

Interplay between Randers metrics and the causal geometry of stationary spacetimes

Miguel Ángel Javaloyes (Universidad de Murcia)

Partially supported by MINECO project MTM2012-34037, Regional J. Andalucía Grant P09-FQM-4496 and Fundación Séneca project 04540/GERM/06, Spain

IPMU (Tokyo), February 10th, 2014



First part: preliminaries on Finsler metrics and stationary spacetimes

- Introduction to Finsler metrics
- Introduction to stationary spacetimes

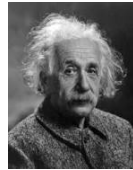
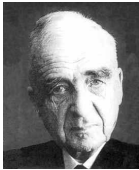
Second part: causality of stationary spacetimes via the Fermat metric

- Fermat Principle in stationary spacetime and Fermat metrics
- Introduction to Causality of a spacetime
- Characterization of causality of a stationary spacetime in terms of the Fermat metric

Third part: almost isometries of Finsler metrics and K -isometries of a stationary spacetime

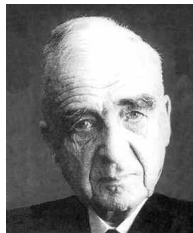
- Almost isometries of quasi-metrics
- Almost isometries of Finsler metrics
- Applications to the study of K -isometries of a stationary spacetime

First part: preliminaries on Finsler metrics and stationary spacetimes



DEFINITION: $F : TM \rightarrow [0, +\infty)$ continuous and

- ① smooth in $TM \setminus \{0\}$
- ② **Positively homogeneous** of degree one
 $F(\lambda v) = \lambda F(v)$ for all $\lambda > 0$
- ③ Fiberwise **strongly convex** square:



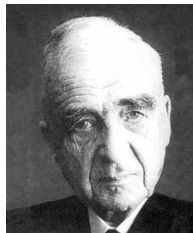
PAUL FINSLER (1894-1970)

$$g_v(w, z) = \frac{1}{2} \frac{\partial^2}{\partial t \partial s} F(v + tw + sz)^2|_{t=s=0} = \frac{1}{2} \text{Hess}(F^2)_v(w, z)$$

for every $w, z \in T_{\pi(v)}M$. Then $g_v(w, w) > 0$ for every $0 \neq w \in T_{\pi(v)}M$.

DEFINITION: $F : TM \rightarrow [0, +\infty)$ continuous and

- ① smooth in $TM \setminus \{0\}$
- ② **Positively homogeneous** of degree one
 $F(\lambda v) = \lambda F(v)$ for all $\lambda > 0$
- ③ Fiberwise **strongly convex** square:



PAUL FINSLER (1894-1970)

$$g_v(w, z) = \frac{1}{2} \frac{\partial^2}{\partial t \partial s} F(v + tw + sz)^2 \Big|_{t=s=0} = \frac{1}{2} \text{Hess}(F^2)_v(w, z)$$

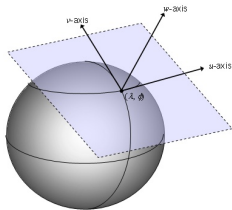
for every $w, z \in T_{\pi(v)}M$. Then $g_v(w, w) > 0$ for every $0 \neq w \in T_{\pi(v)}M$.

It can be showed that this implies:

- F is positive in $TM \setminus \{0\}$ and F^2 is C^1 on TM .
- **Triangle inequality** holds in the fibers

Interpretation of the fundamental tensor

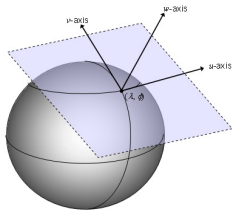
- $S = \text{Indicatrix of } F = \{v \in A : F(v) = 1\}$ (the unit sphere of F).
- The fundamental tensor g_v of F in $v \in S$ coincides with the second fundamental form of S in the hyperplane tangent to S
- $g_v(v, v) = F(v)^2$
- v and $T_v S$ are g_v -orthogonal



Interpretation of the fundamental tensor

- $S = \text{Indicatrix of } F = \{v \in A : F(v) = 1\}$ (the unit sphere of F).
- The fundamental tensor g_v of F in $v \in S$ coincides with the second fundamental form of S in the hyperplane tangent to S
- $g_v(v, v) = F(v)^2$
- v and $T_v S$ are g_v -orthogonal

The indicatrix contains all the geometric information of F



Non-symmetric “distance”

- We can define the **length** of a curve: $\ell_F(\gamma) = \int_a^b F(\dot{\gamma}) ds$
- and then the **distance** between two points:
 $\text{dist}(p, q) = \inf_{\gamma \in C^\infty(p, q)} \ell_F(\gamma)$
- dist is **non-symmetric** because F is **non-reversible**
- the length of a curve $t \rightarrow \gamma(t)$ is different from the length of its reverse $t \rightarrow \gamma(-t)$!!

Non-symmetric “distance”

- We can define the **length** of a curve: $\ell_F(\gamma) = \int_a^b F(\dot{\gamma}) ds$
- and then the **distance** between two points:
 $\text{dist}(p, q) = \inf_{\gamma \in C^\infty(p, q)} \ell_F(\gamma)$
- dist is **non-symmetric** because F is **non-reversible**
- the length of a curve $t \rightarrow \gamma(t)$ is different from the length of its reverse $t \rightarrow \gamma(-t)$!!

We have to distinguish between **forward and backward**:

- balls $B^+(p, r) = \{x \in M : \text{dist}(p, x) < r\}$ and

$$B^-(p, r) = \{x \in M : \text{dist}(x, p) < r\}$$

- Cauchy sequence
- topological completeness
- geodesical completeness

Randers metrics

- Randers metrics in a manifold M is a function $R : TM \rightarrow \mathbb{R}$ defined as:

$$R(v) = \sqrt{h(v, v)} + \omega(v)$$

where h is Riemannian and ω a 1-form with $\|\omega\|_h < 1 \ \forall x \in M$,

- are basic examples of **non-reversible** Finsler metrics: $R(-v) \neq R(v)$.
- Named after the norwegian physicist Gunnar Randers (1914-1992):



Randers, G.: On an asymmetrical metric in the fourspace of General Relativity. Phys. Rev. (2) **59**, 195–199 (1941)



G. RANDERS AND A. EINSTEIN

Zermelo metrics

Given a Riemannian metric g ,
Zermelo metric:

$$Z(v) = \sqrt{\frac{1}{\alpha}g(v, v) + \frac{1}{\alpha^2}g(W, v)^2} - \frac{1}{\alpha}g(W, v),$$

where $\alpha = 1 - g(W, W)$.

It is of Randers type

Geodesics minimize time in the
presence of a wind or current
 W .



Matsumoto metrics


Given a Riemannian metric g , and a one-form β

$$M(v) = \frac{g(v, v)}{\sqrt{g(v, v)} - \beta(v)}$$

defined in

$$A = \{v \in TM : \sqrt{g(v, v)} > 2\beta(v)\}$$

Geodesics minimize time in the presence of a slope

 M. Matsumoto. A slope of a mountain is a Finsler surface with respect to a time measure, J. Math. Kyoto Univ., 29 (1989), pp. 17–25



MOUNT FUJI (NEAR TOKYO)



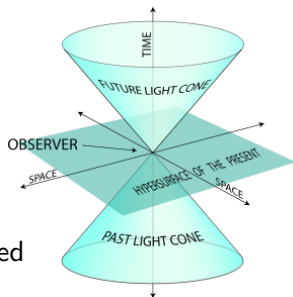
MAKOTO MATSUMOTO (19?? –2005)

Stationary spacetimes

- A Lorentzian manifold (M, g) with index 1 $(+, \dots, +, -)$

- $v \in TM$ is $\begin{cases} \text{timelike if } g(v, v) < 0 \\ \text{lightlike if } g(v, v) = 0 \\ \text{causal if } g(v, v) \leq 0 \\ \text{spacelike if } g(v, v) > 0 \end{cases}$

- A **spacetime** is a Lorentzian manifold endowed with a time-orientation
- The time-orientation is determined by a timelike vector field T
- A causal vector $v \in TM$ is **future-pointing** if $g(v, T) < 0$ (if $g(v, T) > 0$ is past-pointing)

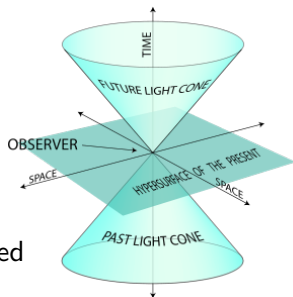


Stationary spacetimes

- A *Lorentzian manifold* (M, g) with index 1 $(+, \dots, +, -)$

- $v \in TM$ is $\begin{cases} \text{timelike if } g(v, v) < 0 \\ \text{lightlike if } g(v, v) = 0 \\ \text{causal if } g(v, v) \leq 0 \\ \text{spacelike if } g(v, v) > 0 \end{cases}$

- A **spacetime** is a Lorentzian manifold endowed with a time-orientation
- The time-orientation is determined by a timelike vector field T
- A causal vector $v \in TM$ is **future-pointing** if $g(v, T) < 0$ (if $g(v, T) > 0$ is past-pointing)

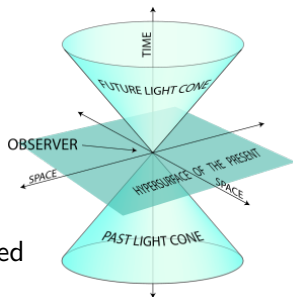


Stationary spacetimes

- A *Lorentzian manifold* (M, g) with index 1
 $(+, \dots, +, -)$

- $v \in TM$ is
$$\begin{cases} \text{timelike if } g(v, v) < 0 \\ \text{lightlike if } g(v, v) = 0 \\ \text{causal if } g(v, v) \leq 0 \\ \text{spacelike if } g(v, v) > 0 \end{cases}$$

- A **spacetime** is a Lorentzian manifold endowed with a time-orientation
- The time-orientation is determined by a timelike vector field T
- A causal vector $v \in TM$ is **future-pointing** if $g(v, T) < 0$ (if $g(v, T) > 0$ is past-pointing)

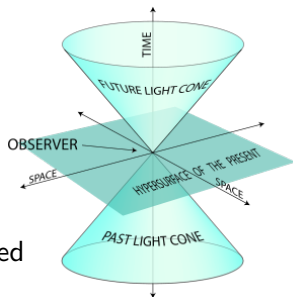


Stationary spacetimes

- A Lorentzian manifold (M, g) with index 1 $(+, \dots, +, -)$

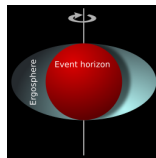
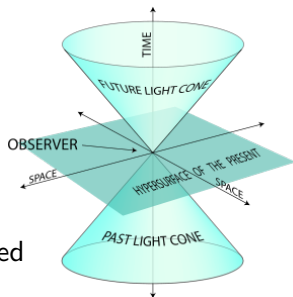
- $v \in TM$ is $\begin{cases} \text{timelike if } g(v, v) < 0 \\ \text{lightlike if } g(v, v) = 0 \\ \text{causal if } g(v, v) \leq 0 \\ \text{spacelike if } g(v, v) > 0 \end{cases}$

- A **spacetime** is a Lorentzian manifold endowed with a time-orientation
- The time-orientation is determined by a timelike vector field T
- A causal vector $v \in TM$ is **future-pointing** if $g(v, T) < 0$ (if $g(v, T) > 0$ is past-pointing)



Stationary spacetimes

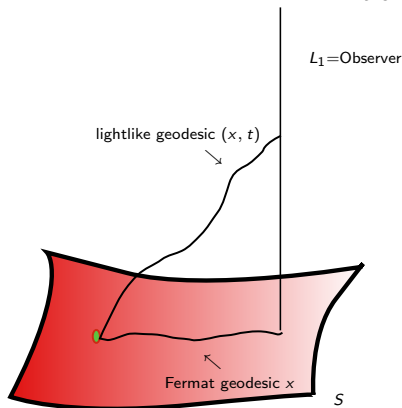
- A *Lorentzian manifold* (M, g) with index 1 $(+, \dots, +, -)$
- $v \in TM$ is $\begin{cases} \text{timelike if } g(v, v) < 0 \\ \text{lightlike if } g(v, v) = 0 \\ \text{causal if } g(v, v) \leq 0 \\ \text{spacelike if } g(v, v) > 0 \end{cases}$
- A **spacetime** is a Lorentzian manifold endowed with a time-orientation
- The time-orientation is determined by a timelike vector field T
- A causal vector $v \in TM$ is **future-pointing** if $g(v, T) < 0$ (if $g(v, T) > 0$ is past-pointing)
- A **stationary spacetime** (M, g) is a Lorentzian manifold endowed with a timelike **Killing vector field**



KERR SPACETIME

Second part: causality of stationary spacetimes via Finsler geometry

$(S \times \mathbb{R}, I)$ is a standard stationary spacetime



S is naturally endowed with a Randers metric F called the **Fermat metric**

Conformally Standard Stationary Spacetimes

- A spacetime (M, g) is **Conformastationary** if it admits a timelike Conformal field K , that is, a timelike vector field satisfying

$$\mathcal{L}_K g = \lambda g,$$

for some function $\lambda : M \rightarrow \mathbb{R}$

- **Standard Conformastationary** means that $M = S \times \mathbb{R}$ and

$$g((v, \tau), (v, \tau)) = \varphi(g_0(v, v) + 2\omega(v)\tau - \tau^2),$$

in $(x, t) \in S \times \mathbb{R}$, where $(v, \tau) \in T_x S \times \mathbb{R}$, $\varphi : S \times \mathbb{R} \rightarrow (0, +\infty)$

- and g_0 is a Riemannian metric on S and ω a 1-form on S .
- In this case, ∂_t is a **timelike conformal field**.

Conformally Standard Stationary Spacetimes

- A spacetime (M, g) is **Conformastationary** if it admits a timelike Conformal field K , that is, a timelike vector field satisfying

$$\mathcal{L}_K g = \lambda g,$$

for some function $\lambda : M \rightarrow \mathbb{R}$

- Standard Conformastationary** means that $M = S \times \mathbb{R}$ and

$$g((v, \tau), (v, \tau)) = \varphi(g_0(v, v) + 2\omega(v)\tau - \tau^2),$$

in $(x, t) \in S \times \mathbb{R}$, where $(v, \tau) \in T_x S \times \mathbb{R}$, $\varphi : S \times \mathbb{R} \rightarrow (0, +\infty)$

- and g_0 is a Riemannian metric on S and ω a 1-form on S .
- In this case, ∂_t is a **timelike conformal field**.

A conformastationary spacetime is **standard** whenever it is distinguishing and the timelike conformal vector field is complete:



M. A. J. AND M. SÁNCHEZ, *A note on the existence of standard splittings for conformally stationary spacetimes*, Classical Quantum Gravity, 25 (2008), pp. 168001, 7.



MIGUEL SÁNCHEZ

Fermat principle in General Relativity

- First established by **Herman Weyl in 1917** for static spacetimes
- The stationary case is considered by **Tulio Levi-Civita in 1927**
- It appears in classical books as **Landau-Lifshitz** “The classical theory of fields” **1962**
- **I. Kovner** gave a version of Fermat principle for an arbitrary spacetime in **1990**
- **Volker Perlick** gave a rigorous proof of this general principle in the same year (**1990**)



H. WEYL
→ (1885-1955)



T. LEVI-CIVITA
→ (1873-1941)



LEV LANDAU
→ (1908-1968)



E. LIFSHITZ
→ (1915-1985)

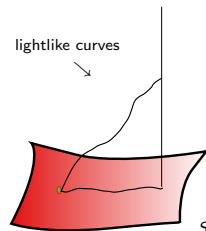


I. KOVNER

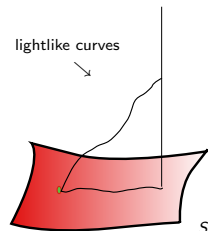


V. PERLICK
→ (BORN IN 1956)

- **Relativistic Fermat Principle:** lightlike pregeodesics are critical points of the arrival time function corresponding to an *observer* in a suitable class of lightlike curves



- **Relativistic Fermat Principle:** lightlike pregeodesics are critical points of the arrival time function corresponding to an *observer* in a suitable class of lightlike curves
- If you consider as observer $s \rightarrow L_1(s) = (x_1, s)$ in $(S \times \mathbb{R}, g)$, given a lightlike curve $\gamma = (x, t)$, the arrival time $AT(\gamma)$ is



$$t(b) = t(a) + \int_a^b \left(\omega(\dot{x}) + \sqrt{g_0(\dot{x}, \dot{x}) + \omega(\dot{x})^2} \right) ds.$$

- because $g_0(\dot{x}, \dot{x}) + 2\omega(\dot{x})\dot{t} - \dot{t}^2 = 0$ ($g(\dot{\gamma}, \dot{\gamma}) = 0$)
- Let us define the Fermat (Finslerian) metric in S as

$$F(v) = \omega(v) + \sqrt{g_0(v, v) + \omega(v)^2},$$

PIERRE DE FERMAT (1601-1665)

Theorem

A curve $s \rightarrow \gamma(s) = (x(s), s)$ is a lightlike pregeodesic of $(S \times \mathbb{R}, g)$ iff $s \rightarrow x(s)$ is a Fermat geodesic with unit speed.

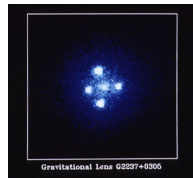
Fermat metric and lightlike geodesics

Theorem

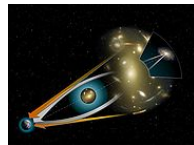
A curve $s \rightarrow \gamma(s) = (x(s), s)$ is a lightlike pregeodesic of $(S \times \mathbb{R}, g)$ iff $s \rightarrow x(s)$ is a Fermat geodesic with unit speed.

Consequences:

- **Gravitational lensing** can be studied from geodesic connectedness in Fermat metric
- Existence of **t -periodic lightlike geodesics** is equivalent to existence of Fermat closed geodesics



EINSTEIN RING



GRAVITATIONAL LENSING

My Collaborators



ERASMO CAPONIO (BARI)



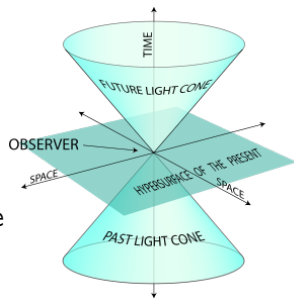
ANTONIO MASIELLO (BARI)



MIGUEL SÁNCHEZ (GRANADA)

Lorentzian Causality

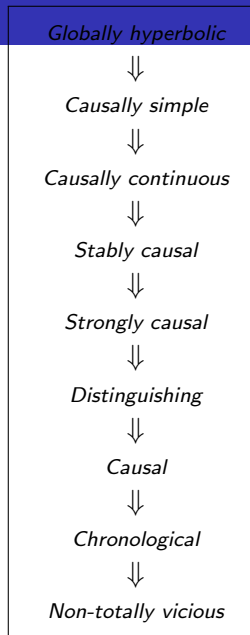
- Causality studies if given two points $p, q \in M$ they are joined by a causal curve
- $p, q \in M$ are **chronologically** related, and write $p \ll q$ if there exists a future-pointing timelike curve γ from p to q
- $p, q \in M$ are **causally** related ($p < q$) if there exists a future-pointing causal curve γ from p to q
- The **chronological future** of $p \in M$ is defined as $I^+(p) = \{q \in M : p \ll q\}$
- The **causal future** of $p \in M$ is defined as $J^+(p) = \{q \in M : p \leq q\}$
- Analogously we define the chronological past $I^-(p)$ and the causal past $J^-(p)$.



The causal ladder

Causal properties classify spacetimes depending on the behaviour of causal cones. A spacetime is:

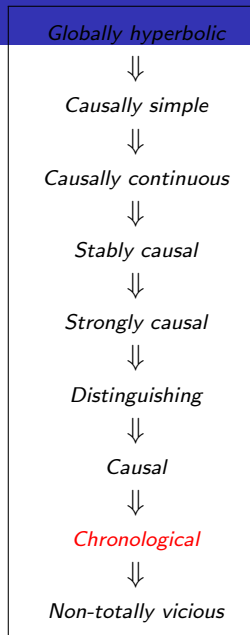
- **Chronological** if $p \notin I^+(p)$ for every $p \in M$.
- **Distinguishing** if $I^+(p) = I^+(q)$ or $I^-(p) = I^-(q)$ implies $p = q$
- **Causally continuous** if it is distinguishing and the Chronological cones $I^\pm(p)$ are continuous in $p \in M$
- **Causally simple** if the causal cones $J^\pm(p)$ are closed for every $p \in M$
- **Globally hyperbolic** if it admits a Cauchy hypersurface (a subset S that meets exactly once every inextendible timelike curve)



The causal ladder

Causal properties classify spacetimes depending on the behaviour of causal cones. A spacetime is:

- **Chronological** if $p \notin I^+(p)$ for every $p \in M$.
- **Distinguishing** if $I^+(p) = I^+(q)$ or $I^-(p) = I^-(q)$ implies $p = q$
- **Causally continuous** if it is distinguishing and the Chronological cones $I^\pm(p)$ are continuous in $p \in M$
- **Causally simple** if the causal cones $J^\pm(p)$ are closed for every $p \in M$
- **Globally hyperbolic** if it admits a Cauchy hypersurface (a subset S that meets exactly once every inextendible timelike curve)



The causal ladder

Causal properties classify spacetimes depending on the behaviour of causal cones. A spacetime is:

- **Chronological** if $p \notin I^+(p)$ for every $p \in M$.
- **Distinguishing** if $I^+(p) = I^+(q)$ or $I^-(p) = I^-(q)$ implies $p = q$
- **Causally continuous** if it is distinguishing and the Chronological cones $I^\pm(p)$ are continuous in $p \in M$
- **Causally simple** if the causal cones $J^\pm(p)$ are closed for every $p \in M$
- **Globally hyperbolic** if it admits a Cauchy hypersurface (a subset S that meets exactly once every inextendible timelike curve)

Globally hyperbolic



Causally simple



Causally continuous



Stably causal



Strongly causal



Distinguishing



Causal



Chronological

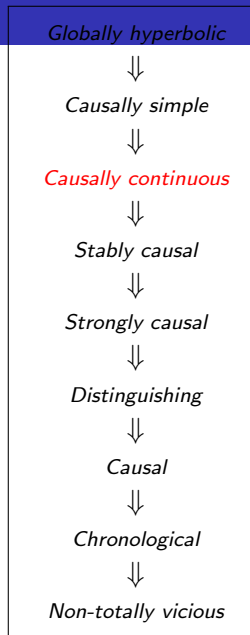


Non-totally vicious

The causal ladder

Causal properties classify spacetimes depending on the behaviour of causal cones. A spacetime is:

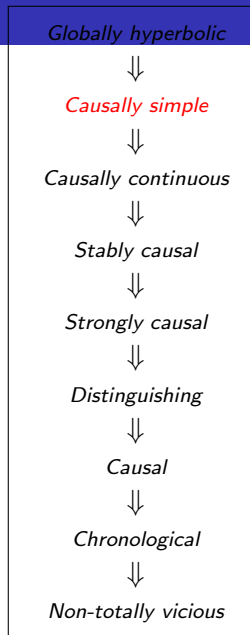
- **Chronological** if $p \notin I^+(p)$ for every $p \in M$.
- **Distinguishing** if $I^+(p) = I^+(q)$ or $I^-(p) = I^-(q)$ implies $p = q$
- **Causally continuous** if it is distinguishing and the Chronological cones $I^\pm(p)$ are continuous in $p \in M$
- **Causally simple** if the causal cones $J^\pm(p)$ are closed for every $p \in M$
- **Globally hyperbolic** if it admits a Cauchy hypersurface (a subset S that meets exactly once every inextendible timelike curve)



The causal ladder

Causal properties classify spacetimes depending on the behaviour of causal cones. A spacetime is:

- **Chronological** if $p \notin I^+(p)$ for every $p \in M$.
- **Distinguishing** if $I^+(p) = I^+(q)$ or $I^-(p) = I^-(q)$ implies $p = q$
- **Causally continuous** if it is distinguishing and the Chronological cones $I^\pm(p)$ are continuous in $p \in M$
- **Causally simple** if the causal cones $J^\pm(p)$ are closed for every $p \in M$
- **Globally hyperbolic** if it admits a Cauchy hypersurface (a subset S that meets exactly once every inextendible timelike curve)



The causal ladder

Causal properties classify spacetimes depending on the behaviour of causal cones. A spacetime is:

- **Chronological** if $p \notin I^+(p)$ for every $p \in M$.
- **Distinguishing** if $I^+(p) = I^+(q)$ or $I^-(p) = I^-(q)$ implies $p = q$
- **Causally continuous** if it is distinguishing and the Chronological cones $I^\pm(p)$ are continuous in $p \in M$
- **Causally simple** if the causal cones $J^\pm(p)$ are closed for every $p \in M$
- **Globally hyperbolic** if it admits a Cauchy hypersurface (a subset S that meets exactly once every inextendible timelike curve)

Globally hyperbolic



Causally simple



Causally continuous



Stably causal



Strongly causal



Distinguishing



Causal



Chronological



Non-totally vicious

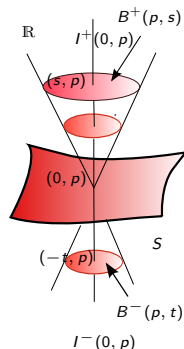
Causality via the Fermat metric

- Let d the non-symmetric distance in S associated to the Fermat metric
- $B^+(x_0, s) = \{p \in S : d(x_0, p) < s\}$ forward balls
- $B^-(x_0, s) = \{p \in S : d(p, x_0) < s\}$ backward balls

Causality via the Fermat metric

- Let d the non-symmetric distance in S associated to the Fermat metric
- $B^+(x_0, s) = \{p \in S : d(x_0, p) < s\}$ forward balls
- $B^-(x_0, s) = \{p \in S : d(p, x_0) < s\}$ backward balls
- Let $(\mathbb{R} \times S, g)$ be a standard stationary spacetime. Then

$$I^\pm(t_0, x_0) = \cup_{s>0} \{t_0 \pm s\} \times B^\pm(x_0, s),$$



Causality through the Fermat metric

Theorem

Let $(S \times \mathbb{R}, g)$ be a standard stationary spacetime.
Then $(S \times \mathbb{R}, g)$ is **causally continuous** and

- (a) the following assertions become equivalent:
 - (i) $(S \times \mathbb{R}, g)$ is **causally simple**,
 - (ii) the associated Finsler manifold (S, F) is convex,
- (b) it is **globally hyperbolic** if and only if $\bar{B}^+(x, r) \cap \bar{B}^-(x, r)$ is compact for every $x \in S$ and $r > 0$.
- (c) a slice $S \times \{t_0\}$, $t_0 \in \mathbb{R}$, is a **Cauchy hypersurface** if and only if the Fermat metric F on S is forward and backward complete.

Globally hyperbolic



Causally simple



Causally continuous



Stably causal



Strongly causal



Distinguishing



Causal



Chronological



Non-totally vicious

Causality through the Fermat metric

Theorem

Let $(S \times \mathbb{R}, g)$ be a standard stationary spacetime.
Then $(S \times \mathbb{R}, g)$ is **causally continuous** and

- (a) the following assertions become equivalent:
 - (i) $(S \times \mathbb{R}, g)$ is **causally simple**,
 - (ii) the associated Finsler manifold (S, F) is convex,
- (b) it is **globally hyperbolic** if and only if $\bar{B}^+(x, r) \cap \bar{B}^-(x, r)$ is compact for every $x \in S$ and $r > 0$.
- (c) a slice $S \times \{t_0\}$, $t_0 \in \mathbb{R}$, is a **Cauchy hypersurface** if and only if the Fermat metric F on S is forward and backward complete.

Globally hyperbolic



Causally simple



Causally continuous



Stably causal



Strongly causal



Distinguishing



Causal



Chronological



Non-totally vicious

Causality through the Fermat metric

Theorem

Let $(S \times \mathbb{R}, g)$ be a standard stationary spacetime.
Then $(S \times \mathbb{R}, g)$ is **causally continuous** and

- (a) the following assertions become equivalent:
 - (i) $(S \times \mathbb{R}, g)$ is **causally simple**,
 - (ii) the associated Finsler manifold (S, F) is convex,
- (b) it is **globally hyperbolic** if and only if $\bar{B}^+(x, r) \cap \bar{B}^-(x, r)$ is compact for every $x \in S$ and $r > 0$.
- (c) a slice $S \times \{t_0\}$, $t_0 \in \mathbb{R}$, is a **Cauchy hypersurface** if and only if the Fermat metric F on S is forward and backward complete.

Globally hyperbolic



Causally simple



Causally continuous



Stably causal



Strongly causal



Distinguishing



Causal



Chronological



Non-totally vicious

Causality through the Fermat metric

Theorem

Let $(S \times \mathbb{R}, g)$ be a standard stationary spacetime.
Then $(S \times \mathbb{R}, g)$ is **causally continuous** and

- (a) the following assertions become equivalent:
 - (i) $(S \times \mathbb{R}, g)$ is **causally simple**,
 - (ii) the associated Finsler manifold (S, F) is convex,
- (b) it is **globally hyperbolic** if and only if $\bar{B}^+(x, r) \cap \bar{B}^-(x, r)$ is compact for every $x \in S$ and $r > 0$.
- (c) a slice $S \times \{t_0\}$, $t_0 \in \mathbb{R}$, is a **Cauchy hypersurface** if and only if the Fermat metric F on S is forward and backward complete.

Globally hyperbolic



Causally simple



Causally continuous



Stably causal



Strongly causal



Distinguishing



Causal



Chronological



Non-totally vicious

Generalized Hopf-Rinow theorem

Theorem (Accurate Hopf-Rinow for Randers metrics)

Let (S, R) a Randers manifold and given a function $f : S \rightarrow \mathbb{R}$ such that $df \leq R$ define



$R_f(x, v) = R(x, v) - df_x(v)$. The following conditions are equivalent:




- (A) the intersection $\bar{B}^+(x, r) \cap \bar{B}^-(x, r)$ of (S, R) is **compact** for every $r > 0$ and $x \in S$
- (B) there exists f such that R_f is **geodesically complete**
- (C) there exists f and $p \in S$ such that the forward and the backward exponentials of R_f are defined in $T_p S$
- (E) there exists f such that **the quasi-metric d_f** associated to R_f is **forward and backward complete**

In such a case, (S, R) is convex.



HEINZ HOPF (1894-1971)

-  E. CAPONIO, M. A. J., AND A. MASIELLO, *On the energy functional on Finsler manifolds and applications to stationary spacetimes*, Math. Ann., 351 (2011), pp. 365–392.
-  E. CAPONIO, M. A. J., AND M. SÁNCHEZ, *On the interplay between Lorentzian Causality and Finsler metrics of Randers type*, Rev. Mat. Iberoamericana, 27 (2011), pp. 919–952.

-  E. CAPONIO, M. A. J., AND A. MASIELLO, *On the energy functional on Finsler manifolds and applications to stationary spacetimes*, Math. Ann., 351 (2011), pp. 365–392.
-  E. CAPONIO, M. A. J., AND M. SÁNCHEZ, *On the interplay between Lorentzian Causality and Finsler metrics of Randers type*, Rev. Mat. Iberoamericana, 27 (2011), pp. 919–952.
- For a review see:
-  M. A. J., *Conformally standard stationary spacetimes and Fermat metrics*, Proceedings of GeLoGra 2011.

Last part: Almost Isometries. My collaborators



LEANDRO LICHTENFELZ

NOTRE DAME UNIVERSITY (USA)



PAOLO PICCIONE

UNIVERSIDADE DE SAO PAULO (BRASIL)

Almost isometries of quasi-metrics

Definition

Given a set X , we say that a function $d : X \times X \rightarrow \mathbb{R}$ is a *quasi-metric* if

- (i) $d(x, y) \geq 0$ for every $x, y \in X$ and $d(x, y) = 0$ if and only if $x = y$,
- (ii) $d(x, y) + d(y, z) \geq d(x, z)$ (triangle inequality).

Almost isometries of quasi-metrics

Definition

Given a set X , we say that a function $d : X \times X \rightarrow \mathbb{R}$ is a *quasi-metric* if

- (i) $d(x, y) \geq 0$ for every $x, y \in X$ and $d(x, y) = 0$ if and only if $x = y$,
- (ii) $d(x, y) + d(y, z) \geq d(x, z)$ (triangle inequality).

As a consequence of the lack of symmetry, there are two kinds of balls:

- $B_d^+(x, r) = \{y \in X : d(x, y) < r\}$ (forward balls)
- $B_d^-(x, r) = \{y \in X : d(y, x) < r\}$ (backward balls)

respectively, for $x \in X$ and $r > 0$.

Almost isometries of quasi-metrics

Definition

Given a set X , we say that a function $d : X \times X \rightarrow \mathbb{R}$ is a *quasi-metric* if

- (i) $d(x, y) \geq 0$ for every $x, y \in X$ and $d(x, y) = 0$ if and only if $x = y$,
- (ii) $d(x, y) + d(y, z) \geq d(x, z)$ (triangle inequality).

As a consequence of the lack of symmetry, there are two kinds of balls:

- $B_d^+(x, r) = \{y \in X : d(x, y) < r\}$ (forward balls)
- $B_d^-(x, r) = \{y \in X : d(y, x) < r\}$ (backward balls)

respectively, for $x \in X$ and $r > 0$.

Definition

A pair (X, d) will be called a *quasi-metric space* endowed with the topology induced by the family $B_d^+(x, r) \cap B_d^-(x, r)$, $x \in M$ and $r > 0$.

Let us observe that this topology coincides with the topology generated by (the balls of) the *symmetrized metric* $\tilde{d}(x, y) = \frac{1}{2}(d(x, y) + d(y, x))$.

Quasi-metrics spaces have been studied by many mathematicians:

- Fréchet 1909, Hausdorff 1914, Mazurkiewicz 1930, Wilson 1931, Busemann 1944
- and also by a spanish mathematician: **Julio Rey Pastor** 1940



REY PASTOR (1888-1962)

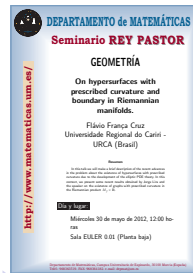
Quasi-metrics spaces have been studied by many mathematicians:

- Fréchet 1909, Hausdorff 1914, Mazurkiewicz 1930, Wilson 1931, Busemann 1944
- and also by a spanish mathematician: **Julio Rey Pastor** 1940

Our seminar in the university of Murcia is called “Rey Pastor” after him



REY PASTOR (1888-1962)



Quasi-metrics and the triangular function

In a quasi-metric space we can define the **length of a continuous curve** $\alpha : [a, b] \subseteq \mathbb{R} \rightarrow X$ as

$$\ell(\alpha) = \sup_{\mathcal{P}} \sum_{i=1}^r d(\alpha(s_i), \alpha(s_{i+1})),$$

where \mathcal{P} is the set of partitions $a = s_1 < s_2 < \dots < s_{r+1} = b$, $r \in \mathbb{N}$.

Quasi-metrics and the triangular function

In a quasi-metric space we can define the **length of a continuous curve** $\alpha : [a, b] \subseteq \mathbb{R} \rightarrow X$ as

$$\ell(\alpha) = \sup_{\mathcal{P}} \sum_{i=1}^r d(\alpha(s_i), \alpha(s_{i+1})),$$

where \mathcal{P} is the set of partitions $a = s_1 < s_2 < \dots < s_{r+1} = b$, $r \in \mathbb{N}$.

- We say that α is **rectifiable** when $\ell(\alpha)$ is finite.
- Moreover, we say that a curve γ in X from p to q is a **minimizing geodesic** if $\ell(\gamma) = d(p, q)$.

Quasi-metrics and the triangular function

In a quasi-metric space we can define the **length of a continuous curve** $\alpha : [a, b] \subseteq \mathbb{R} \rightarrow X$ as

$$\ell(\alpha) = \sup_{\mathcal{P}} \sum_{i=1}^r d(\alpha(s_i), \alpha(s_{i+1})),$$

where \mathcal{P} is the set of partitions $a = s_1 < s_2 < \dots < s_{r+1} = b$, $r \in \mathbb{N}$.

- We say that α is **rectifiable** when $\ell(\alpha)$ is finite.
- Moreover, we say that a curve γ in X from p to q is a **minimizing geodesic** if $\ell(\gamma) = d(p, q)$.

Definition

Let us define the **triangular function** $T : X \times X \times X \rightarrow [0, +\infty[$ of a quasi-metric space (X, d) as $T(x, y, z) = d(x, y) + d(y, z) - d(x, z)$ for every $x, y, z \in X$.

Evidently, T is continuous.

Proposition

A curve $\alpha : [a, b] \subseteq \mathbb{R} \rightarrow X$ is a minimizing geodesic of a quasi-metric space (X, d) iff $T(\alpha(s_1), \alpha(s_2), \alpha(s_3)) = 0$ for every $a \leq s_1 < s_2 < s_3 \leq b$.

Almost isometries

Proposition

A curve $\alpha : [a, b] \subseteq \mathbb{R} \rightarrow X$ is a minimizing geodesic of a quasi-metric space (X, d) iff $T(\alpha(s_1), \alpha(s_2), \alpha(s_3)) = 0$ for every $a \leq s_1 < s_2 < s_3 \leq b$.

Definition

Let (X_1, d_1) and (X_2, d_2) be two quasi-metric spaces. A bijection $\varphi : X_1 \rightarrow X_2$ is an *almost isometry* if it preserves the triangular function, that is,

$$T_2(\varphi(x), \varphi(y), \varphi(z)) = T_1(x, y, z)$$

for every $x, y, z \in X_1$, where T_1 and T_2 are the triangular functions associated respectively to (X_1, d_1) and (X_2, d_2) .

Almost isometries

Proposition

A curve $\alpha : [a, b] \subseteq \mathbb{R} \rightarrow X$ is a minimizing geodesic of a quasi-metric space (X, d) iff $T(\alpha(s_1), \alpha(s_2), \alpha(s_3)) = 0$ for every $a \leq s_1 < s_2 < s_3 \leq b$.

Definition

Let (X_1, d_1) and (X_2, d_2) be two quasi-metric spaces. A bijection $\varphi : X_1 \rightarrow X_2$ is an *almost isometry* if it preserves the triangular function, that is,

$$T_2(\varphi(x), \varphi(y), \varphi(z)) = T_1(x, y, z)$$

for every $x, y, z \in X_1$, where T_1 and T_2 are the triangular functions associated respectively to (X_1, d_1) and (X_2, d_2) .

Corollary

Almost isometries preserve minimizing geodesics.

Proposition

Given quasi-metric spaces (X_1, d_1) and (X_2, d_2) , a bijection $\varphi : X_1 \rightarrow X_2$ is an almost isometry iff $\exists f : X_2 \rightarrow \mathbb{R}$ such that for every $x, y \in X_1$:

$$d_2(\varphi(x), \varphi(y)) = d_1(x, y) + f(\varphi(x)) - f(\varphi(y)) \quad (1)$$

Almost isometries

Proposition

Given quasi-metric spaces (X_1, d_1) and (X_2, d_2) , a bijection $\varphi : X_1 \rightarrow X_2$ is an almost isometry iff $\exists f : X_2 \rightarrow \mathbb{R}$ such that for every $x, y \in X_1$:

$$d_2(\varphi(x), \varphi(y)) = d_1(x, y) + f(\varphi(x)) - f(\varphi(y)) \quad (1)$$

Proof.

\Rightarrow (the converse is straightforward)

- Fix a point $x_0 \in X_1$ and define $f : X_2 \rightarrow \mathbb{R}$ as $f(z) = d_2(z, \varphi(x_0)) - d_1(\varphi^{-1}(z), x_0)$ for every $z \in X_2$.

Almost isometries

Proposition

Given quasi-metric spaces (X_1, d_1) and (X_2, d_2) , a bijection $\varphi : X_1 \rightarrow X_2$ is an almost isometry iff $\exists f : X_2 \rightarrow \mathbb{R}$ such that for every $x, y \in X_1$:

$$d_2(\varphi(x), \varphi(y)) = d_1(x, y) + f(\varphi(x)) - f(\varphi(y)) \quad (1)$$

Proof.

\Rightarrow (the converse is straightforward)

- Fix a point $x_0 \in X_1$ and define $f : X_2 \rightarrow \mathbb{R}$ as $f(z) = d_2(z, \varphi(x_0)) - d_1(\varphi^{-1}(z), x_0)$ for every $z \in X_2$.
- Given $x, y \in X_1$, as φ preserves the triangular function, we have

$$\begin{aligned} d_1(x, y) + d_1(y, x_0) - d_1(x, x_0) \\ = d_2(\varphi(x), \varphi(y)) + d_2(\varphi(y), \varphi(x_0)) - d_2(\varphi(x), \varphi(x_0)), \end{aligned}$$

which is equivalent to (1).

Almost isometries

Some observations:

- In metric spaces, almost isometries are always isometries
- If $\varphi : (X_1, d_1) \rightarrow (X_2, d_2)$ is an almost isometry, then

$$\varphi : (X_1, \tilde{d}_1) \rightarrow (X_2, \tilde{d}_2)$$

is an isometry, where

$$\tilde{d}_1(x, y) = \frac{1}{2}(d_1(x, y) + d_1(y, x)),$$

$$\tilde{d}_2(x, y) = \frac{1}{2}(d_2(x, y) + d_2(y, x)).$$

- Moreover, φ is a homeomorphism and the functions $f : X_2 \rightarrow \mathbb{R}$ are continuous

Almost isometries

Notation:

- $\text{Iso}(X, d)$ is the **group of isometries** of (X, d)
- $\widetilde{\text{Iso}}(X, d)$ is the group of almost isometries of (X, d) . It will be called the *extended isometry group* of (X, d) .

Almost isometries

Notation:

- $\text{Iso}(X, d)$ is the **group of isometries** of (X, d)
- $\widetilde{\text{Iso}}(X, d)$ is the group of almost isometries of (X, d) . It will be called the **extended isometry group** of (X, d) .

Proposition

- With the above notation, $\widetilde{\text{Iso}}(X, d)$ and $\text{Iso}(X, d)$ are **topological groups** endowed with the compact-open topology.
- If the topology induced by d is locally compact, then $\widetilde{\text{Iso}}(X, d)$ and $\text{Iso}(X, d)$ are **locally compact**.

Almost isometries

Notation:

- $\text{Iso}(X, d)$ is the **group of isometries** of (X, d)
- $\widetilde{\text{Iso}}(X, d)$ is the group of almost isometries of (X, d) . It will be called the **extended isometry group** of (X, d) .

Proposition

- With the above notation, $\widetilde{\text{Iso}}(X, d)$ and $\text{Iso}(X, d)$ are **topological groups** endowed with the compact-open topology.
- If the topology induced by d is locally compact, then $\widetilde{\text{Iso}}(X, d)$ and $\text{Iso}(X, d)$ are **locally compact**.

Proof.

The proof follows from the inclusions:

$$\text{Iso}(X, d) \subseteq \widetilde{\text{Iso}}(X, d) \subseteq \text{Iso}(X, \tilde{d}).$$

Local almost isometries

Definition

Let (X_1, d_1) and (X_2, d_2) be two quasi-metric spaces. A map $\varphi : X_1 \rightarrow X_2$ is a *local almost isometry* if $\forall x \in X_1, \exists U \subseteq X_1, V \subseteq X_2$ open subsets, with $x \in U$, such that $\varphi|_U : U \rightarrow V$ is an almost isometry.

Definition

Let (X_1, d_1) and (X_2, d_2) be two quasi-metric spaces. A map $\varphi : X_1 \rightarrow X_2$ is a *local almost isometry* if $\forall x \in X_1, \exists U \subseteq X_1, V \subseteq X_2$ open subsets, with $x \in U$, such that $\varphi|_U : U \rightarrow V$ is an almost isometry.

- define d_I as the infimum of the lengths of curves between two points. We say that (X, d) is a *length space* when $d_I = d$.
- We say that a quasi-metric space is *weakly finitely compact* if $B^+(x, r) \cap B^-(x, r)$ are precompact $\forall x \in X$ and $r > 0$.

Local almost isometries

Definition

Let (X_1, d_1) and (X_2, d_2) be two quasi-metric spaces. A map $\varphi : X_1 \rightarrow X_2$ is a **local almost isometry** if $\forall x \in X_1, \exists U \subseteq X_1, V \subseteq X_2$ open subsets, with $x \in U$, such that $\varphi|_U : U \rightarrow V$ is an almost isometry.

- define d_l as the infimum of the lengths of curves between two points. We say that (X, d) is a **length space** when $d_l = d$.
- We say that a quasi-metric space is **weakly finitely compact** if $B^+(x, r) \cap B^-(x, r)$ are precompact $\forall x \in X$ and $r > 0$.

Theorem

Let $\varphi : (X_1, d_1) \rightarrow (X_2, d_2)$ be a local almost isometry. Assume that (X_1, d_1) and (X_2, d_2) are length spaces, d_1 is weakly finitely compact and X_2 is locally arc-connected and simply connected. Then φ is an almost isometry.

Almost isometries of Finsler metrics

Let us define the symmetrized Finsler metric of F as

$$\hat{F}(v) = \frac{1}{2}[F(v) + F(-v)]$$

for every $v \in TM$. The sum of Finsler metrics is a Finsler metric:



M. A. J. AND M. SÁNCHEZ, *On the definition and examples of Finsler metrics*, to appear in Ann. Sc. Nor. Pisa

Almost isometries of Finsler metrics

Let us define the symmetrized Finsler metric of F as

$$\hat{F}(v) = \frac{1}{2}[F(v) + F(-v)]$$

for every $v \in TM$. The sum of Finsler metrics is a Finsler metric:



M. A. J. AND M. SÁNCHEZ, *On the definition and examples of Finsler metrics*, to appear in Ann. Sc. Nor. Pisa

Lemma

If $\varphi : (M_1, F_1) \rightarrow (M_2, F_2)$ is an almost isometry then

$$\varphi : (M_1, \hat{F}_1) \rightarrow (M_2, \hat{F}_2)$$

is an isometry and φ is smooth.

Proof.

- To see that φ is an isometry prove that preserves the length of curves
- φ is smooth because it is an isometry of a Riemannian average metric

Proposition

- *If \exists an almost isometry $\varphi : (M_1, F_1) \rightarrow (M_2, F_2)$, then there exists a smooth $f : M_2 \rightarrow \mathbb{R}$ such that $\varphi^*(F_1) = F_2 + df$.*
- *Conversely, if $\varphi^*(F_1) = F_2 + df$, the map φ is an almost isometry.*

Proposition

- *If \exists an almost isometry $\varphi : (M_1, F_1) \rightarrow (M_2, F_2)$, then there exists a smooth $f : M_2 \rightarrow \mathbb{R}$ such that $\varphi^*(F_1) = F_2 + df$.*
- *Conversely, if $\varphi^*(F_1) = F_2 + df$, the map φ is an almost isometry.*

Proposition

Let (M, F) be a Finsler manifold. Then the extended isometry group $\widetilde{\text{Iso}}(M, F)$ is a closed subgroup of $\text{Iso}(M, \hat{F})$. In particular, $\widetilde{\text{Iso}}(M, F)$ is a Lie group.

Proposition

- *If \exists an almost isometry $\varphi : (M_1, F_1) \rightarrow (M_2, F_2)$, then there exists a smooth $f : M_2 \rightarrow \mathbb{R}$ such that $\varphi^*(F_1) = F_2 + df$.*
- *Conversely, if $\varphi^*(F_1) = F_2 + df$, the map φ is an almost isometry.*

Proposition

Let (M, F) be a Finsler manifold. Then the extended isometry group $\widetilde{\text{Iso}}(M, F)$ is a closed subgroup of $\text{Iso}(M, \hat{F})$. In particular, $\widetilde{\text{Iso}}(M, F)$ is a Lie group.

Proof.

Use that $\widetilde{\text{Iso}}(M, F) \subset \text{Iso}(M, \hat{F})$



Corollary

Let (M, R) be a Randers manifold and $\varphi : M \rightarrow M$ an almost isometry for R . Then φ is an isometry for h .

Corollary

Let (M, R) be a Randers manifold and $\varphi : M \rightarrow M$ an almost isometry for R . Then φ is an isometry for h .

Proof.

Just observe that the symmetrized Finsler metric of R is given by $\hat{R}(v) = \sqrt{h(v, v)}$ for $v \in TM$. □

K -conformal maps

- Let (M, g) be a spacetime and K a Killing field
- We say that a diffeomorphism $\psi : M \rightarrow M$ is **K -conformal** if
 - It is conformal, $\psi_*(g) = \lambda g$, $\lambda \neq 0$, (ψ_* is the pushforward) and
 - preserves K , $\psi_*(K) = K$
- Now consider a normalized standard stationary spacetime $(S \times \mathbb{R}, g)$ with

$$g((v, \tau), (v, \tau)) = g_0(v, v) + 2\omega(v)\tau - \tau^2$$

$v \in TS$ and $\tau \in \mathbb{R}$.

K -conformal maps

- Let (M, g) be a spacetime and K a Killing field
- We say that a diffeomorphism $\psi : M \rightarrow M$ is **K -conformal** if
 - It is conformal, $\psi_*(g) = \lambda g$, $\lambda \neq 0$, (ψ_* is the pushforward) and
 - preserves K , $\psi_*(K) = K$
- Now consider a normalized standard stationary spacetime $(S \times \mathbb{R}, g)$ with

$$g((v, \tau), (v, \tau)) = g_0(v, v) + 2\omega(v)\tau - \tau^2$$

$v \in TS$ and $\tau \in \mathbb{R}$.

Theorem

If $\psi : (S \times \mathbb{R}, g) \rightarrow (S \times \mathbb{R}, g)$ is a K -conformal map, then

$$\psi(x, t) = (\varphi(x), t + f(x)),$$

and $\varphi_*(F) = F + df$ and $\varphi : (S, h) \rightarrow (S, h)$ is an isometry, where

$$h(v, v) = g_0(v, v) + \omega(v)^2.$$

K -conformal maps

Theorem

If $\psi : (S \times \mathbb{R}, g) \rightarrow (S \times \mathbb{R}, g)$ is a K -conformal map, then $\psi(x, t) = (\varphi(x), t + f(x))$ and $\varphi_*(F) = F + df$ and $\varphi : (S, h) \rightarrow (S, h)$ is an isometry, where

$$h(v, v) = g_0(v, v) + \omega(v)^2.$$

Proof.



K -conformal maps

Theorem

If $\psi : (S \times \mathbb{R}, g) \rightarrow (S \times \mathbb{R}, g)$ is a K -conformal map, then $\psi(x, t) = (\varphi(x), t + f(x))$ and $\varphi_*(F) = F + df$ and $\varphi : (S, h) \rightarrow (S, h)$ is an isometry, where

$$h(v, v) = g_0(v, v) + \omega(v)^2.$$

Proof.

- K -conformal implies that ψ maps orbits of ∂_t to orbits of ∂_t , that is, $\psi(x, t) = (\varphi(x), t + f(x))$
- As ψ is conformal, maps lightlike pregeodesics to lightlike pregeodesics
- Then Fermat metric maps Fermat pregeodesics to Fermat pregeodesics and $\ell_{\varphi_*(F)}(\gamma) = \ell_F(\gamma) + f(\gamma(1)) - f(\gamma(0))$
- This means that $\varphi_*(F)$ and $F + df$ have the same geodesics and therefore they are equal



K -conformal maps

Lemma

$\text{Conf}_K(M, g)$ (here $M = \mathbb{R} \times S$) is a closed subgroup of $\text{Conf}(M, g)$. Moreover the one-parameter subgroup \mathcal{K} generated by K is closed and normal in $\text{Conf}_K(M, g)$.

Proof.



K -conformal maps

Lemma

$\text{Conf}_K(M, g)$ (here $M = \mathbb{R} \times S$) is a closed subgroup of $\text{Conf}(M, g)$. Moreover the one-parameter subgroup \mathcal{K} generated by K is closed and normal in $\text{Conf}_K(M, g)$.

Proof.

- First part is obvious in the C^1 topology.
- If $\psi \in \text{Conf}_K(M, g)$ then $\psi(x, t) = (\varphi(x), t + f(x))$ with $\varphi \in \widetilde{\text{Iso}}(S, F)$
- Moreover, $\psi^{-1}(x, t) = (\varphi^{-1}(x), t - f(\varphi^{-1}(x)))$
- Then if $K^T : M \rightarrow M$ is given by $K^T(x, t) = (x, t + T)$, it follows that $\psi \circ K^T \circ \psi^{-1} = K^T$ (\mathcal{K} is normal)



K -conformal maps

Proposition

The map $\pi : \text{Conf}_K(M, g) \rightarrow \widetilde{\text{Iso}}(S, F)$ defined as $\pi(\psi) = \varphi$ is a *Lie group homomorphism* and $\bar{\pi} : \text{Conf}_K(M, g)/\mathcal{K} \rightarrow \widetilde{\text{Iso}}(S, F)$ is an *isomorphism*.

Proof.

K -conformal maps

Proposition

The map $\pi : \text{Conf}_K(M, g) \rightarrow \widetilde{\text{Iso}}(S, F)$ defined as $\pi(\psi) = \varphi$ is a **Lie group homomorphism** and $\bar{\pi} : \text{Conf}_K(M, g)/\mathcal{K} \rightarrow \widetilde{\text{Iso}}(S, F)$ is an **isomorphism**.

Proof.

- We just have to prove that $\bar{\pi}$ is one-to-one.
- **Injective:** if ψ_1 and ψ_2 project on the same almost isometry map φ , then by last Prop. $\psi_1(x, t) = (\varphi(x), t + f(x) + c_1)$ and $\psi_2(x, t) = (\varphi(x), t + f(x) + c_2)$, $\psi_2 \circ \psi_1^{-1} = K^{c_2 - c_1}$ and $[\psi_1] = [\psi_2]$
- **Surjective:** given an almost isometry φ , we construct the map

$$\psi(x, t) = (\varphi(x), t + f(x))$$

Clearly, it preserves ∂_t . By Fermat principle, it maps lightlike pregeodesics to lightlike pregeodesics, then it preserves the lightcone and it must be conformal (by Dajcker-Nomizu [83]).

Corollary

Given a manifold S , for a generic set of data (g_0, ω) , the stationary metric $g = g(g_0, \omega)$ on $M = S \times \mathbb{R}$ has discrete K -conformal group $\text{Conf}_K(M, g)/\mathcal{K}$.

Corollary





Given a manifold S , for a generic set of data (g_0, ω) , the stationary metric $g = g(g_0, \omega)$ on $M = S \times \mathbb{R}$ has discrete K -conformal group $\text{Conf}_K(M, g)/\mathcal{K}$.

Corollary

If S is compact, then $\text{Conf}_K(S \times \mathbb{R}, g)/\mathcal{K}$ and $\widetilde{\text{Iso}}(S, F)$ are compact Lie groups.


Other references

About behavior of geodesics:




-  L. BILIOTTI AND M. Á. JAVALOYES. t -periodic light rays in conformally stationary spacetimes via Finsler geometry. *Houston J. Math.*, 37(1):127–146, 2011.
-  E. CAPONIO, M. Á. JAVALOYES, AND A. MASIELLO. Finsler geodesics in the presence of a convex function and their applications. *J. Phys. A*, 43(13):135207, 15, 2010.
-  E. CAPONIO, M. Á. JAVALOYES, AND A. MASIELLO. Morse theory of causal geodesics in a stationary spacetime via Morse theory of geodesics of a Finsler metric. *Ann. Inst. H. Poincaré Anal. Non Linéaire*, 27(3):857–876, 2010.
-  E. CAPONIO, A. GERMINARIO, AND M. SÁNCHEZ. Geodesics on convex regions of stationary spacetimes and Finslerian Randers spaces. *arXiv:1112.3892v1 [math.DG]*, 2011.

Other references

More about causality:

-  J. L. FLORES, J. HERRERA, AND M. SÁNCHEZ. Gromov, Cauchy and causal boundaries for Riemannian, Finslerian and Lorentzian manifolds. *Mem. Amer. Math. Soc.*

Other aspects:

-  G. W. GIBBONS, C. A. R. HERDEIRO, C. M. WARNICK, AND M. C. WERNER, Stationary metrics and optical Zermelo-Randers-Finsler geometry, *Phys. Rev. D*, 79 (2009), pp. 044022, 21.
-  A. DIRMEIER, M. PLAUE, AND M. SCHERFNER. Growth conditions, Riemannian completeness and Lorentzian causality. *J. Geom. Phys.*, 62(3):604–612, 2011.
-  M. C. WERNER, Gravitational lensing in the Kerr-Randers optical geometry, *Gen. Relativity Gravitation* 44 (2012), 3047–3057.

Open problems and work in progress

- Compute explicitly some extended isometry group
- to study the case in that K is not timelike



Thanks a lot for your attention