

Quantum Nature of D-branes

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1. Introduction

One of important directions in string theory is to reveal quantum nature of the gravity.

In this talk, we focus on **D0-branes** in type IIA superstring theory.

We take into account

- Higher derivative corrections
- Hawking radiation
- dual Matrix model analysis

to reveal quantum nature of D0-branes.

Plan of talk

1. Introduction
2. Higher Derivative Corrections in M-theory
3. Smeared black 0-brane
4. Hawking Radiation of a D0-brane from Smeared black 0-brane
5. Matrix Model Analysis of Smeared Black 0-brane
6. Summary

2. Higher Derivative Corrections in M-theory

$$\delta_{\text{SUSY}} S_{\text{SUGRA}} = 0$$



$$\delta_{\text{SUSY}} (S_{\text{SUGRA}} + S_{\text{Quantum}}) = 0$$



want to know

First we review the construction of **11 dimensional supergravity**.

The massless fields consists of vielbein $e^a{}_\mu$, Majorana gravitino ψ_μ , and 3-form field $A_{\mu\nu\rho}$.

The building blocks of the Lagrangian are

$$R^{ab}{}_{\mu\nu} : [L]^{-2}, \quad F_{\mu\nu\rho\sigma} : [L]^{-1}, \quad D_\mu : [L]^{-1}, \quad \bar{\psi}_\mu \gamma^{\rho_1 \dots \rho_n} \psi_\nu = [L]^{-1}, \quad \psi_{\mu\nu} = 2D_{[\mu} \psi_{\nu]} : [L]^{-\frac{3}{2}}$$

Then generic form of the Lagrangian is given by

$$\mathcal{L} = [eR] + [eF^2] + [e\epsilon_{11}AF^2]_1 + [e\psi\psi]_2 + [eF\psi\psi]_4 + \mathcal{O}(\psi^4)$$

Here we used following abbreviations.

$$[eF\psi\psi]_4 = c_1 e F^{\mu\nu\rho\sigma} \bar{\psi}_\mu \gamma_{\rho\sigma} \psi_\nu + c_2 e F_{\alpha\beta\gamma\delta} \bar{\psi}_\mu \gamma^{\mu\nu\alpha\beta\gamma\delta} \psi_\nu + c_3 e F_{\alpha\beta\gamma\delta} \bar{\psi}_\mu \gamma^{\alpha\beta\gamma\delta} \psi^\mu + c_4 e F_{\alpha\beta\gamma}{}^\mu \bar{\psi}_{(\mu} \gamma^{\nu\alpha\beta\gamma} \psi_{\nu)}$$

$$[e\epsilon_{11}AF^2]_1 = c_5 e \epsilon_{11}^{\mu_1 \dots \mu_{11}} A_{\mu_1 \mu_2 \mu_3} F_{\mu_4 \dots \mu_7} F_{\mu_8 \dots \mu_{11}}$$

Generic form of the local supersymmetry transformation is given by

$$\begin{aligned}\delta e^\mu_a &= -\bar{\epsilon}\gamma^\mu\psi_a, \\ \delta\psi_\mu &= d_1 D_\mu\epsilon + d_2 F_{\mu jkl}\gamma^{jkl}\epsilon + d_3 F_{ijkl}\gamma^{ijkl}\epsilon + \mathcal{O}(\psi^2), \\ \delta A_{\mu\nu\rho} &= d_4 \bar{\epsilon}\gamma_{[\mu\nu}\psi_{\rho]} + d_5 \bar{\epsilon}\gamma_{\mu\nu\rho\sigma}\psi^\sigma\end{aligned}$$

By imposing local supersymmetry, 10 unknown coefficients c_i and d_i are uniquely determined.

$$\begin{array}{ccccc} [eR] & [e\psi\psi]_2 & [eF\psi\psi]_4 & [eF^2] & [e\epsilon_{11}AF^2]_1 \\ \swarrow & \searrow & \swarrow & \searrow & \swarrow \\ \downarrow & \downarrow & \downarrow & \downarrow & \downarrow \\ \boxed{[eR\bar{\epsilon}\psi] \quad [eF\bar{\epsilon}D\psi] \quad [eDF\bar{\epsilon}\psi] \quad [eF^2\bar{\epsilon}\psi]} = 0 \end{array}$$

⇒

$$\begin{array}{ccccc} c_1 = -\frac{1}{8}, & c_2 = -\frac{1}{96}, & c_3 = 0, & c_4 = 0, & c_5 = -\frac{1}{(144)^2}, \\ d_1 = 2, & d_2 = -\frac{1}{18}, & d_3 = \frac{1}{144}, & d_4 = -3, & d_5 = 0 \end{array}$$

Higher derivative corrections in string theory are investigated in various ways

- String scattering amplitude Gross, Witten; Gross, Sloan
- Non linear sigma model Grisaru, Zanon
- Superfield method
- Duality
- Noether's method ... and so on

By combining all these results, we find that corrections start from α'^3 order, and **a part of bosonic terms in type IIA** is written as

$$\begin{aligned}
 \mathcal{L} = & e^{-2\phi} R + \dots && \text{SUGRA} \\
 & + \frac{\zeta(3)\alpha'^3}{2^8 \cdot 4!} e^{-2\phi} \left(t_8 t_8 R^4 + \frac{1}{4 \cdot 2!} \epsilon_{10} \epsilon_{10} R^4 \right) + \dots && \text{tree} \\
 & + \frac{\pi^2 \alpha'^3}{3 \cdot 2^8 \cdot 4!} g_s^2 \left(t_8 t_8 R^4 - \frac{1}{4 \cdot 2!} \epsilon_{10} \epsilon_{10} R^4 - \frac{1}{6} \epsilon_{10} t_8 B R^4 \right) + \dots && \text{1-loop}
 \end{aligned}$$

t_8 : tensor with 8 indices , ϵ_{10} : 10D antisymmetric tensor, B : NS 2-form field

Let us discuss how to obtain higher derivative corrections via supersymmetry.

The complete structure of higher derivative terms can be determined by local supersymmetry.

Hyakutake, Ogushi (2005)

Local supersymmetry transformation (neglect flux dependence):

$$\delta e^a{}_\mu = \bar{\epsilon} \gamma^a \psi_\mu, \quad \delta \psi_\mu = 2D_\mu \epsilon, \quad \delta A_{\mu\nu\rho} = -3\bar{\epsilon} \gamma_{[\mu\nu} \psi_{\rho]}$$

Cancellation (neglect flux dependence):

$$\begin{array}{cccc}
 [R^4]_7, & [AR^4]_2, & [R^3\psi\psi]_{92}, & [R^2\psi_2 D\psi_2]_{25} \\
 \swarrow \searrow \swarrow \searrow \swarrow \searrow \swarrow \searrow & & & \\
 \boxed{[R^4\epsilon\psi]_{116}, \quad [R^2DR\epsilon\psi_2]_{88}, \quad [R^3\epsilon D\psi_2]_{40}} & & & = 0
 \end{array}$$

Solution is given by

$$a \left(t_8 t_8 R^4 + \frac{1}{4!} \epsilon_{11} \epsilon_{11} R^4 \right) + b \left(t_8 t_8 R^4 - \frac{1}{4!} \epsilon_{11} \epsilon_{11} R^4 - \frac{1}{6} \epsilon_{11} t_8 A R^4 \right)$$

So far we have just considered the cancellation of variations which do not depend on the 4-form field strength F . Then next step is to examine the cancellation of variations which linearly depend on F .

To do this we add following terms,

$$[eR^3F^2]_{30}, \quad [eR^2(DF)^2]_{24},$$

$$[eR^3F\bar{\psi}\psi]_{447}, \quad [eR^2F\bar{\psi}_2\psi_2]_{190}, \quad [eR^2DF\bar{\psi}\psi_2]_{614}, \quad [eRDF\bar{\psi}_2D\psi_2]_{113}$$

Under the local supersymmetry, these transform into

$$\boxed{[eR^2DRF\bar{\epsilon}\psi]_{1563}, \quad [eR^3F\bar{\epsilon}\psi_2]_{513}, \quad [eR^3DF\bar{\epsilon}\psi]_{995},}$$

$$[eRDRDF\bar{\epsilon}\psi_2]_{371}, \quad [eR^2DF\bar{\epsilon}D\psi_2]_{332}, \quad [eR^2DDF\bar{\epsilon}\psi_2]_{151}} = 0$$

Then we have 4169 equations among 1544 variables. **Solution** becomes

$$b\left(t_8t_8R^4 - \frac{1}{4!}\epsilon_{11}\epsilon_{11}R^4 - \frac{1}{6}\epsilon_{11}t_8AR^4\right) + ([R^3F^2] \oplus [R^2(DF)^2])$$



uniquely determined!

Hyakutake (2007)

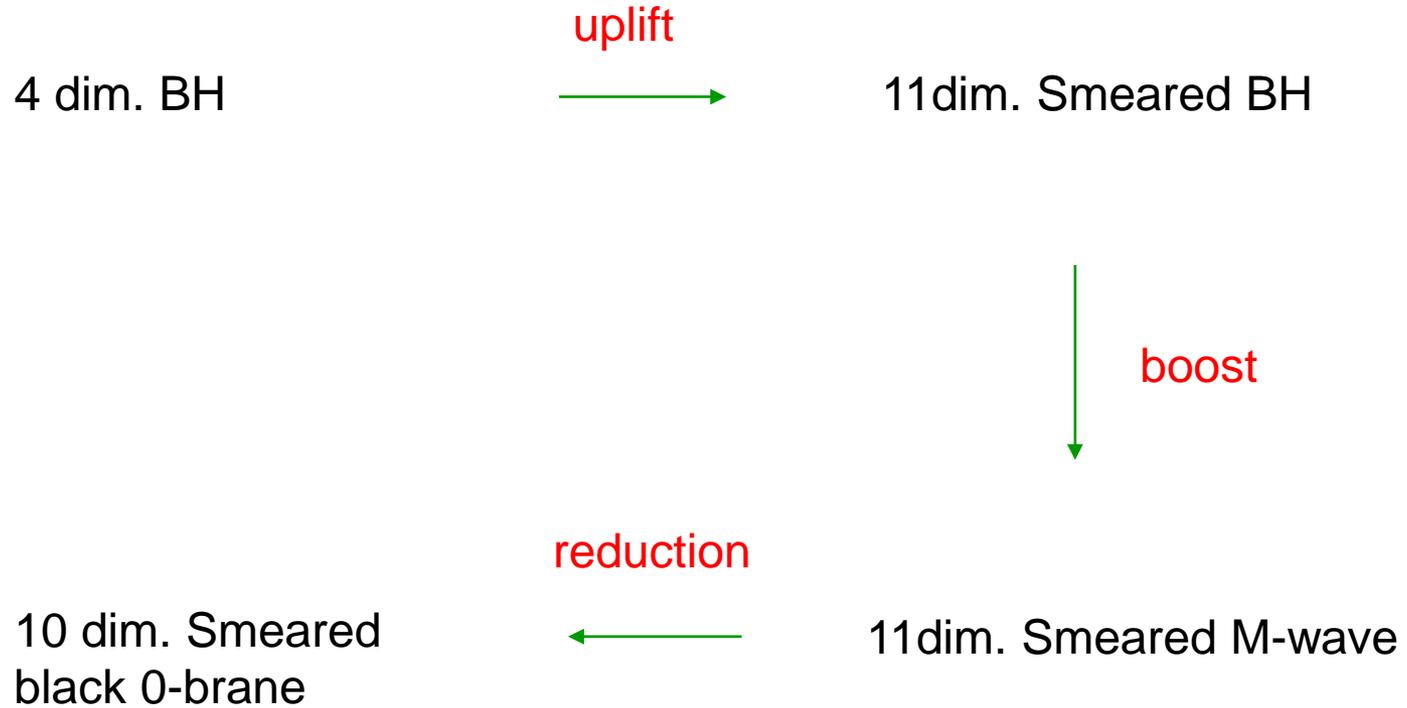
In summary, the higher derivative corrections in M-theory becomes

$$\begin{aligned}
 S_{11} &= \frac{1}{2\kappa_{11}^2} \int d^{11}x e \left\{ R + \gamma \left(t_8 t_8 R^4 - \frac{1}{4!} \epsilon_{11} \epsilon_{11} R^4 \right) \right\} + \dots \\
 &= \frac{1}{2\kappa_{11}^2} \int d^{11}x e \left\{ R + 24\gamma \left(R_{abcd} R_{abcd} R_{efgh} R_{efgh} - 64 R_{abcd} R_{aefg} R_{bcdh} R_{efgh} \right. \right. \\
 &\quad \left. \left. + 2 R_{abcd} R_{abef} R_{cdgh} R_{efgh} + 16 R_{acbd} R_{aebf} R_{cgdh} R_{efgh} \right. \right. \\
 &\quad \left. \left. - 16 R_{abcd} R_{aefg} R_{befh} R_{cdgh} - 16 R_{abcd} R_{aefg} R_{bfch} R_{cdgh} \right) \right\} + \dots
 \end{aligned}$$

$$\gamma \sim \ell_p^6$$

This action contains enough information to deal with quantum corrections to geometrical solution, such as **M-wave** in 11 dimensions, or **black 0-brane** in 10 dimensions.

3. Smearred Black 0-brane



Let us review the construction of smeared black 0-brane.

First we **uplift** 4 dim. BH solution into 11 dimensions.

$$ds_{11}^2 = \underbrace{-F dt^2 + F^{-1} dr^2 + r^2 d\Omega_2^2}_{\text{4 dim. part}} + \underbrace{dy_m^2 + dz^2}_{\text{add 7 dim. directions}}, \quad F = 1 - \frac{r_h}{r}$$

$m = 4, \dots, 9$

Next we **boost** the smeared BH along \mathcal{Z} direction.

$$\begin{aligned} ds_{11}^2 &= -F(\cosh \beta dt + \sinh \beta dz)^2 + (\sinh \beta dt + \cosh \beta dz)^2 + F^{-1} dr^2 + r^2 d\Omega_2^2 + dy_m^2 \\ &= -H^{-1} F dt^2 + F^{-1} dr^2 + r^2 d\Omega_2^2 + dy_m^2 + H \left(dz + (1 - H^{-1}) \frac{\cosh \beta}{\sinh \beta} dt \right)^2 \\ H &= 1 + \frac{r_h \sinh^2 \beta}{r} \end{aligned}$$

Finally **reduce** \mathcal{Z} direction.

$$\begin{aligned} ds_{10}^2 &= -H^{-\frac{1}{2}} F dt^2 + H^{\frac{1}{2}} (F^{-1} dr^2 + r^2 d\Omega_2^2 + dy_m^2), \\ e^\phi &= H^{\frac{3}{4}}, \quad C^{(1)} = (1 - H^{-1}) \frac{\cosh \beta}{\sinh \beta} dt \end{aligned}$$

Smeared black 0-brane

11dim. Smearred
Quantum BH



boost

11dim. Smearred
Quantum M-wave

reduction



10 dim. Smearred
Quantum black 0-brane

Let us construct a **smearing quantum BH solution** in 11 dimensions.

The ansatz is given by

$$ds_{11}^2 = -G_1^{-1} F_1 dt^2 + F_1^{-1} dr^2 + r^2 d\Omega_2^2 + G_2(dy_m^2 + dz^2)$$

EOMs are messy but, up to linear order of γ , we can solve them as

$$F_1(x) = 1 - \frac{1}{x} + \frac{\gamma}{r_h^6} \left[\frac{448}{x} \left\{ c + \sum_{n=1}^9 \frac{1}{nx^n} + \frac{216}{7x^8} - \frac{215}{7x^9} \right\} \right], \quad x = \frac{r}{r_h}$$

$$G_1(x) = 1 + \frac{\gamma}{r_h^6} \left[896 \left\{ \sum_{n=1}^9 \frac{n+1}{nx^n} + \frac{12}{x^9} \right\} \right],$$

$$G_2(x) = 1 + \frac{\gamma}{r_h^6} \left[256 \sum_{n=1}^9 \frac{1}{nx^n} \right]$$

Here c is an integral constant, but it can be absorbed by redefinition of r_h .

There are two more integral constants, but they are fixed by demanding asymptotic flatness and no singularity at $r = r_h$.

Now we **boost** the smeared BH along \mathcal{Z} direction.

$$\begin{aligned}
 ds_{11}^2 &= -G_1^{-1} F_1 (\cosh \beta dt + \sinh \beta dz)^2 + G_2 (\sinh \beta dt + \cosh \beta dz)^2 \\
 &\quad + F_1^{-1} dr^2 + r^2 d\Omega_8^2 + G_2 dy_m^2 \\
 &= -H_1^{-1} F_1 dt^2 + F_1^{-1} dr^2 + r^2 d\Omega_8^2 + G_2 dy_m^2 \\
 &\quad + H_2 \left(dz + \left(1 - H_2^{-\frac{1}{2}} H_3^{-\frac{1}{2}} \right) \frac{\cosh \beta}{\sinh \beta} dt \right)^2
 \end{aligned}$$

$$H_1 = G_1 G_2^{-1} H_2, \quad H_2 = G_2 + (G_2 - G_1^{-1} F_1) \sinh^2 \beta, \quad H_3 = G_2^{-2} H_2$$

And **reduce** \mathcal{Z} direction. We obtain **smeared quantum black 0-brane solution**.

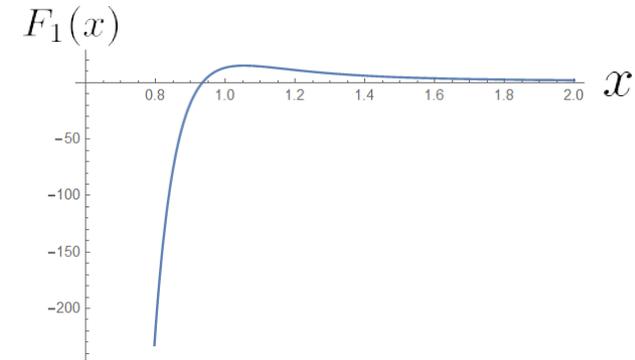
$$\begin{aligned}
 ds_{10}^2 &= -H_1^{-1} H_2^{\frac{1}{2}} F_1 dt^2 + H_2^{\frac{1}{2}} \left(F_1^{-1} dr^2 + r^2 d\Omega_2^2 + G_2 dy_m^2 \right), \\
 e^\phi &= H_2^{\frac{3}{4}}, \quad C^{(1)} = \left(1 - H_2^{-\frac{1}{2}} H_3^{-\frac{1}{2}} \right) \frac{\cosh \beta}{\sinh \beta} dt
 \end{aligned}$$

Thermodynamics of the smeared quantum black 0-brane

Horizon

$$F_1\left(\frac{r_{\text{horizon}}}{r_h}\right) = 0 \quad \text{at the horizon}$$

$$r_{\text{horizon}} = r_h - \frac{\gamma}{r_h^5} f_1(1)$$



Temperature

$$T = \frac{1}{4\pi} H_1^{-1/2} \left. \frac{dF_1}{dr} \right|_{r_{\text{horizon}}} = \frac{1}{4\pi r_h} \sqrt{\frac{\alpha}{1+\alpha}} \left\{ 1 - \gamma \frac{128-c}{r_h^6} \right\} \quad \alpha = \frac{1}{\sinh^2 \beta}$$

Entropy

Black hole entropy is evaluated by using Wald's formula.

$$S = -2\pi V_6 \int_{\text{horizon}} d\Omega_2 dz \sqrt{h} \frac{\partial \mathcal{S}}{\partial R_{\mu\nu\rho\sigma}} N_{\mu\nu} N_{\rho\sigma}$$

By inserting the solution obtained so far, the entropy is calculated as

$$S = \frac{4\pi V_6}{2\kappa_{10}^2} 4\pi r_h^2 \sqrt{\frac{1+\alpha}{\alpha}} \left\{ 1 + \gamma \frac{2(1088-c)}{r_h^6} \right\}$$

Mass and Charge

Mass is calculated by using ADM mass formula.

$$M = \frac{4\pi V_6}{2\kappa_{10}^2} r_h \left\{ 2 + \frac{1}{\alpha} - \gamma \frac{2(-896 + c) + \alpha^{-1}(-2048 + c)}{r_h^6} \right\},$$
$$Q = \frac{4\pi V_6}{2\kappa_{10}^2} \frac{\sqrt{1 + \alpha}}{\alpha} r_h \left\{ 1 + \gamma \frac{2048 - c}{r_h^6} \right\}$$

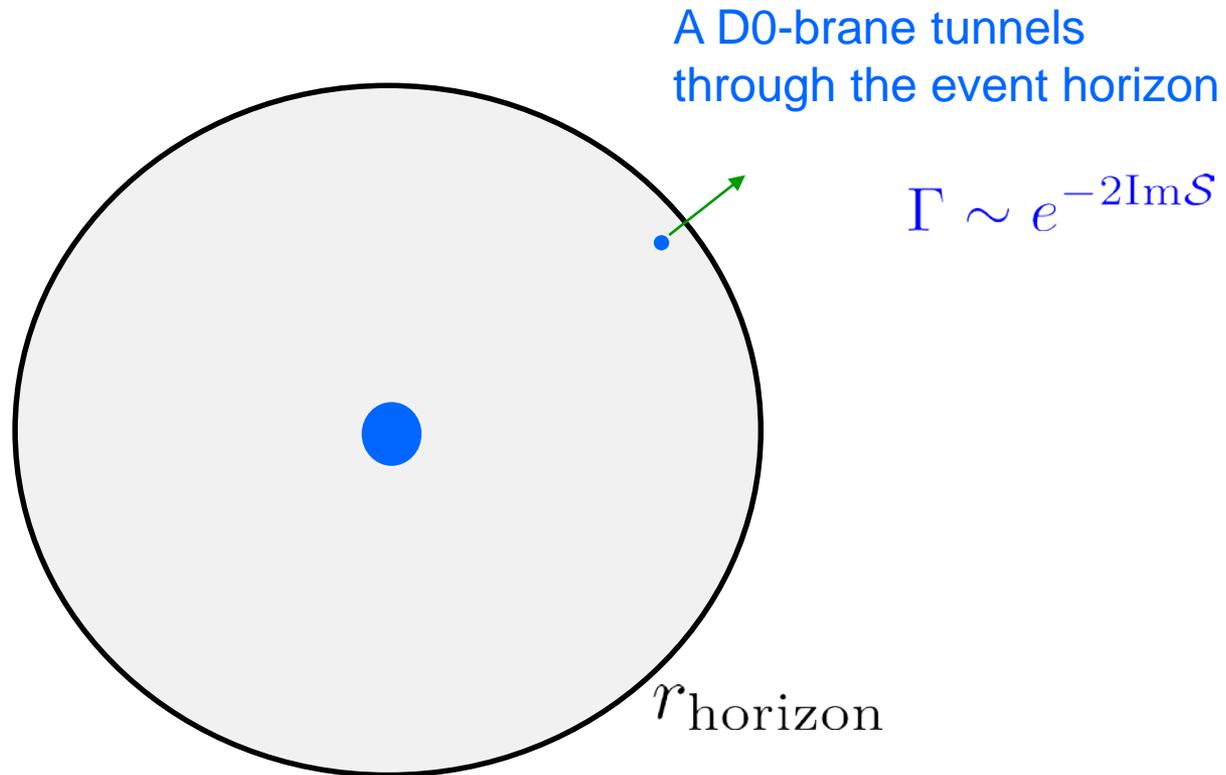
Charge is not renormalized if we choose $c = 2048$.

1st Law of thermodynamics

$$\delta M = T\delta S + \Phi\delta Q$$

Electric potential : $\Phi = C_t^{(1)} \Big|_{\text{horizon}} = 1/\sqrt{1 + \alpha}$

4. Hawking Radiation of a D0-brane from Smearred Black 0-brane



Smearing quantum black 0-brane

Tunneling effect can be evaluated by using WKB approximation.

Parikh-Wilczek(1998)

Thus we want to evaluate the imaginary part of the action in the background of smeared quantum black 0-brane.

Lagrangian for a D0-brane, moving along the radial direction, is given by

$$\mathcal{S} = -T_0 \int dt e^{-\phi} \sqrt{-\det (g_{\mu\nu} \partial_t X^\mu \partial_t X^\nu)} - T_0 \int dt C_t^{(1)}$$
$$\mathcal{L} = -T_0 e^{-\phi} H_1^{-\frac{1}{2}} H_2^{\frac{1}{4}} F_1^{\frac{1}{2}} \sqrt{1 - H_1 F_1^{-2} \dot{r}^2} - T_0 C_t^{(1)}$$

Then conjugate momentum and Hamiltonian are written as

$$p_r = \frac{\partial \mathcal{L}}{\partial \dot{r}} = T_0 H_1^{-\frac{1}{2}} H_2^{-\frac{1}{2}} F_1^{\frac{1}{2}} \frac{H_1 F_1^{-2} \dot{r}}{\sqrt{1 - H_1 F_1^{-2} \dot{r}^2}}$$
$$\mathcal{H}(r, p_r) = H_1^{-\frac{1}{2}} H_2^{-\frac{1}{2}} F_1^{\frac{1}{2}} \sqrt{T_0^2 + H_2 F_1 p_r^2} + T_0 C_t^{(1)}$$

Next we consider Hamilton-Jacobi equation.

$$-\frac{\partial \mathcal{S}}{\partial t} = \mathcal{H}(r, p_r), \quad p_r = \frac{\partial \mathcal{S}}{\partial r}$$

To solve this, we assume

$$S = -\underline{\delta E} t + W(r)$$

energy of test D0-brane

Then we obtain

$$W(r) = \int dr H_1^{\frac{1}{2}} F_1^{-1} \sqrt{(\delta E - T_0 C_t^{(1)})^2 - T_0^2 H_1^{-1} H_2^{-1} F_1}$$

$$\sim \frac{1}{2\pi T} \int \frac{d\rho}{\rho} \sqrt{(\delta E - T_0 C_t^{(1)})^2 - T_0^2 H_1^{-1} H_2^{-1} F_1} \quad \text{pole at horizon}$$

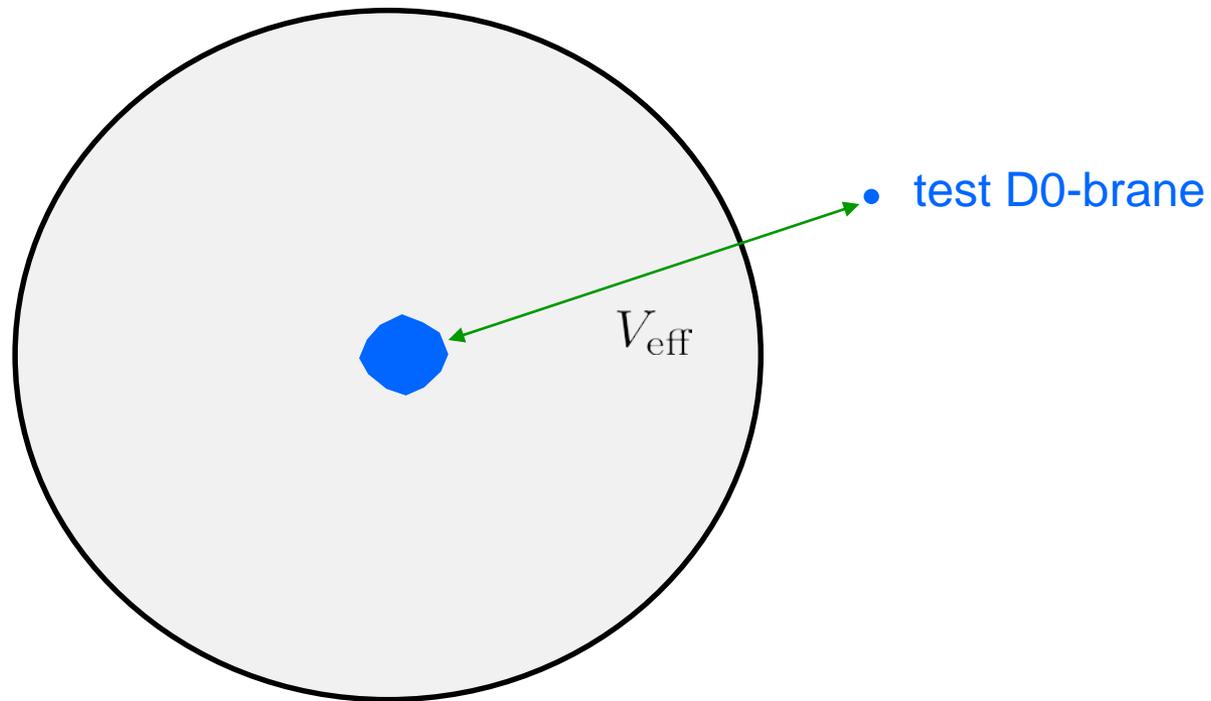
$$d\rho = H_2^{\frac{1}{4}} F_1^{-\frac{1}{2}} dr, \quad H_1^{-1} H_2^{\frac{1}{2}} F_1 \sim \frac{1}{4} H_1^{-1} \left(\frac{dF_1}{dr} \right)^2 \Big|_{r_{\text{horizon}}} \rho^2$$

The emission rate for a D0-brane is estimated as

$$P = e^{-2 \text{Im} W} = e^{-\frac{\delta E - T_0 \Phi}{T}} = e^{-\delta S}$$

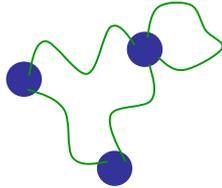
5. Matrix Model Analysis of Smearred Fuzzy Sphere

fuzzy sphere = Non-commutative configuration of D0-branes



Smeared fuzzy sphere via D0-branes

N D0-branes are dynamical due to oscillations of open strings



massless modes : $N \times N$ matrices

$$(A_t(t), \Phi_i(t), \theta(t)) \quad i = 1, \dots, 9$$

Action for N D0-branes is obtained by requiring global supersymmetry with 16 supercharges.

➔ (1+0) dimensional supersymmetric $U(N)$ gauge theory

$$\mathcal{S}_0 = \frac{1}{g_{\text{YM}}^2} \int dt \text{tr} \left(\frac{1}{2} D_t \Phi_i D_t \Phi^i + \frac{1}{4} [\Phi_i, \Phi_j]^2 + \frac{i}{2} \theta^T D_t \theta + \frac{1}{2} \theta^T \gamma^i [\Phi_i, \theta] \right)$$

$$D_t \Phi_i = \partial_t \Phi_i - i[A_t, \Phi_i]$$

EOMs become

$$\ddot{\Phi}_i = [\Phi^j, [\Phi_i, \Phi_j]], \quad [\Phi^i, D_t \Phi_i] = 0 \quad \theta = 0$$

Ansatz for the fuzzy sphere is written as

$$\Phi_a = \frac{f(t)}{2\pi\ell_s^2} \frac{\Sigma_a}{2}, \quad \left[\frac{\Sigma_a}{2}, \frac{\Sigma_b}{2} \right] = i\epsilon_{abc} \frac{\Sigma_c}{2}$$

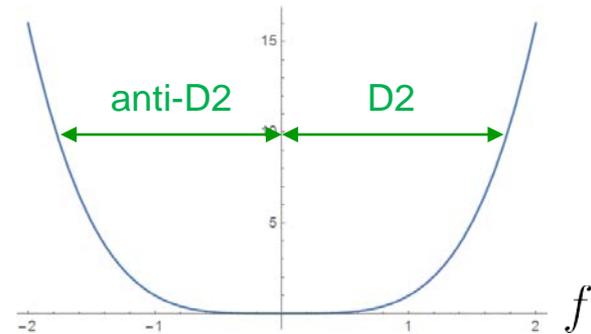
$$a, b, c = 1, 2, 3$$

Then EOMs become

$$\ddot{f} = -\frac{2}{(2\pi\ell_s^2)^2} f^3$$

➔ Oscillation in f^4 potential

Collins, Tucker (1976)



Radius of the fuzzy sphere is estimated as

$$R(t) = \sqrt{\frac{1}{N} \text{tr}(X_1^2 + X_2^2 + X_3^2)} = \frac{f(t)}{2} \sqrt{N^2 - 1}$$

$$X_i = 2\pi\ell_s^2 \Phi_i$$

Now we consider the fuzzy sphere smeared into (x_4, \dots, x_9) directions, which are represented as

$$\Phi_a = \frac{f(t)}{2\pi\ell_s^2} \frac{\Sigma_a}{2} \otimes \mathbf{1}_Z, \quad \Phi_m = \mathbf{1}_N \otimes P_m$$

$$a = 1, 2, 3 \quad m = 4, \dots, 9$$

We want to evaluate an effective potential between smeared fuzzy sphere and a test D0-brane.

Kabat, Taylor (1997)

Euclidean action is given by

$$\mathcal{S}_E = \frac{1}{2g_{\text{YM}}^2} \int dt \text{tr} \left(D_\tau \Phi_i D_\tau \Phi^i - \frac{1}{2} [\Phi_i, \Phi_j]^2 + (\dot{A}_\tau - i[B_i, \Phi^i])^2 \right. \\ \left. + \theta^T D_\tau \theta - \theta^T \gamma^i [\Phi_i, \theta] - D_\mu \bar{C} D^\mu C \right)$$

Let us consider the fluctuation around fuzzy sphere background and evaluate 1-loop effective potential.

$$B_i = \begin{pmatrix} b_i & \leftarrow \theta \\ 0 & x_i \end{pmatrix} \begin{array}{l} \leftarrow \text{smeared fuzzy sphere} \\ \leftarrow \text{test D0} \end{array}$$

$$A_\tau = \begin{pmatrix} 0 & a \\ a^\dagger & 0 \end{pmatrix}, \quad \Phi_i = B_i + \begin{pmatrix} 0 & \phi_i \\ \phi_i^\dagger & 0 \end{pmatrix}, \quad \theta = \begin{pmatrix} 0 & \psi \\ \psi^\dagger & 0 \end{pmatrix}, \quad C = \begin{pmatrix} 0 & c \\ c^\dagger & 0 \end{pmatrix}$$

Mass squared for fluctuations : $\Omega^2 = M_0 + M_1$

$$K_k = b_k - x_k \mathbf{1}_{N'} \quad k = 1, \dots, 9$$

$$K^2 = K_k^2$$

boson

$$\Omega_b^2 = K^2 \mathbf{1}_{10} + M_{1b}$$

$$M_{1b} = 2i \begin{pmatrix} 0 & \dot{K}_j \\ -\dot{K}_i & -i[K_i, K_j] \end{pmatrix} \equiv 2i \begin{pmatrix} 0 & F_{\tau j} \\ F_{i\tau} & F_{ij} \end{pmatrix}$$

fermion

$$\Omega_f^2 = K^2 \mathbf{1}_{16} + M_{1f}$$

$$M_{1f} = \gamma^i \dot{K}_i + \frac{1}{2} \gamma^{ij} [K_i, K_j] \equiv \frac{i}{2} \gamma^{\mu\nu} F_{\mu\nu}$$

ghost

$$\Omega_g^2 = K^2$$

The effective potential is given by

$$\begin{aligned}
 V_{\text{eff}} &= \text{tr}(\Omega_b) - \frac{1}{2}\text{tr}(\Omega_f) - 2\text{tr}(\Omega_g) \\
 &= -\frac{1}{2\sqrt{\pi}} \int_0^\infty \frac{d\tau}{\tau^{3/2}} \text{tr}(e^{-\tau\Omega_b^2}) + \frac{1}{4\pi} \int_0^\infty \frac{d\tau}{\tau^{3/2}} \text{tr}(e^{-\tau\Omega_f^2}) + \frac{1}{\pi} \int_0^\infty \frac{d\tau}{\tau^{3/2}} \text{tr}(e^{-\tau\Omega_g^2})
 \end{aligned}$$

Each term can be evaluated perturbatively like

$$\text{tr}(e^{-\tau(M_0+M_1)}) = \text{tr}(e^{-\tau M_0} U(\tau))$$

$$U(\tau) \equiv e^{\tau M_0} e^{-\tau(M_0+M_1)} = \mathbf{1} - \int_0^\tau dz_1 M_1(z_1) + \int_0^\tau dz_1 M_1(z_1) \int_0^{z_1} dz_2 M_1(z_2) - \dots$$

$$M_1(\tau) \equiv e^{\tau M_0} M_1 e^{-\tau M_0}$$

$$= \text{tr}(e^{-\tau M_0}) - \int_0^\tau dz_1 \text{tr}(e^{-\tau M_0} M_1(z_1)) + \dots$$

Non-trivial contributions arises at M_1^4 .

The effective action of $\mathcal{O}(F^4)$ is expressed as

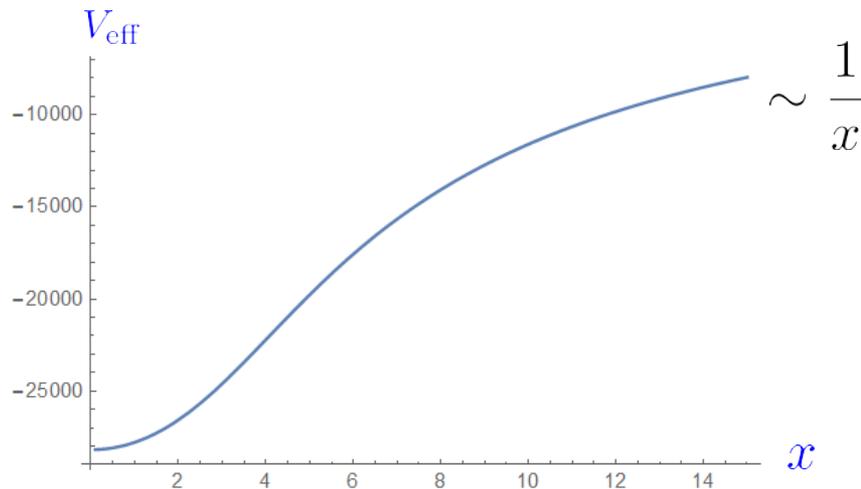
$$V_{\text{eff}} = -\frac{1}{2\sqrt{\pi}} \int_0^\infty \frac{d\tau}{\tau^{3/2}} \int_0^\tau dz_1 \int_0^{z_1} dz_2 \int_0^{z_2} dz_3 \int_0^{z_3} dz_4$$

$$\text{tr}_{N'} \left[e^{-\tau K^2} \left\{ 8F^\mu{}_\nu(z_1)F^\nu{}_\rho(z_2)F^\rho{}_\sigma(z_3)F^\sigma{}_\mu(z_4) + 16F_{\mu\nu}(z_1)F^{\mu\lambda}(z_2)F^{\nu\sigma}(z_3)F_{\lambda\sigma}(z_4) \right. \right.$$

$$\left. \left. - 4F_{\mu\nu}(z_1)F^{\mu\nu}(z_2)F_{\rho\sigma}(z_3)F^{\rho\sigma}(z_4) - 2F_{\mu\nu}(z_1)F_{\rho\sigma}(z_2)F^{\mu\nu}(z_3)F^{\rho\sigma}(z_4) \right\} \right]$$

It is possible to evaluate the above potential by inserting the background.

The result for $N = 2$ smeared fuzzy sphere with $R \sim 5$ becomes :



6. Summary

Summary :

- We have constructed the smeared black 0-brane solution including quantum corrections.
- The radiation of a D0-brane from smeared black 0-brane is estimated.
- Dual matrix analysis for the smeared fuzzy sphere has been done.

Future directions :

- Numerical analysis for the smeared fuzzy sphere
- Test of the gauge/gravity duality