \times S^5$
with RR flux N_c



(3+1)-dimensional
 $\mathcal{N}=4$ $SU(N_c)$
super-Yang-Mills

(Maldacena '97)

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holography



classical gravity

with **small curvatures**

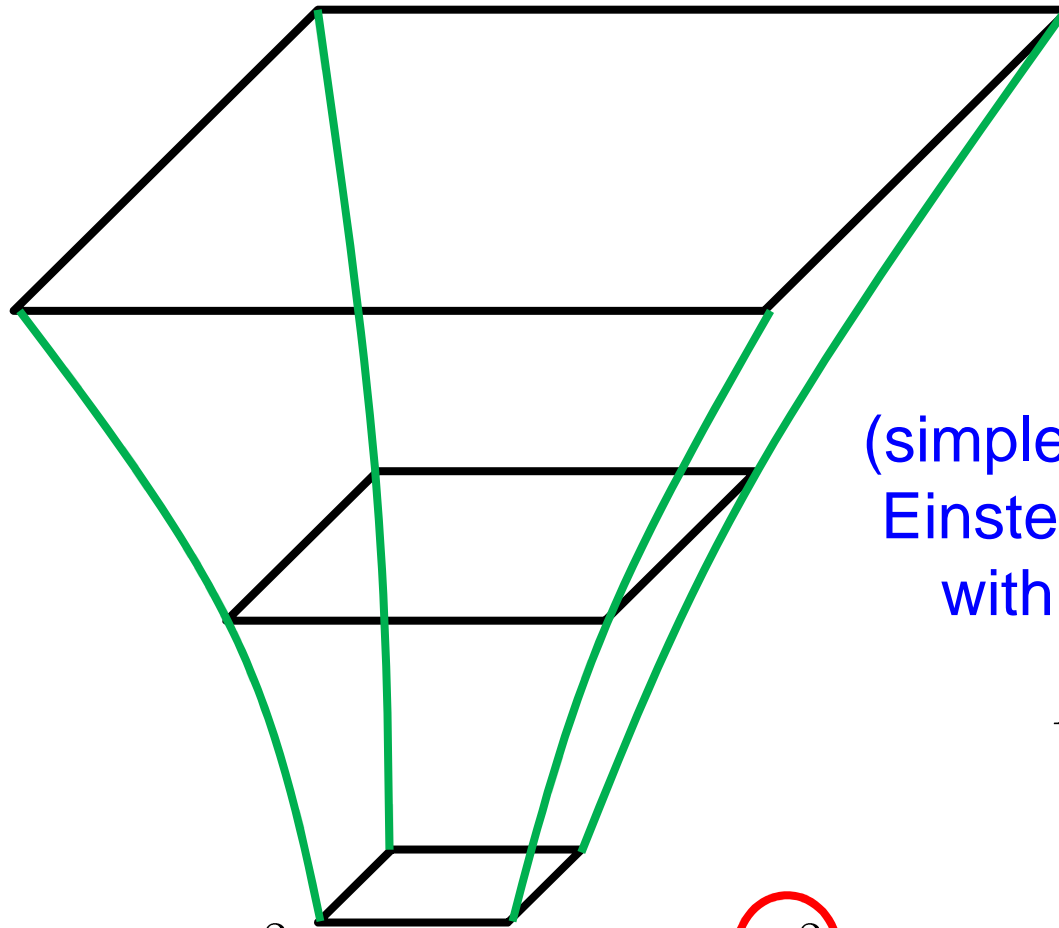
large central charge ($N_c \rightarrow \infty$)

strong coupling ($\lambda \rightarrow \infty$)

anti-de Sitter space:

$$R \sim -\frac{1}{L^2}$$

maximally symmetric geometry with negative curvature



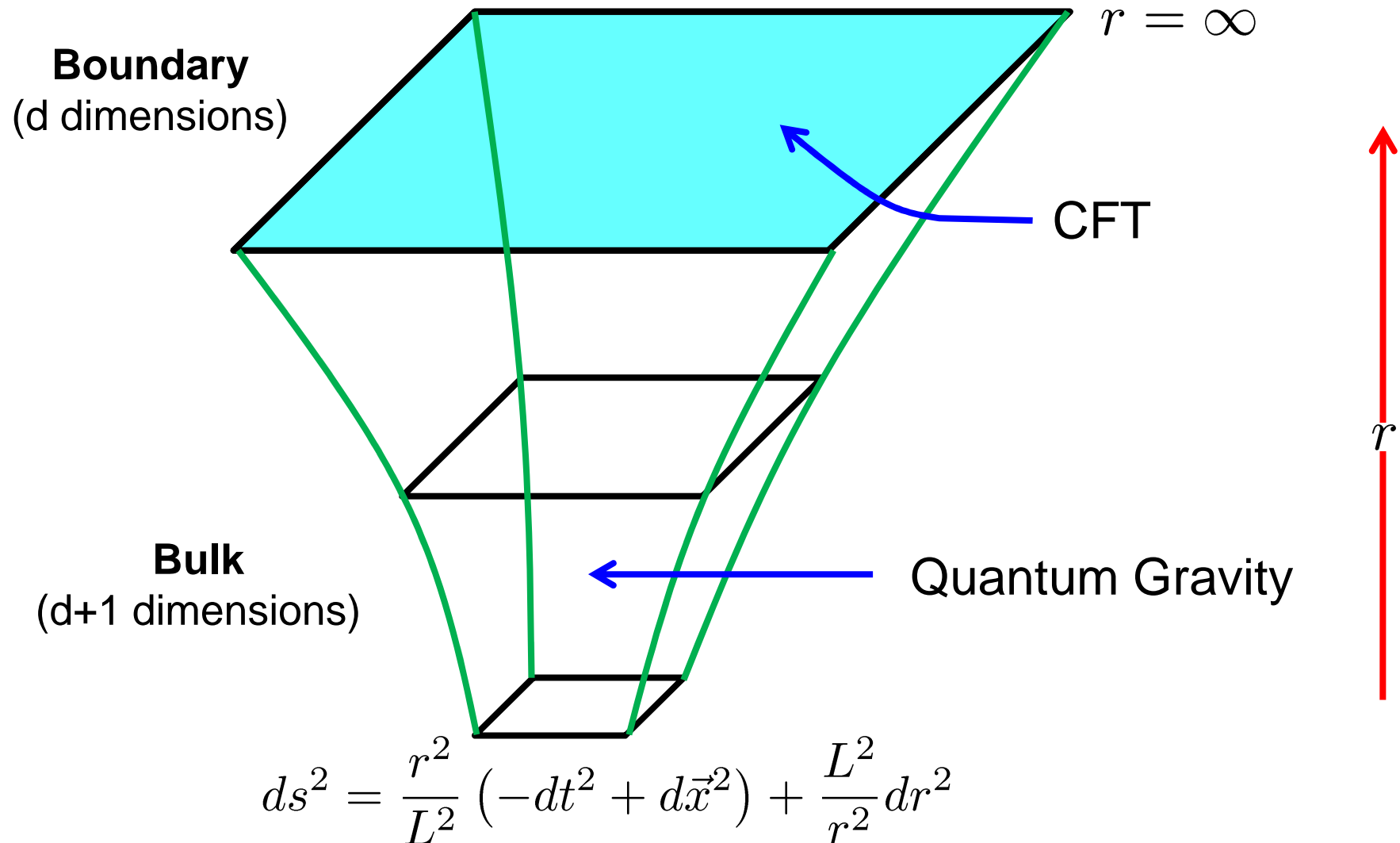
(simplest) solution of
Einstein's equations
with negative Λ :

$$R_{ab} = -\frac{d}{L^2}$$

$$ds^2 = \frac{r^2}{L^2} (-dt^2 + d\vec{x}^2) + \frac{L^2}{r^2} dr^2$$

anti-de Sitter space:

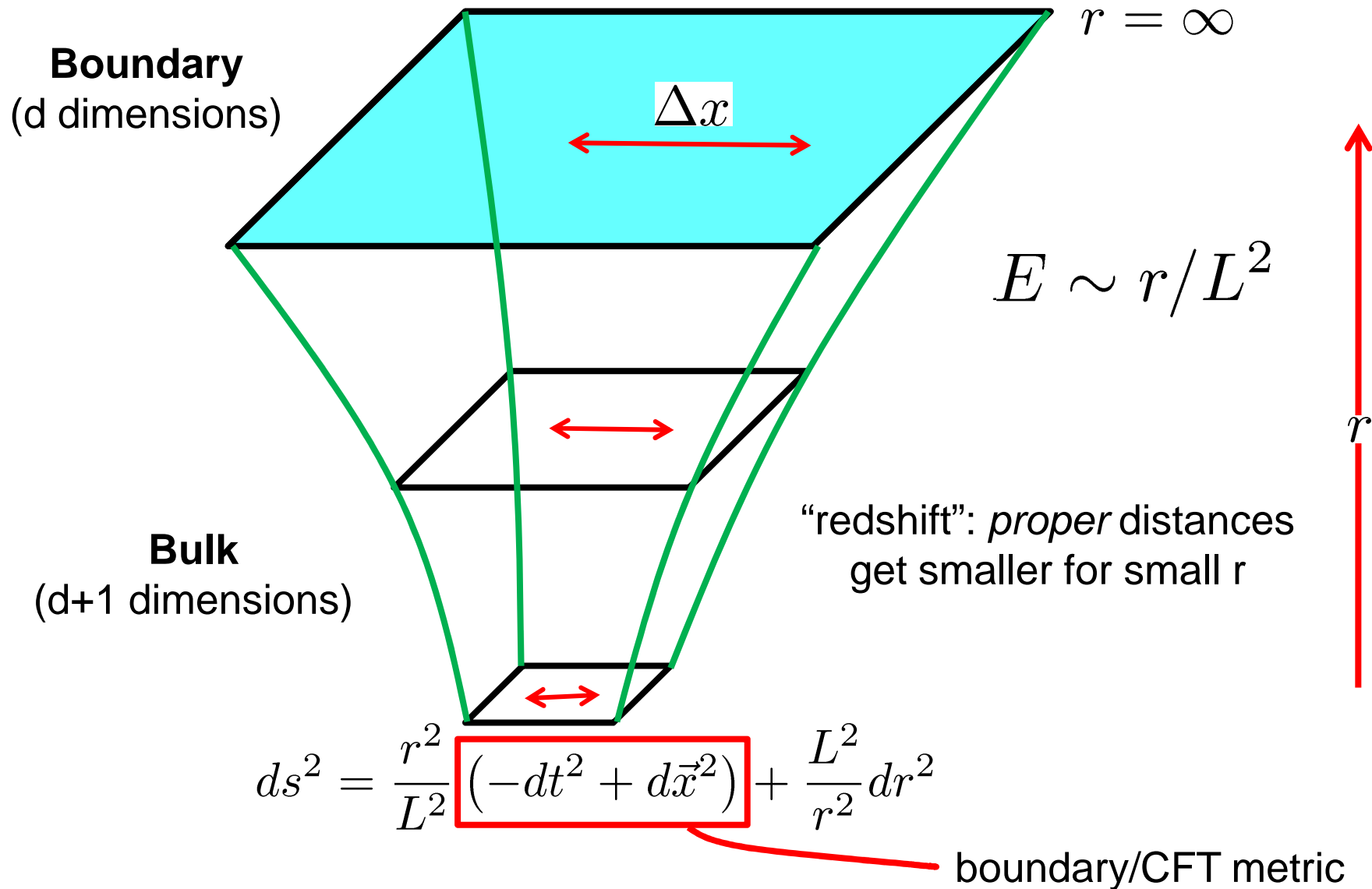
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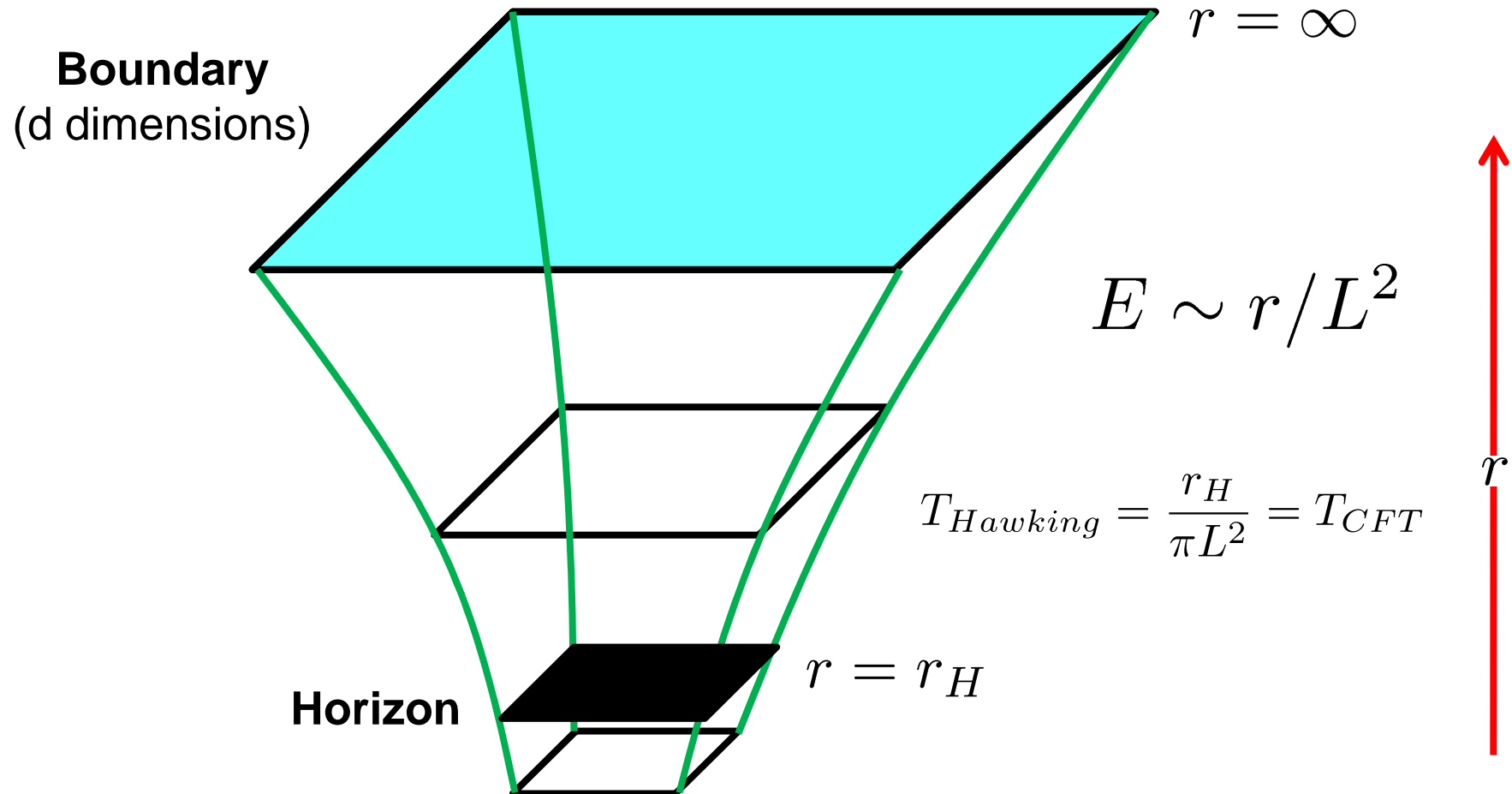
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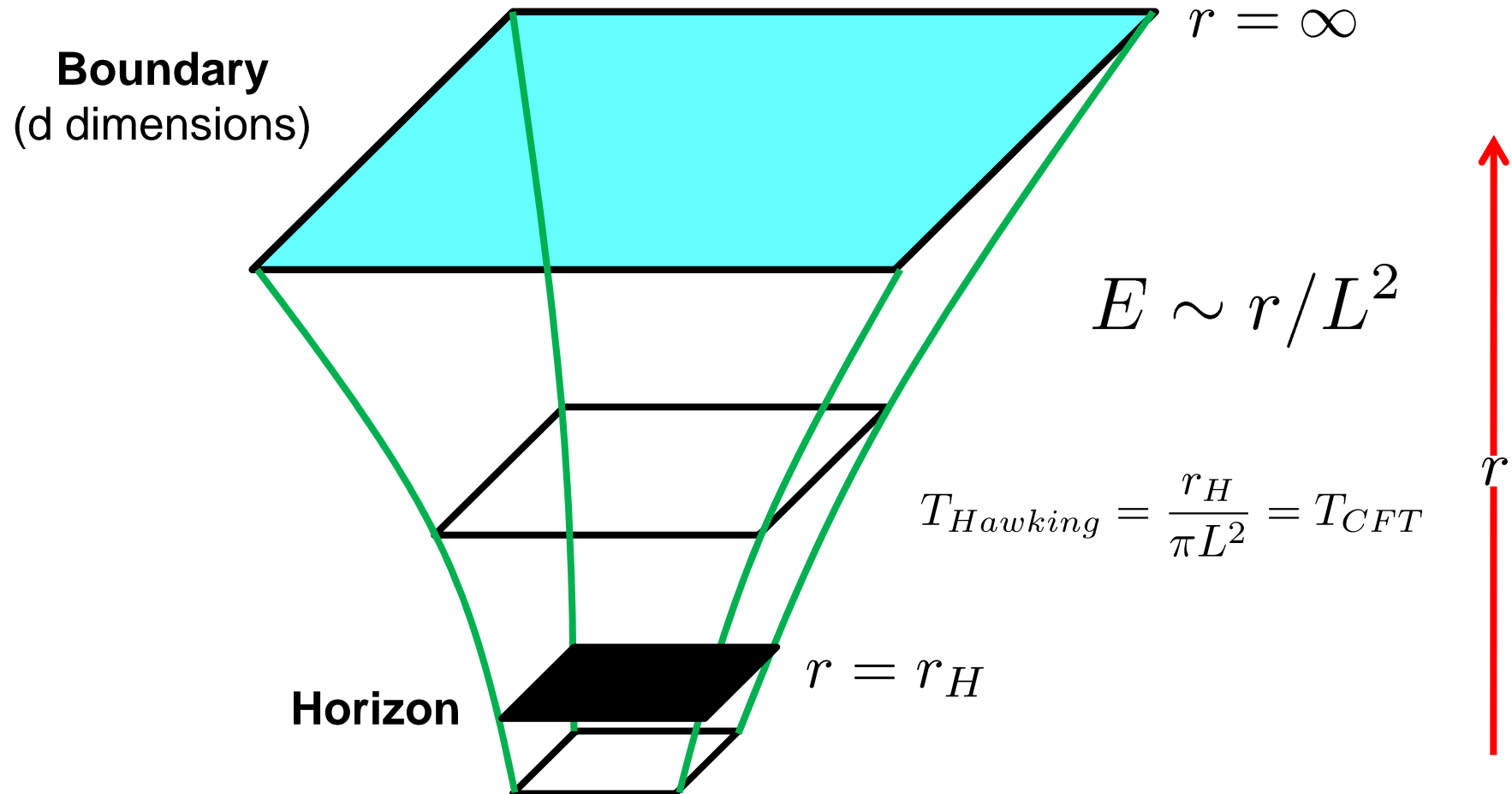


AdS Black Hole = CFT Thermal State

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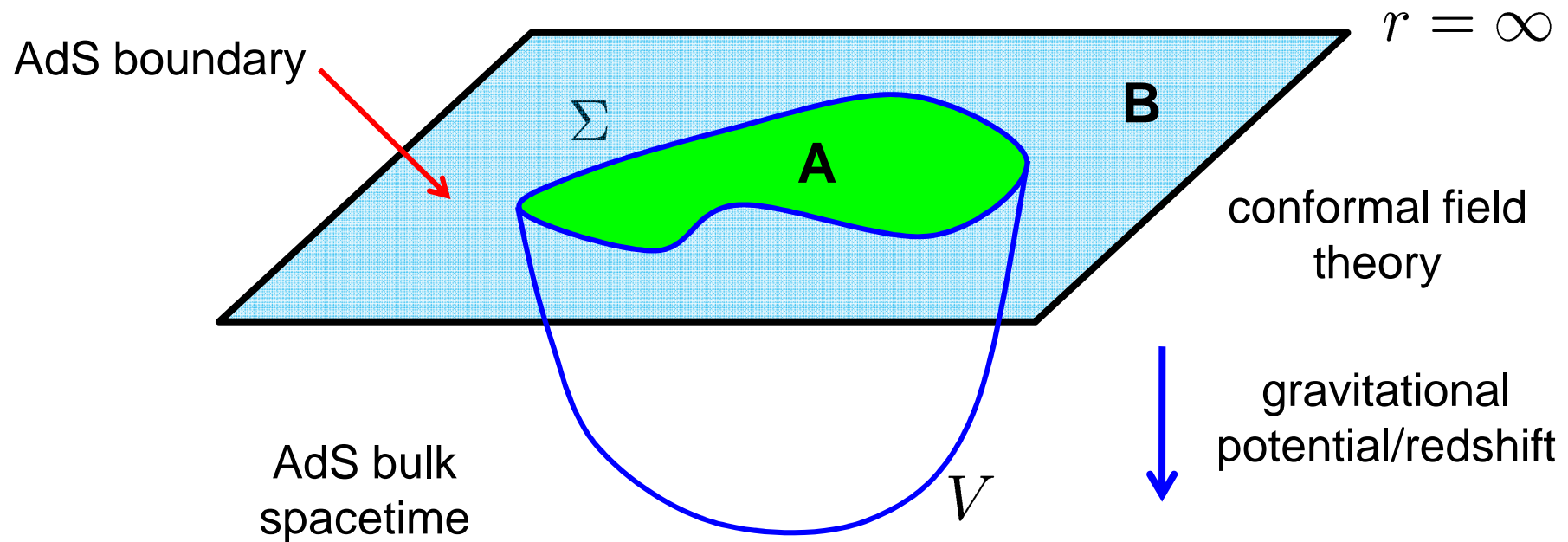
$$R \sim -\frac{1}{L^2}$$



Horizon Entropy = Entropy of Thermal State

$$S = \frac{A_{\text{horizon}}}{4G_N}$$

Holographic Entanglement Entropy:



$$S(A) = \min_{\partial V = \Sigma} \frac{A_V}{4G_N}$$

looks like
BH entropy!

Holographic Entanglement Entropy:

$$S(A) = \min_{\partial V = \Sigma} \frac{A_V}{4G_N} \quad \text{conjecture}$$

Extensive consistency tests:

1) leading contribution yields “area law” $S = \tilde{c}_0 \frac{A_\Sigma}{\delta^{d-2}} + \dots$

. . . . (more tests)

6) for general even d, connection to central charges of CFT

(Hung, RCM & Smolkin, arXiv:1101.5813)

7) derivation of holographic EE for spherical entangling surfaces

(Casini, Huerta & RCM, arXiv:1102.044)

(see also: RCM & Sinha, arXiv:1011.5819)

6) for general even d, connection to central charges of CFT

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- trace anomaly in CFT (with even d) defines central charges

$$d=4: \quad \langle T_\mu^\mu \rangle = \frac{\mathbf{c}}{16\pi^2} I_4 - \frac{\mathbf{a}}{16\pi^2} E_4$$

$$I_4 = C_{\mu\nu\rho\sigma} C^{\mu\nu\rho\sigma} \quad \text{and} \quad E_4 = R_{\mu\nu\rho\sigma} R^{\mu\nu\rho\sigma} - 4R_{\mu\nu} R^{\mu\nu} + R^2$$

- universal/logarithmic contribution to entanglement entropy determined by central charges using trace anomaly, eg,

$$S_{uni} = \log(R/\delta) \frac{1}{2\pi} \int_\Sigma d^2x \sqrt{h} \left[\mathbf{c} \left(C^{ijkl} \tilde{g}_{ik}^\perp \tilde{g}_{jl}^\perp - K_a^{ib} K_b^{ia} + \frac{1}{2} K_a^{ia} K_b^{ib} \right) - \mathbf{a} \mathcal{R} \right]$$

(Solodukhin)

- R&T proposal for holographic EE reproduce precisely this result
- extends to certain higher curvature theories (eg, GB gravity)

$$S = \min_{\partial V = \Sigma} \frac{2\pi}{\ell_p^3} \int_V d^3x \sqrt{h} [1 + \lambda L^2 \mathcal{R}]$$

7) derivation of holographic EE for spherical entangling surfaces

(Casini, Huerta & RCM, arXiv:1102.044)

(see also: RCM & Sinha, arXiv:1011.5819)

- holographic translation for standard calculation of EE is difficult
- new calculation for special case:

CFT in d-dim. flat space and choose $\Sigma = S^{d-2}$ with radius R

→ by conformal mapping relate to thermal entropy
on $\mathcal{H} = R \times H^{d-1}$ with $\mathcal{R} \sim 1/R^2$ and $T=1/2\pi R$

$$S_{EE} = S_{thermal}$$

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$$S_{EE} = S_{thermal} = S_{horizon}$$

→ thermal bath in CFT = black hole in AdS

- can calculate holographic EE for any bulk gravity theory

universal contributions:

$$S = \cdots + (-)^{\frac{d}{2}-1} 4 a_d^* \log(2R/\delta) + \cdots \quad \text{for even } d$$
$$\cdots + (-)^{\frac{d-1}{2}} 2\pi a_d^* + \cdots \quad \text{for odd } d$$

Conclusions:

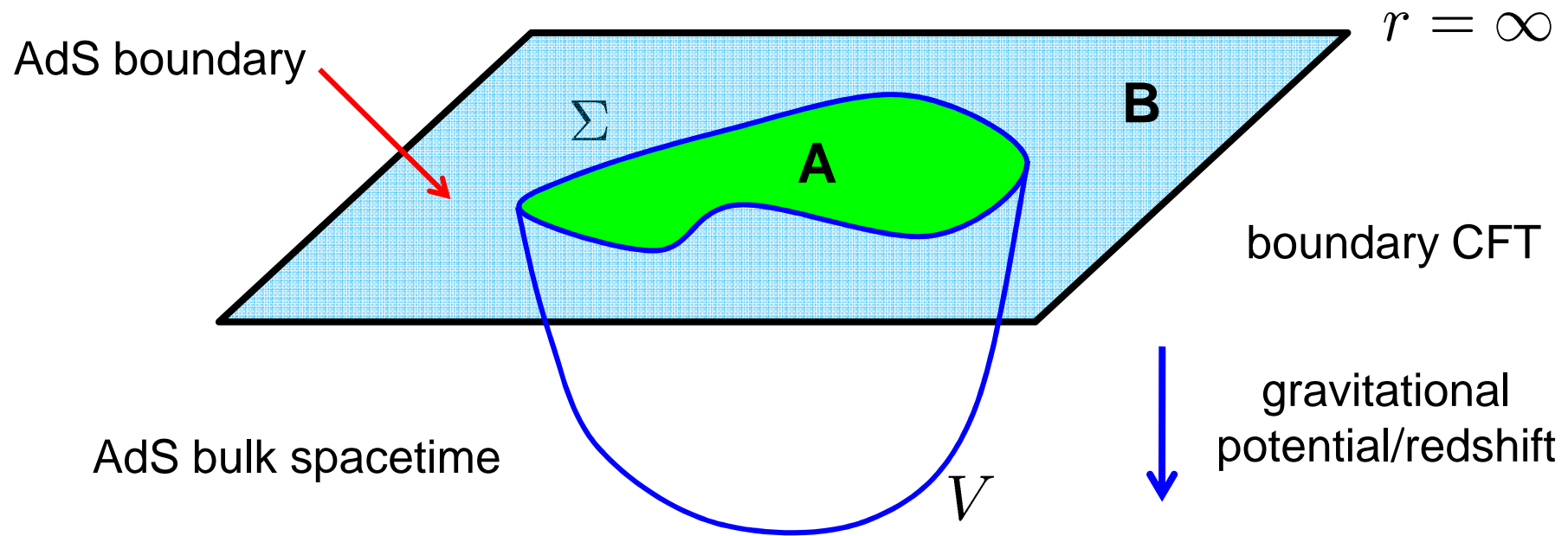
- AdS/CFT correspondence (gauge/gravity duality) has proven an excellent tool to study strongly coupled gauge theories
- **holographic entanglement entropy** is part of an interesting dialogue has opened between string theorists and physicists in a variety of fields (eg, condensed matter, nuclear physics, . . .)
- potential to learn new lessons about general properties of entanglement entropy that have application beyond the context of AdS/CFT correspondence
- potential to learn new lessons about general properties of quantum gravity or string theory

Lots to explore!

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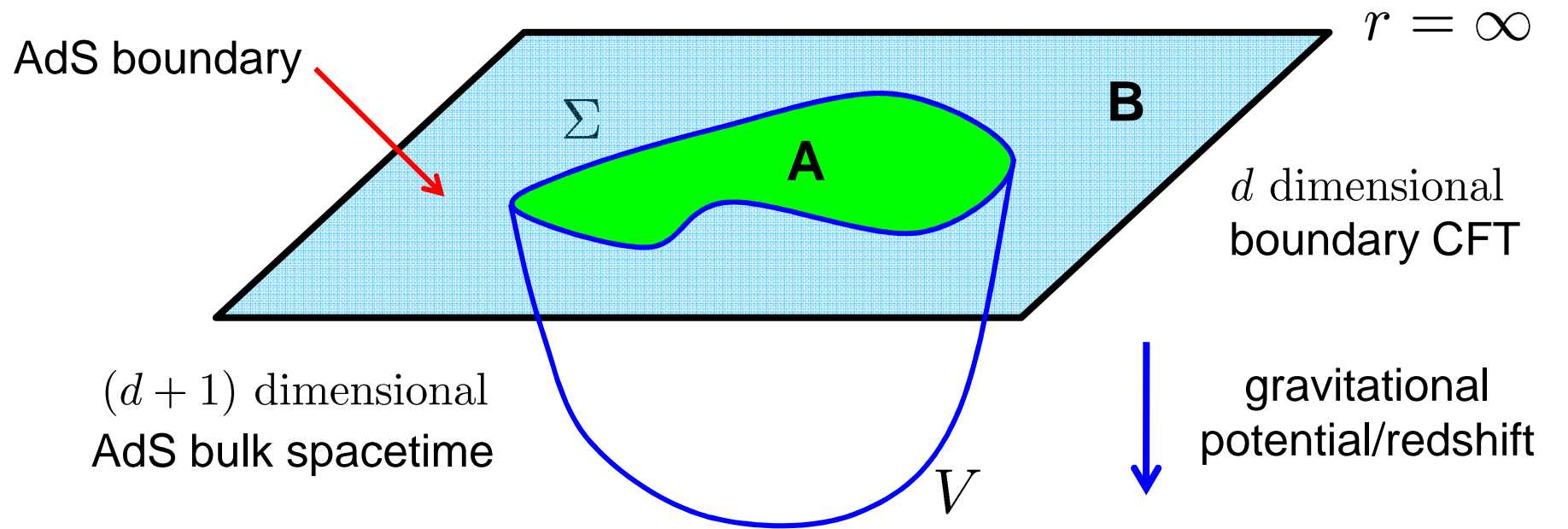
Holographic Entanglement Entropy:



$$S(A) = \min_{\partial V = \Sigma} \frac{A_V}{4G_N}$$

looks like
BH entropy!

Holographic Entanglement Entropy:

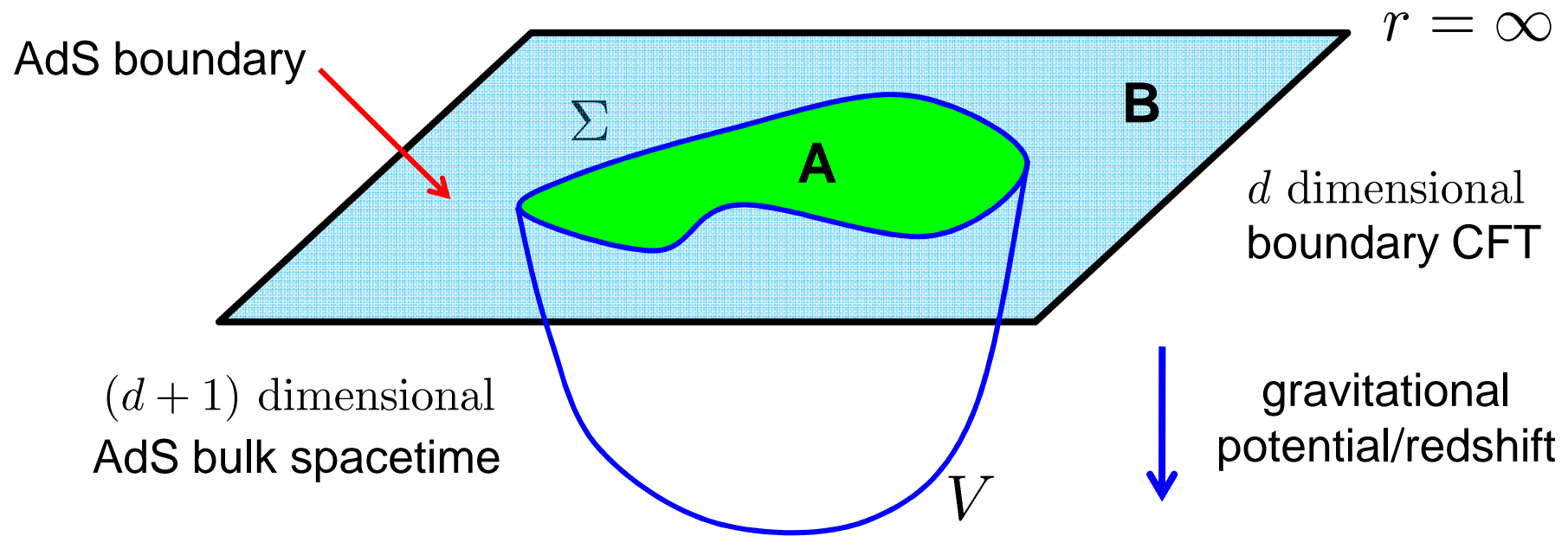


$$S(A) = \underbrace{\text{ext}_{V \sim A}}_{\text{technicalities}} \frac{A_V}{4G_N}$$

$(d-1)$ dimensional

looks like BH entropy!

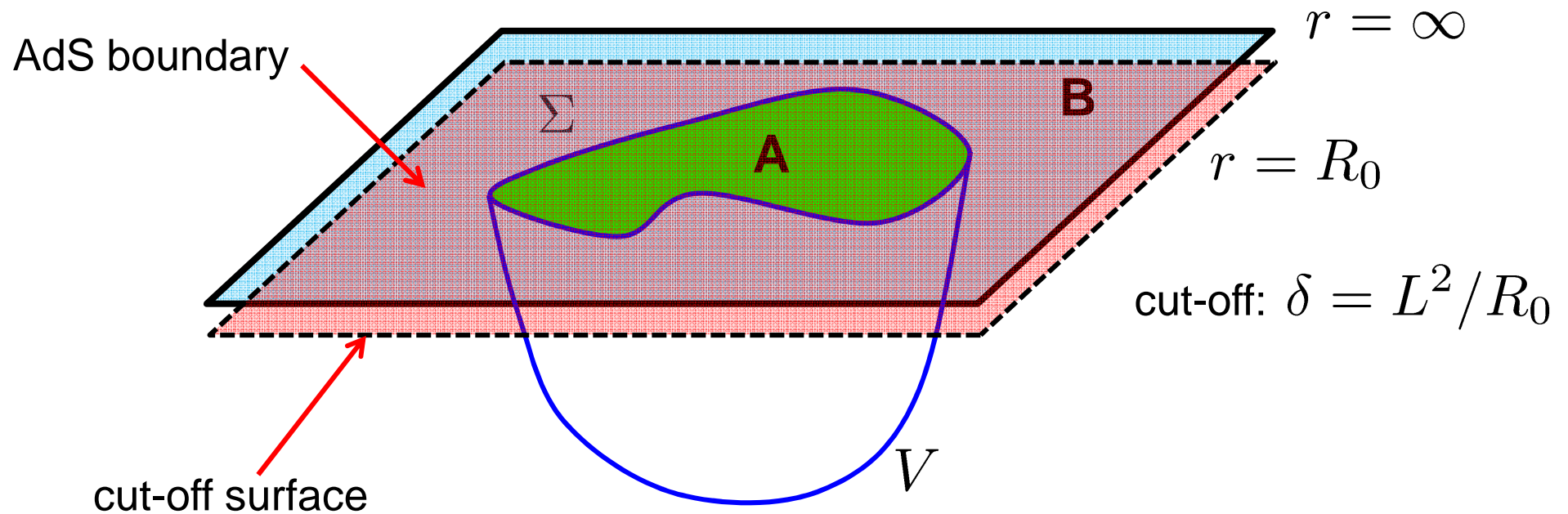
Holographic Entanglement Entropy:



$$S(A) = \text{ext}_{V \sim A} \frac{A_V}{4G_N} = \infty!!$$

- “UV divergence” because area integral extends to $r = \infty$

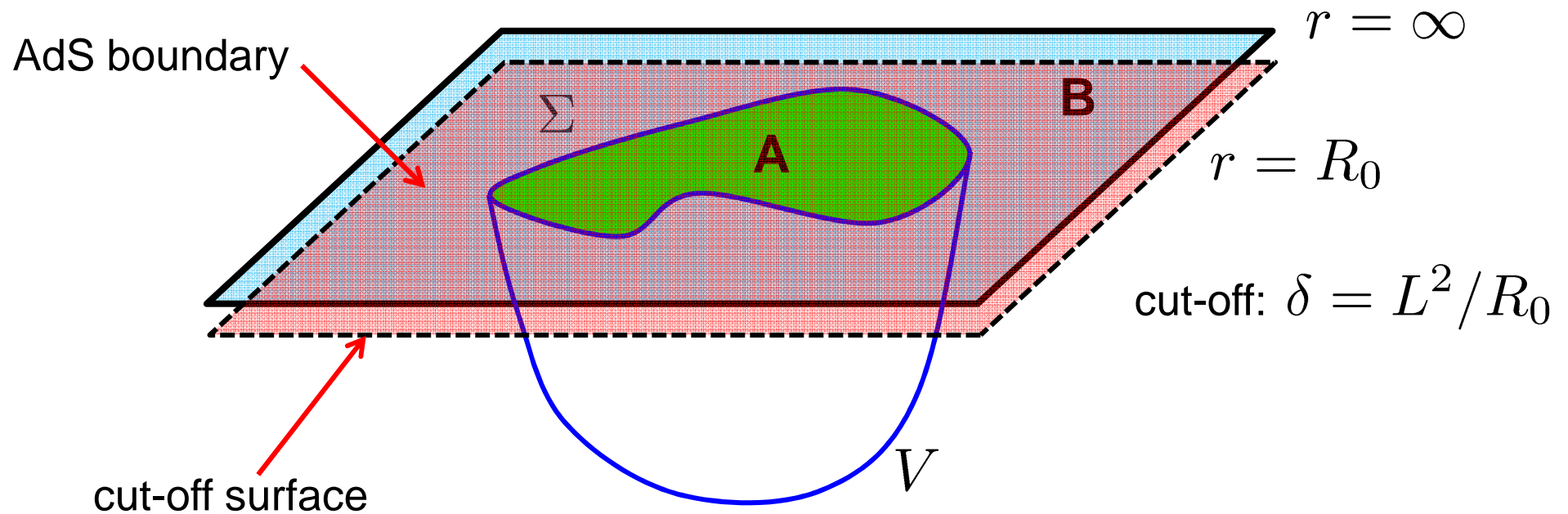
Holographic Entanglement Entropy:



$$S(A) = \text{ext}_{V \sim A} \frac{A_V}{4G_N} \simeq \frac{L^{d-1}}{G_N} \frac{\mathcal{A}_\Sigma}{\delta^{d-2}} + \dots$$

- “UV divergence” because area integral extends to $r = \infty$
- finite result by stopping radial integral at large radius: $r = R_0$
 → short-distance cut-off in boundary theory: $\delta = L^2/R_0$

Holographic Entanglement Entropy:



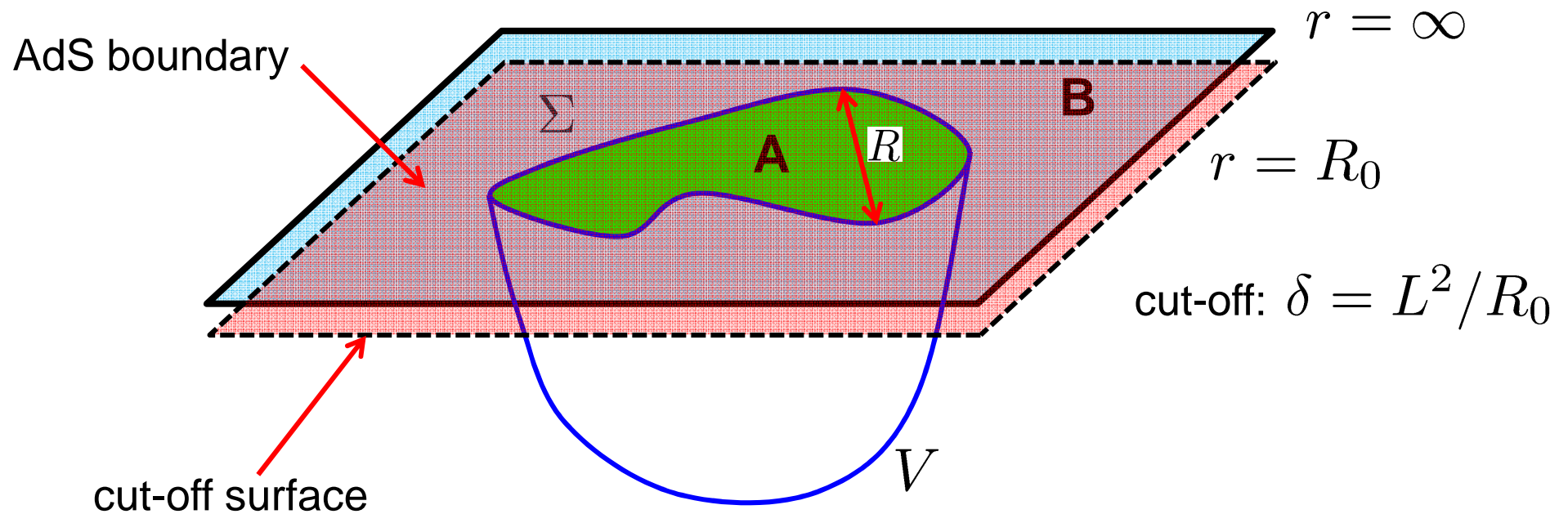
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central charge
(counts dof)

$$(L/\ell_{Planck})^{d-1}$$

"Area Law"

Holographic Entanglement Entropy:



general expression (as desired):

$$S(A) \simeq c_0 (R/\delta)^{d-2} + c_2 (R/\delta)^{d-4} + \dots$$

$$\left\{ \begin{array}{ll} +c_{d-2} \log(R/\delta) + \dots & (d \text{ even}) \\ + \underbrace{c_{d-2} + \dots}_{\text{universal contributions}} & (d \text{ odd}) \end{array} \right.$$

universal contributions

Holographic Entanglement Entropy:

$$S(A) = \text{ext}_{V \sim A} \frac{A_V}{4G_N}$$

conjecture

Extensive consistency tests:

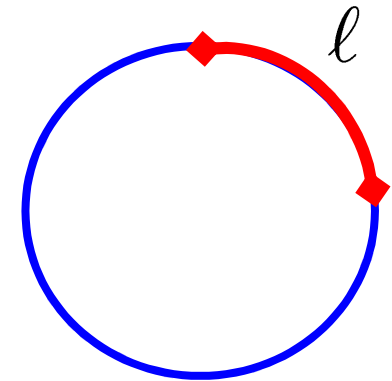
1) leading contribution yields “area law”

$$S \simeq \frac{L^{d-1}}{G_N} \frac{\mathcal{A}_\Sigma}{\delta^{d-2}} + \dots$$

2) recover known results of Calabrese & Cardy
for d=2 CFT

$$S = \frac{c}{3} \log \left(\frac{C}{\pi \delta} \sin \frac{\pi \ell}{C} \right)$$

(also result for thermal ensemble)



C = circumference

Holographic Entanglement Entropy:

$$S(A) = \underset{V \sim A}{\text{ext}} \frac{A_V}{4G_N}$$

conjecture

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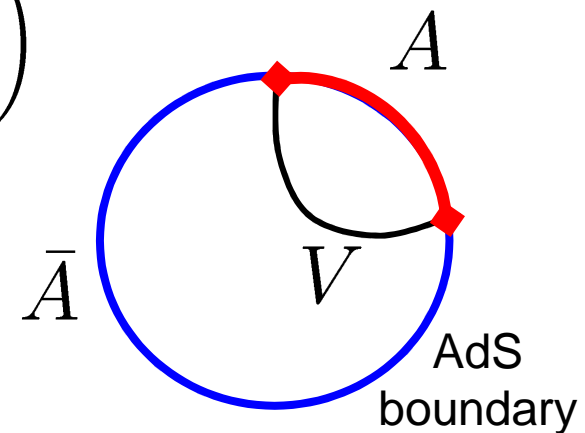
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(also result for thermal ensemble)

3) $S(A) = S(\bar{A})$ in a pure state

→ A and \bar{A} both yield same bulk surface V

(not pure state → horizon in bulk; $S(A) \neq S(\bar{A})$ for thermal state)



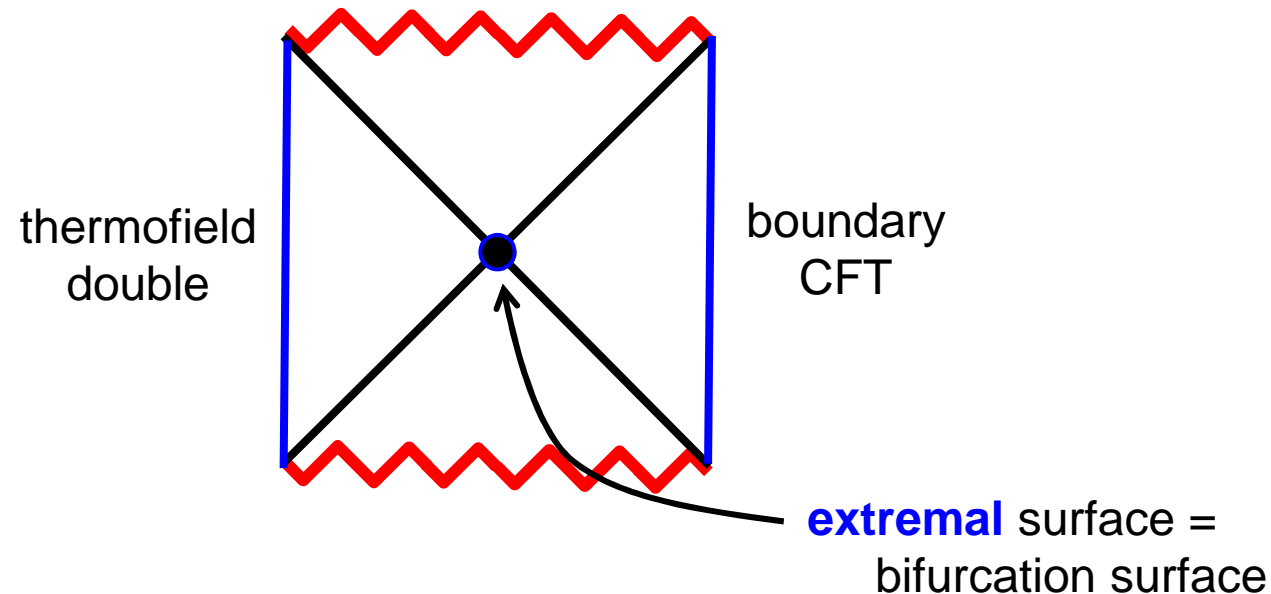
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$$S(A) = \text{ext}_{V \sim A} \frac{A_V}{4G_N}$$

conjecture

Extensive consistency tests:

- 4) Entropy of eternal black hole =
entanglement entropy of boundary CFT & thermofield double
(Headrick)



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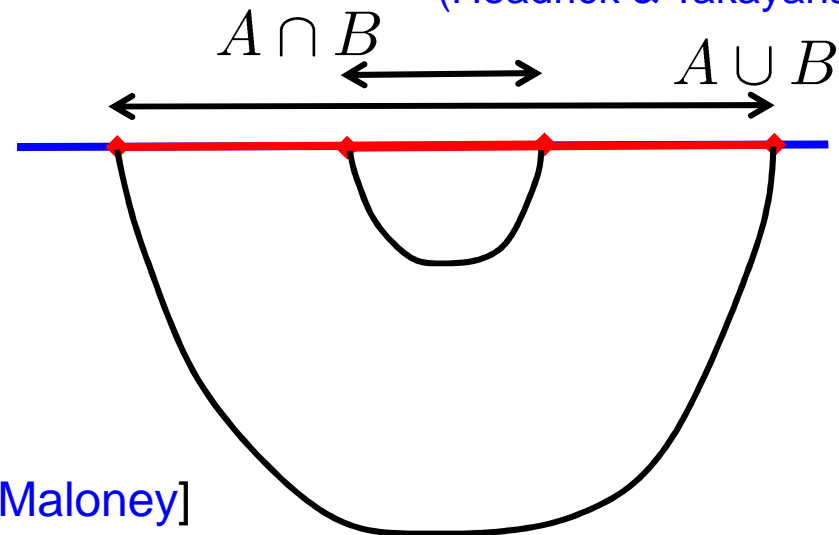
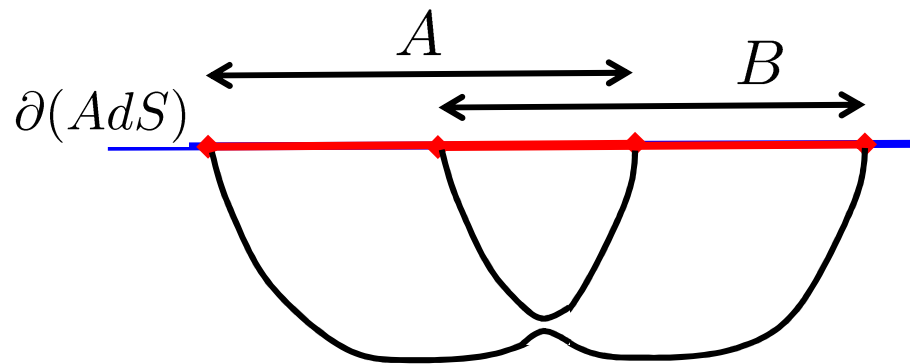
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5) sub-additivity: $S(A \cup B) + S(A \cap B) \leq S(A) + S(B)$

(Headrick & Takayanagi)



[“all” other inequalities: [Hayden, Headrick & Maloney](#)]

Holographic Entanglement Entropy:

$$S(A) = \underset{V \sim A}{\text{ext}} \frac{A_V}{4G_N} \quad \text{conjecture}$$

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Central charges and trace anomaly:

$$d=2: \quad \langle T_\mu{}^\mu \rangle = -\frac{\mathbf{c}}{12} R$$

$$d=4: \quad \langle T_\mu{}^\mu \rangle = \frac{\mathbf{c}}{16\pi^2} I_4 - \frac{\mathbf{a}}{16\pi^2} E_4$$

$$I_4 = C_{\mu\nu\rho\sigma} C^{\mu\nu\rho\sigma} \quad \text{and} \quad E_4 = R_{\mu\nu\rho\sigma} R^{\mu\nu\rho\sigma} - 4R_{\mu\nu} R^{\mu\nu} + R^2$$

- in higher (even) dimensions, number of central charges grows

$$\langle T_\mu{}^\mu \rangle = \sum \mathbf{B}_i (\text{Weyl invariants})_i - 2(-)^{d/2} \mathbf{A} (\text{Euler density})_d$$

(Deser & Schwimmer)

- universal contribution to entanglement entropy determined using trace anomaly (for even d)

(Holzhey, Larsen & Wilczek; Calabrese & Cardy; Takayanagi & Ryu; Schwimmer & Theisen)

$$S_{univ} = \log(R/\delta) \, 2\pi \int_\Sigma d^{d-2}x \, \sqrt{h} \, \frac{\partial \langle T_\lambda{}^\lambda \rangle}{\partial R^{\mu\nu}{}_{\rho\sigma}} \hat{\varepsilon}^{\mu\nu} \hat{\varepsilon}_{\rho\sigma}$$

(RCM & Sinha)

- **partial result!** needs rotational symmetry on entangling surface Σ

Central charges and trace anomaly:

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- in higher (even) dimensions, numbers of central charges grows
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$$d=2: \quad S = \frac{\mathbf{c}}{3} \log \left(\frac{C}{\pi \delta} \sin \frac{\pi \ell}{C} \right) \quad (\text{Holzhey, Larsen \& Wilczek; Calabrese \& Cardy})$$

d=4:

$$S_{uni} = \log(R/\delta) \frac{1}{2\pi} \int_\Sigma d^2x \sqrt{h} \left[\mathbf{c} \left(C^{ijkl} \tilde{g}_{ik}^\perp \tilde{g}_{jl}^\perp - \underline{K_a^{ib} K_b^{ia} + \frac{1}{2} K_a^{ia} K_b^{ib}} \right) - \mathbf{a} \mathcal{R} \right]$$

corrections for general (smooth) Σ (Solodukhin)

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- in higher dimensions, numbers of central charges grows
- universal contribution to entanglement entropy determined using trace anomaly (for even d)
- central charges identified in AdS/CFT using holographic trace anomaly: (Henningson & Skenderis)
e.g., for (boundary) d=4: $a = c = \pi^2 L^3 / \ell_P^3$
- for general d, central charges $\propto (L/\ell_P)^{d-1}$
- for Einstein gravity, all central charges equal for any d
- distinguishing central charges requires higher curvature gravity

Holographic Entanglement Entropy:

- consider more general gravity theory in AdS:

$$I = \int d^{d+1}x \sqrt{-g} \mathcal{L}(g^{ab}, R^{ab}_{cd}, \nabla_e R^{ab}_{cd}, \dots, matter)$$

- how do we evaluate holographic entanglement entropy?

→ take direction from tests of R&T prescription

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$$S(A) = \text{ext}_{V \sim A} \frac{A_V}{4G_N}$$

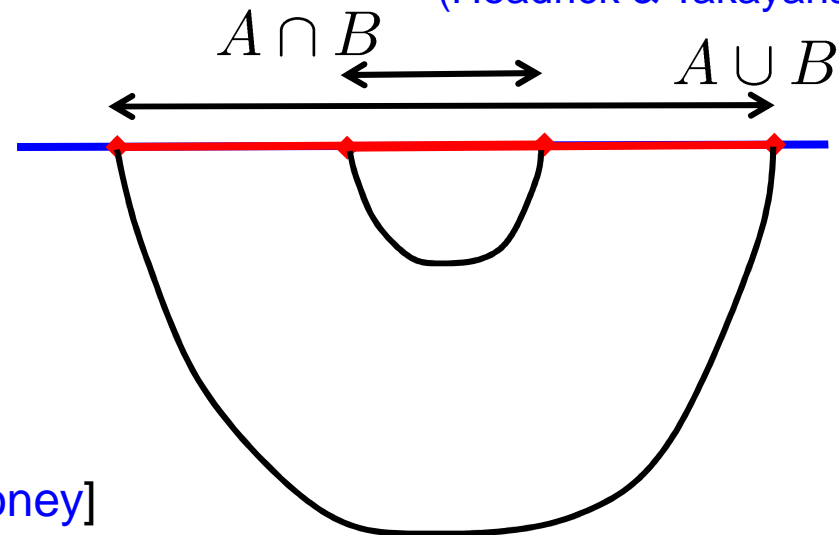
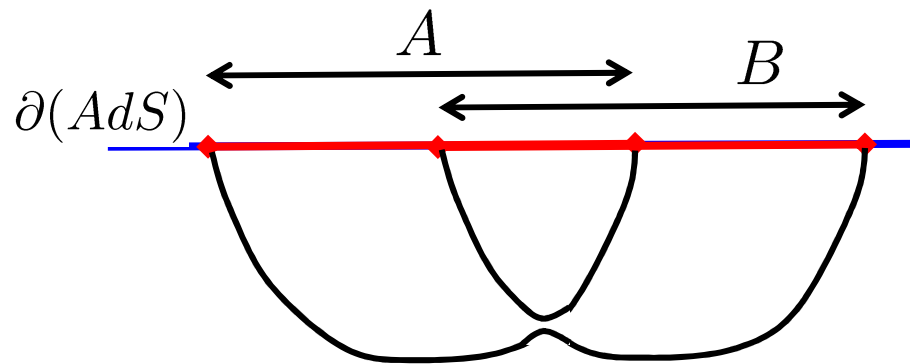
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[other inequalities: Hayden, Headrick & Maloney]

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$$I = \int d^{d+1}x \sqrt{-g} \mathcal{L}(g^{ab}, R^{ab}_{cd}, \nabla_e R^{ab}_{cd}, \dots, \text{matter})$$

- natural **conjecture**: extremize Wald's entropy formula

$$S = -2\pi \int d^{d-1}x \sqrt{h} \frac{\partial \mathcal{L}}{\partial R^{\mu\nu}_{\rho\sigma}} \hat{\varepsilon}^{\mu\nu} \hat{\varepsilon}_{\rho\sigma}$$

- focus on universal term for d=4:

(Solodukhin)

$$S_{uni} = \log(R/\delta) \frac{1}{2\pi} \int_{\Sigma} d^2x \sqrt{h} \left[\mathbf{c} \left(C^{ijkl} \tilde{g}_{ik}^{\perp} \tilde{g}_{jl}^{\perp} - K_a^{ib} K_b^{ia} + \frac{1}{2} K_a^{ia} K_b^{ib} \right) - \mathbf{a} \mathcal{R} \right]$$

- holographic calculation following above conjecture yields

$$S_{uni} = \log(R/\delta) \frac{1}{2\pi} \int_{\Sigma} d^2x \sqrt{h} \left[\mathbf{a} \left(C^{ijkl} \tilde{g}_{ik}^{\perp} \tilde{g}_{jl}^{\perp} - K_a^{ib} K_b^{ia} + \frac{1}{2} K_a^{ia} K_b^{ib} \right) - \mathbf{a} \mathcal{R} \right]$$

→ **conjecture** wrong 🙄

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- triumph of R&T prescription in Einstein gravity!! (**c** = **a**)
- for general gravity action, conjecture is wrong
- **there is nothing wrong with Wald's formula!!**

→ to proceed further, focus on special gravity actions

(Hung, Myers & Smolkin)

Holographic Entanglement Entropy:

- consider special case of Gauss-Bonnet gravity:

$$I = \frac{1}{2\ell_p^3} \int d^5x \sqrt{-g} \left[\frac{12}{L^2} + R + L^2 \frac{\lambda}{2} \underbrace{(R^{abcd} R_{abcd} - 4R_{ab} R^{ab} + R^2)}_{\text{4d Euler density}} \right]$$

4d Euler density

- higher curvature but eom are still **second order!!** (Lovelock)
- studied in detail for stringy gravity in 1980's

(Zwiebach; Boulware & Deser; Wheeler; Myers & Simon;)

- interest recently in AdS/CFT studies – a toy model with $c \neq a$

(eg, Brigante, Liu, Myers, Shenker, Yaida, de Boer, Kulaxizi, Parnachev, Camanho, Edelstein, Buchel, Sinha, Paulos, Escobedo, Smolkin, Cremonini, Hofman,)

- black hole entropy:

(Jacobson & Myers)

$$S_{JM} = \frac{2\pi}{\ell_p^3} \int d^3x \sqrt{h} [1 + \lambda L^2 \mathcal{R}]$$

- **not** precisely same as Wald entropy – agree when K_{ab}^i vanish

(Hung, Myers & Smolkin)

Holographic Entanglement Entropy: (deBoer, Kulaxizi & Parnachev)

- consider special case of Gauss-Bonnet gravity:

$$I = \frac{1}{2\ell_p^3} \int d^5x \sqrt{-g} \left[\frac{12}{L^2} + R + L^2 \frac{\lambda}{2} \left(\underline{R^{abcd} R_{abcd} - 4R_{ab} R^{ab} + R^2} \right) \right]$$

4d Euler density

- second conjecture: extremize JM entropy formula

$$S_{JM} = \frac{2\pi}{\ell_p^3} \int d^3x \sqrt{h} \left[1 + \lambda L^2 \mathcal{R} \right]$$

- again consider universal term for d=4:

(Solodukhin)

$$S_{univ} = \log(\ell/\delta) \frac{1}{2\pi} \int_{\Sigma} d^2x \sqrt{h} \left[\mathbf{c} \left(C^{ijkl} \tilde{g}_{ik}^{\perp} \tilde{g}_{jl}^{\perp} - K_a^{ib} K_b^{ia} + \frac{1}{2} K_a^{ia} K_b^{ib} \right) - \mathbf{a} \mathcal{R} \right]$$

- holographic calculation following above conjecture yields

$$S_{univ} = \log(\ell/\delta) \frac{1}{2\pi} \int_{\Sigma} d^2x \sqrt{h} \left[\mathbf{c} \left(C^{ijkl} \tilde{g}_{ik}^{\perp} \tilde{g}_{jl}^{\perp} - K_a^{ib} K_b^{ia} + \frac{1}{2} K_a^{ia} K_b^{ib} \right) - \mathbf{a} \mathcal{R} \right]$$



passes nontrivial test



(Hung, Myers & Smolkin)

Holographic Entanglement Entropy: (deBoer, Kulaxizi & Parnachev)

- consider special case of Gauss-Bonnet gravity:

$$I = \frac{1}{2\ell_p^3} \int d^5x \sqrt{-g} \left[\frac{12}{L^2} + R + L^2 \frac{\lambda}{2} \underbrace{(R^{abcd} R_{abcd} - 4R_{ab} R^{ab} + R^2)}_{\text{4d Euler density}} \right]$$

4d Euler density

- second **conjecture**: extremize JM entropy formula

$$S_{JM} = \frac{2\pi}{\ell_p^3} \int d^3x \sqrt{h} [1 + \lambda L^2 \mathcal{R}]$$

- ✓ reproduces universal term for any smooth surface in d=4
- ✓ partial results for d=6 (geometries with rotational symmetry;
found new curvature corrections when $K_{ab}^i = 0$)
- ✓ extends to general Lovelock theories for d≥6
- ? still no general result for completely general gravity action ?
→ with sufficient symmetry, Wald entropy seems correct
- ? curious instability to adding handles for $\lambda > 0$? (Ogawa & Takayanagi)

Holographic Entanglement Entropy:

$$S(A) = \underset{V \sim A}{\text{ext}} \frac{A_V}{4G_N} \quad \text{conjecture}$$

Extensive consistency tests:

4) Entropy of eternal black hole =
entanglement entropy of boundary CFT & thermofield double
(Headrick)

5) strong subadditivity: $S(A \cup B) + S(A \cap B) \leq S(A) + S(B)$
(Headrick & Takayanagi)

6) for general even d , connection to central charges of CFT
(Hung, RCM & Smolkin, arXiv:1101.5813)

7) derivation of holographic EE for spherical entangling surfaces
(Casini, Huerta & RCM, arXiv:1102.044)
(see also: RCM & Sinha, arXiv:1011.5819)

Calculating Entanglement Entropy:

$$S_{EE} = -\text{Tr} [\rho_A \log \rho_A]$$

- “standard” approach relies on **replica trick** and calculating Renyi entropy first and taking $n \rightarrow 1$ limit

$$S_n = \frac{1}{1-n} \log \text{Tr} [\rho_A^n] \qquad S_{EE} = \lim_{n \rightarrow 1} S_n$$

- **replica trick** involves path integral of QFT in **singular** n -fold cover of background spacetime
- problematic in holographic framework
 - produce singularity in dual gravity description
(resolved by quantum gravity/string theory?)

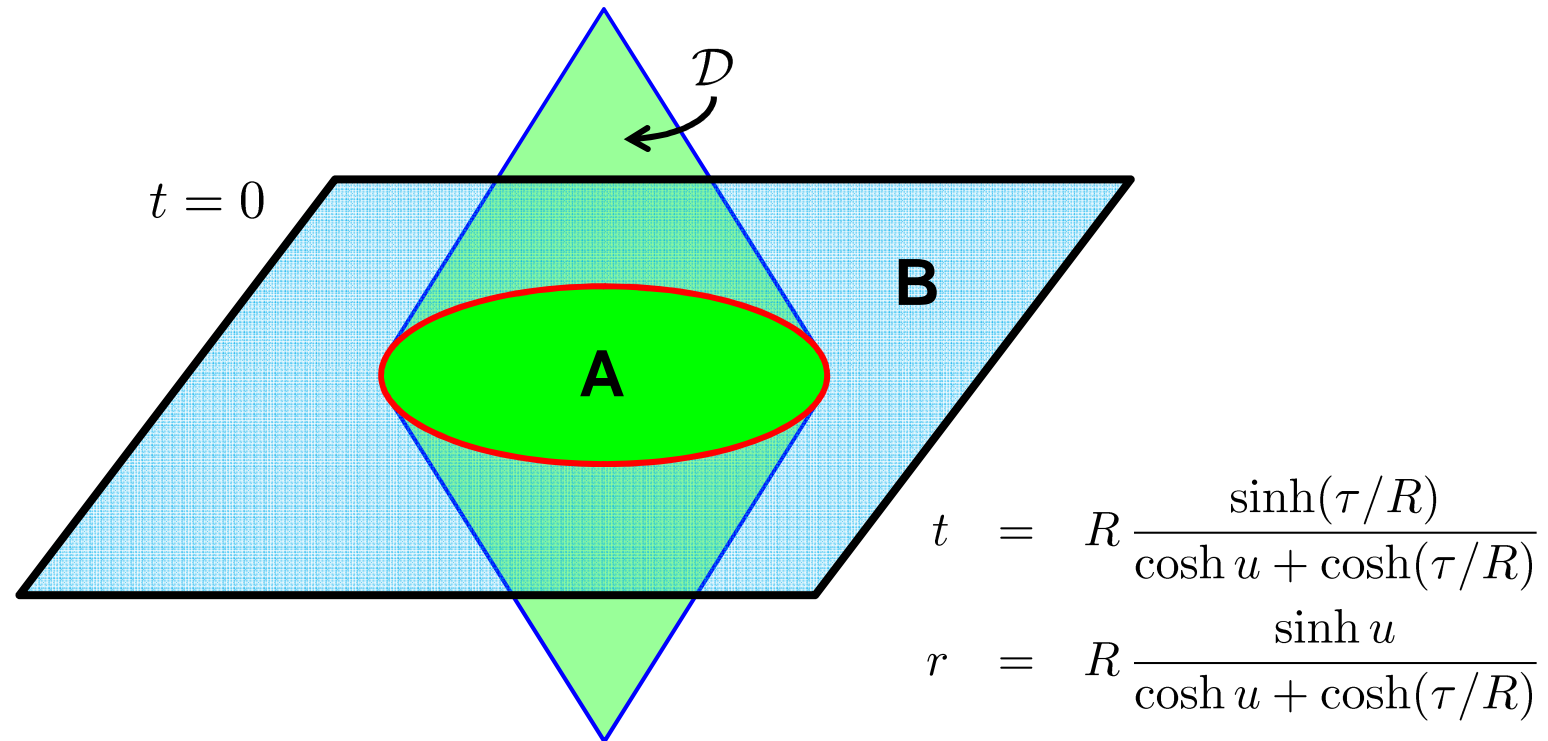
(Fursaev; Headrick)

- **need another calculation with simpler holographic translation**

Calculating Entanglement Entropy:

(Casini, Huerta & RCM)

- take **CFT** in d-dim. flat space and choose $\Sigma = S^{d-2}$ with radius R
→ entanglement entropy: $S_{EE} = -Tr [\rho_A \log \rho_A]$

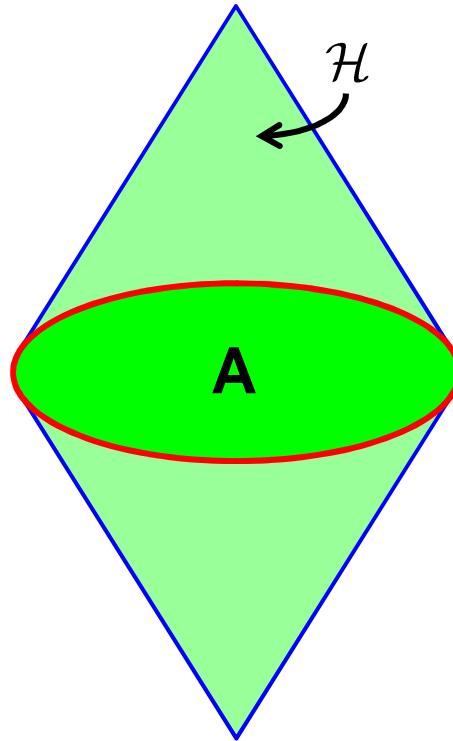


- density matrix ρ_A describes physics in entire causal domain \mathcal{D}
- conformal mapping: $\mathcal{D} \rightarrow \mathcal{H} = R \times H^{d-1}$

General result for any CFT

(Casini, Huerta & RCM)

- take CFT in d-dim. flat space and choose S^{d-2} with radius R
→ entanglement entropy: $S_{EE} = -Tr [\rho_A \log \rho_A]$



$$t = R \frac{\sinh(\tau/R)}{\cosh u + \cosh(\tau/R)}$$
$$r = R \frac{\sinh u}{\cosh u + \cosh(\tau/R)}$$

- conformal mapping: $\mathcal{D} \rightarrow \mathcal{H} = R \times H^{d-1}$

curvature scale: $1/R$

temperature: $T=1/2\pi R$!!

- for CFT: $\rho_{thermal} = U \rho_A U^{-1} \longrightarrow \boxed{S_{EE} = S_{thermal}}$

General result for any CFT

(Casini, Huerta & RCM)

- take CFT in d-dim. flat space and choose S^{d-2} with radius R
 - entanglement entropy: $S_{EE} = -Tr [\rho_A \log \rho_A]$
 - by conformal mapping relate to thermal entropy on $\mathcal{H} = R \times H^{d-1}$ with $\mathcal{R} \sim 1/R^2$ and $T=1/2\pi R$

$$S_{EE} = S_{thermal}$$

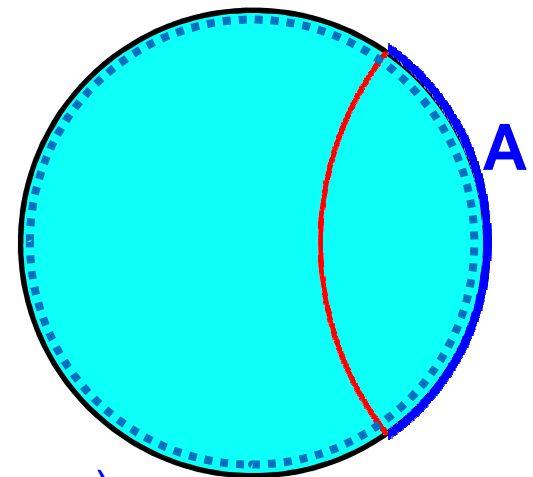
AdS/CFT correspondence:

- thermal bath in CFT = black hole in AdS

$$S_{EE} = S_{thermal} = S_{horizon}$$

- only need to find appropriate black hole
 - topological BH with hyperbolic horizon which intersects A on AdS boundary

(Aminneborg et al; Emparan; Mann; . . .)

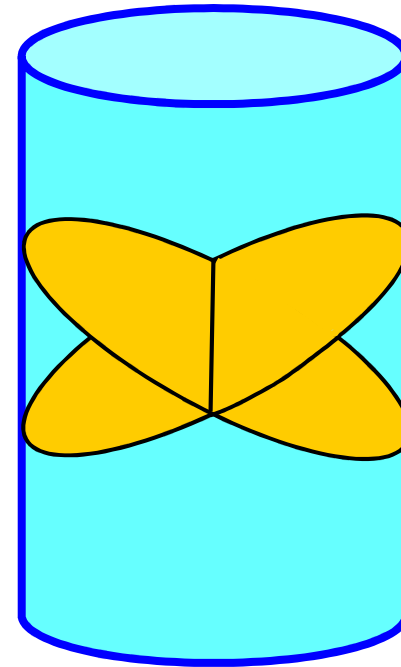


$$S_{EE} = S_{thermal} = S_{horizon}$$

- desired “black hole” is a hyperbolic foliation of empty AdS space

$$ds^2 = \frac{L^2}{z^2} (dz^2 - dt^2 + d\vec{x}^2) d\tau^2 + \rho^2 d\Sigma_2^{d-1} \quad \longrightarrow \quad T = \frac{1}{2\pi R}$$

- “Rindler coordinates” of AdS space



$$S_{EE} = S_{thermal} = S_{horizon}$$

- desired “black hole” is a hyperbolic foliation of empty AdS space

$$ds^2 = \frac{L^2 d\rho^2}{(\rho^2 - L^2)} - \frac{\rho^2 - L^2}{R^2} d\tau^2 + \rho^2 d\Sigma_2^{d-1} \quad \longrightarrow \quad T = \frac{1}{2\pi R}$$

- apply Wald’s formula (for any gravity theory) for horizon entropy:

$$\begin{aligned} S &= -2\pi \int d^{d-1}x \sqrt{h} \frac{\partial \mathcal{L}}{\partial R^{\mu\nu}_{\rho\sigma}} \hat{\varepsilon}^{\mu\nu} \hat{\varepsilon}_{\rho\sigma} \\ &= \frac{2\pi}{\pi^{d/2}} \Gamma(d/2) \frac{a_d^*}{R^{d-1}} V(H^{d-1}) \end{aligned}$$

(RCM & Sinha)

where a_d^* = central charge for “A-type trace anomaly”

for even d

= entanglement entropy defines effective central charge

for odd d

$$S_{EE} = S_{thermal} = S_{horizon}$$

- desired “black hole” is a hyperbolic foliation of empty AdS space

$$ds^2 = \frac{L^2 d\rho^2}{(\rho^2 - L^2)} - \frac{\rho^2 - L^2}{R^2} d\tau^2 + \rho^2 d\Sigma_2^{d-1} \quad \longrightarrow \quad T = \frac{1}{2\pi R}$$

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$$S = \frac{2\pi}{\pi^{d/2}} \Gamma(d/2) \frac{a_d^*}{R^{d-1}} V(H^{d-1})$$

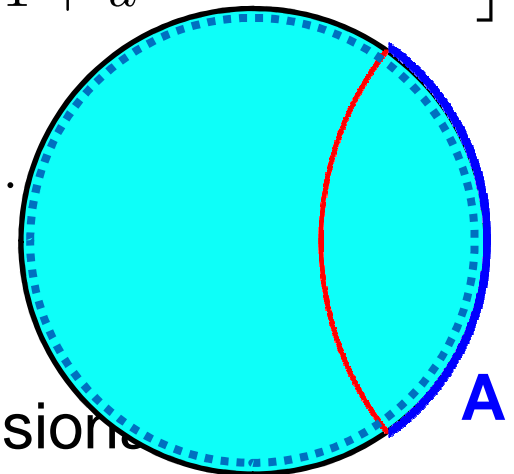
intersection with standard
regulator surface: $z_{min} = \delta$

$$S = a_d^* \frac{4\pi^{\frac{d-3}{2}}}{(d-2)\Gamma(\frac{d-1}{2})} \left(\frac{R}{\delta}\right)^{d-2} + \dots$$



“area law” for d-dimensions

$$ds^2 = R^2 \left[\frac{du^2}{1+u^2} + u^2 d\Omega_2^{d-2} \right]$$



$$S_{EE} = S_{thermal} = S_{horizon}$$

- desired “black hole” is a hyperbolic foliation of empty AdS space

$$ds^2 = \frac{L^2 d\rho^2}{(\rho^2 - L^2)} - \frac{\rho^2 - L^2}{R^2} d\tau^2 + \rho^2 d\Sigma_2^{d-1} \quad \longrightarrow \quad T = \frac{1}{2\pi R}$$

- apply Wald’s formula (for any gravity theory) for horizon entropy:

$$S = \frac{2\pi}{\pi^{d/2}} \Gamma(d/2) \frac{a_d^*}{R^{d-1}} V(H^{d-1})$$

$$ds^2 = R^2 \left[\frac{du^2}{1+u^2} + u^2 d\Omega_2^{d-2} \right]$$

universal contributions:

$$S = \dots + (-)^{\frac{d}{2}-1} 4 a_d^* \log(2R/\delta) + \dots \quad \text{for even } d$$

$$\dots + (-)^{\frac{d-1}{2}} 2\pi a_d^* + \dots \quad \text{for odd } d$$

- discussion extends to case with background: $R^{1,d-1} \rightarrow R \times S^{d-1}$

Holographic Renyi entropy:

- turn to Renyi entropy (close cousin of entanglement entropy)

$$S_n = \frac{1}{1-n} \log \text{Tr} [\rho_A^n] \qquad S_{EE} = \lim_{n \rightarrow 1} S_n$$

- universal contribution (for even d)

$$S_n = \cdots + \text{constant} \times \log(R/\delta) + \cdots$$

Holographic Renyi entropy:

- turn to Renyi entropy (close cousin of entanglement entropy)

$$S_n = \frac{1}{1-n} \log \text{Tr} [\rho_A^n] \qquad S_{EE} = \lim_{n \rightarrow 1} S_n$$

- universal contribution (for even d)

$$\text{d=2: } S_n = \dots + \frac{c}{6} \left(1 + \frac{1}{n} \right) \log (R/\delta) + \dots$$

(Calabrese & Cardy)

- (almost) no calculations for $d > 2$

Holographic Renyi entropy:

- turn to Renyi entropy (close cousin of entanglement entropy)

$$S_n = \frac{1}{1-n} \log \text{Tr} [\rho_A^n] \qquad S_{EE} = \lim_{n \rightarrow 1} S_n$$

- “standard” calculation involves **singular** n-fold cover of spacetime
→ problematic for translation to dual AdS gravity
- our previous derivation lead to thermal density matrix

$$\rho_A = U^{-1} \frac{e^{-H/T_0}}{\text{Tr} [e^{-H/T_0}]} U \qquad \text{with} \quad T_0 = \frac{1}{2\pi R}$$

$$\text{→} \quad \text{Tr} [\rho_A^n] = \frac{\text{Tr} [e^{-nH/T_0}]}{\text{Tr} [e^{-H/T_0}]^n} \quad \text{←} \quad \text{partition function at new temperature, } T = T_0/n$$

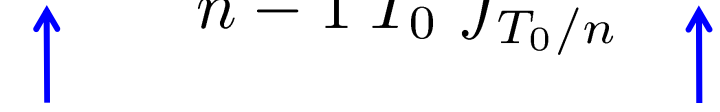
Holographic Renyi entropy:

- turn to Renyi entropy (close cousin of entanglement entropy)

$$S_n = \frac{1}{1-n} \log \text{Tr} [\rho_A^n] \qquad S_{EE} = \lim_{n \rightarrow 1} S_n$$

- “standard” calculation involves **singular** n-fold cover of spacetime
→ problematic for translation to dual AdS gravity
- with bit more work, find convenient formula:

$$S_n = \frac{n}{n-1} \frac{1}{T_0} \int_{T_0/n}^{T_0} S(T) dT \quad \text{where} \quad T_0 = \frac{1}{2\pi R}$$



Renyi entropy for spherical Σ thermal entropy on hyperbolic space H^{d-1}

- in holographic framework, need to know topological black hole solutions for arbitrary temperature

Holographic Renyi entropy:

- Renyi entropy of CFT for spherical entangling surface:

$$S_n = \frac{n}{n-1} \frac{1}{T_0} \int_{T_0/n}^{T_0} S(T) dT \quad \text{where} \quad T_0 = \frac{1}{2\pi R}$$

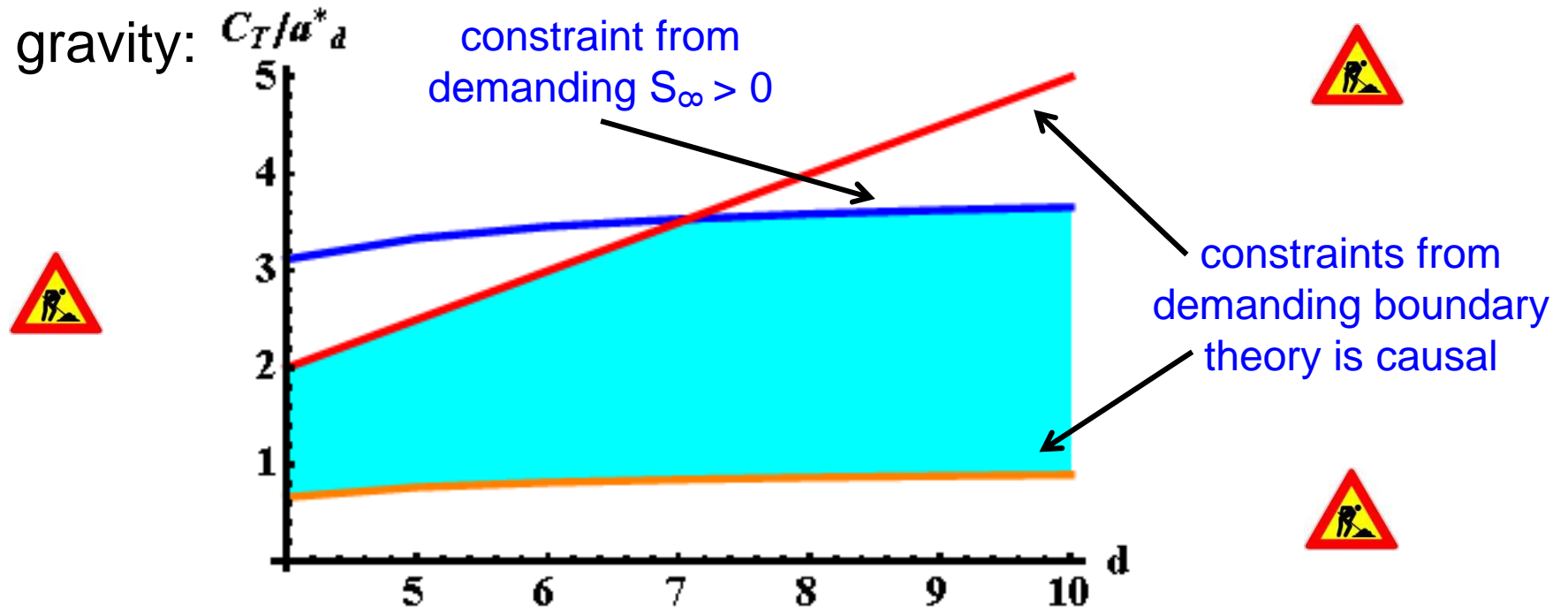
- need to know topological black holes for arbitrary temperature
- focus on gravity theories where we can calculate: Einstein, Gauss-Bonnet, Lovelock, quasi-topological,
- for example, with GB gravity and (boundary) d=4:

$$S_n = \frac{n}{n-1} \frac{V(H^3)}{4\pi} \frac{3c-a}{3a-c} (1-x^2) \left[(5a-c)x^2 - (13a-5c) + 4a \frac{2ax^2 - (a-c)}{(3a-c)x^2 - (a-c)} \right]$$

$$\text{where } 0 = x^3 - \frac{3a-c}{5a-c} \left(\frac{x^2}{n} + x \right) + \frac{1}{n} \frac{a-c}{5a-c}$$

- no elegant result as was found for $d=2$ CFT, ie, S_n depends on both central charges and dependence on n does not factor out
- further work (with quasi-topological gravity) shows the universal coefficient depends on more data from the boundary CFT than central charges appearing in the trace anomaly (eg, t_4)
- **preliminary** work indicates positivity of Renyi entropies may constrain gravitational couplings in higher curvature models

GB gravity: C_T/a^*_d



Conclusions:

- AdS/CFT correspondence (gauge/gravity duality) has proven an excellent tool to study strongly coupled gauge theories
- holographic entanglement entropy is part of an interesting dialogue has opened between string theorists and physicists in a variety of fields (eg, condensed matter, nuclear physics, . . .)
- potential to learn lessons about issues in boundary theory
eg, readily calculate Renyi entropies for wide class of theories in higher dimensions
- potential to learn lessons about issues in bulk gravity theory
eg, holographic entanglement entropy may give new insight into quantum gravity or emergent spacetime

(eg, van Raamsdonk)

Lots to explore!