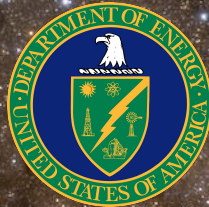


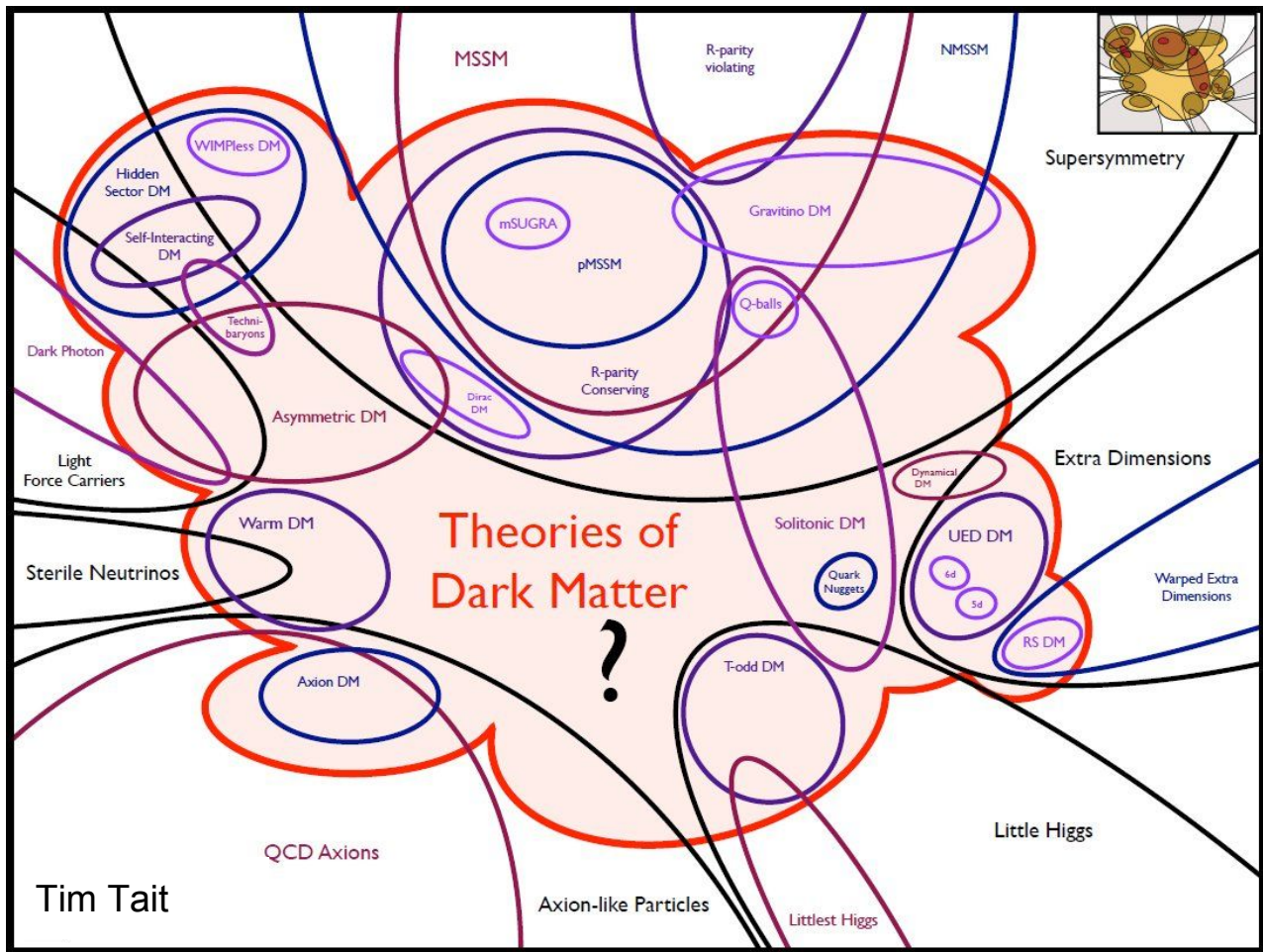
Primordial Black Hole Dark Matter

New Venues in Formation and Detection

Volodymyr Takhistov

University of California, Los Angeles
(UCLA)





PBH DM

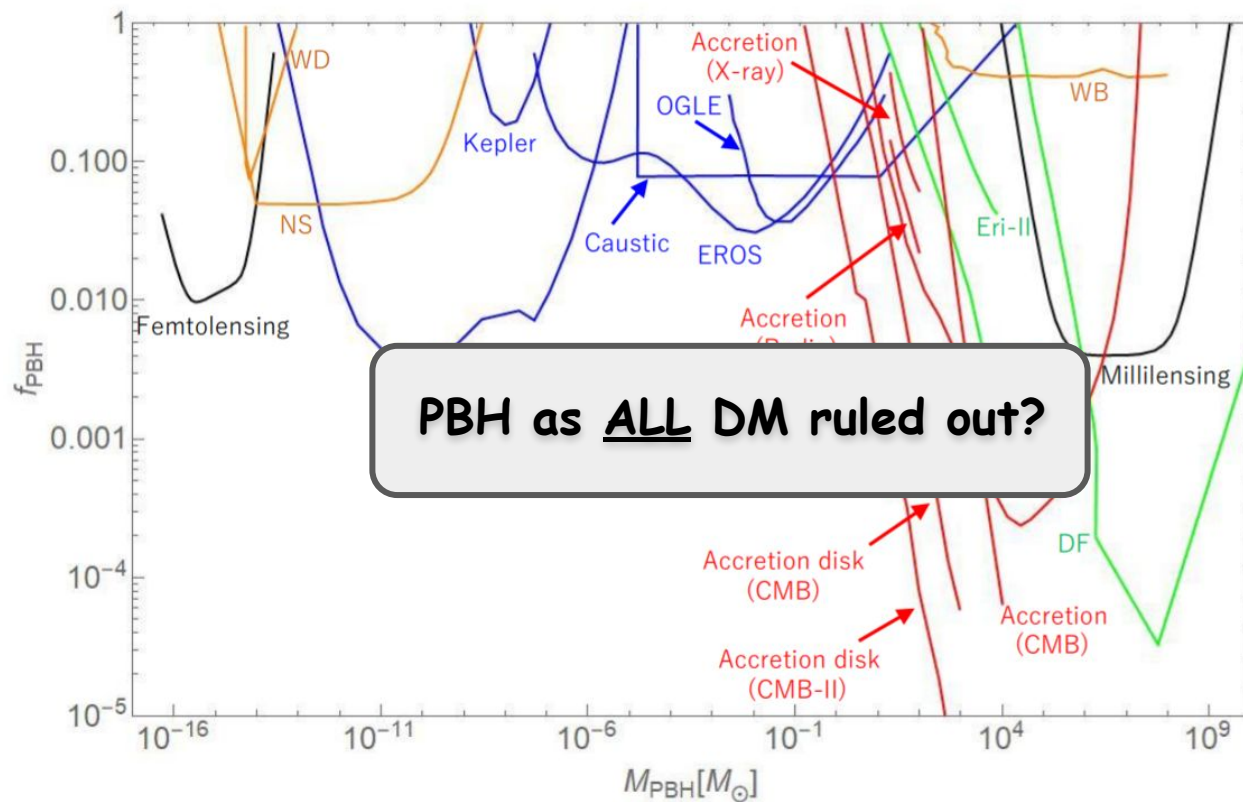
- Black holes

- astrophysical → old stars
- primordial → early Universe [Zeldovich, Novikov, 1967; Hawking, 1971; Carr, Hawking, 1974]

- Why get excited about PBH DM ?

- no signs of particle DM
- GW astronomy [Sasaki, Thorne+, 1997; Bird+, 2016; Garcia-Bellido+, 2015...]
- already appear in standard cosmology (but unlikely)
- generic in many BSM models
- help solve astronomy puzzles (e.g. seed massive BHs)
[Kawasaki, Kusenko, Yanagida, 2012]

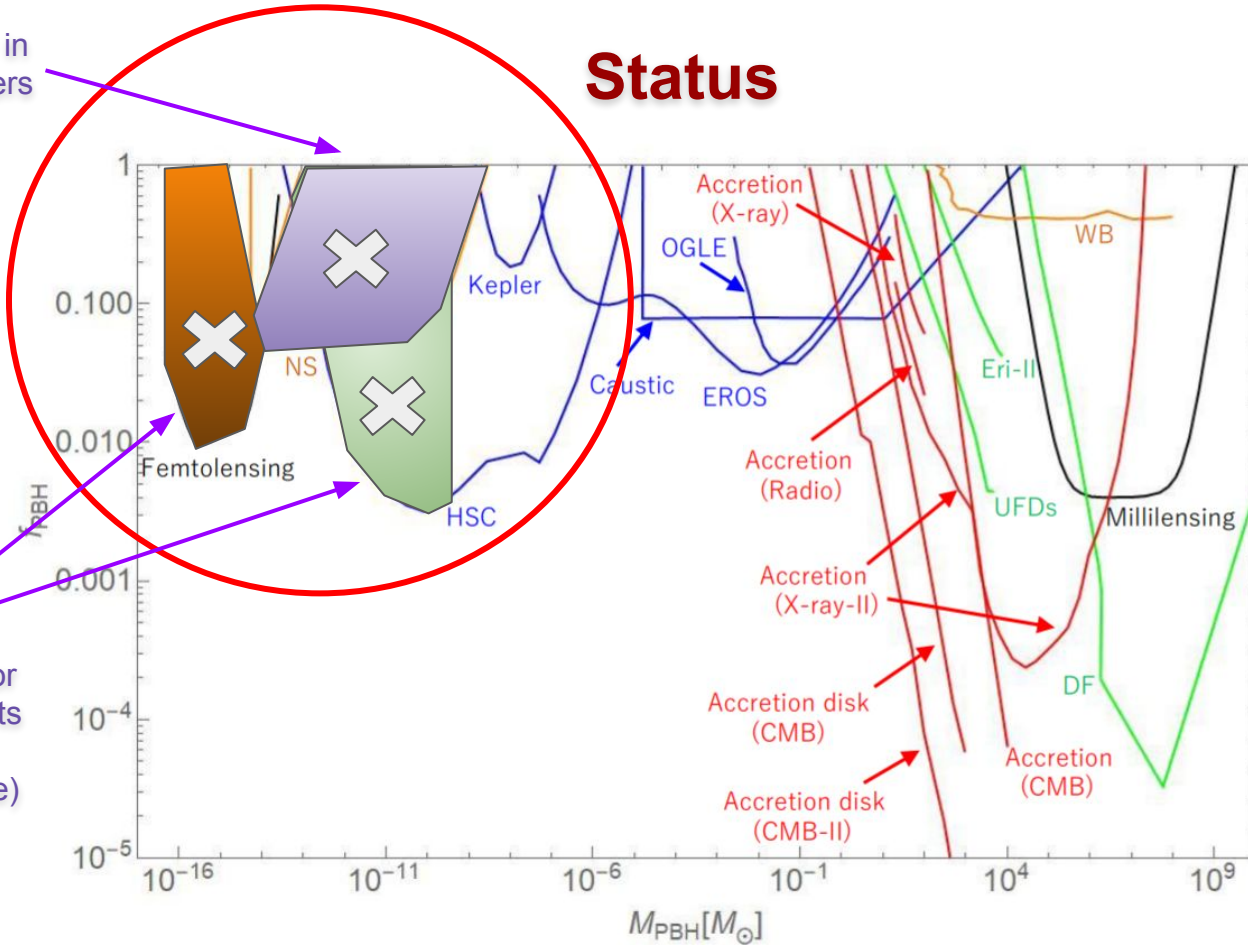
Status



[Sasaki+, 2017]

Status

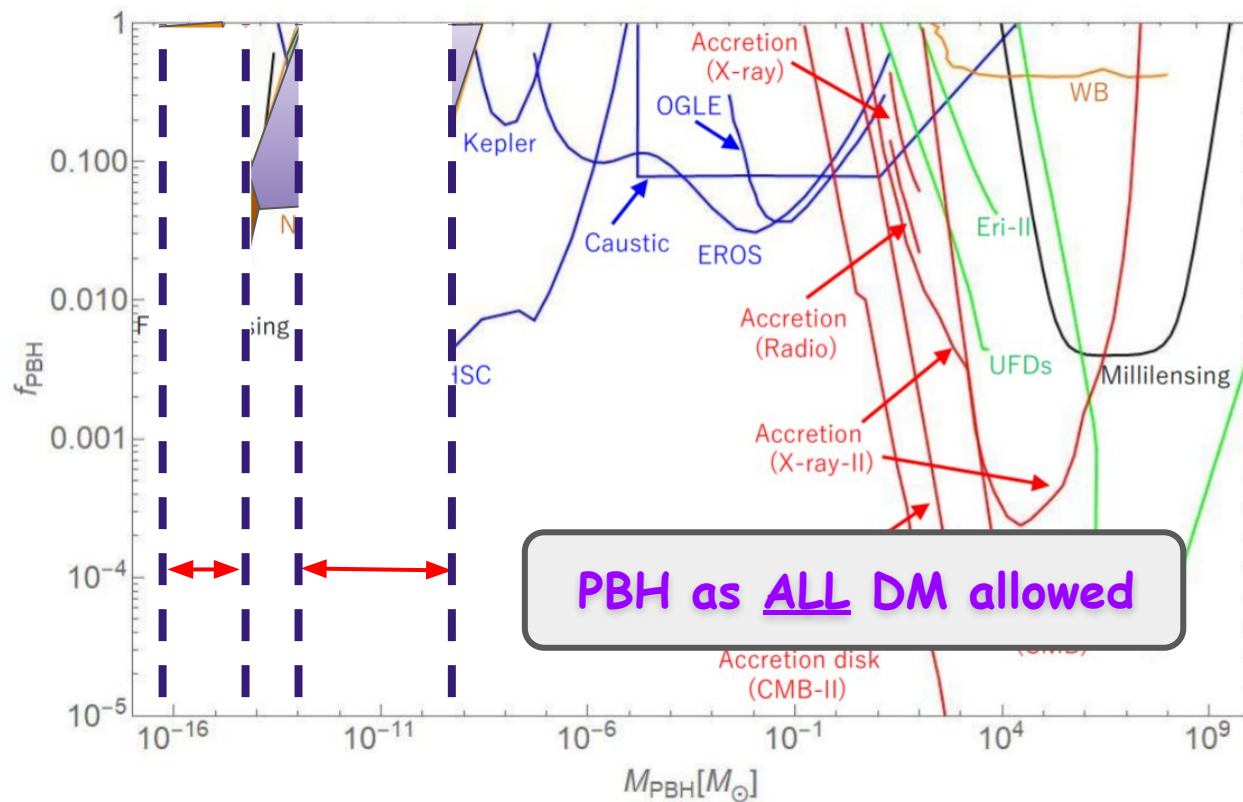
too much DM in
globular clusters



didn't account for
all lensing effects
(wave optics /
finite source size)

[Sasaki+, 2017]

Status



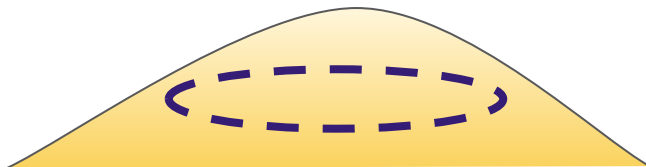
[Sasaki+, 2017]

How do PBHs form?

PBH formation: radiation-dominated era

- THE “standard scenario”: large perturbations ($\delta \sim 1$) enter horizon $\rightarrow$ collapse

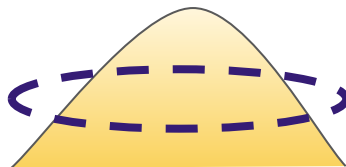
[Kawasaki, Sasaki, Yanagida ...]



PBH formation: radiation-dominated era

- THE “standard scenario”: large perturbations ($\delta \sim 1$) enter horizon $\rightarrow$ collapse

[Kawasaki, Sasaki, Yanagida ...]



PBH formation: radiation-dominated era

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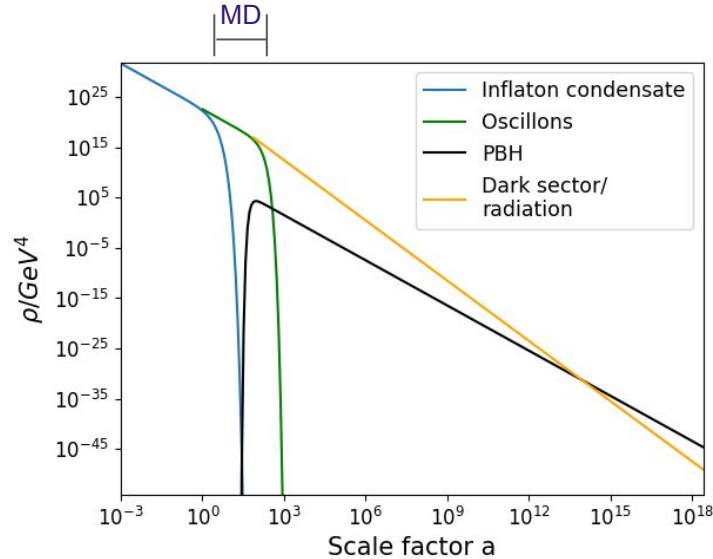


- Need to tune inflaton potential
- Sensitive to restrictions on scalar field behavior
 - Example: “string swampland conjectures” [Kawasaki, VT, 2018]

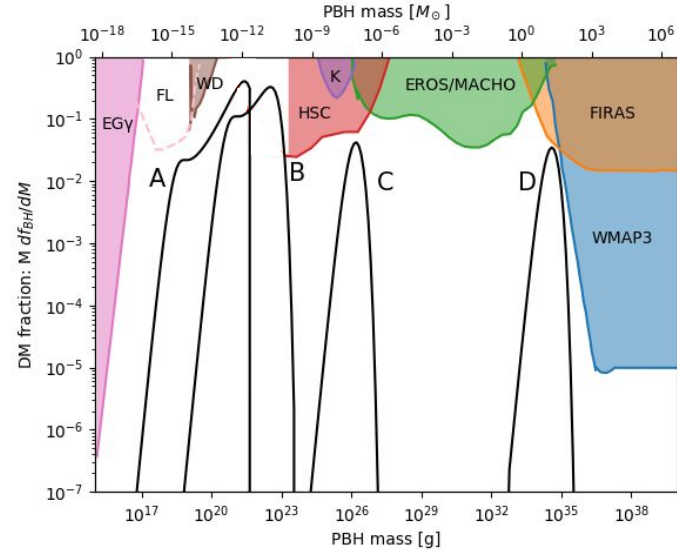
PBH formation: matter-dominated era

- In MD perturbations grow $\delta \sim a(t)$, even small ones can become large
- Simple general mechanism: scalar field fragmentation
 - **A**: spectator field (charged: Q-balls) [Cotner, Kusenko, PRL 2016]
 - **B**: inflating field (real: oscillons) [Cotner, Kusenko, VT, PRD 2018]
- Basic idea
 - after inflation oscillating field breaks apart
 - statistics govern lump distribution (big lumps $\rightarrow$ big fluctuations)
 - region densities grow with time $\rightarrow$ some collapse to BH

PBH formation: matter-dominated era



[Cotner, Kusenko, VT, PRD 2018]



* New analytic tools to streamline fragmentation computations to be released soon
[Cotner, Kusenko, VT, *prep.*]

How to test
open parameter space?

Compact stars as PBH laboratories

- PBHs can be effectively captured by NS or WD in DM-rich environments (e.g. Galactic Center)
- Captured PBH settle and grow inside, destroying the star host

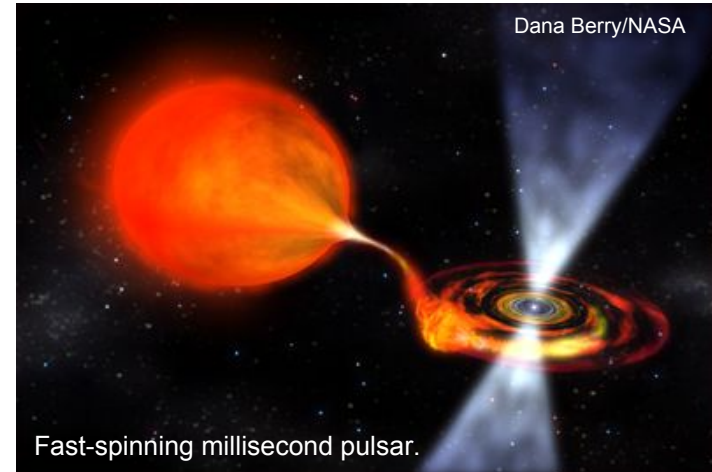
→ r-process nucleosynthesis, 511 keV, FRBs

[Fuller, Kusenko, VT, PRL 2017]

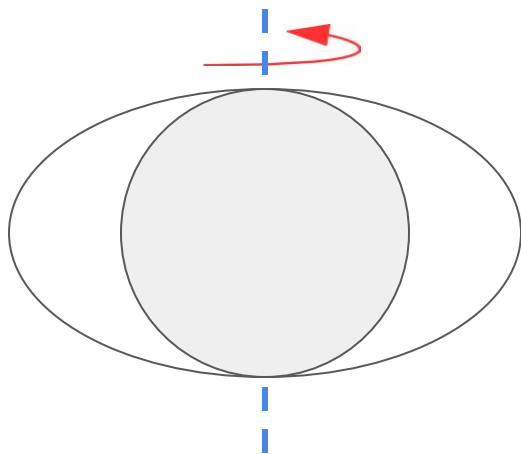
+ *Viewpoint Highlight* by H.-T. Janka

→ solar-mass BHs, GRBs, microquasars

[VT, PLB 2017; VT, 2017]



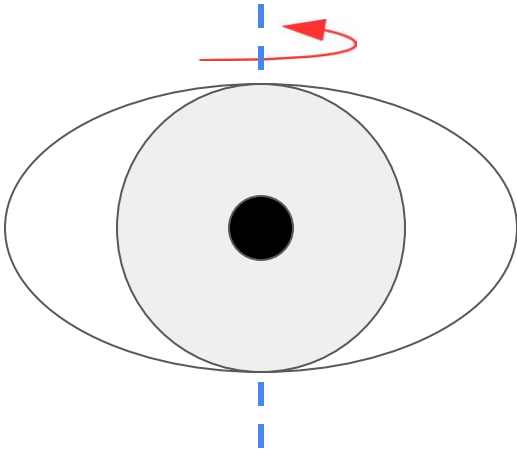
PBH inside MSP



- MSP near mass-shedding limit → stretched spheroid (analytic Roche lobe model)

[Fuller, Kusenko, VT, PRL 2017]

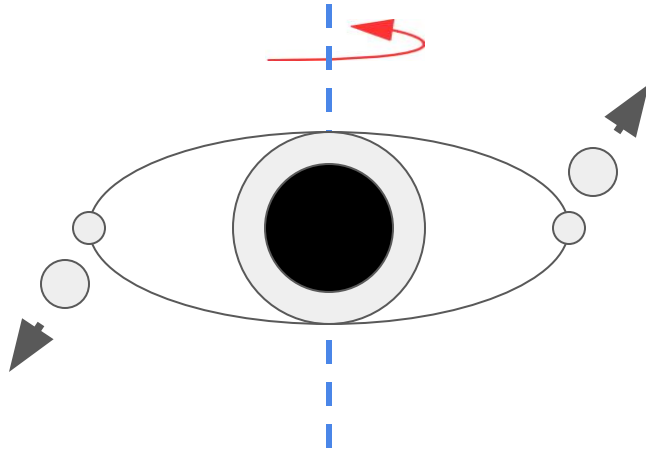
PBH inside MSP



- MSP near mass-shedding limit → stretched spheroid (analytic Roche lobe model)
 - Add PBH: star consumed from inside, contracts → angular momentum transferred out
- * viscosity [Kouvaris, Tinyakov, 2015] & magnetic stresses ensure rigid rotation

[Fuller, Kusenko, VT, PRL 2017]

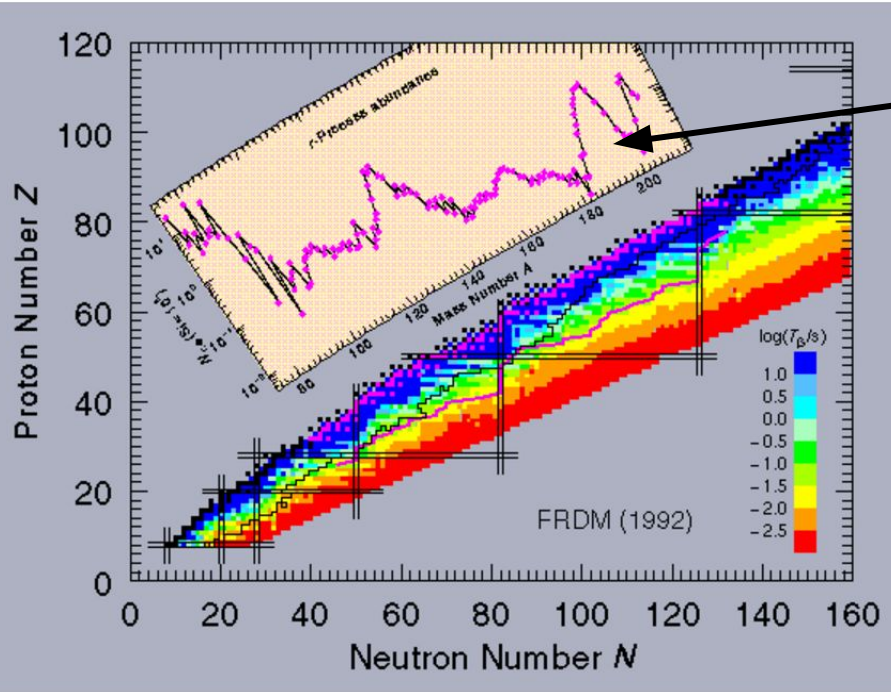
PBH inside MSP



- **MSP near mass-shedding limit** → stretched spheroid (analytic Roche lobe model)
- **Add PBH:** star consumed from inside, contracts → angular momentum transferred out
 - * viscosity [Kouvaris, Tinyakov, 2015] & magnetic stresses ensure rigid rotation
- Matter exceeds escape velocity at equator → **neutron-rich ejecta**

[Fuller, Kusenko, VT, PRL 2017]

R-process: nucleosynthesis



- Neutrons **rapidly** captured before β -decay
- Dominant producer of heavy astro-elements
- Need neutron-rich setting, site **unknown**
 - **SN?** problem (neutrinos)
 - **NS-NS?** Make all? (evolution, kilonova)

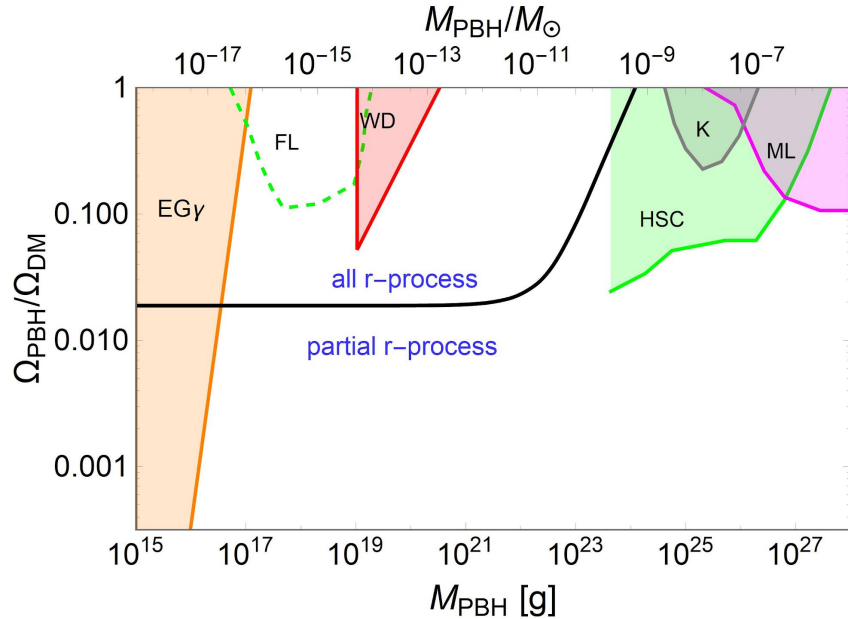
Image: Los Alamos, Nuclear Data Group

R-process: observations

- Milky Way (total) $\sim 10^4 M_{\odot}$
- Ultra-faint dwarfs (UFDs): 1 (Reticulum II) out of 10 shows large abundance
→ consistent with “rare event” [Ji+, Nature 2016]

R-process: PBH-NS

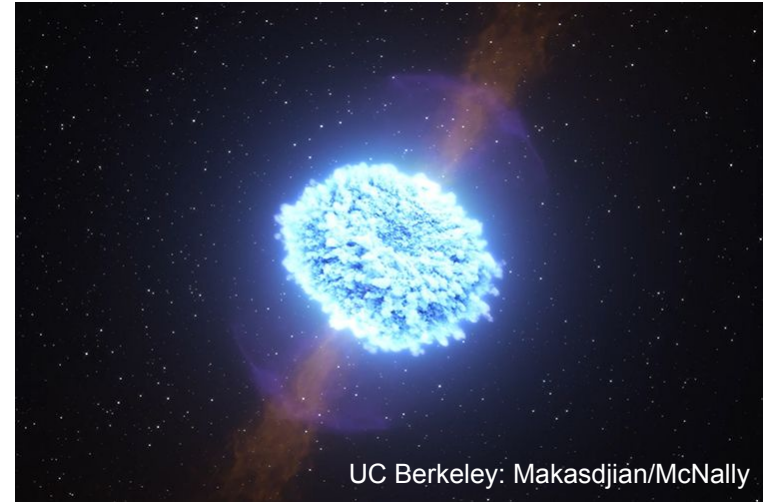
- PBH-NS emits $\sim 0.1M_{\odot}$
(10x as NS-NS!)
- Consistent with abundances in
Milky Way & UFDs ✓



[Fuller, Kusenko, VT, PRL 2017]

PBH-NS laboratories: long kilonova

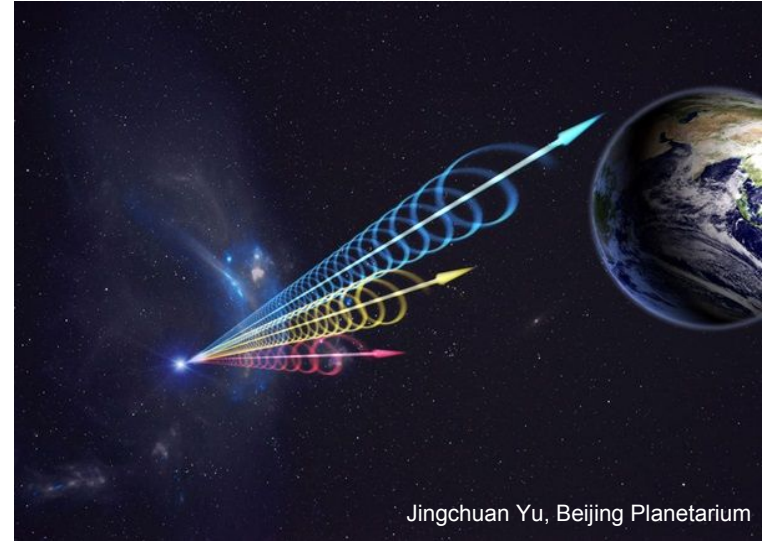
- **Kilonova:** afterglow from ejecta radioactive decays
- **PBH-NS vs. mergers**
 - more ejecta → longer kilonova
 - no merger GWs



[Fuller, Kusenko, VT, PRL 2017]

PBH-NS laboratories: FRBs

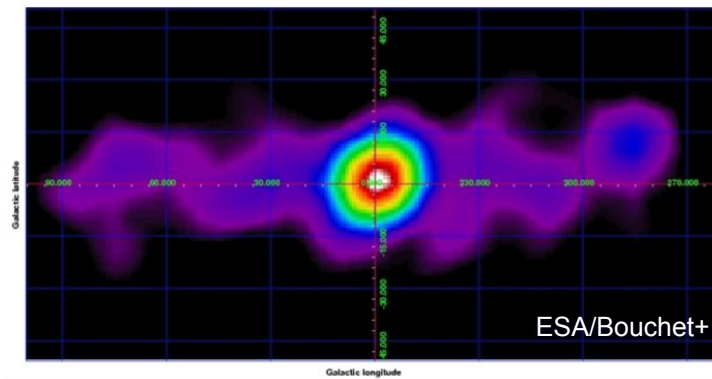
- Fast Radio Bursts
 - origin unknown
 - ~100 found, 1 repeater
- Release of NS magnetic field energy on PBH-NS consumption timescales (ms)
→ non-repeating FRB



[Fuller, Kusenko, VT, PRL 2017]

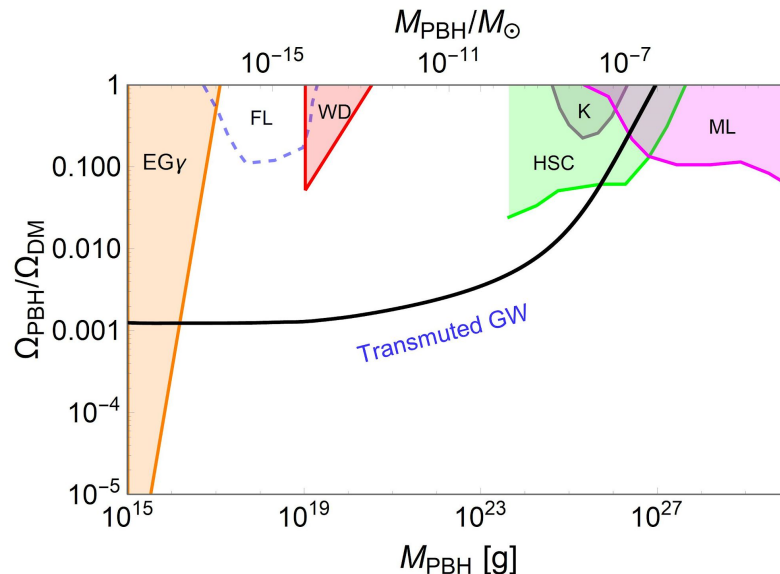
PBH-NS laboratories: 511 keV

- Galactic Center 511 keV γ -rays
 - consistent with e^+ annihilation
 - e^+ origin unknown
 - need e^+ production rate $10^{50}/\text{yr}$
 - need e^+ energies $< 3 \text{ MeV}$
- Temperatures & production rates from PBH-NS ejecta consistent with GC signal
- 511 keV signal discovered in Reticulum II [Siegert+, 2016]
→ strengthens connection



PBH-NS/WD laboratories: solar-mass BHs

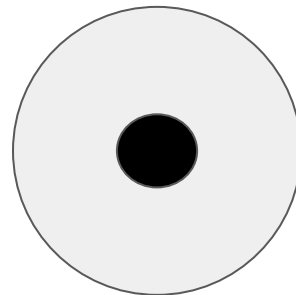
- No astrophysical BHs $\lesssim 2M_{\odot}$
- PBH + NS/WD
→ “transmuted” solar-mass BHs
- New “transmuted” binary signals
(e.g. double kilonova)
- Evade constraints on solar-mass PBHs



[VT, PLB 2017]

PBH-NS laboratories: GRBs

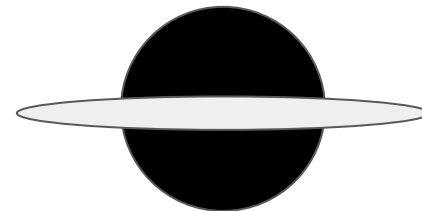
- “Standard” Gamma-ray Burst progenitor: BH + disk
→ disk accreted, binding energy released
- Can form from PBH-NS → GRB without GW



[VT, 2017]

PBH-NS laboratories: GRBs

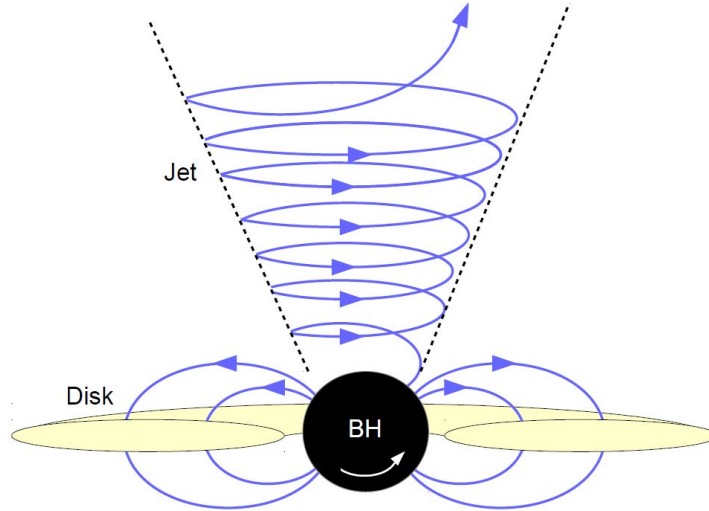
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[VT, 2017]

PBH-NS laboratories: GRBs

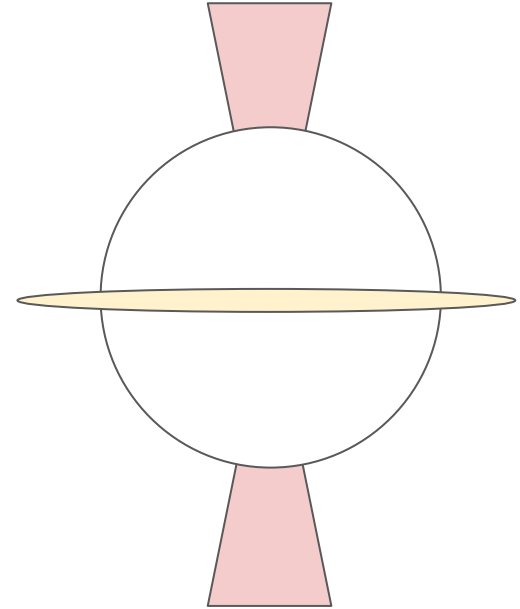
- GRB jet formation



Blandford-Znajek
Mechanism

PBH-WD laboratories: microquasars

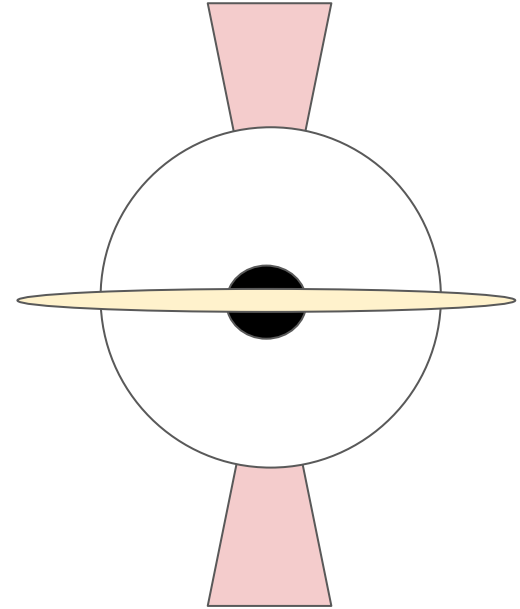
- Accreting WD jets non-relativistic $L_{\text{jet}} \sim \frac{1}{R}$



[VT, 2017]

PBH-WD laboratories: microquasars

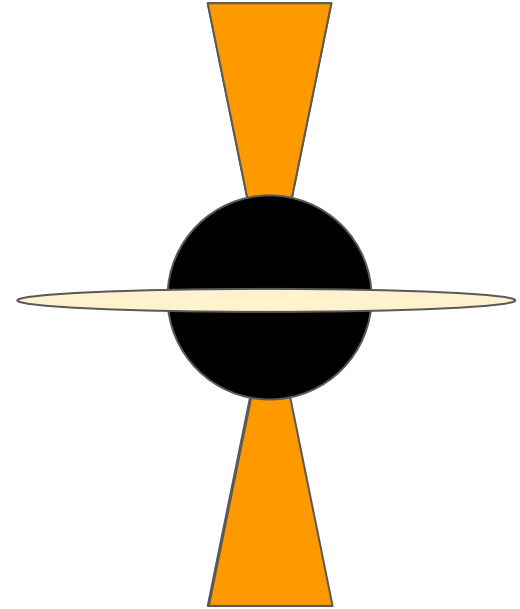
- Accreting WD jets non-relativistic $L_{\text{jet}} \sim \frac{1}{R}$
- WD + PBH



[VT, 2017]

PBH-WD laboratories: microquasars

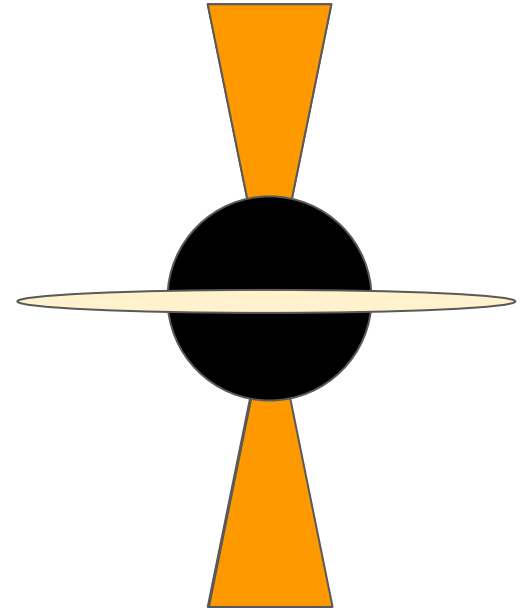
- Accreting WD jets non-relativistic $L_{\text{jet}} \sim \frac{1}{R}$
- WD + PBH $\rightarrow$ solar-mass BH accretor
 - radius $\downarrow$, luminosity $\uparrow$



[VT, 2017]

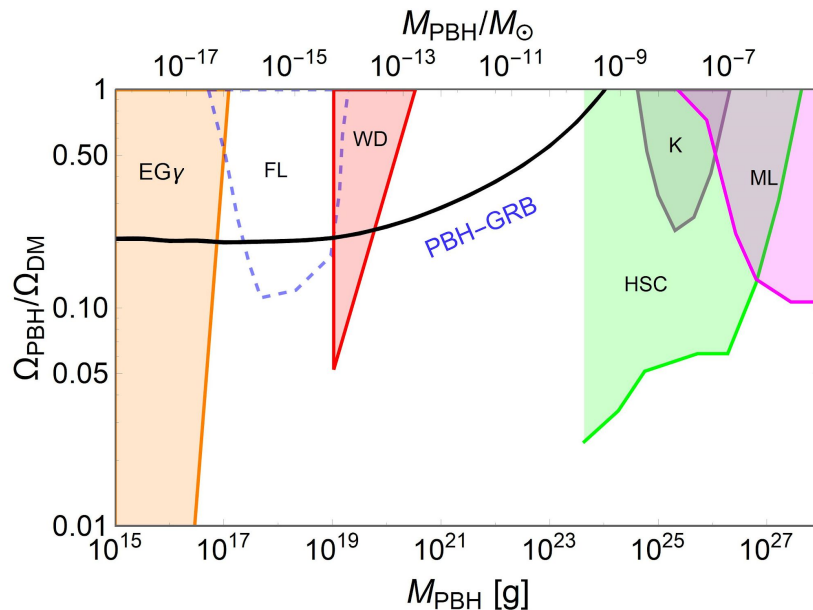
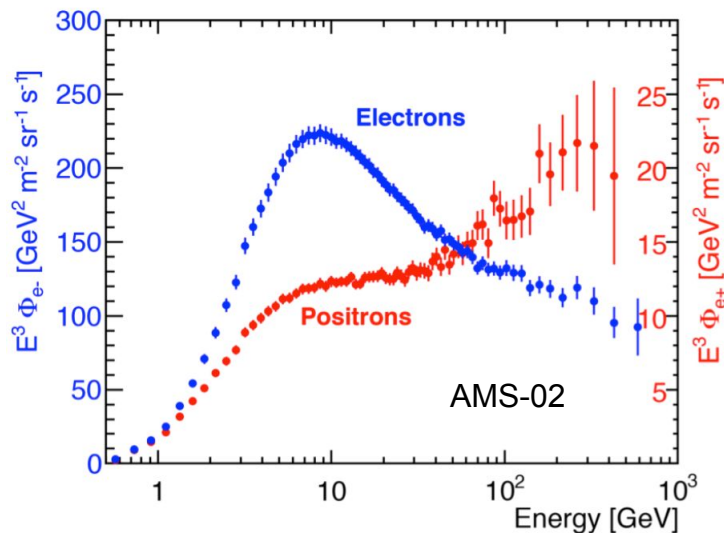
PBH-WD laboratories: microquasars

- Accreting WD jets non-relativistic $L_{\text{jet}} \sim \frac{1}{R}$
- WD + PBH $\rightarrow$ solar-mass BH accretor
 - radius $\downarrow$, luminosity $\uparrow$
- Continuous relativistic jet microquasar



[VT, 2017]

PBH-NS/WD laboratories: jet positrons



- High-E positrons from PBH-GRB jets can contribute to excess
→ astro-DM connection

Where do we stand?

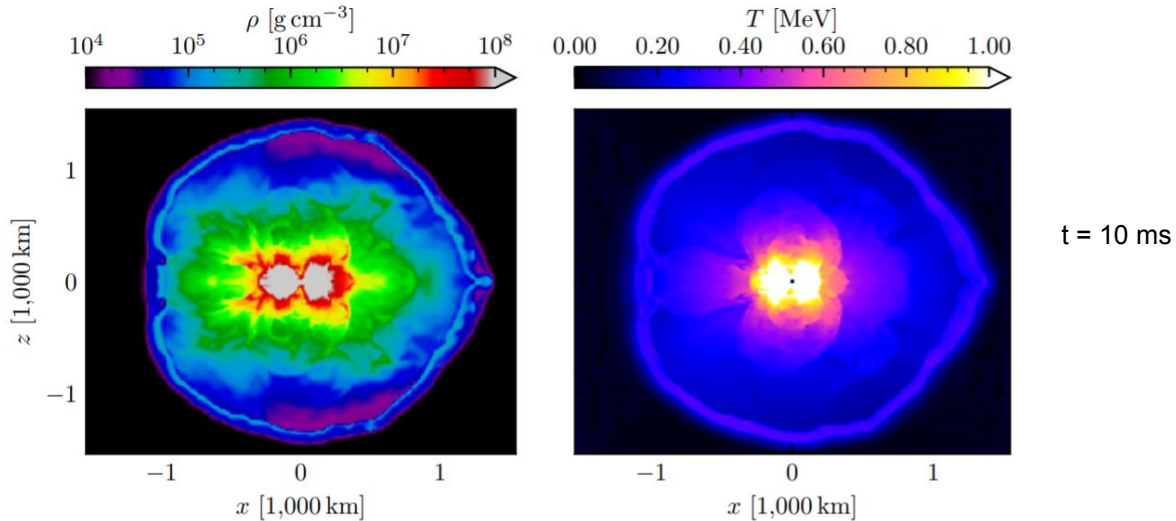
- Renaissance era in PBH research
 - strong synergy with emerging field of multi-messenger astronomy
- Simple general formation mechanism: scalar field fragmentation
- Compact stars as PBH laboratories
 - probe open parameter space
 - many novel astrophysical signatures
 - possible solutions to astronomy puzzles

Another Excitement

**Positrons and 511 keV
emission as tracers of
neutron star mergers**

[Fuller, Kusenko, Radice, VT, 2018]

- After merger, expanding ejecta heated by nuclear processes to $O(\text{MeV})$
→ thermal equilibrium population of e^+e^-



- Some positrons escape, annihilate → 511 keV signal

- A new multi-messenger signal component of ALL NS-NS/NS-BH mergers
→ **new tracer of merger history**
- Rates match observation of 511 keV signal from GC ✓
- 511 keV signal discovered in Reticulum II [Siebert+, 2016] → **confirms connection**
- **Doesn't rely on "new" hypothetical physics !!!**