



DAWN

*The Epoch of Reionization
with the Prime Focus Spectrograph*

Sune Toft,
Director
Cosmic Dawn Center (DAWN)



Danmarks
Grundforskningsfond
Danish National
Research Foundation



UNIVERSITY OF
COPENHAGEN



- 
1. *The Cosmic DAWN Center*
 2. *The Cosmic DAWN Survey*
 3. *COSMOS2019*

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DAWN – A new International research center

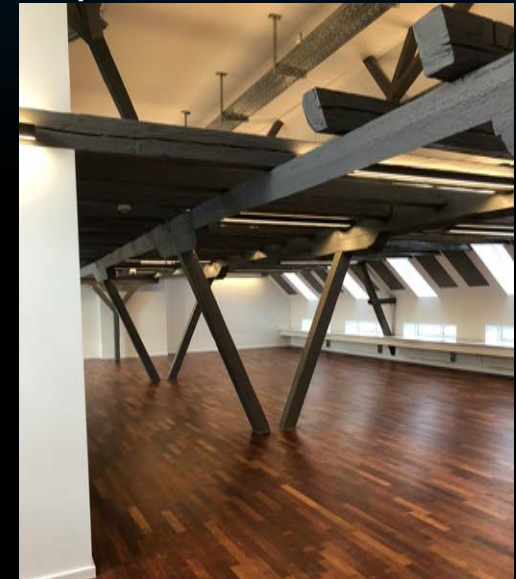
April signing



June opening

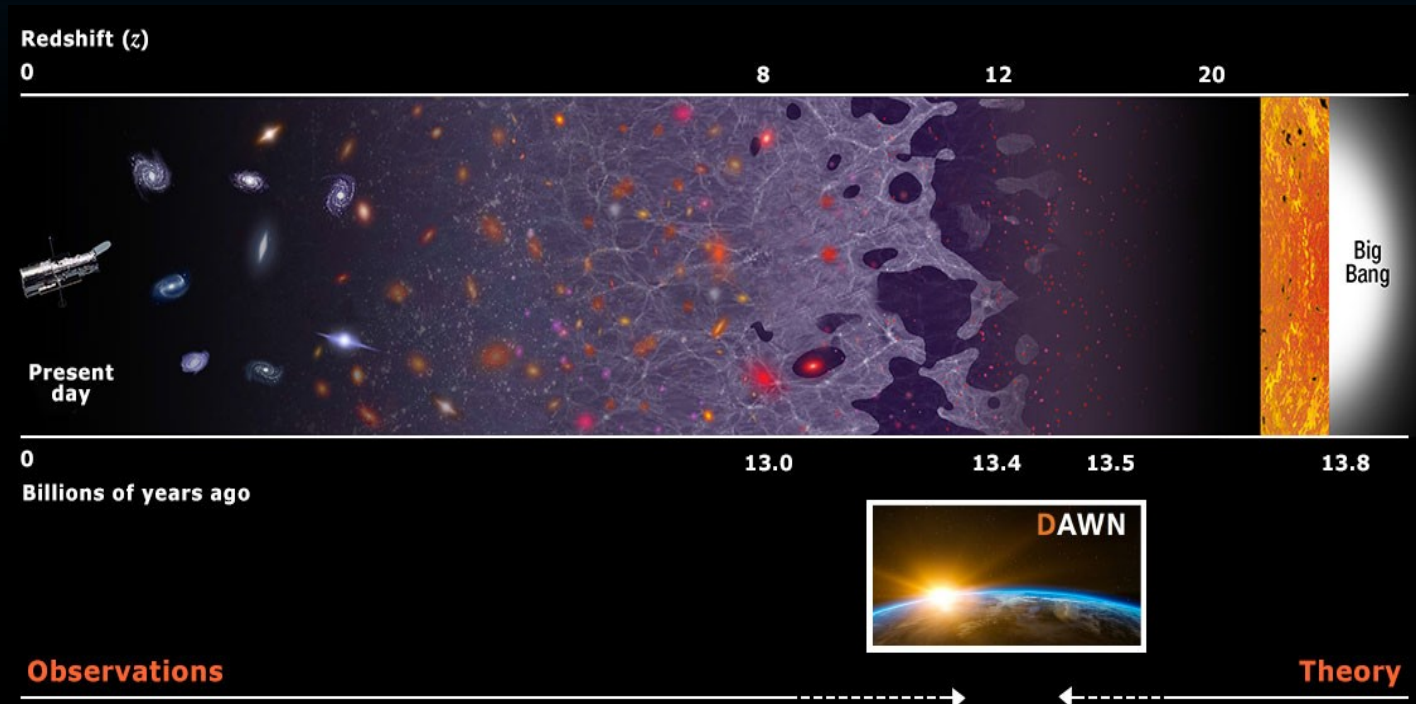


Sep/Oct move-in



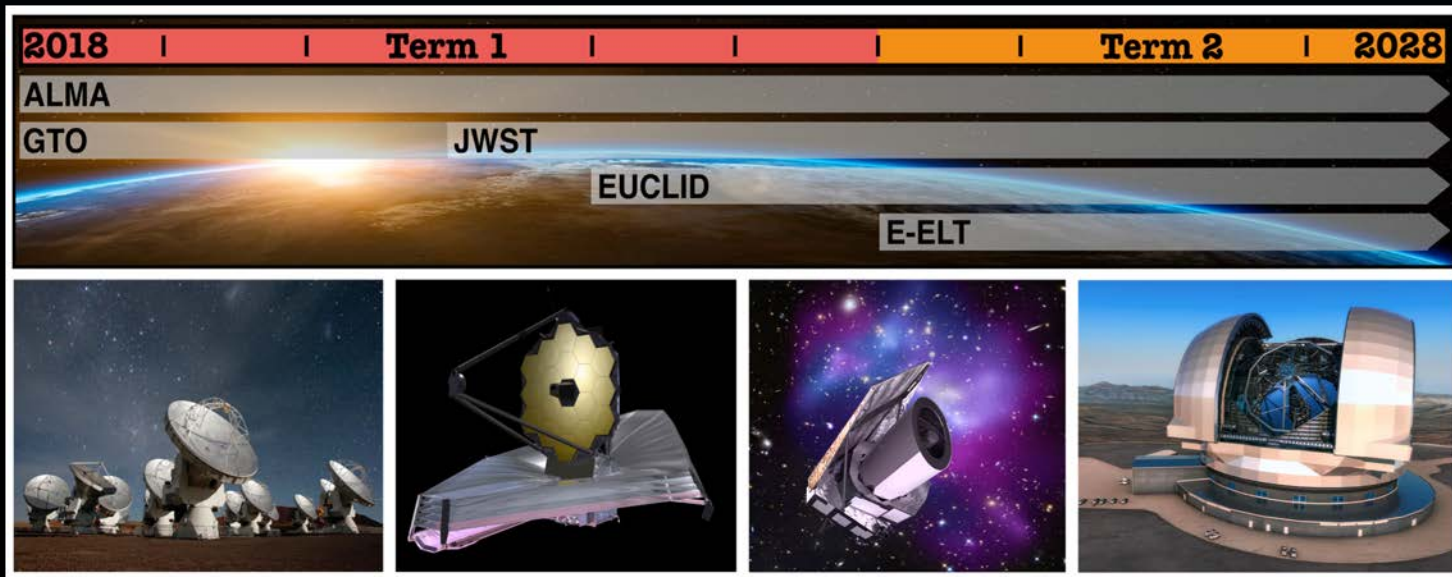
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At the Niels Bohr Institute



Gabe Brammer



Darach Watson



Sune Toft



Thomas Greve



Georgios Magdis



Peter Jakobsen



Charles Steinhardt



Allan Hornstrup



Hans Ulrik
Nørregaard Nielsen

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International Associates



Peter Capak
(IPAC/Caltech)



Kate Whitaker
(Uconn)



Fabian Walter
(MPIA)



Karina Caputi
(Kapteyn/Groningen)



Pascal Oesch
(Geneva)



Claudia Lagos
(U. Western Australia)



Desika Narayanan
(U. Florida)



Kristian Finlator
(New Mexico State U.)

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DAWN Fellowship and PhD Programs

2018 Fellows



Daniel Ceverino



Kimihiko Nakajima

2018 PhD students



John Weaver



Vasilii Kokorev



Meghana Killi

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International Summer Research Program

10-15 students from the US, 11 weeks over the summer

-Caltech SURF@DAWN program
-NSF/IRES grant (\$300K)



International Research Experience for Students
DAWN-IRES Scholars Program

Summer Research in Astrophysics in Copenhagen, Denmark

We welcome applicants for the inaugural DAWN Scholars Program in Copenhagen, Denmark. The Cosmic Dawn Center spans the full observational and theoretical realm of "cosmic dawn", the time period within a few billion years of the Big Bang. Joining the vibrant research atmosphere, students will be provided a once in a lifetime opportunity to perform research relevant to current open questions in astrophysics.

When and where does the program run? This 11 week internship runs at the Cosmic Dawn Center in Copenhagen Denmark June through mid-August 2019. The program provides a \$6500 stipend and travel to/from Denmark and to the annual American Astronomical Society meeting in January 2020.

Who is eligible to apply? Because our program is NSF funded, we can only accept students who are US citizens or greencard holders. Both undergrad and grad students within their first two years of a masters or PhD program are eligible. We strongly encourage women and members of underrepresented groups to apply.

How do I apply? Applications should be completed through the online form available at <https://dawn.uconn.edu/apply>.

APPLICATION DEADLINE: FEB 1, 2019
dawn.uconn.edu/apply



Tivoli Gardens (right) is the second oldest operational amusement park located in the heart of the historic city of Copenhagen, the bike capital of the world. Nyhavn port (bottom, left) and Copenhagen more generally is big enough to be a metropolis with shopping, culture and nightlife, yet small enough to be intimate, safe and easy to navigate.

*Come join us as a **DAWN Scholar** to explore the rich uncharted territory of the **first stars, galaxies and black holes** in our Universe!*



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Scientific Meetings

June 2018

- DAWN summit I (30)
- CPHDUST2018 (132)
- COSMOS team meeting (100)
- Ultra Vista Team Meeting (10)
- Hawaii-Two-0 Team meeting (10)

October 2018

- JWST/NIRSpec GTO meeting (23)
- Quenching workshop (10)
- Buffalo survey team meeting (40)
- XXM Time allocation committee (5)



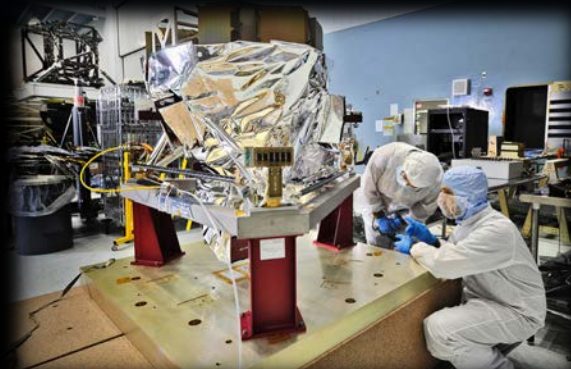
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JWST Guaranteed time observations teams (>1000h)

Near Infrared Spectrograph

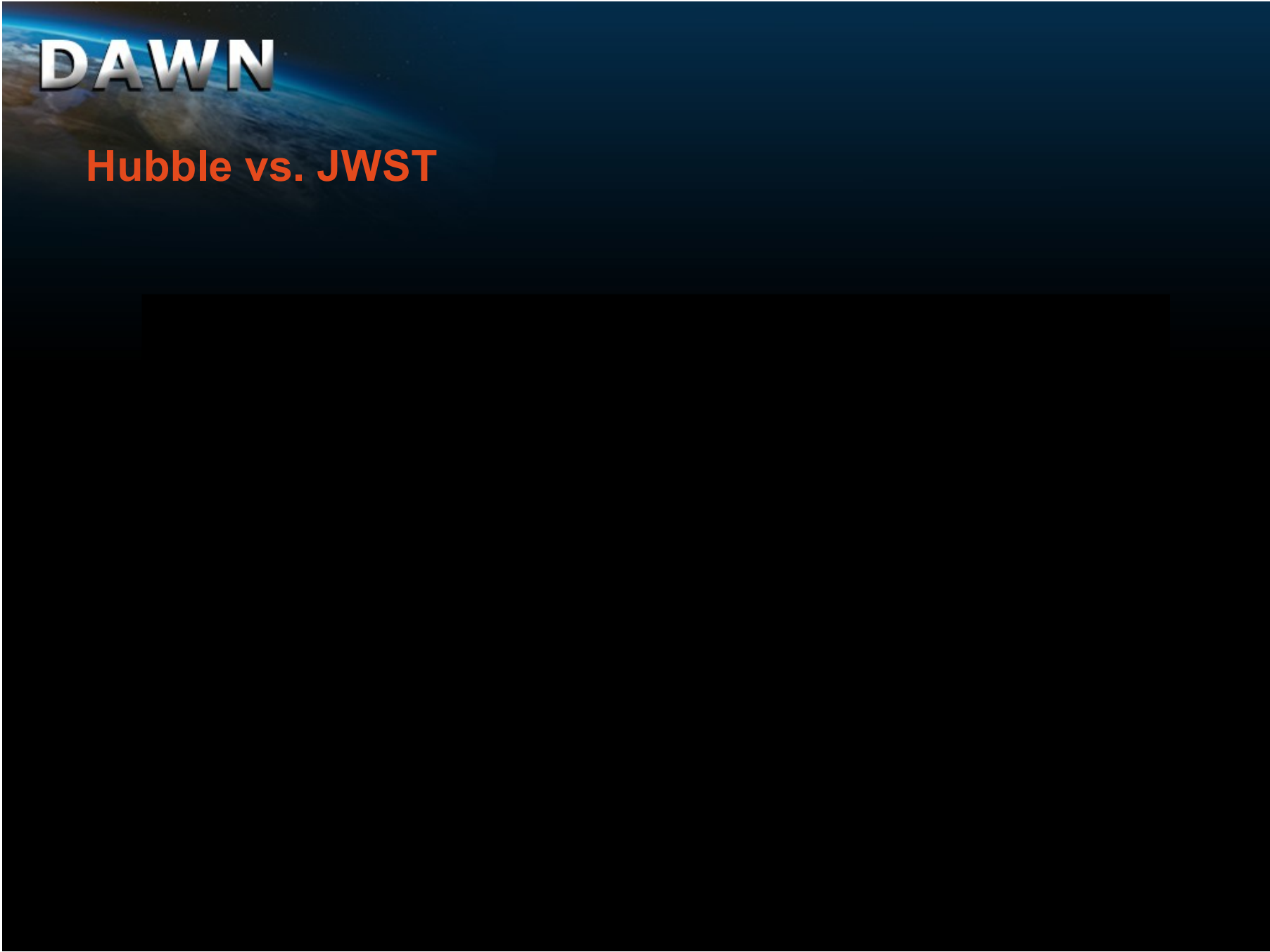


Mid Infrared Instrument



Near Infrared Imager and Slitless Spectrograph



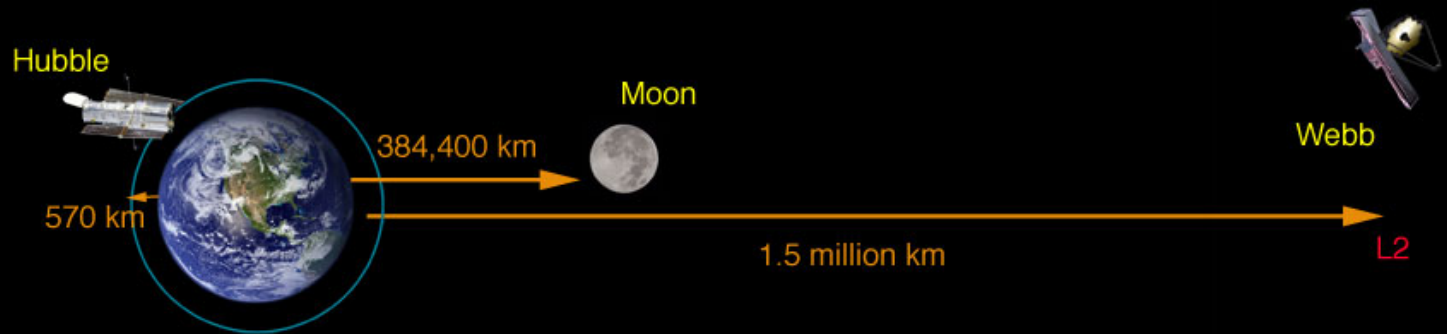
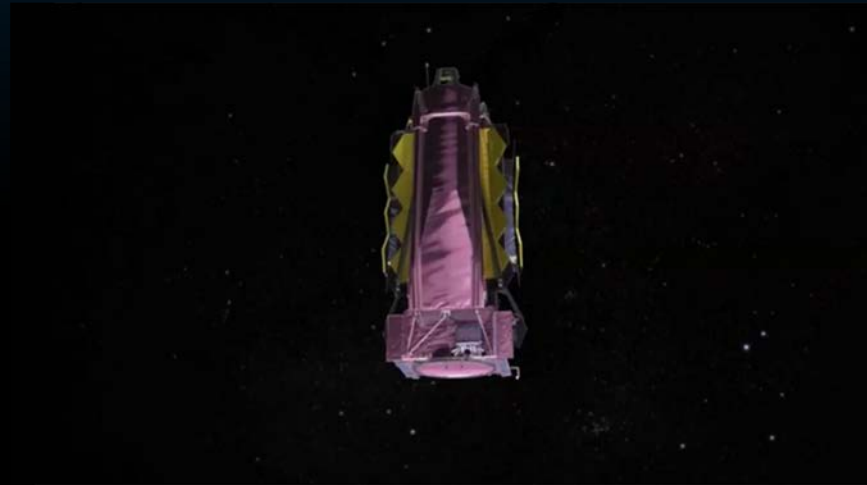


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Hubble vs. JWST

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Hubble vs. JWST



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Finding the first galaxies



Pascal Oesch
(geneva)

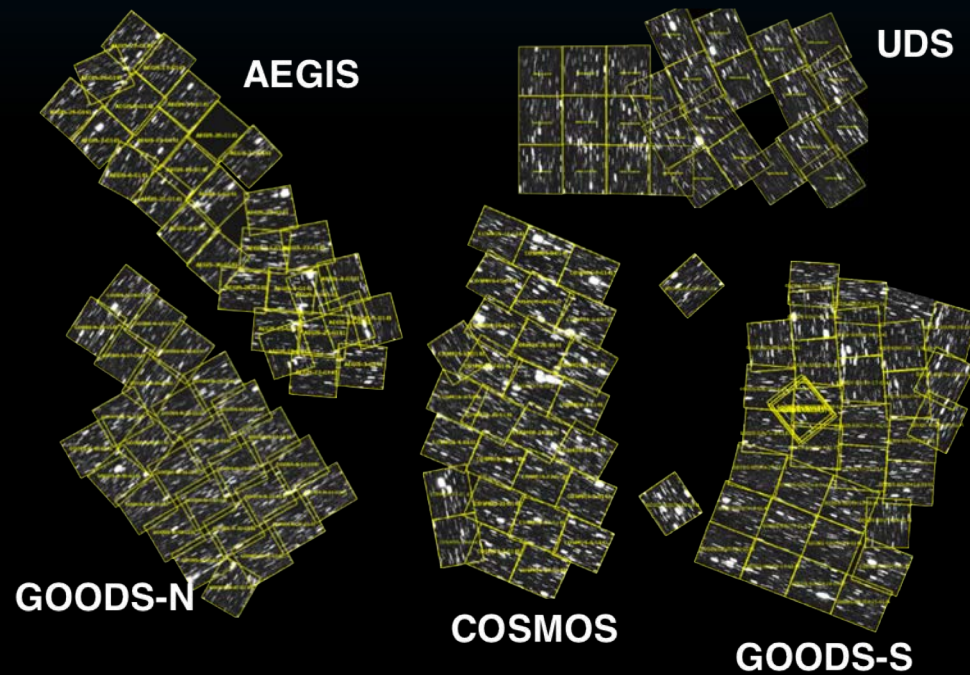
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Grizzly Hubble Archive Project

Total Area 0.25 deg²



Gabriel
Brammer



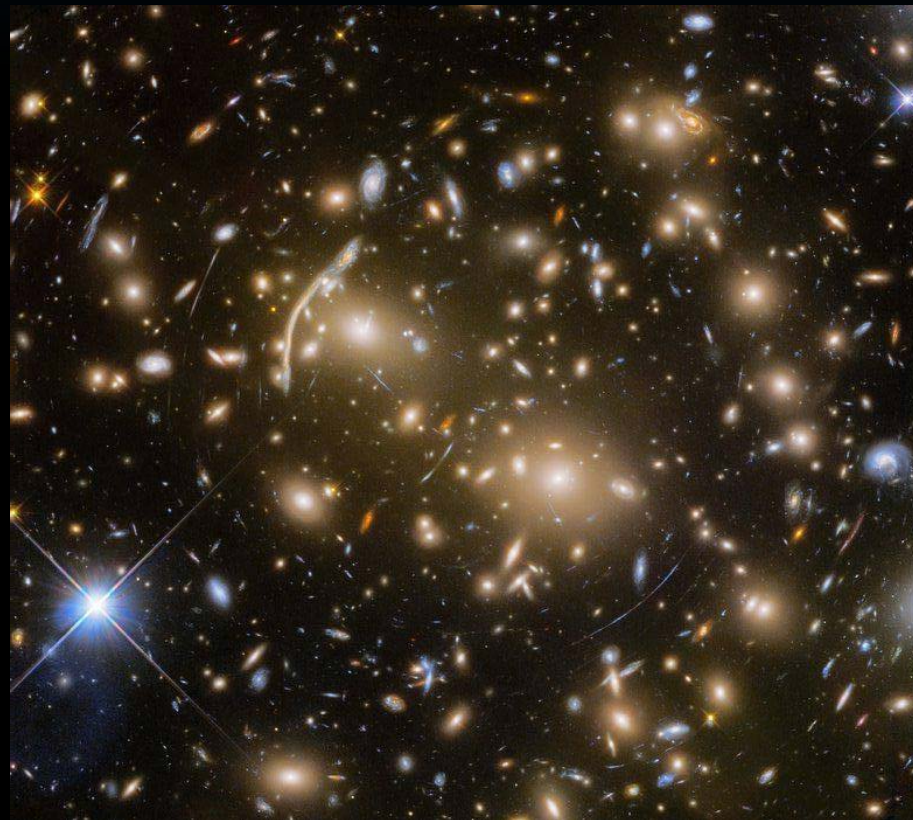
>5 times larger area in Hubble archive never
systematically explored

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**BUFFALO – New Hubble program (2018-19):
Beyond Ultra Deep Fields and Legacy observations
(100 orbits)**



Charles
Steinhardt
(CoI)

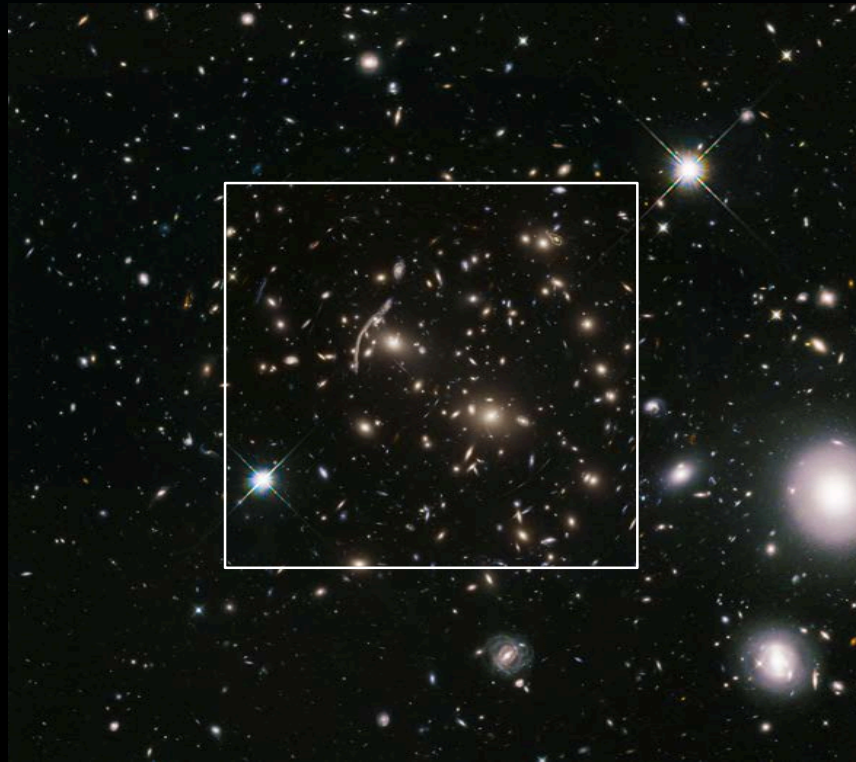


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BUFFALO –Increasing discovery space for Cosmic Dawn Galaxies with new sightlines

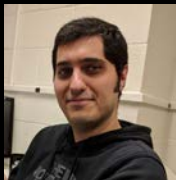
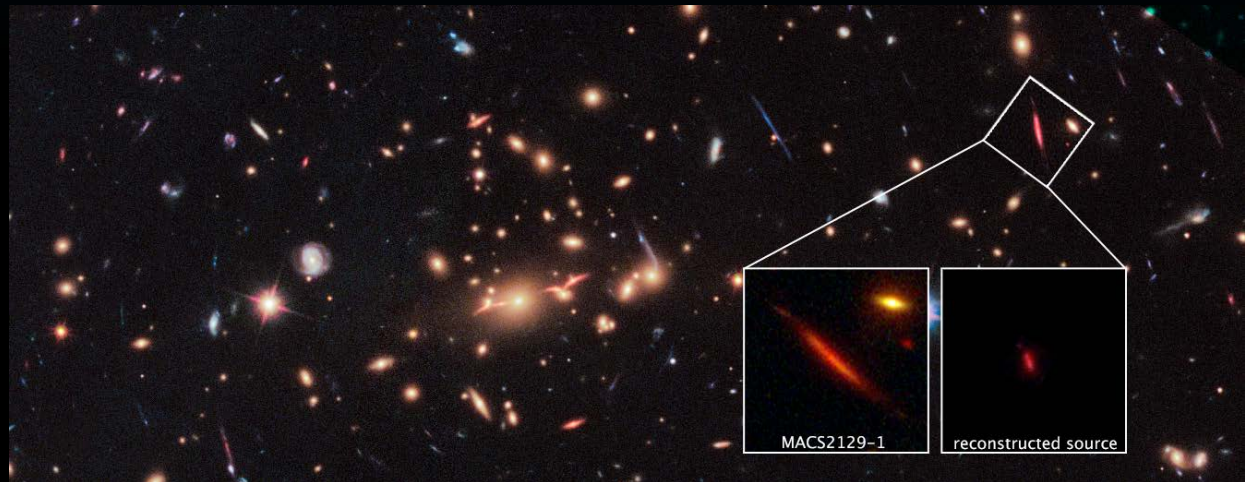


Charles
Steinhardt
(CoI)



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Requiem – New Hubble Grism Program (2019) to resolve the earliest dead galaxies (60 orbits)



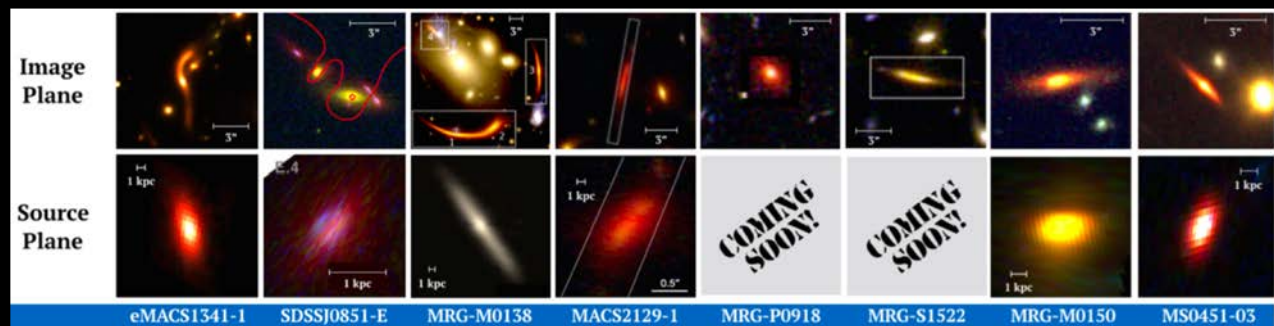
Mohammed Akhshik



Kate Whitaker
(CoI)

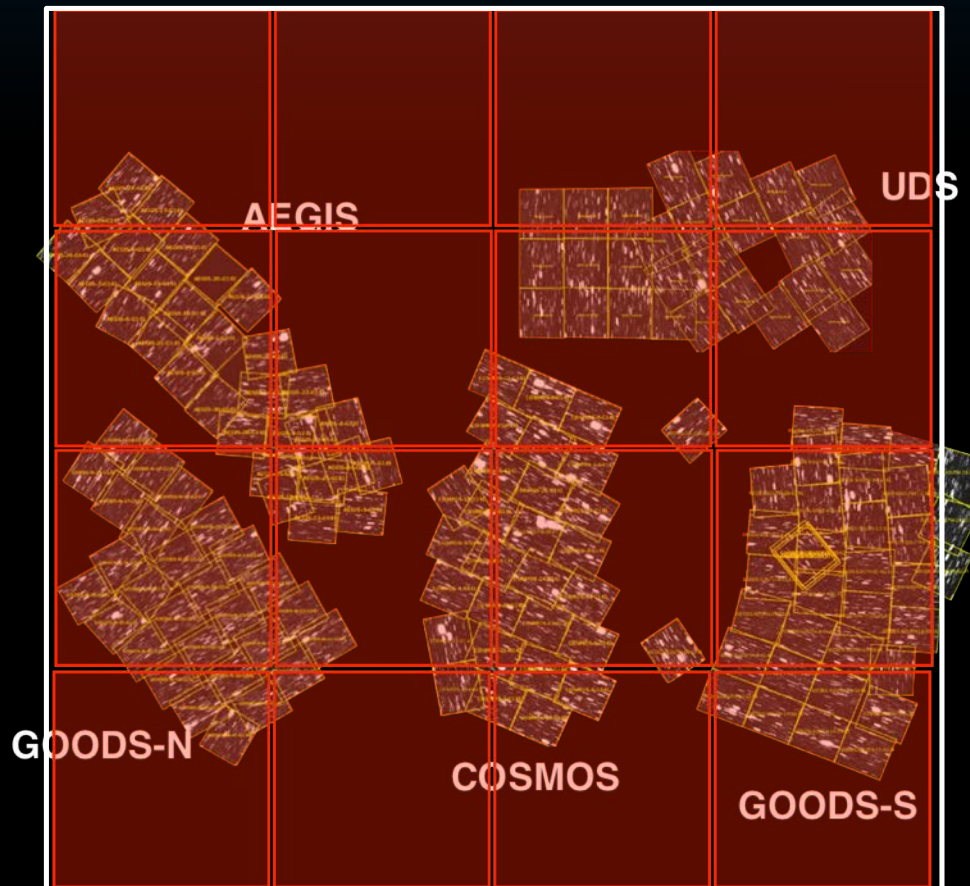


Gabriel Brammer
(CoI)



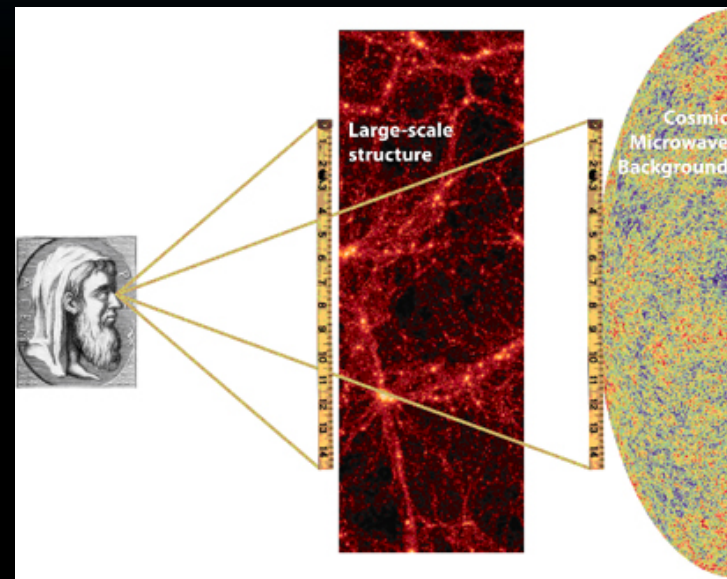
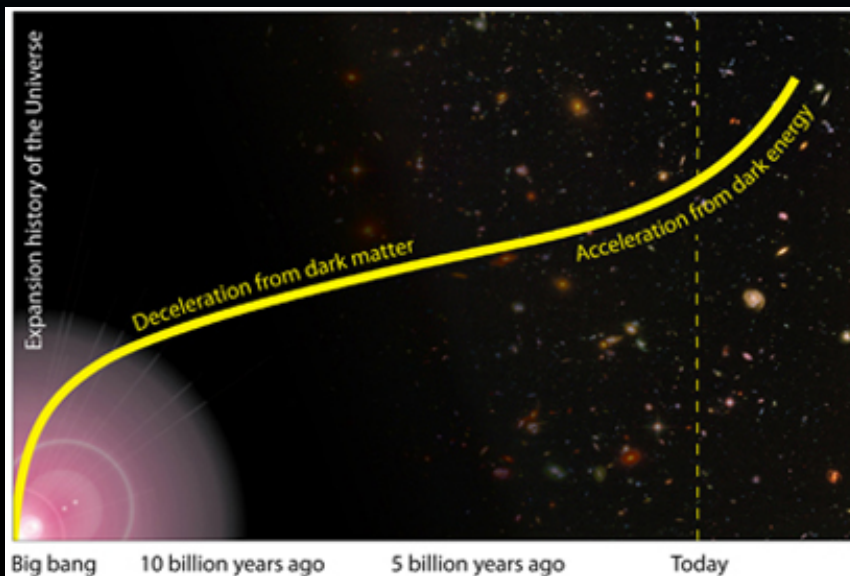
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Euclid ~ Hubble with much wider field of view



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Euclid: Understanding Dark Matter and Dark Energy



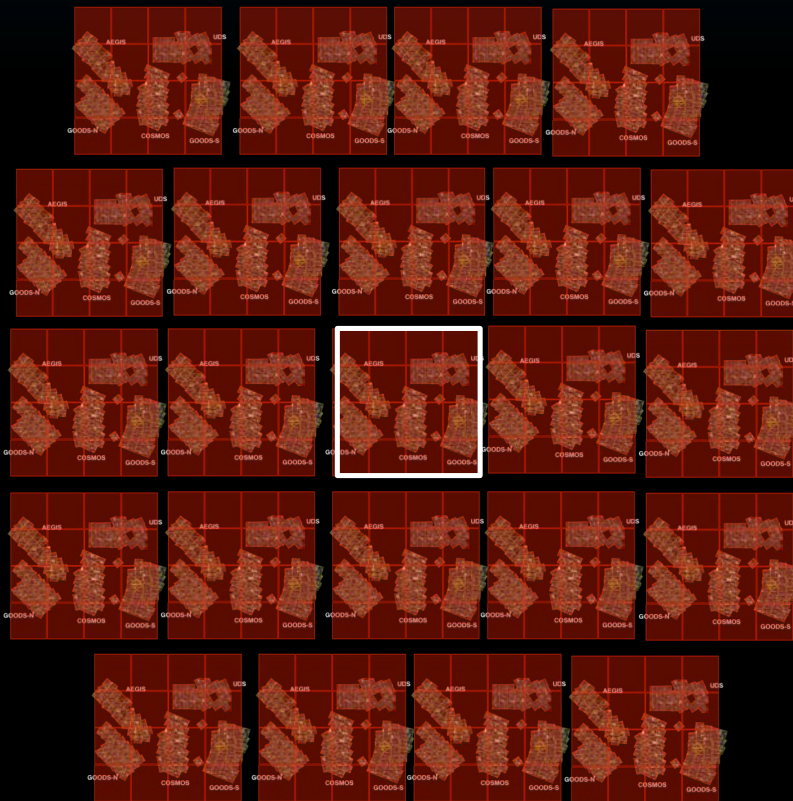
Mapping the cosmic web with billions of galaxies back to Cosmic noon ($z=2$) will reveal nature of dark matter and dark energy through weak lensing and clustering



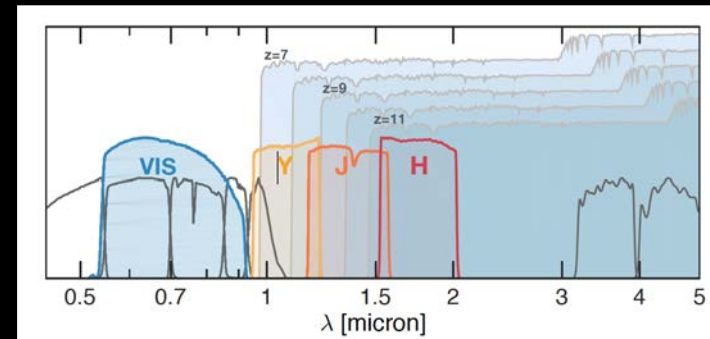
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Euclid Deep Fields – 40 deg²

4x



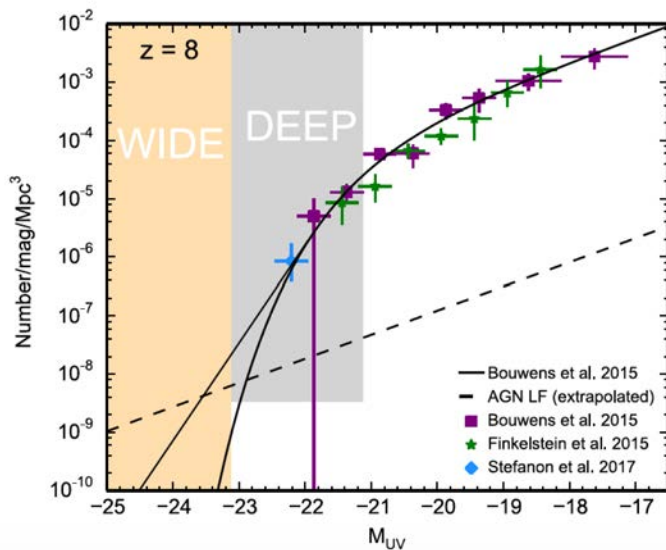
- North Ecliptic Pole (10 deg²)
- Chandra Deep Field South 10 deg²)
- Akari Deep Field South (20 deg², TBC)



High resolution Y+J+HK imaging to 26th mag. Essential for high redshift galaxy selection

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Thousands of bright $z > 8$ Galaxies

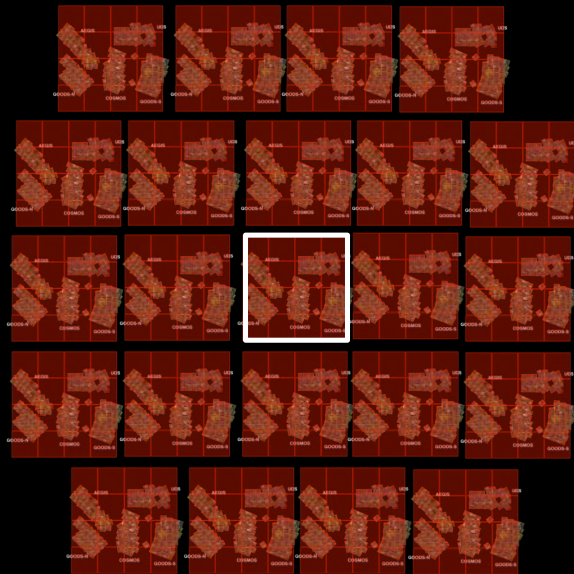


Redshift	LBG (Schechter LF)	LBG (DPL LF)	AGN
6	46000	57000	200
7	14000	14000	48
8	2300	1900	12
9	410	630	3
10	27	220	1
11	0-100??		

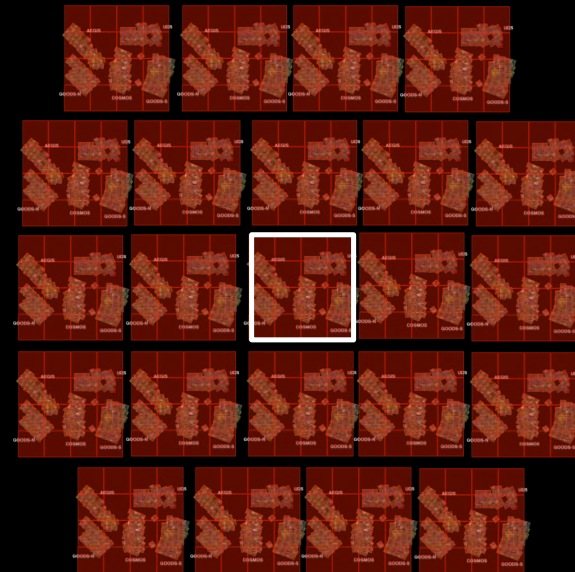
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20 Square degrees observable from Hawaii (NEP,CDFS)

North Ecliptic Pole (10 deg²)



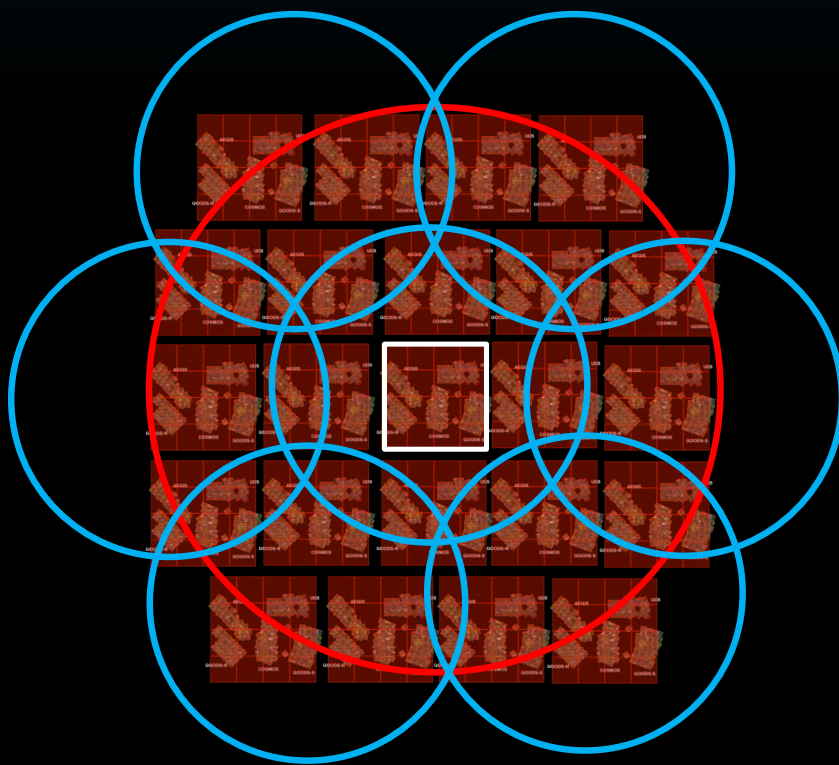
Chandra Deep Field South (10 deg²)



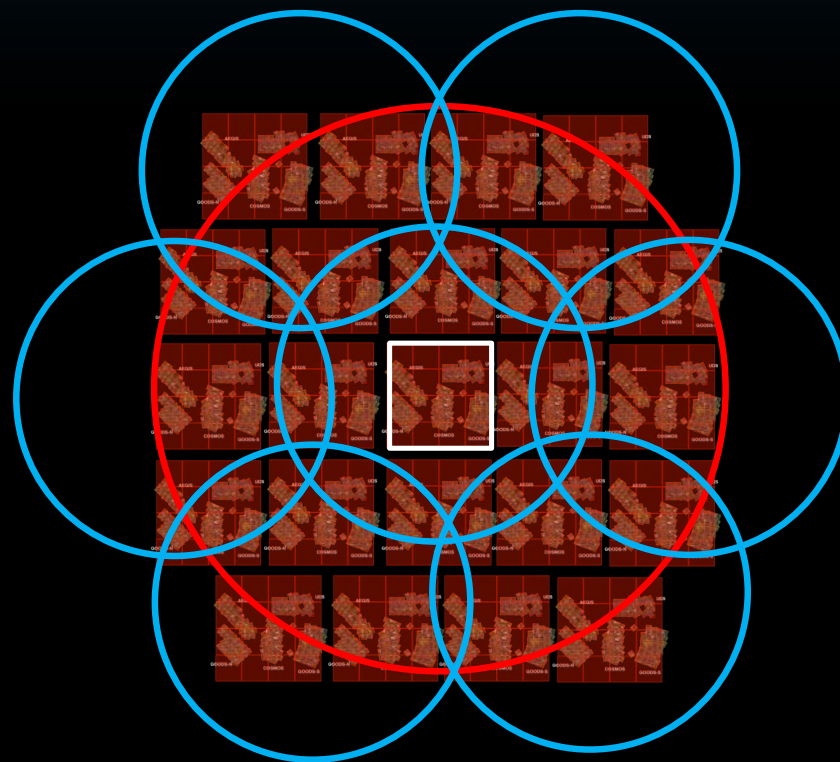
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Cosmic Dawn Survey (2018-2020)

North Ecliptic Pole (10 deg²)



Chandra Deep Field South (10 deg²)

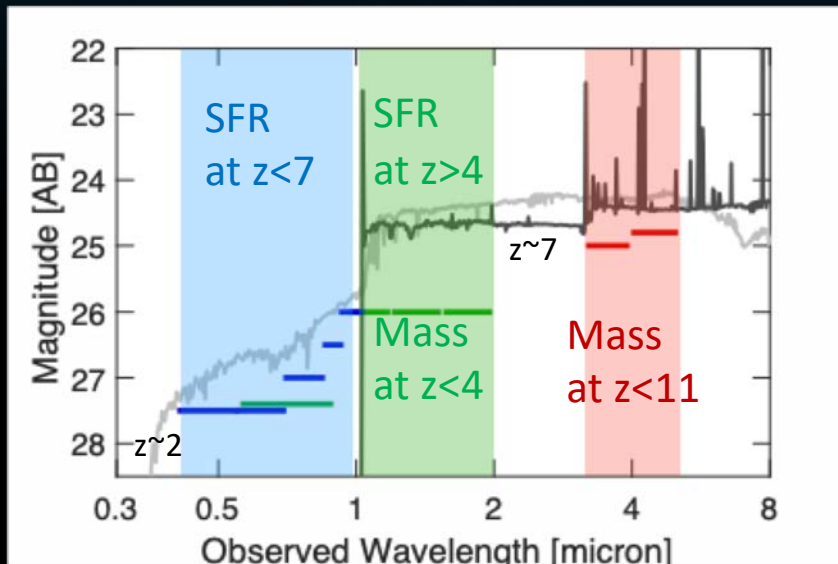


○ Subaru/HSC (30n, PI:Sanders)

○ Spitzer IRAC (6000h, PI:Capak)

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Cosmic Dawn Survey of Euclid Deep Fields (2019-2021)



Cosmic Dawn Survey will be 100 times larger than existing HST+Spitzer extragalactic surveys (CANDELS)

Unbiased sample with reliable photo-z, Stellar mass, SFR for all $M > M^*$ galaxies to $z=8$, and brighter galaxies to $z>10$

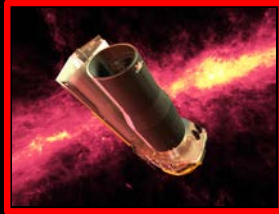
Full picture of large scale structure over 95% of the history of the Universe



Hawaii-Two-0
HSC (30n)
Keck (10n)



Euclid deep



Spitzer WFIRST
Euclid Legacy
survey (6000h)



LMT – LSS survey
1.1, 1.4, 2.1 mm
0.25 mJy RMS



5500 $7 < z < 8$
1000 $8 < z < 9$
200 $9 < z < 10$

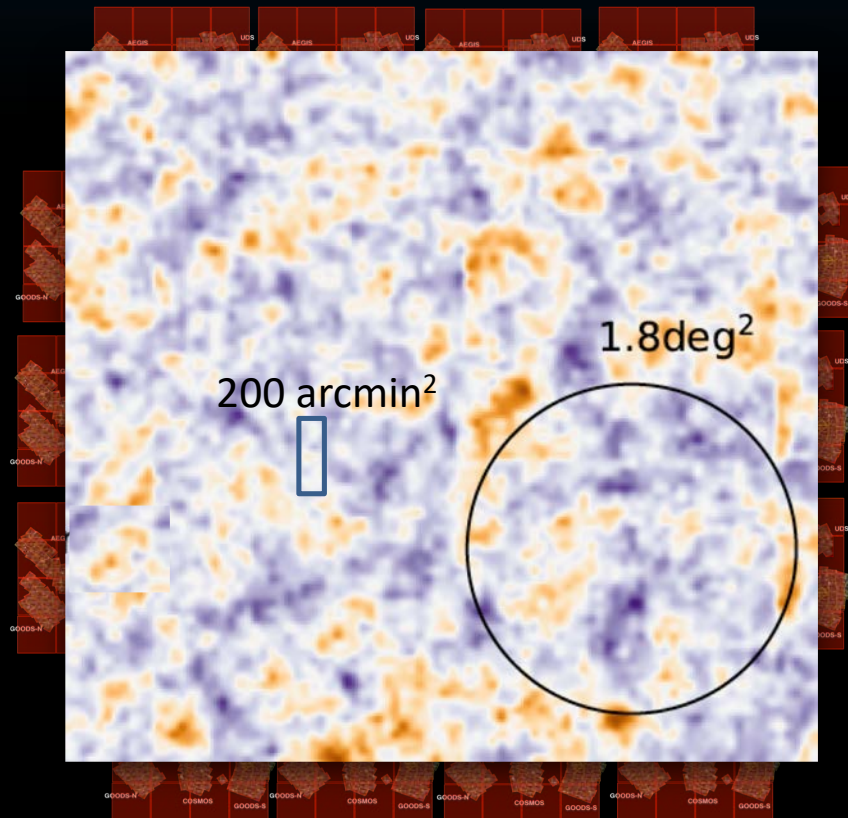
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Characterize large scale structure at Cosmic Dawn

$z=5$ dark matter density map (Springel+ 2005)

Large area needed to map out large scale structure to find and characterize highest density peaks of the Universe during Cosmic Dawn

These pinpoint the formation sites of the first galaxies



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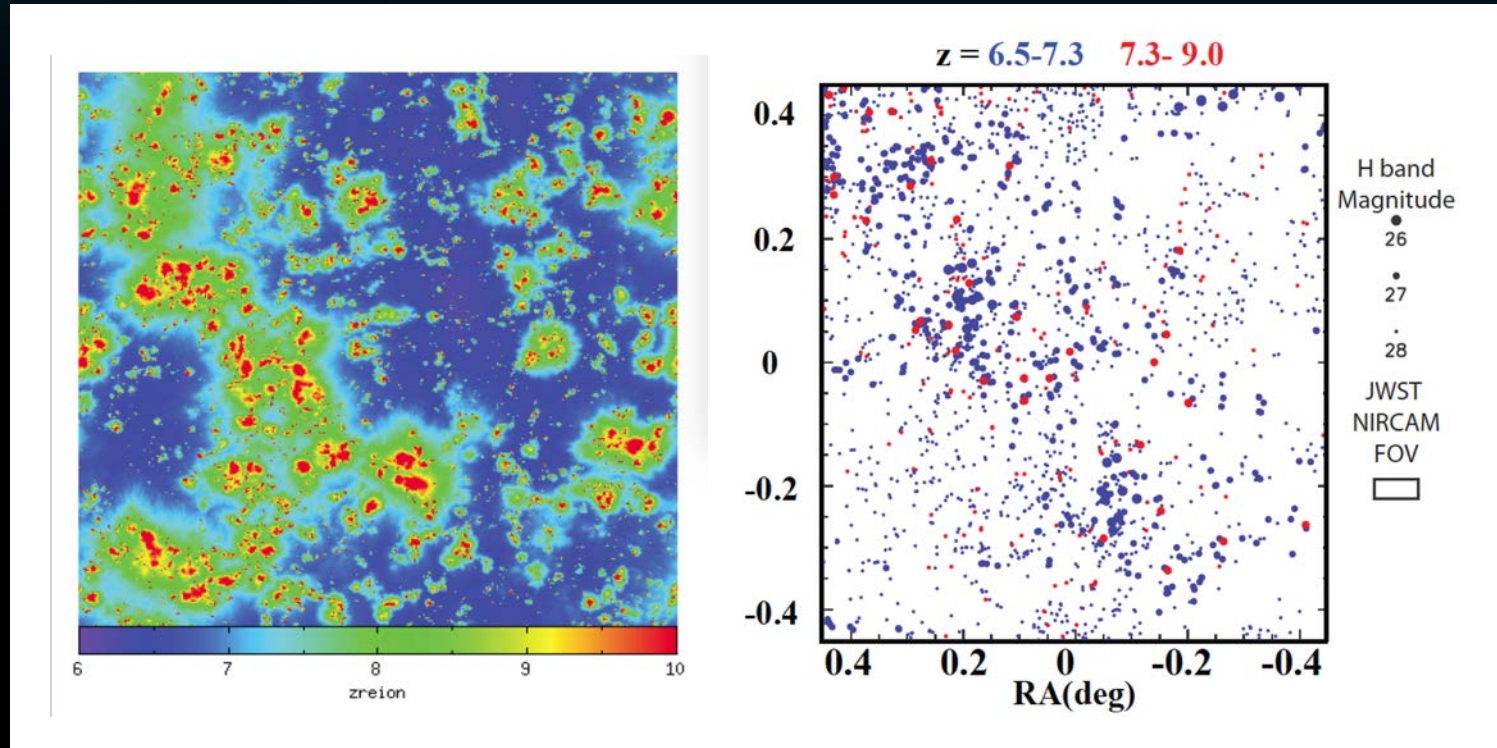
Reionization – the last phase transition of the universe



www.eso.org

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How did Reionization start and Proceed?



- Simulations suggest it happens on degree scales (Trac+2008)
- Starts at $z > 10$ in bubbles around the highest density peaks, traced well by the brightest galaxies

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ORIGINS – European Research Council Synergy Grant



Sune Toft
(DAWN)



David Sanders
(Hawaii)



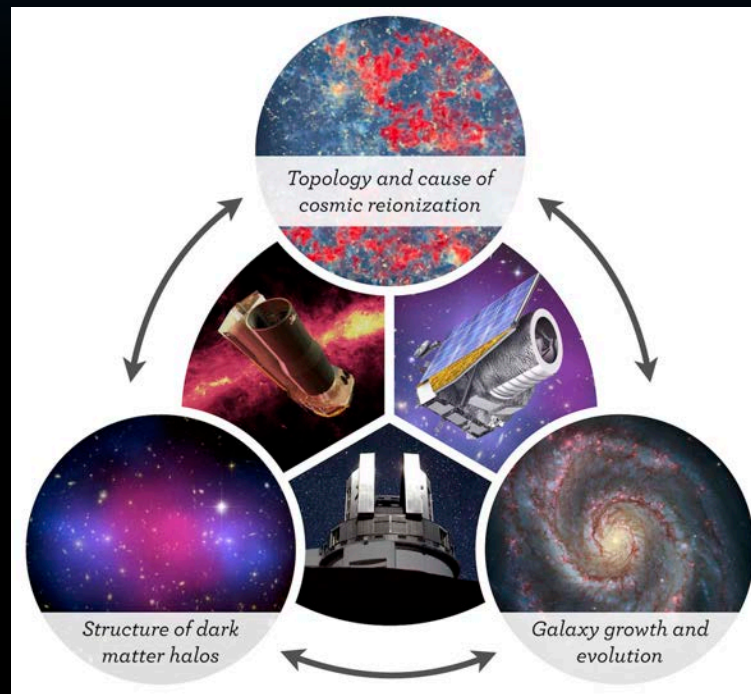
Peter Capak
(IPAC/Caltech)



Luigi Guzzo
(Milan)



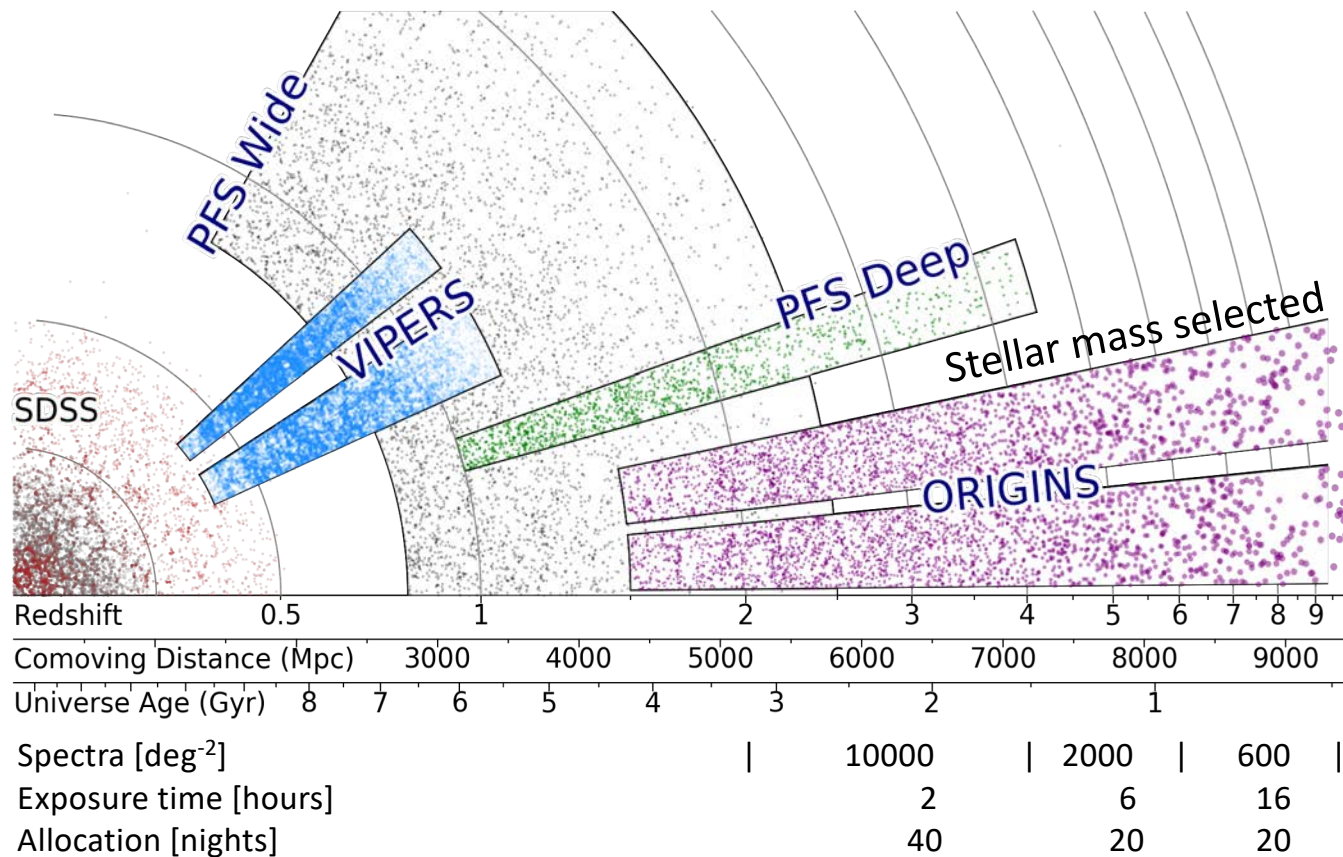
Pascal Oesch
(Geneva)



Proposed partnership between Japan, Hawaii and Europe to extend the PFS survey deep into Epoch of Reionization

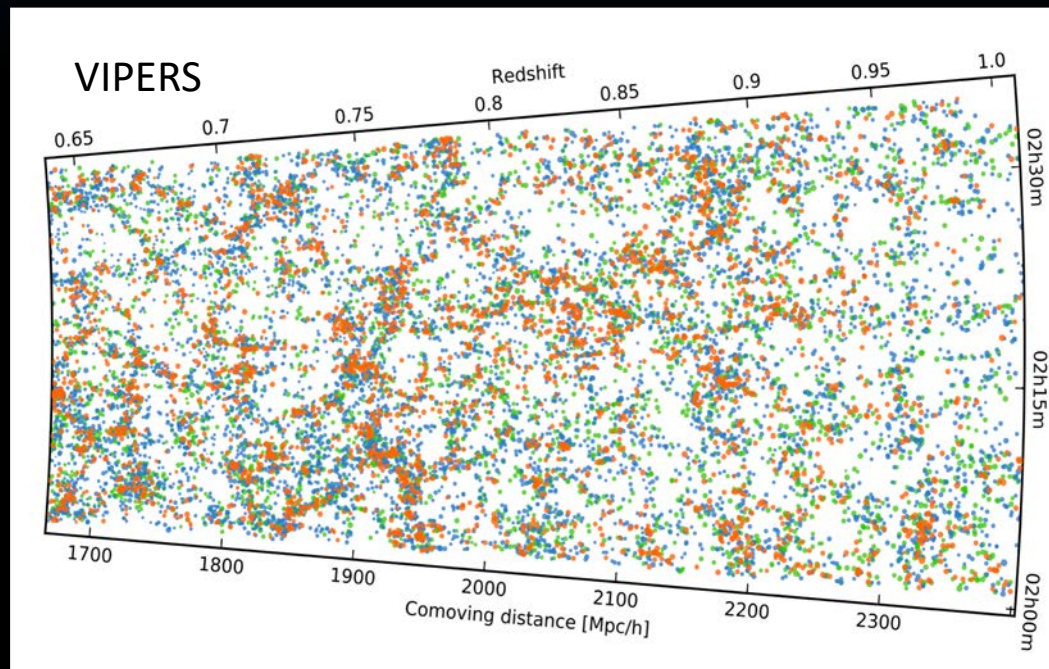
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Expand PFS survey to map LSS through EoR



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Characterize large scale structure through EoR

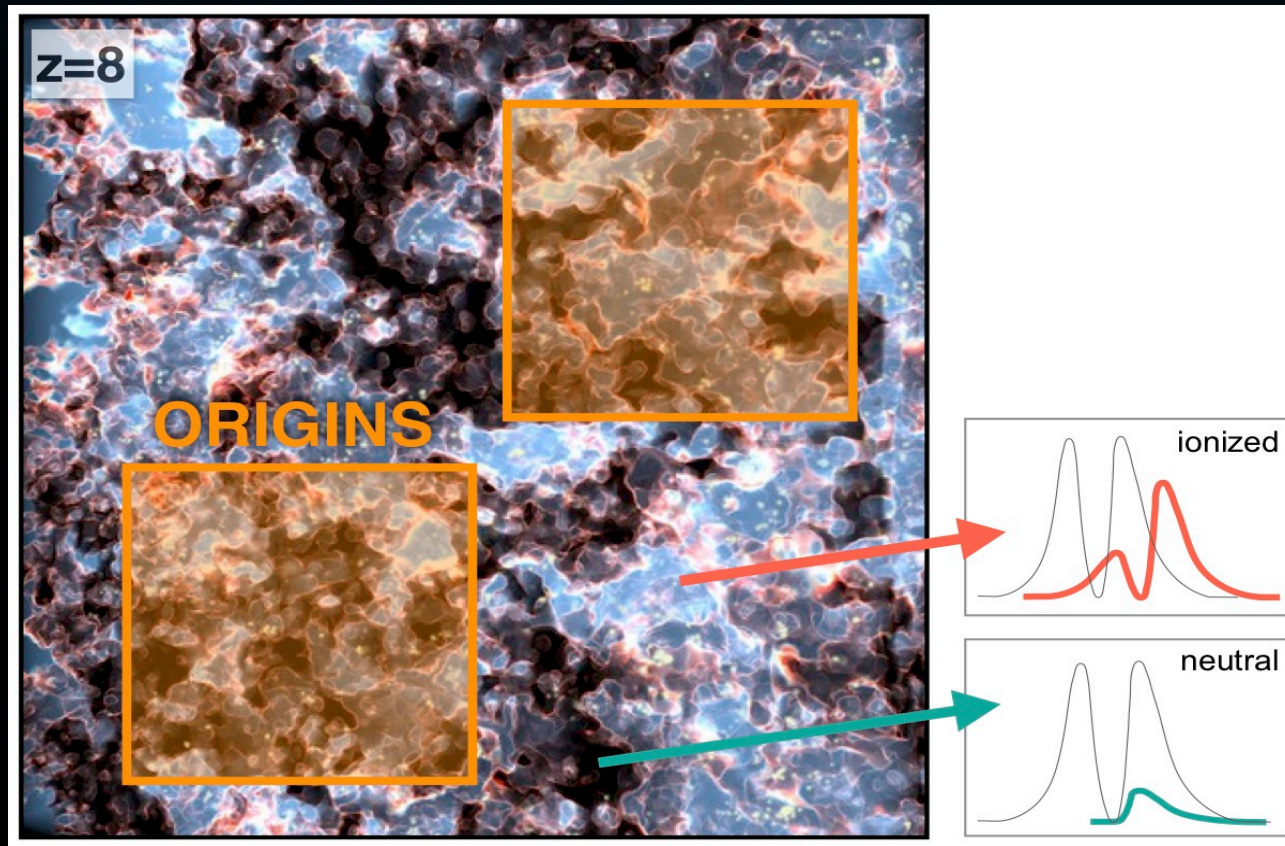


- Test of gravity and Cosmology to $z=4$ (clustering + Redshift space distortions)
- Co-evolution of Dark matter and galaxies to $z=10$ (clustering and abundance matching)

Unleash strong synergy between major Cosmology and galaxy evolution surveys of the next decade

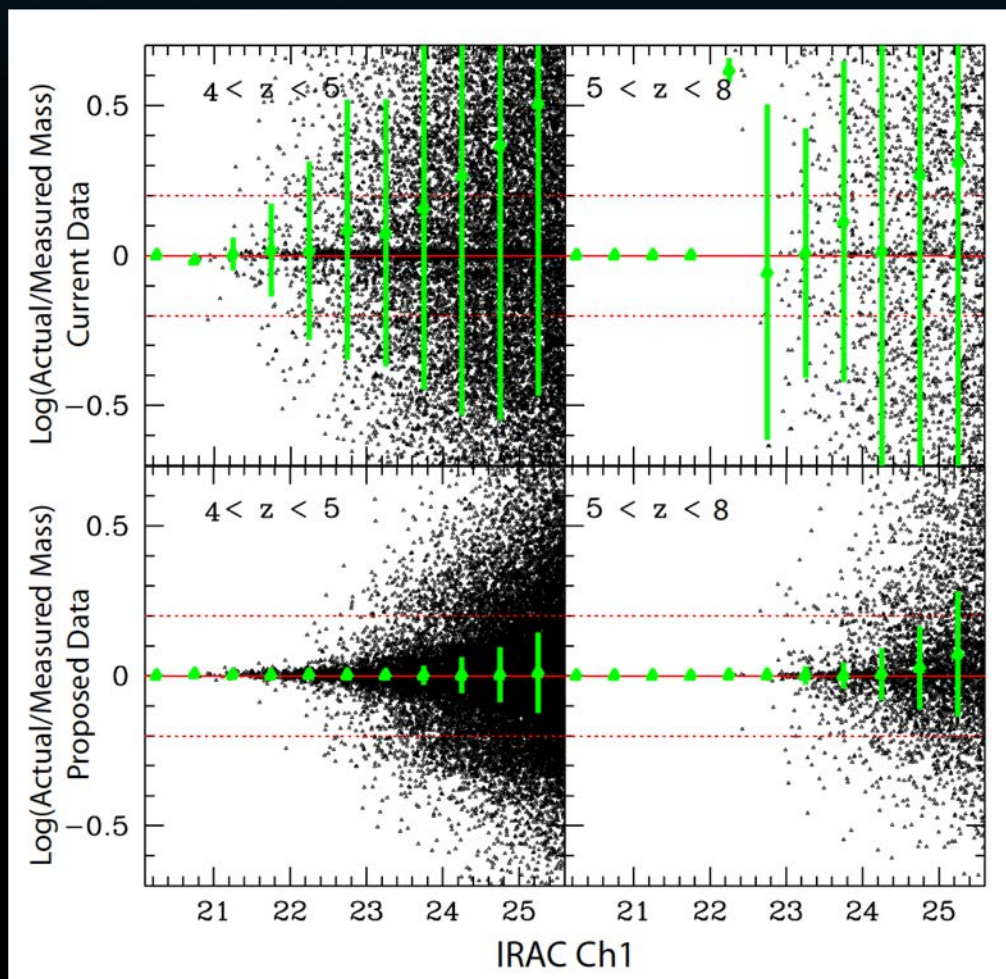
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Topology and sources of Cosmic Reionization



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Importance of Spitzer for stellar masses at $z > 4$

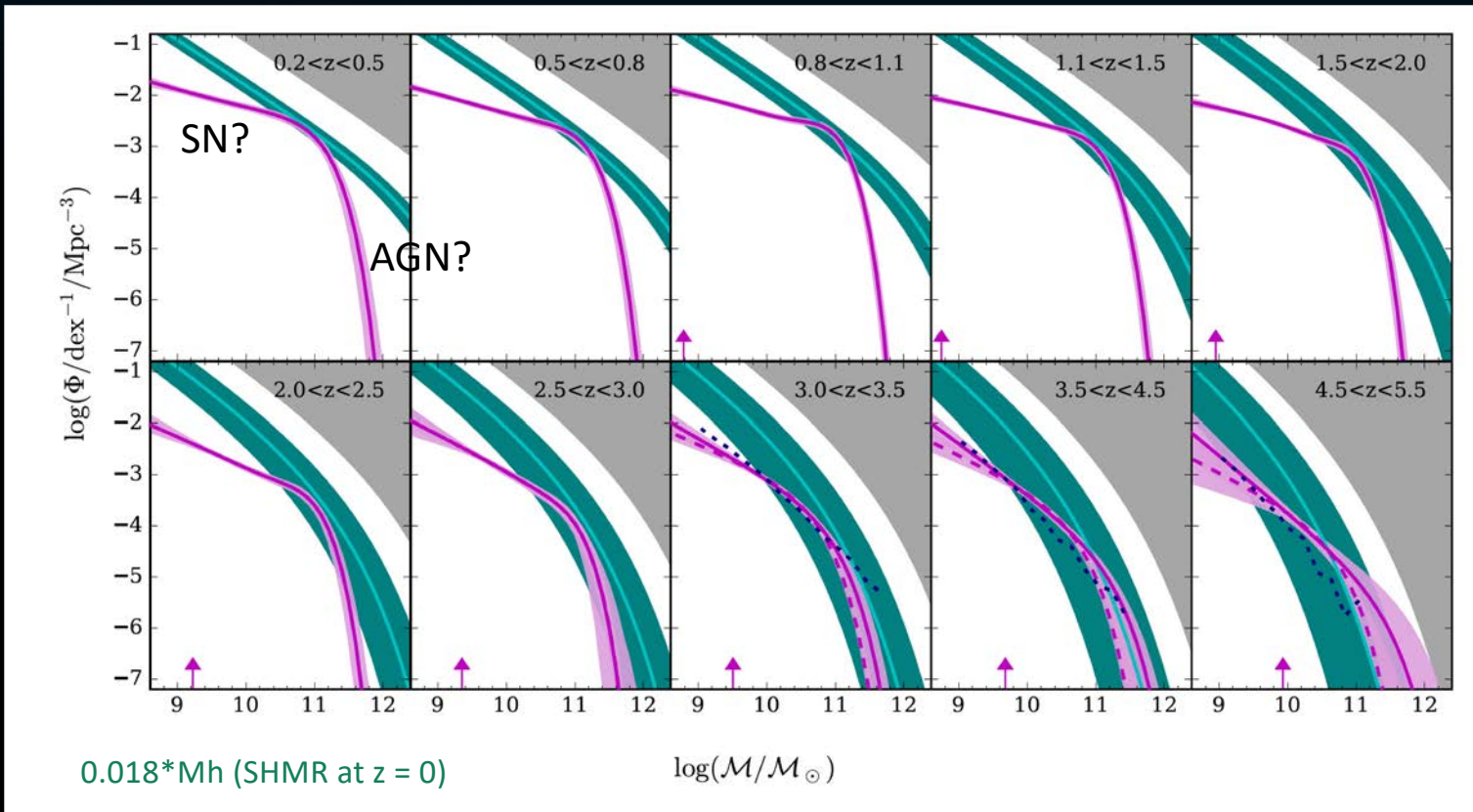


Euclid + HSC

Euclid + HSC + Spitzer

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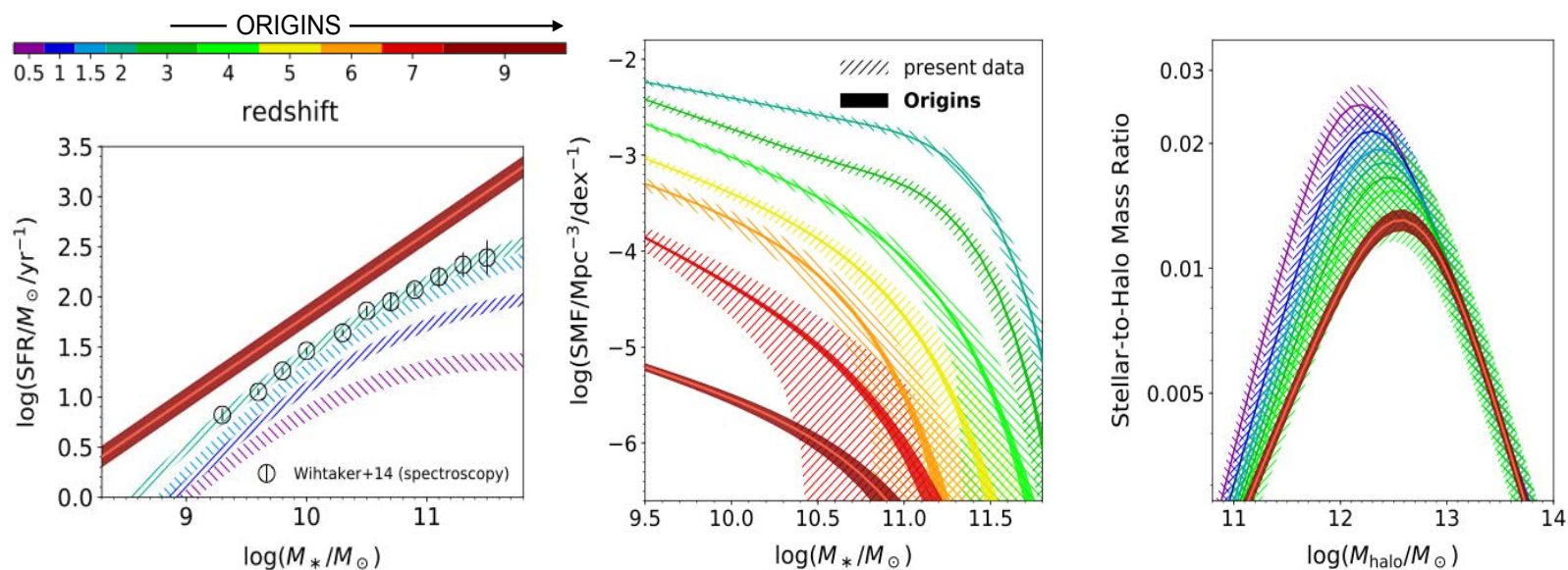
Probing the onset of Feedback Mechanisms



(Davidzon+2017)

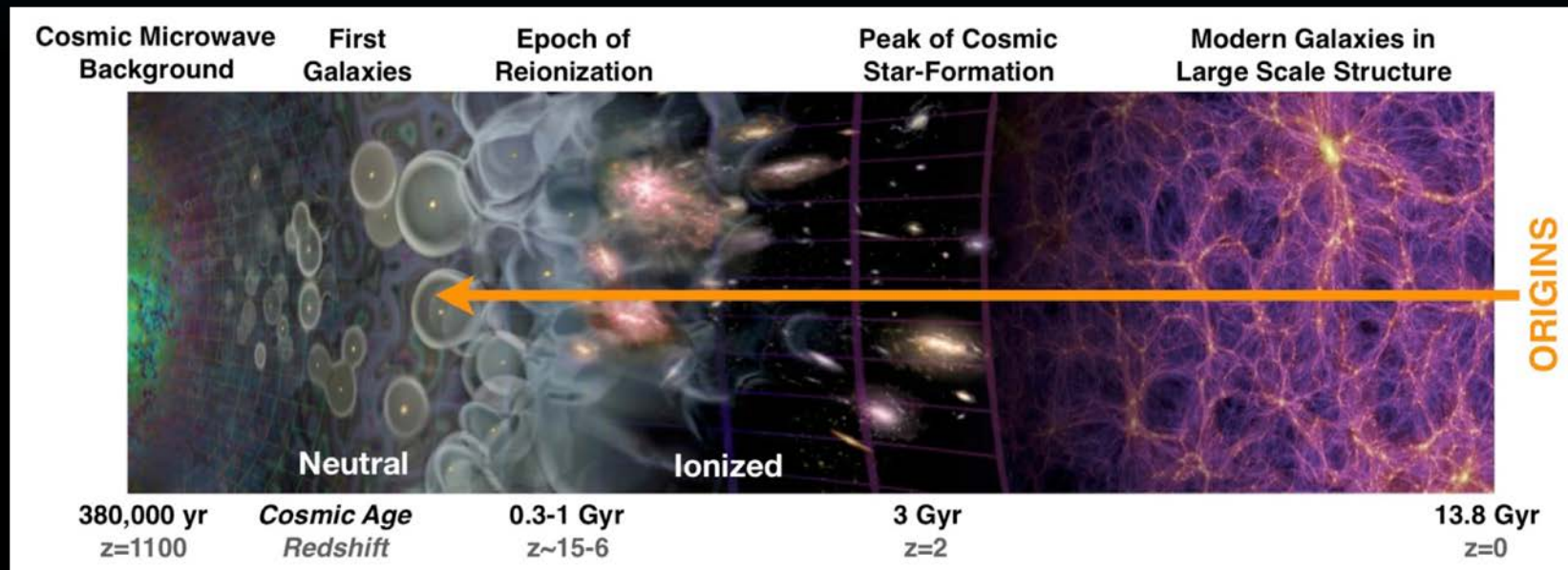
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Main sequence, stellar to halo mass function to $z=10$



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A holistic understand of the coevolution of galaxies and DM over 95% of Cosmic Time



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COSMOS2015

HSC:Hawaii+Japan



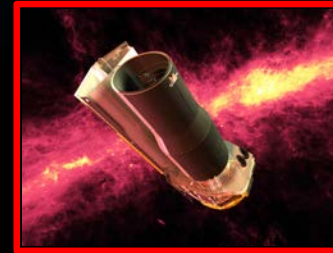
+

UltraVISTA DR2



+

Spitzer/SPLASH

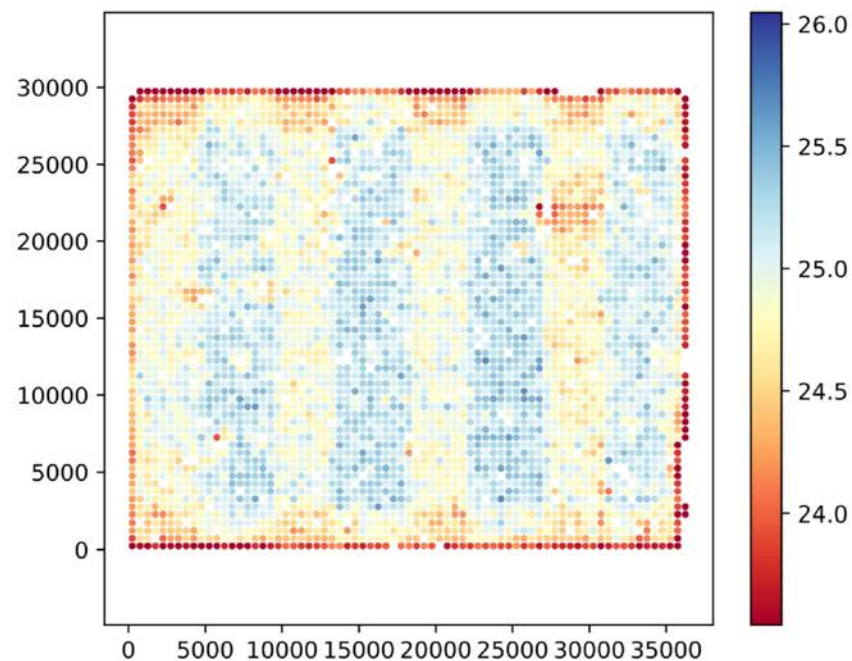


COSMOS2015 has been fantastically successful thanks to the great combination of optical Subaru data and deep infrared and near-infrared data

COSMOS2019: New UltraVISTA data (0.9 mag deeper); New u^* data (CLAUDS, 3-4 times deeper); all IRAC data (6 times deeper); New HSC data; better photometric extraction; better astrometric precision

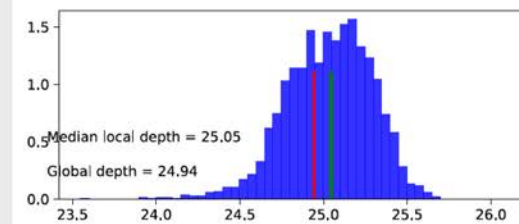
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Deeper/Uniform NIR UltraVISTA Data (DR4)



- 5 sigma, 2" aperture

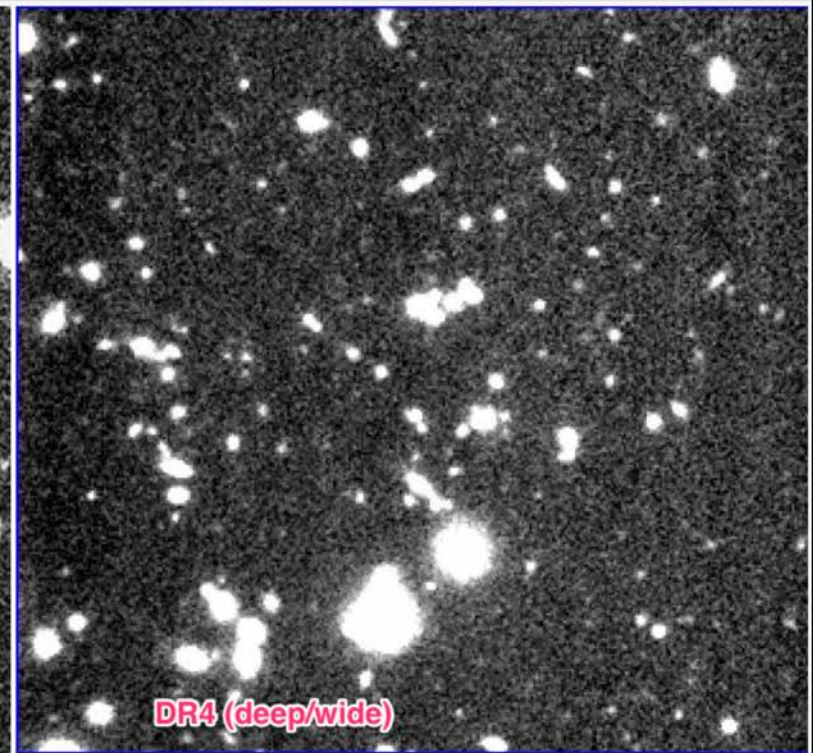
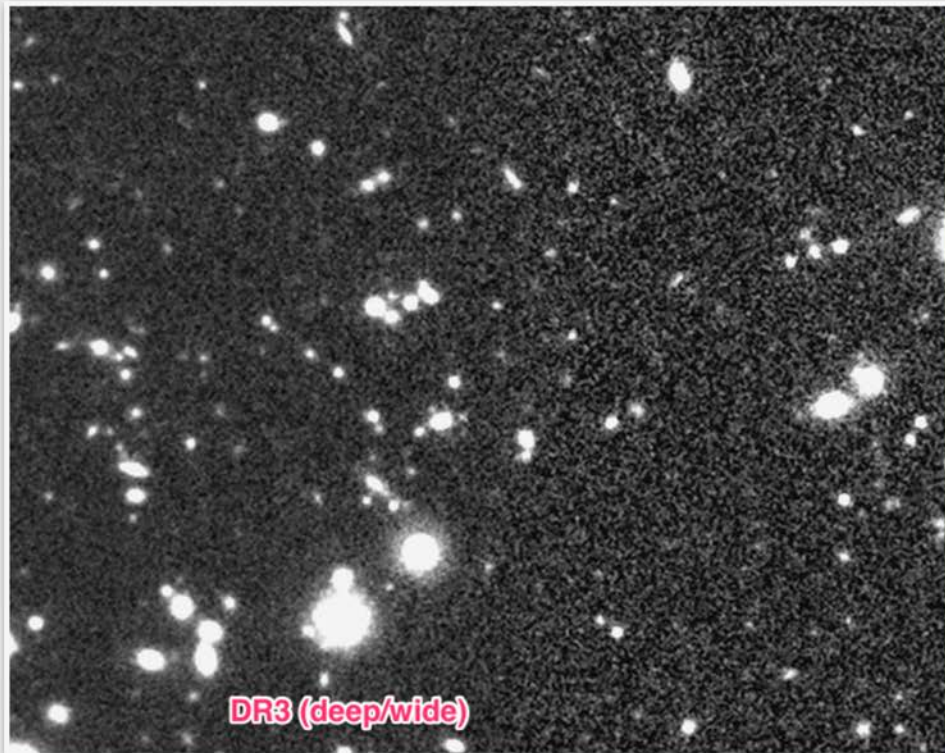
0.9 mag. in deep
increase compared to
DR3; 10% in YJ in
ultra-deep



courtesy R. Bowler

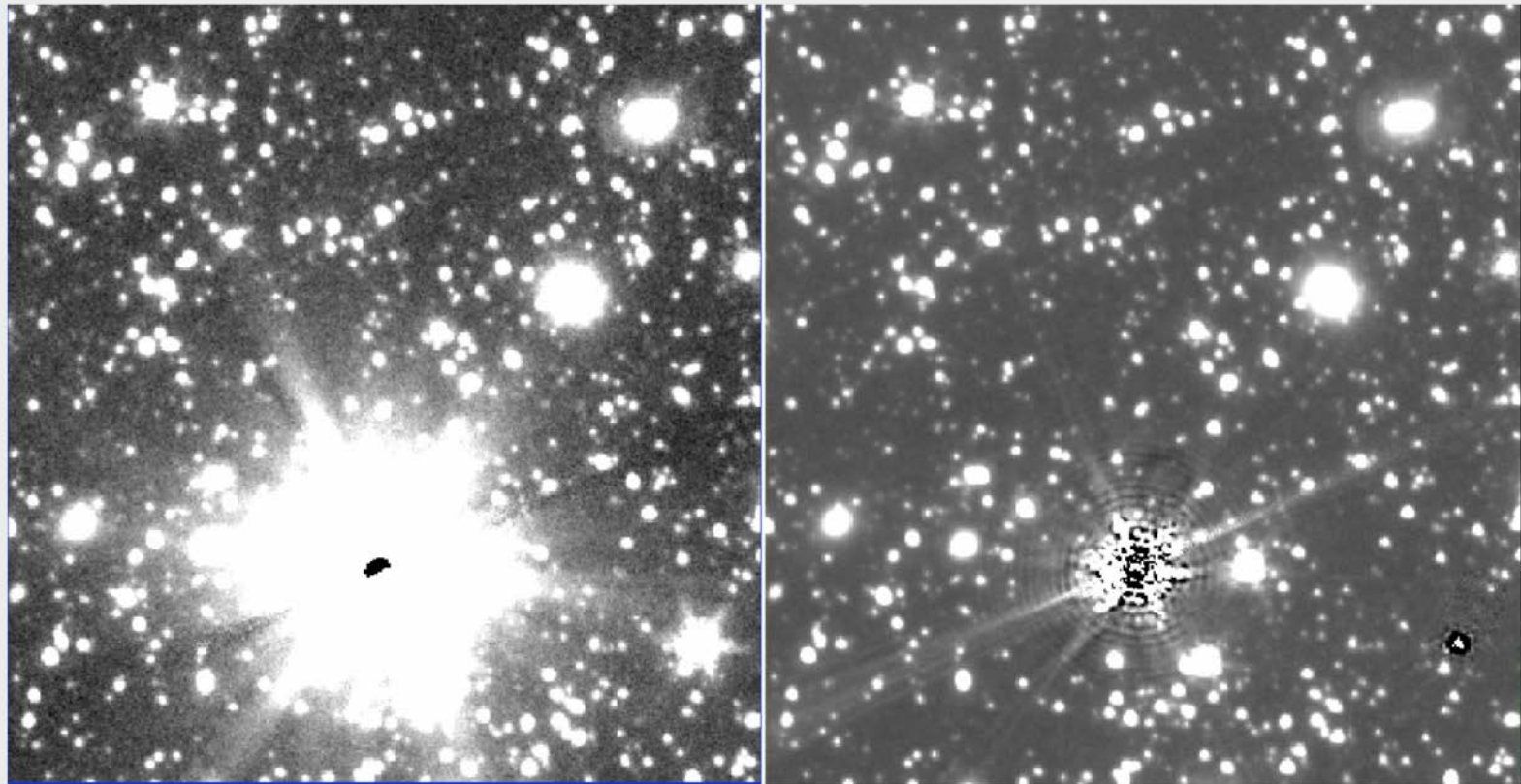
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Deeper/Uniform NIR UltraVISTA Data



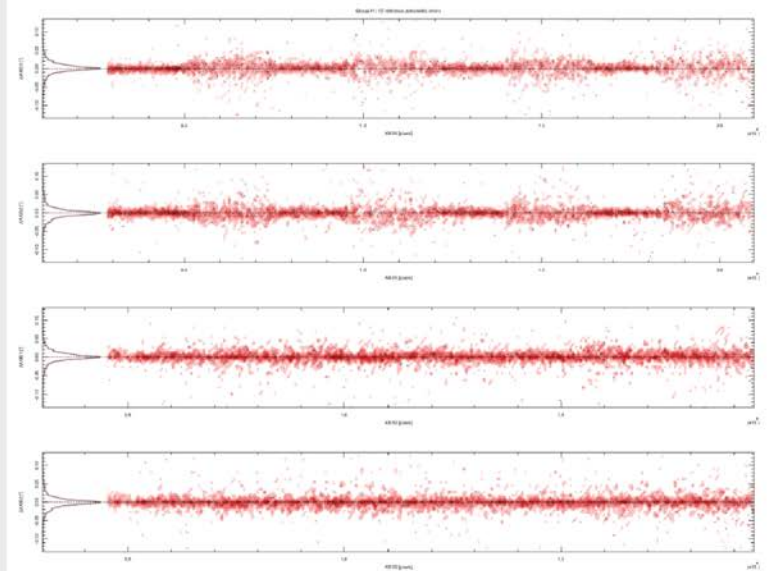
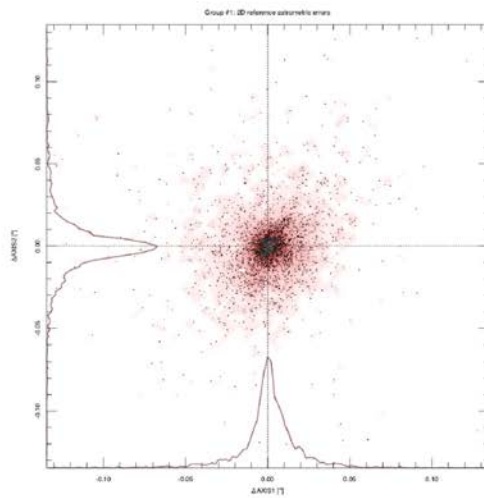
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Six times deeper Spitzer Data (4h/pix -> 26h/pix)



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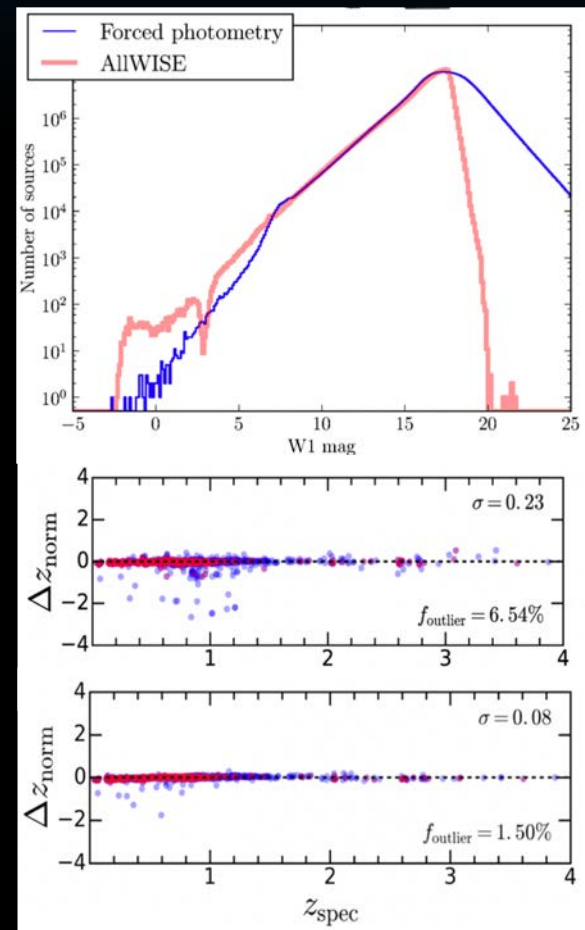
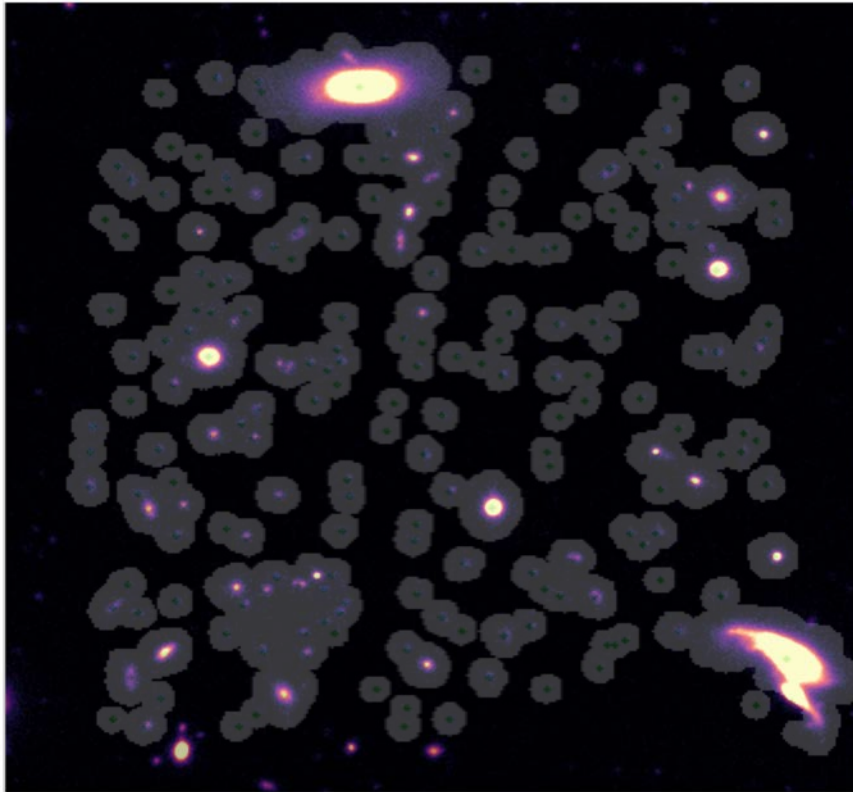
Better Astrometry



Compared to Gaia DR1. RMS better than 1/3 pixel! ->
Smaller apertures for photo-zed -> better precision

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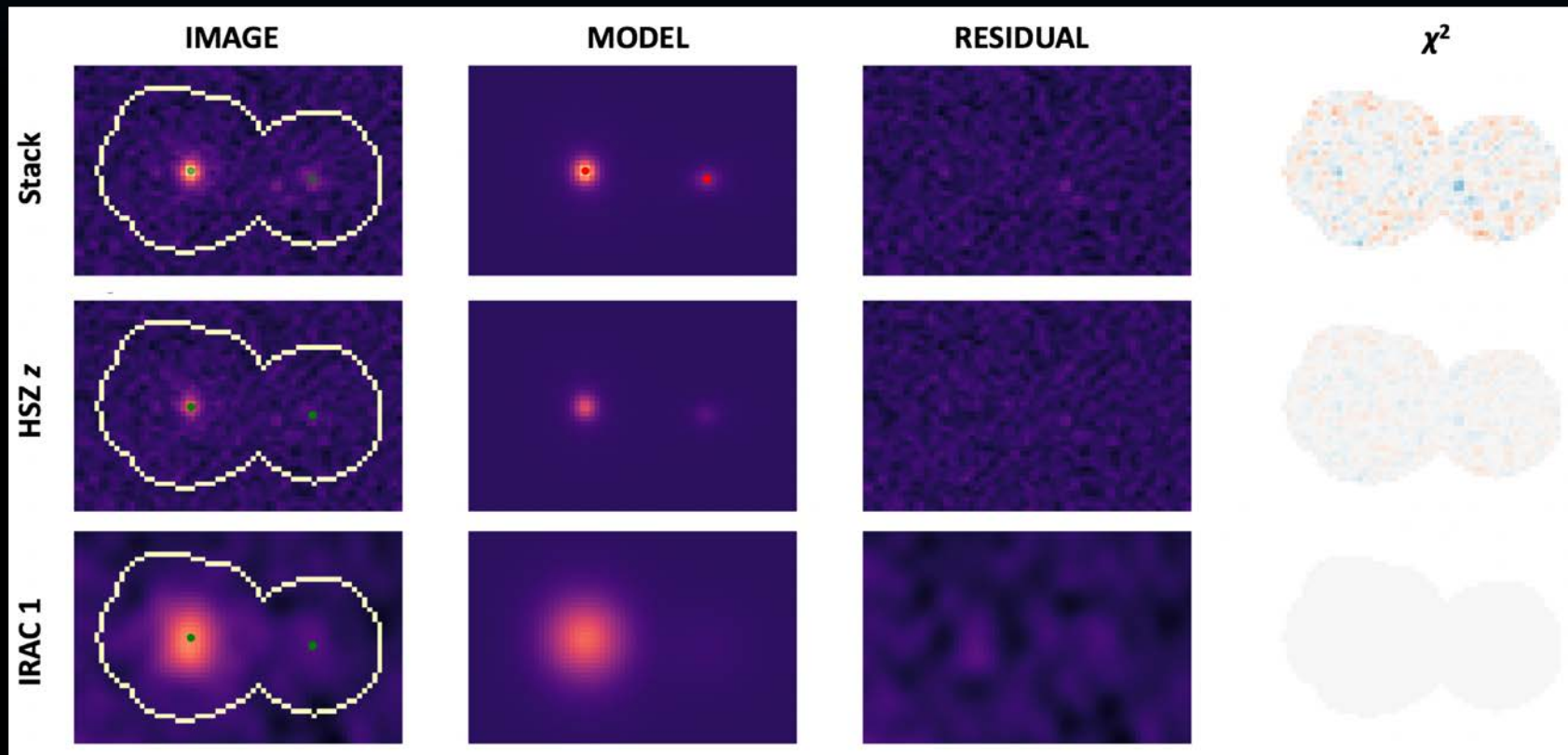
Better Photometry – Forced with Tractor



(Lang+16, Nyland+17)

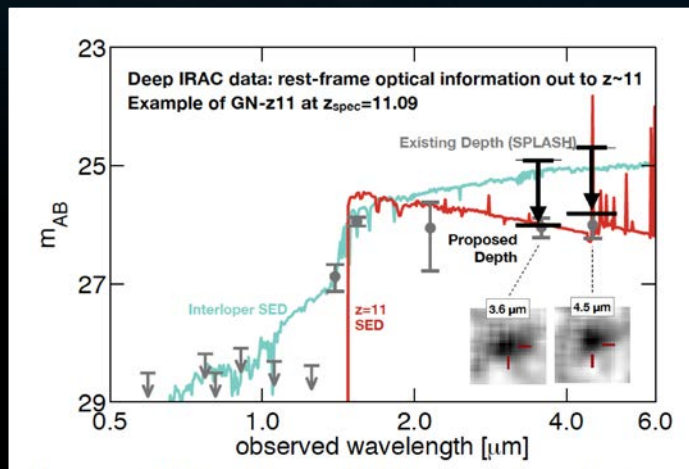
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Better Photometry – Forced with Tractor



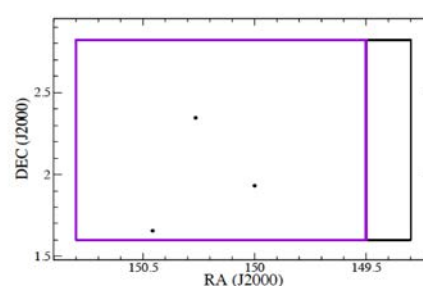
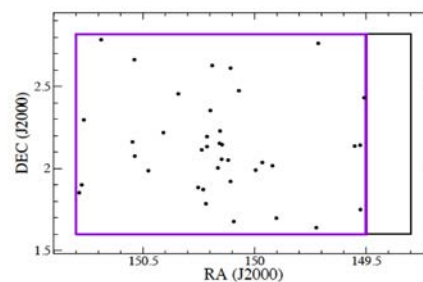
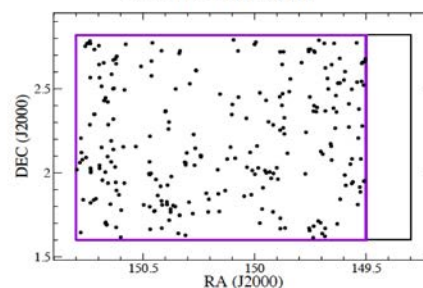
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COSMOS2019 Proposed Collaboration



COSMOS 2019 –A successful collaboration between Subaru/HSC and the COSMOS team ?

COSMOS 2015



COSMOS 2019

